

Holography and Motttness: a Discrete Marriage

Thanks to: NSF, EFRC
(DOE)

M. Edalati



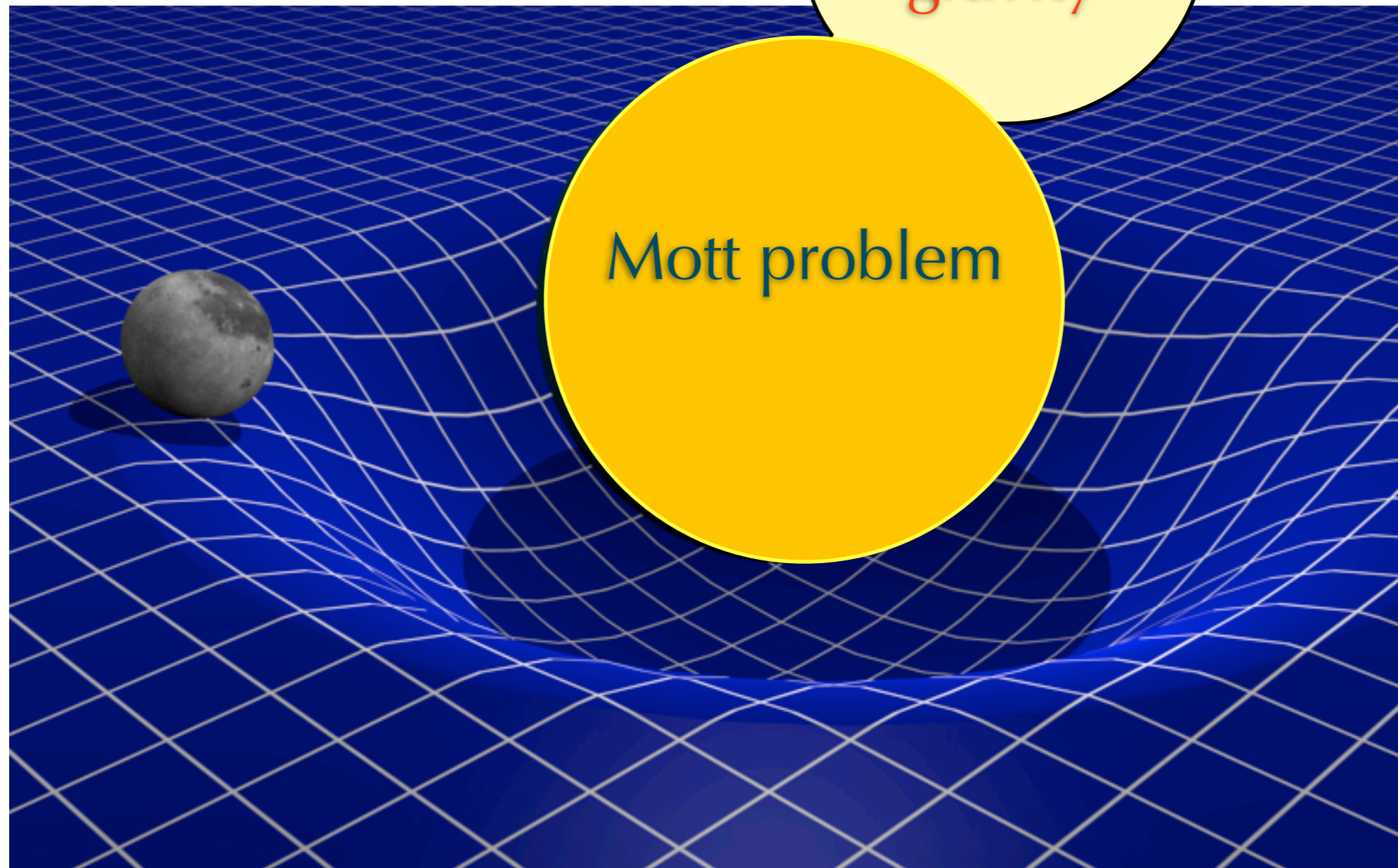
Ka Wai Lo



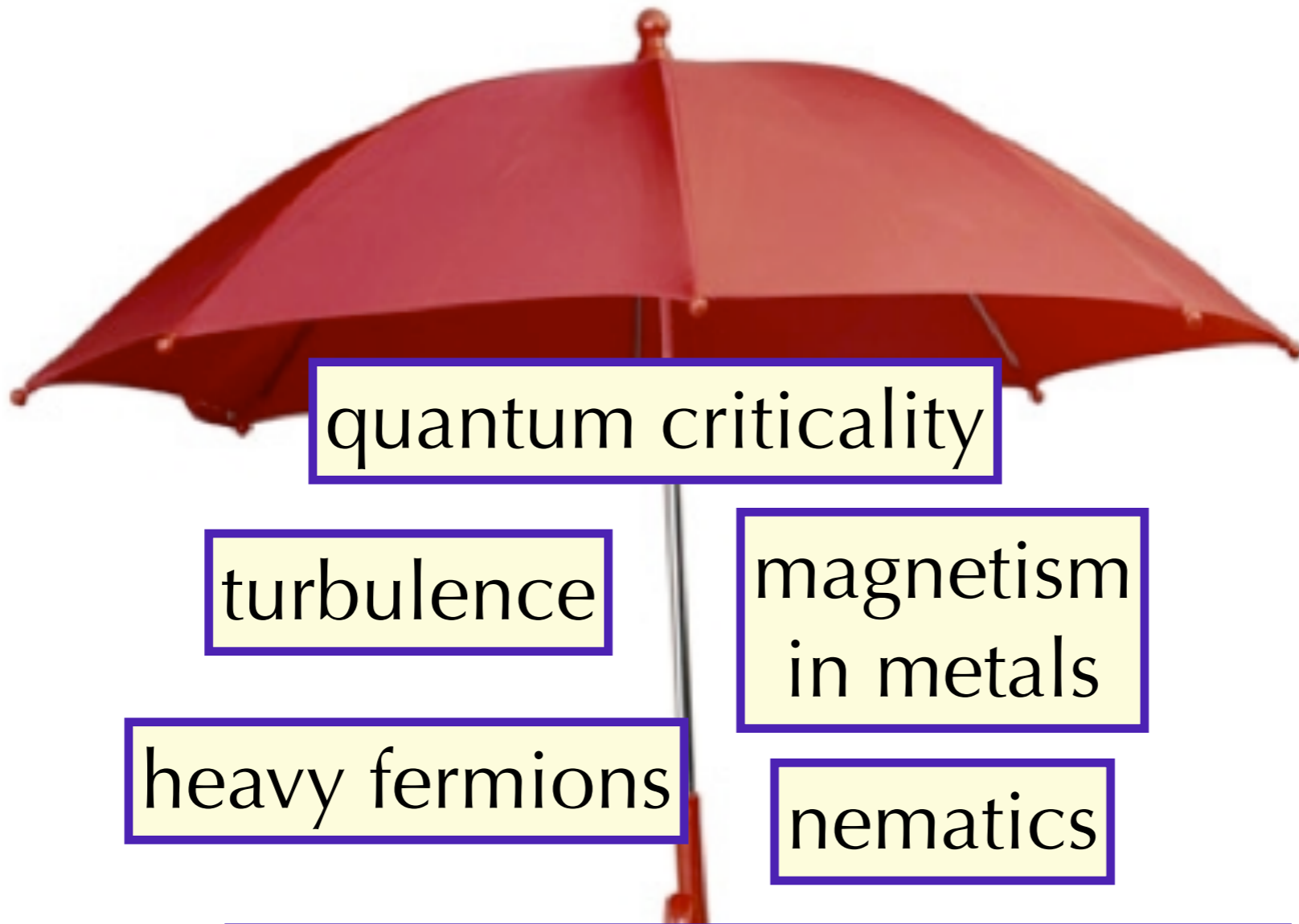
R. G. Leigh



Seungmin Hong



Unsolved condensed matter problems



quantum criticality

turbulence

magnetism
in metals

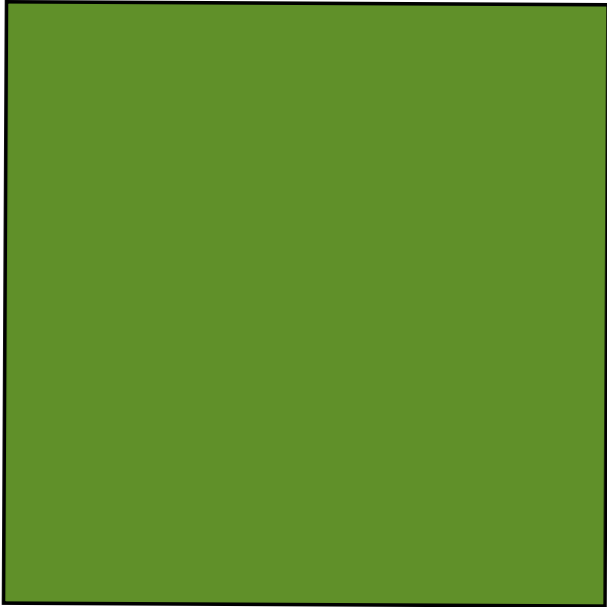
heavy fermions

nematics

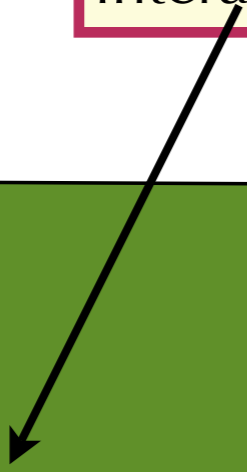
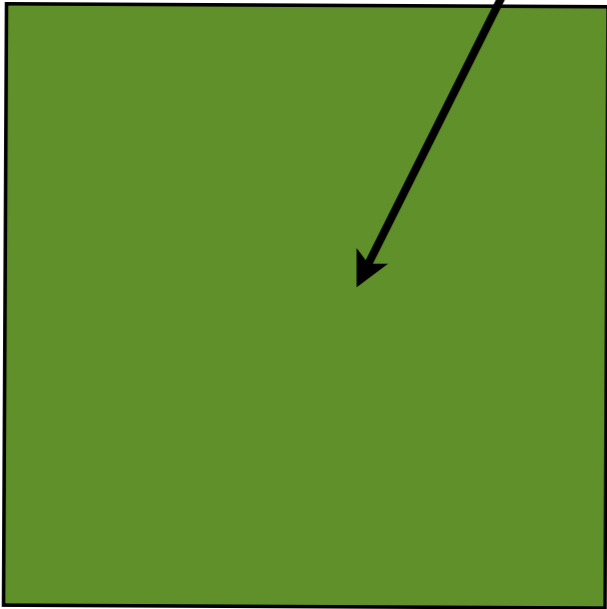
high T_c (cuprate, organics,...)

physics at strong coupling

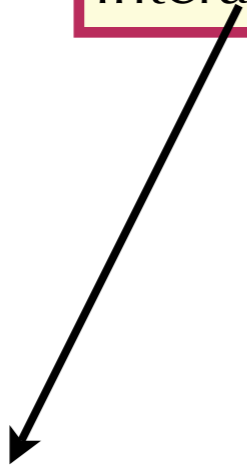
What computational tools do we have for strongly correlated electron systems?



interacting

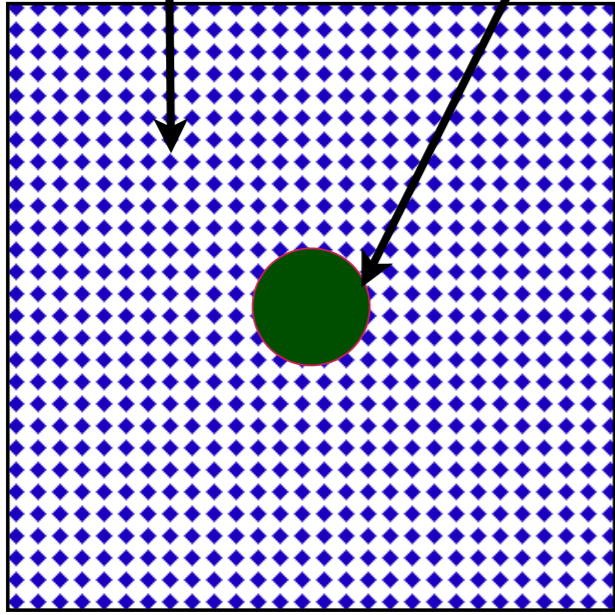


interacting



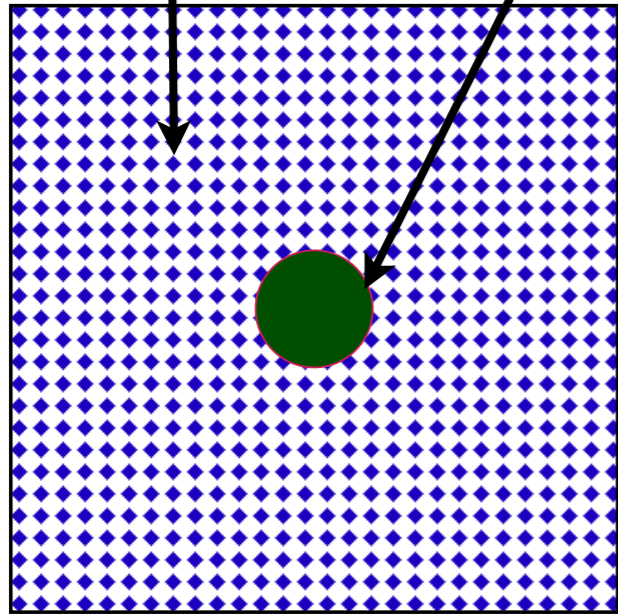
non-interacting

interacting



non-interacting

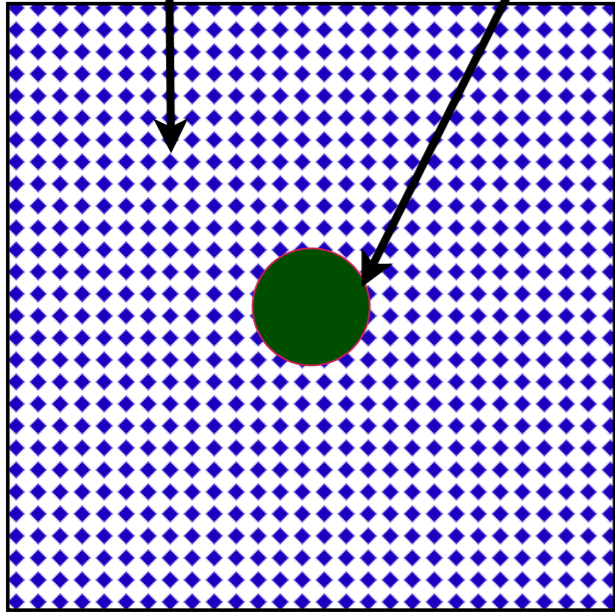
interacting



DMFT

non-interacting

interacting



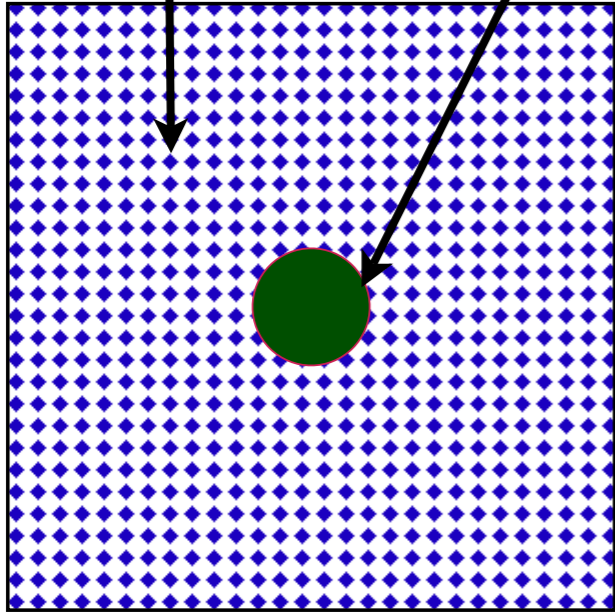
DMFT

the only difference between this and a theory is that this is not a theory



non-interacting

interacting



DMFT

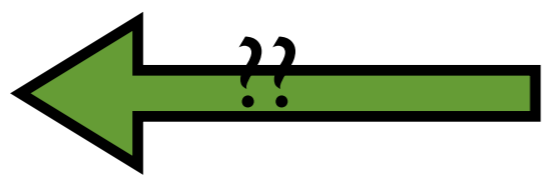
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why does this work at long wavelengths?



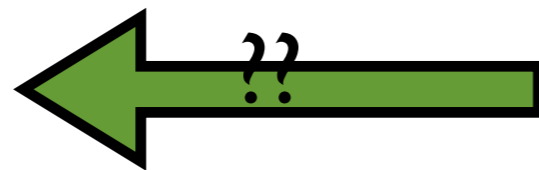


IR



UV
QFT

IR

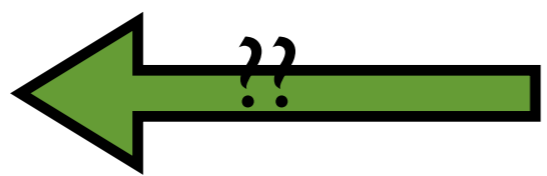


UV

QFT

Wilsonian
program
(fermions:
new degrees of
freedom)

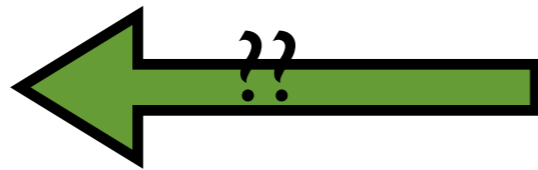
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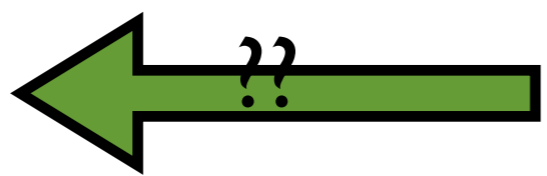
UV
QFT

coupling constant

$$g = 1/ego$$

~~Wilsonian
program
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IR



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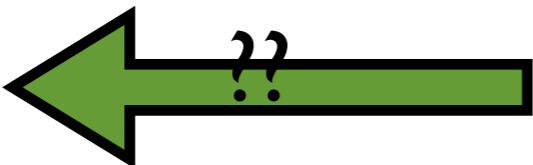
$$\frac{dg(E)}{d\ln E} = \beta(g(E))$$

locality in energy

~~Wilsonian
program
(fermions:
new degrees of
freedom~~

implement E-scaling with an extra dimension

IR



UV
QFT

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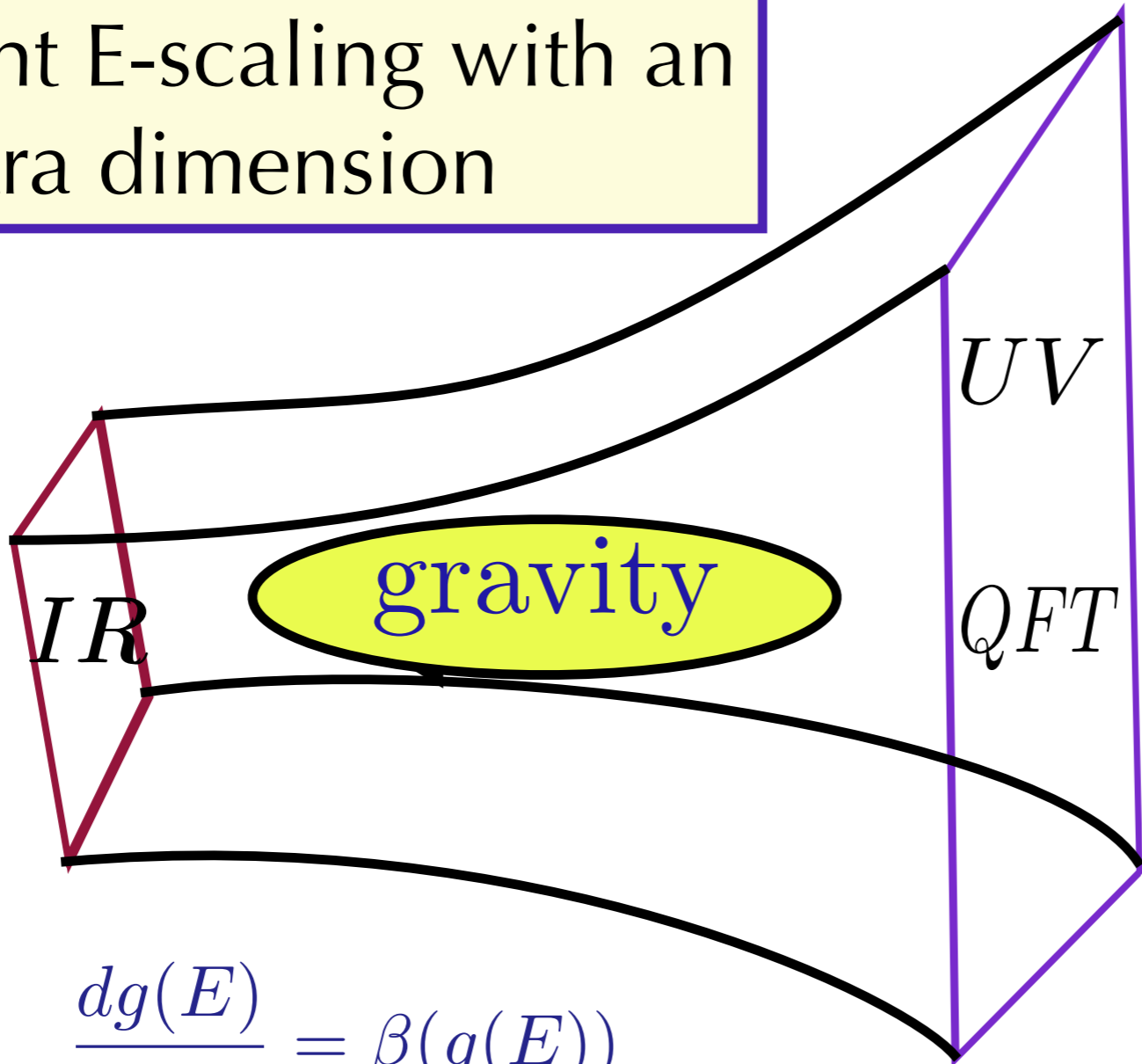
$$\frac{dg(E)}{d \ln E} = \beta(g(E))$$

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gauge-gravity duality
(Maldacena, 1997)

implement E-scaling with an
extra dimension



coupling constant

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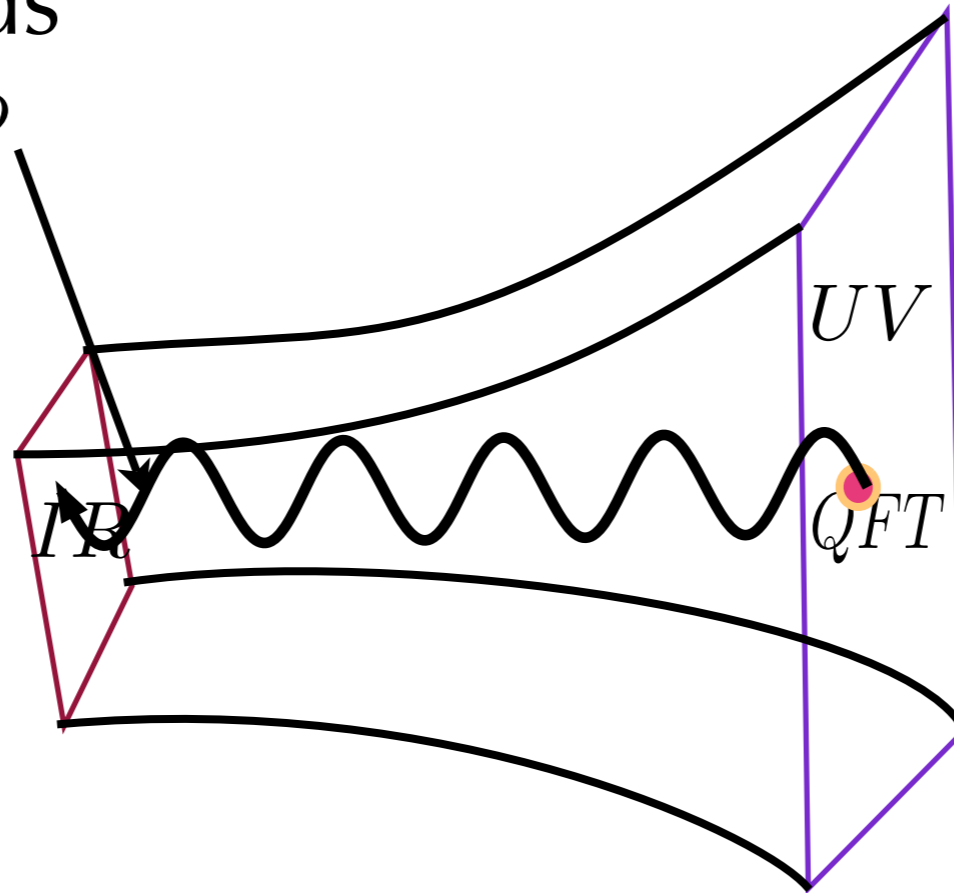
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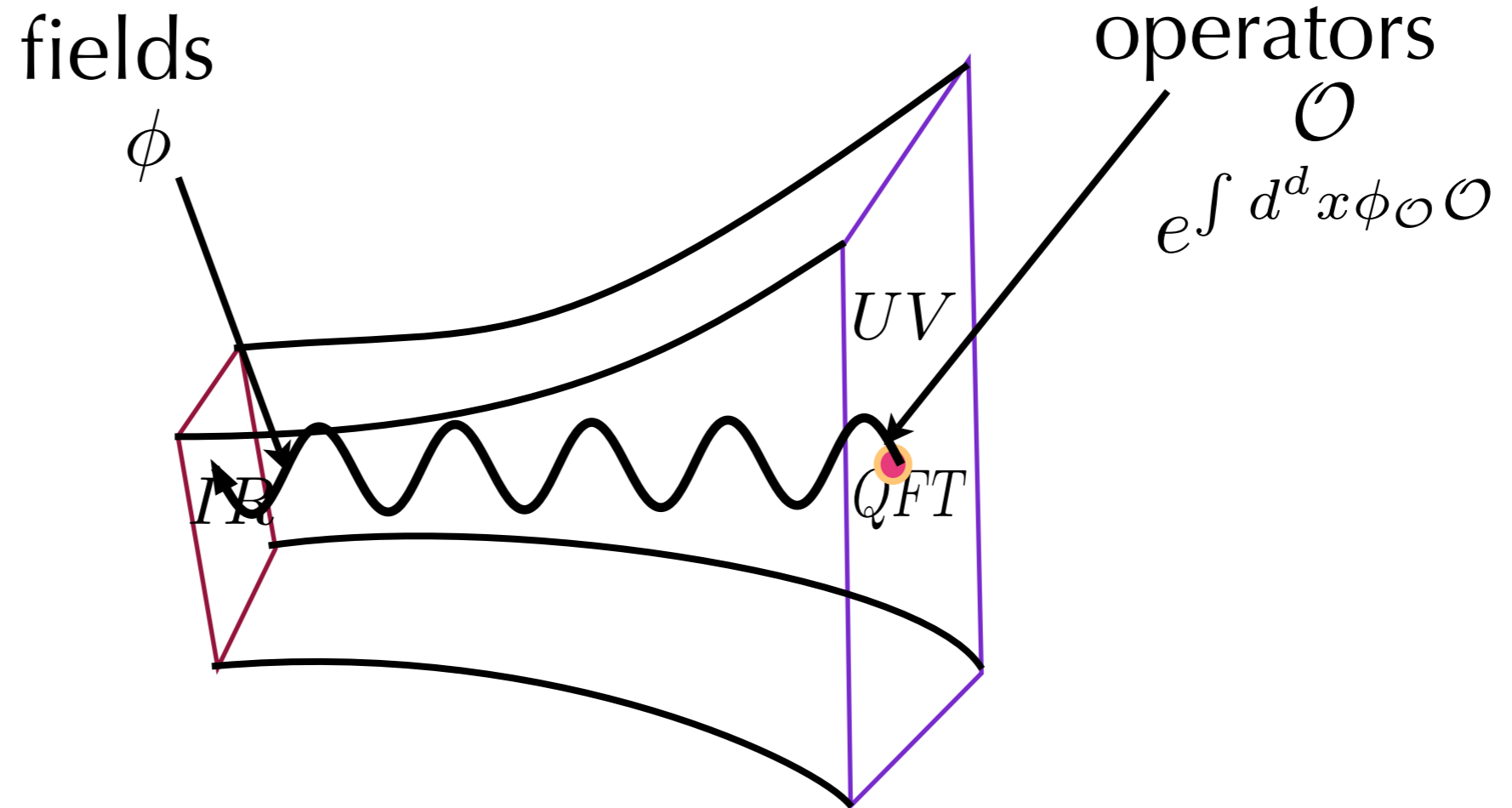
fields

ϕ



$$e^{\int d^d x \phi \mathcal{O}}$$

Claim: $Z_{\text{QFT}} = e^{-S_{\text{ADS}}^{\text{on-shell}}(\phi(\phi_{\partial\text{ADS}} = J_{\mathcal{O}}))}$



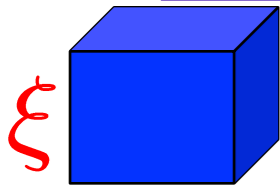
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What holography
does for you?

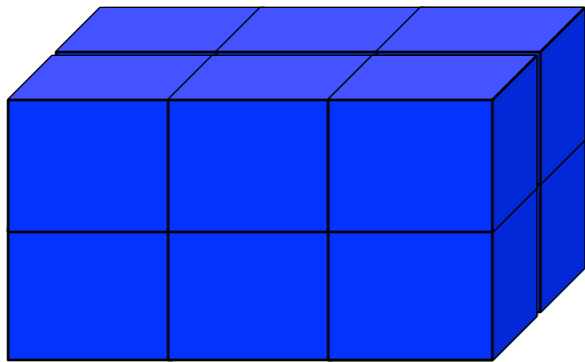
What holography
does for you?

Landau-Wilson

Hamiltonian



$$\xi_t \propto \xi^z$$



long-wavelengths

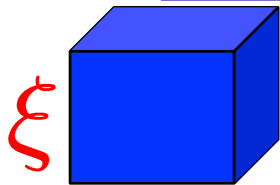
RG equations

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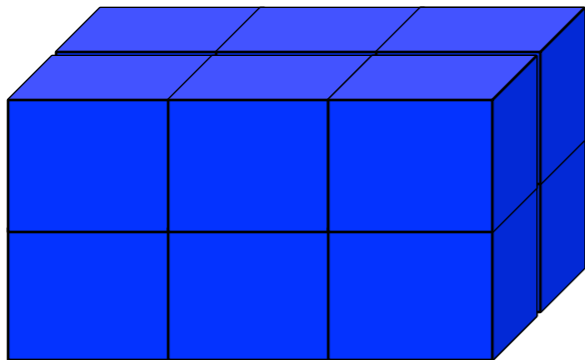
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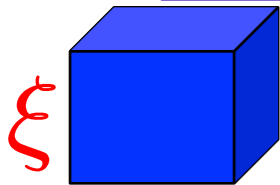
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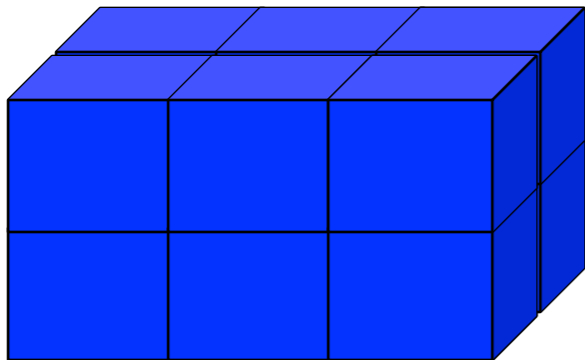
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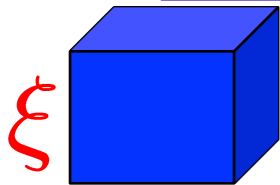
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RG=GR

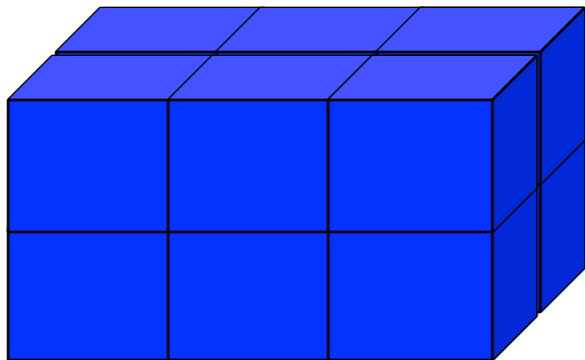
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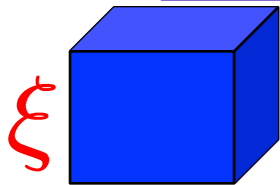
RG=GR

strong-coupling is
easy

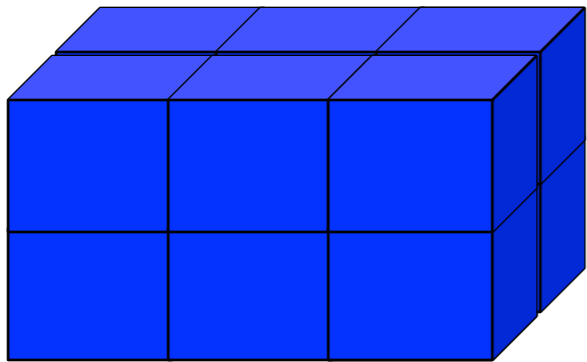
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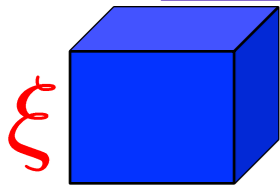
strong-coupling is
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microscopic UV
model not easy
(need *M*-theory)

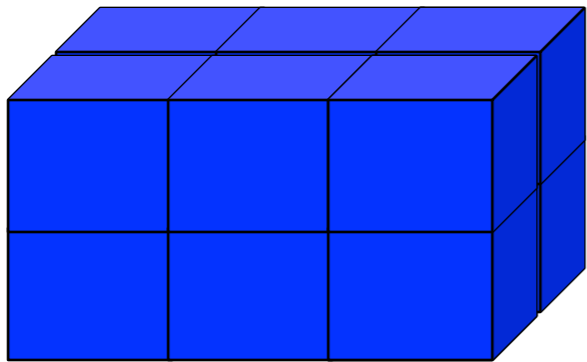
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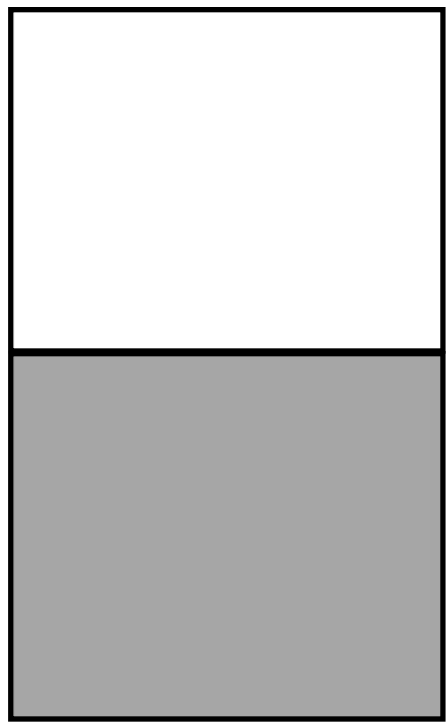
strong-coupling is
easy

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so what
(currents,
symmetries)

Can holography solve the Mott
problem?

Half-filled band



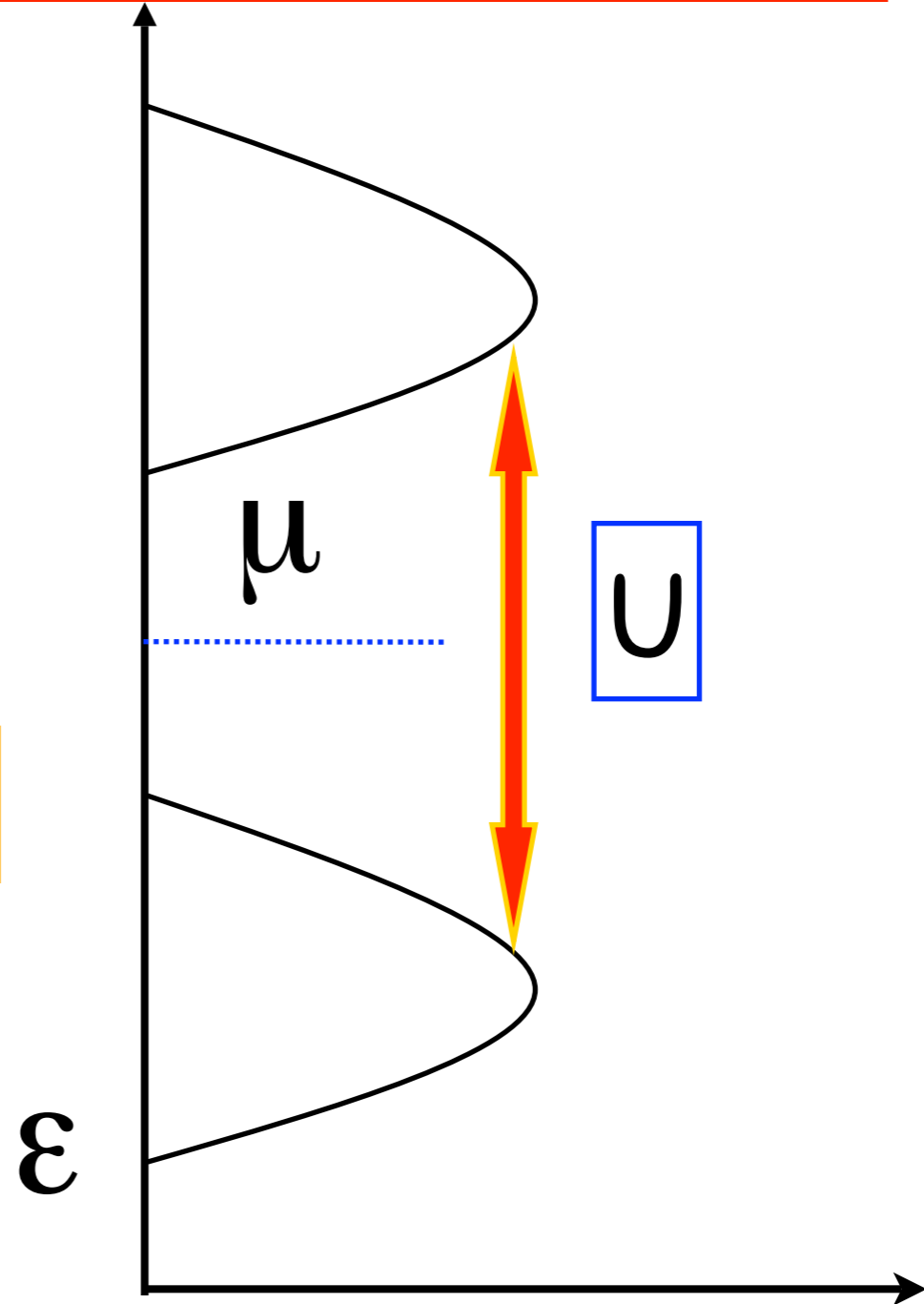
Free electrons

$$U \gg t$$

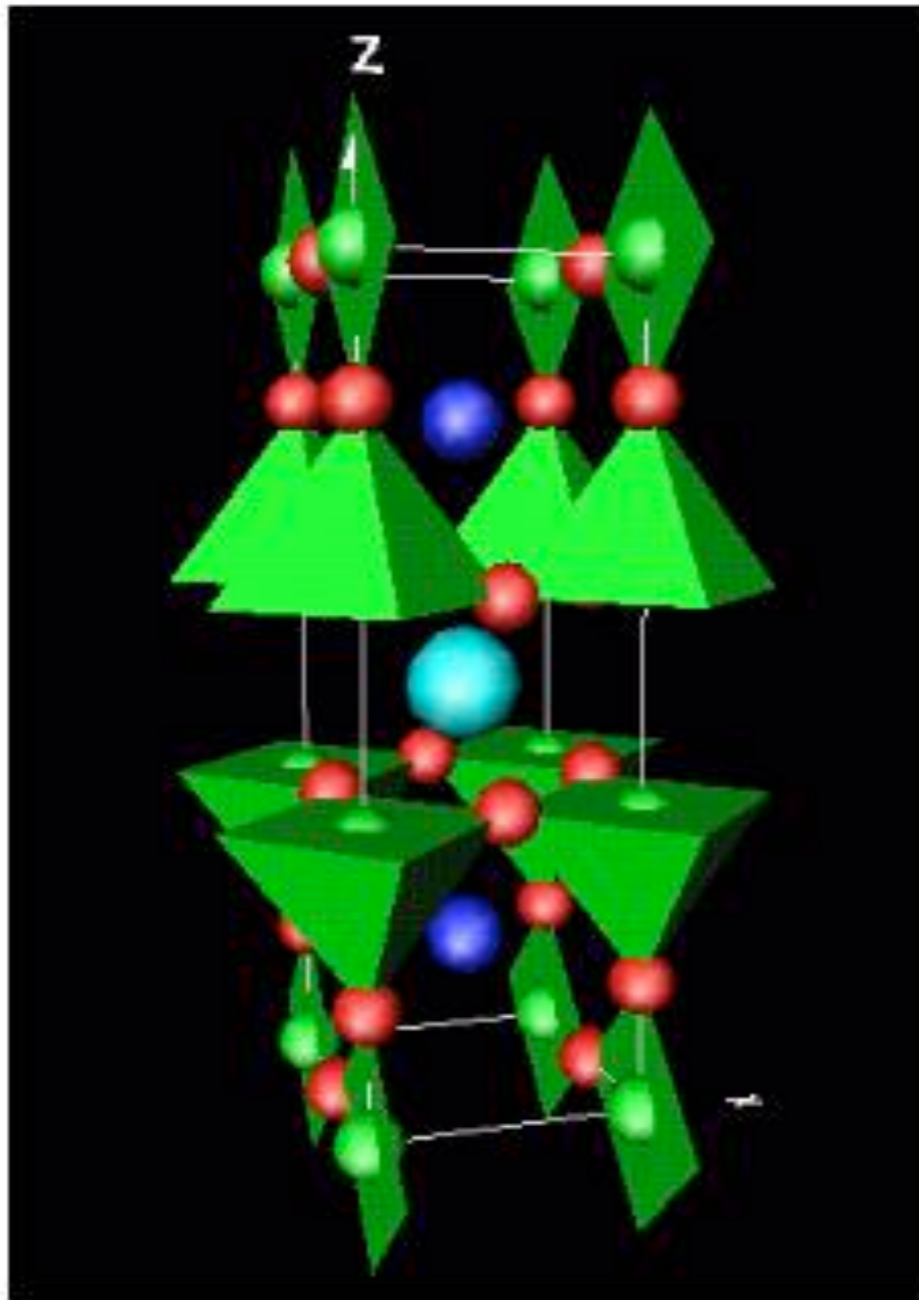


charge gap

gap with no symmetry breaking!!



Why is the Mott problem important?



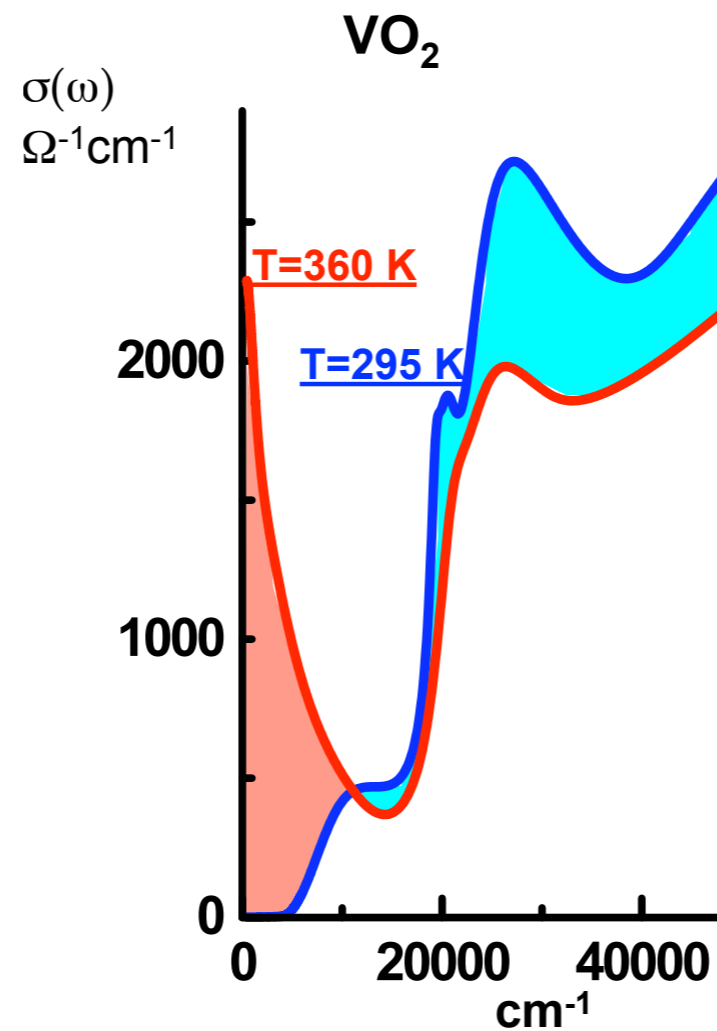
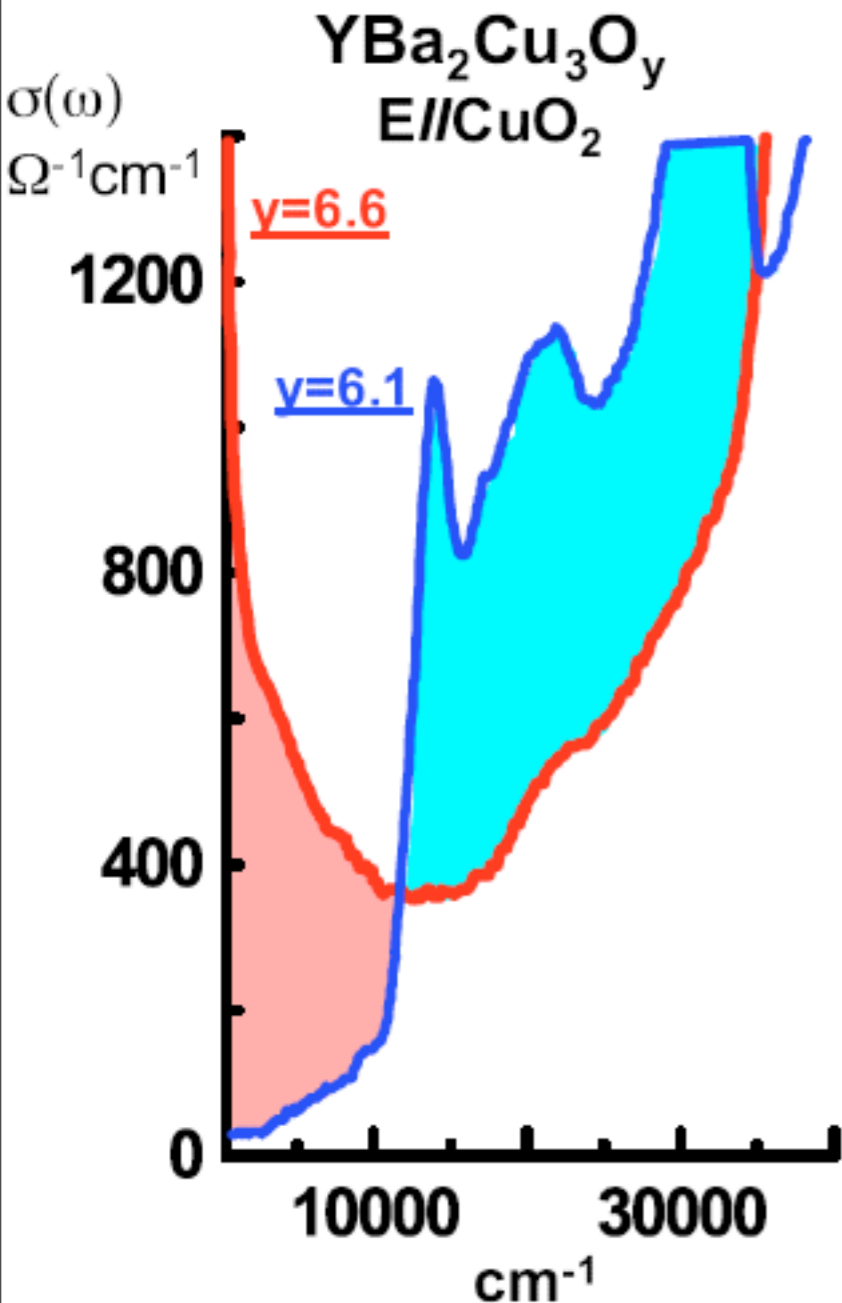
$$U/t = 10 \gg 1$$

interactions dominate:
Strong Coupling Physics

$\text{Y Ba}_2\text{Cu}_3\text{O}_7$
Cuprate Superconductors

Experimental facts: Motttness

$\Delta = 0.6eV > \Delta_{\text{dimerization}}$ (Mott, 1976) $\frac{\Delta}{T_{\text{crit}}} \approx 20$



transfer
of spectral
weight to
high energies
beyond any ordering
scale

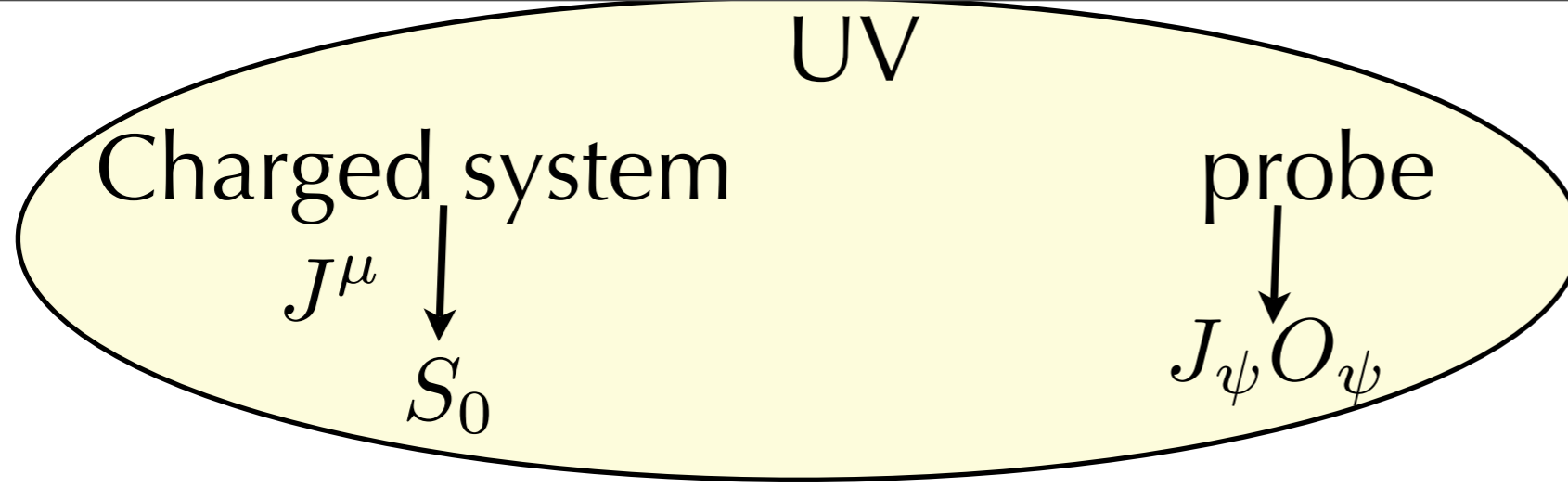
Recall, $eV = 10^4 K$

*M. M. Qazilbash, K. S. Burch, D. Whisler,
D. Shrekenhamer, B. G. Chae, H. T. Kim,
and D. N. Basov PRB 74, 205118 (2006)*

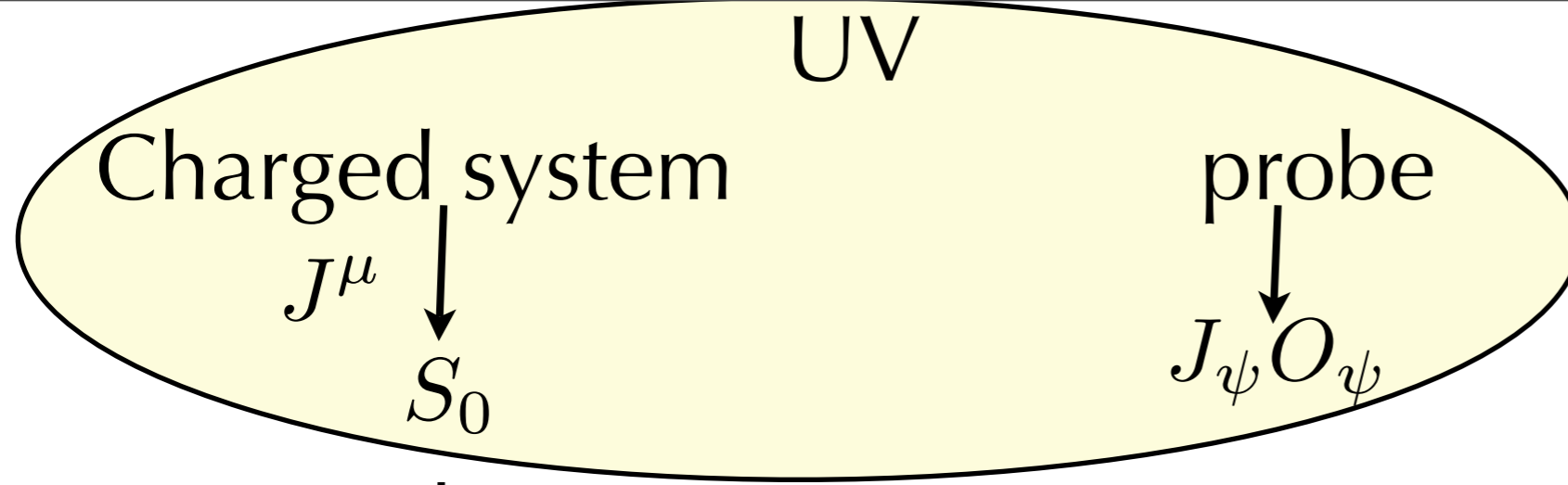
What bulk gravitational theory gives rise to a gap in $\text{Im}G$
without
'spontaneous' symmetry breaking?



dynamically generated gap: Mott gap
(for probe fermions)



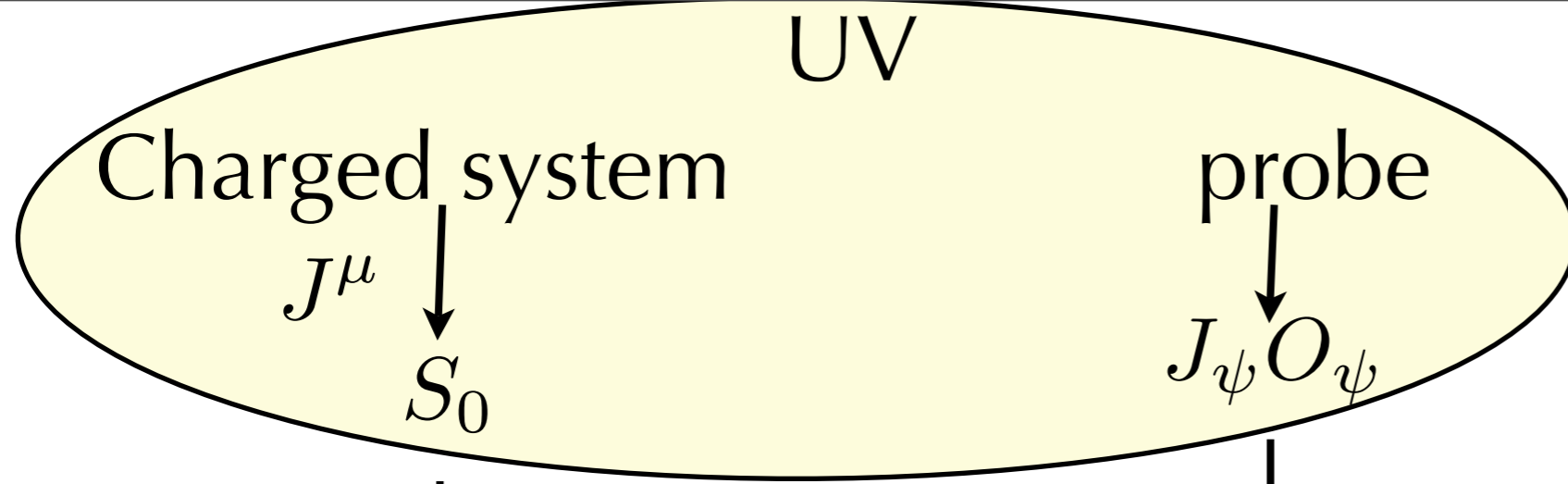
what has been
done?
MIT, Leiden,
McMaster,...



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RN-AdS

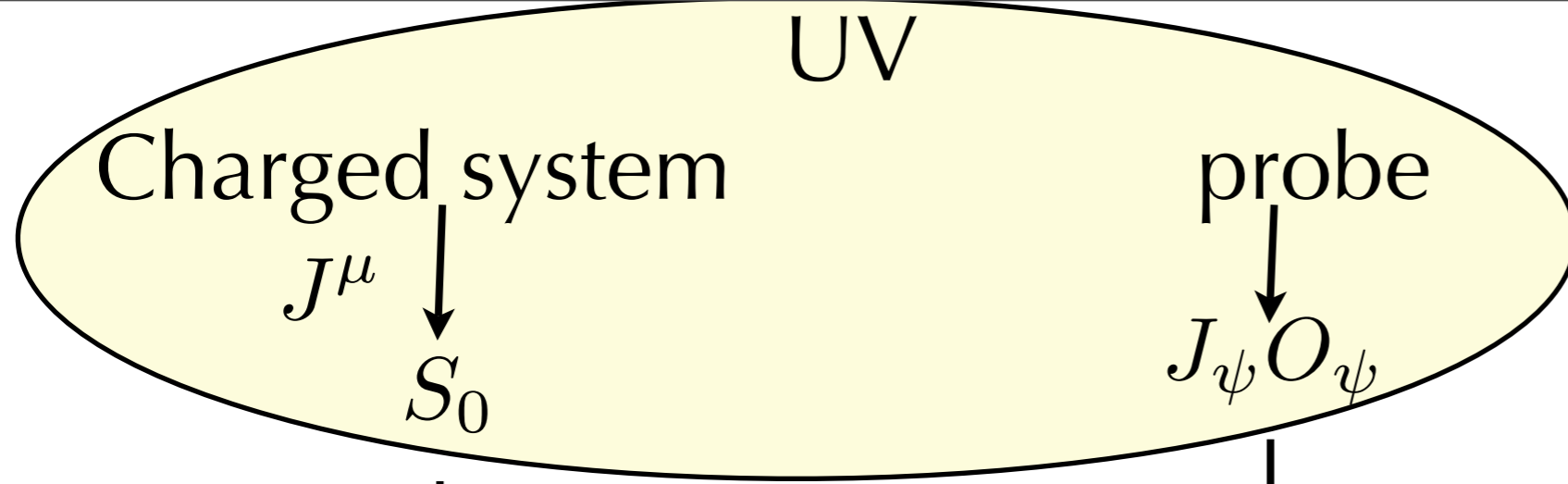
ds^2, A_t



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MIT, Leiden,
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RN-AdS
 ds^2, A_t

ψ Dirac Eq.



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MIT, Leiden, McMaster,...

RN-AdS
 ds^2, A_t

ψ Dirac Eq.

in-falling boundary conditions

$$\psi(r \rightarrow \infty) \approx ar^m + br^{-m}$$

Retarded Green function: $G = \frac{b}{a} = f(\text{UV } (k_F), \text{IR } (q, m))$

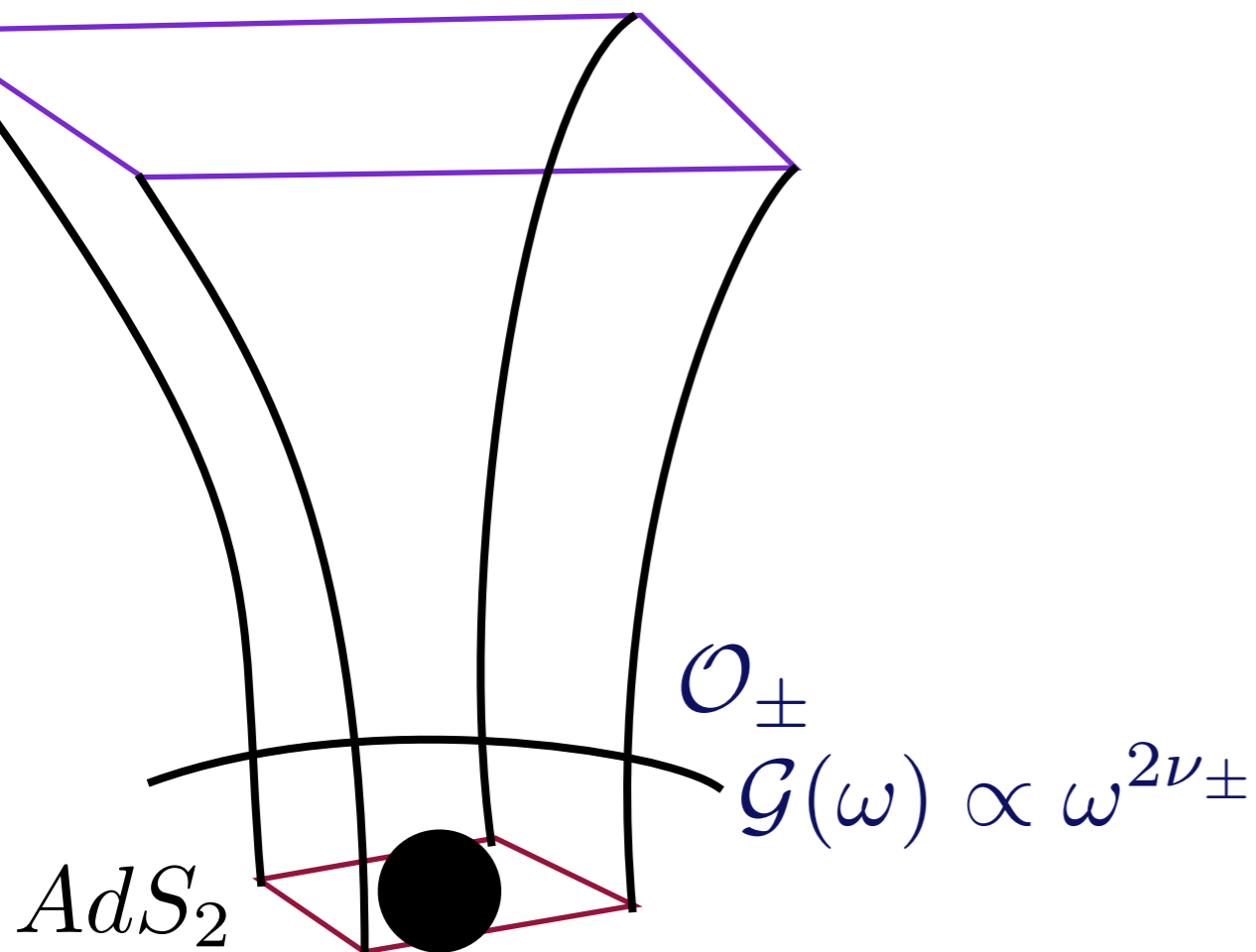
a=0 defines FS

$$ds^2 = \frac{L_2^2}{\zeta^2} (-d\tau^2 + d\zeta^2) + \frac{r_0^2}{R^2} d\vec{x}^2$$

$$\zeta = \omega L_2^2 / (r - r_0)$$

boundary physics:
 $AdS_2 \times R^2$

$$\mathcal{O} \quad G_R(k, \omega) = f(\mathcal{G}(\omega))$$

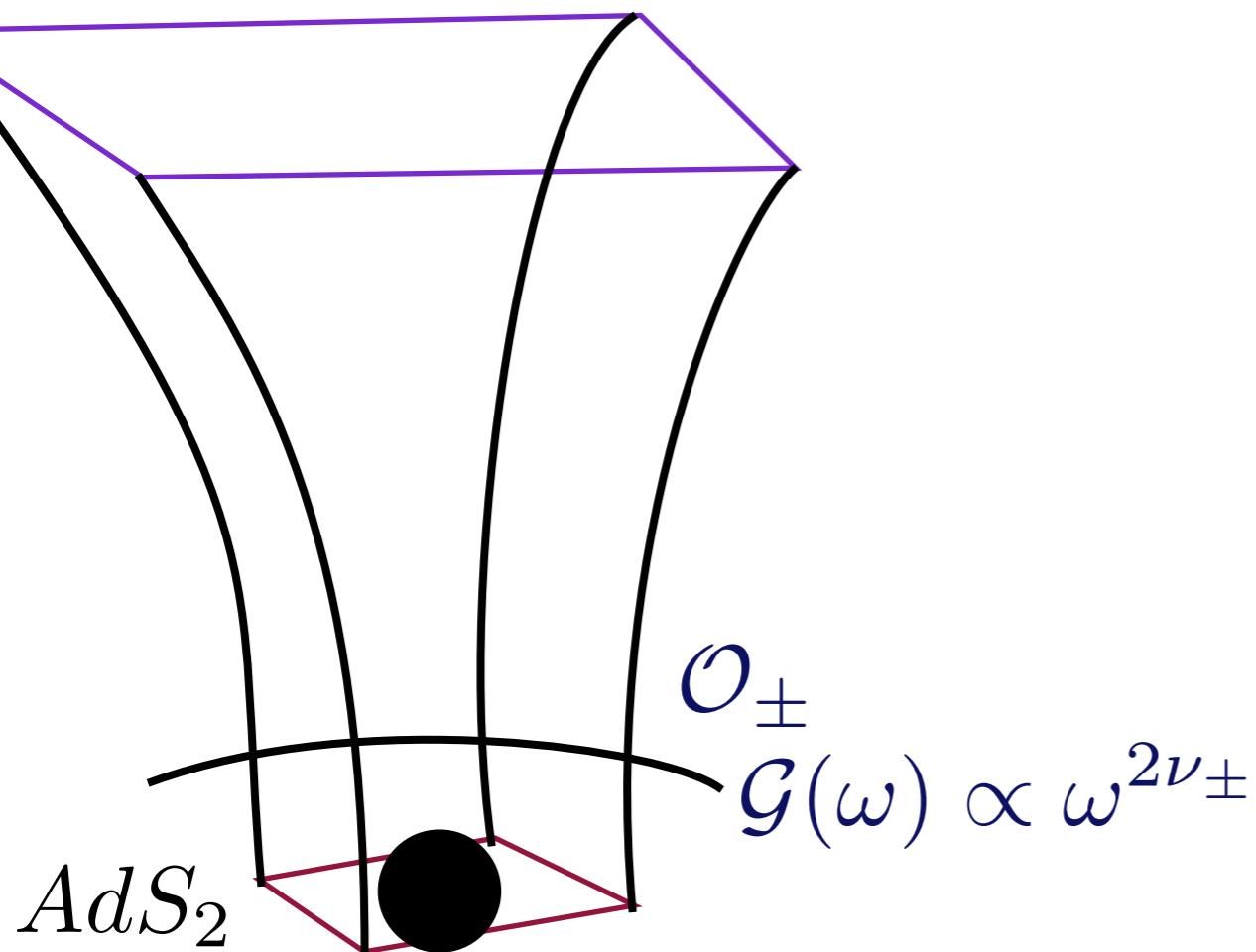


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$$\{\vec{x}, \pi_{\zeta}\} \rightarrow \{\vec{x}, \lambda\tau, \lambda\zeta\}$$

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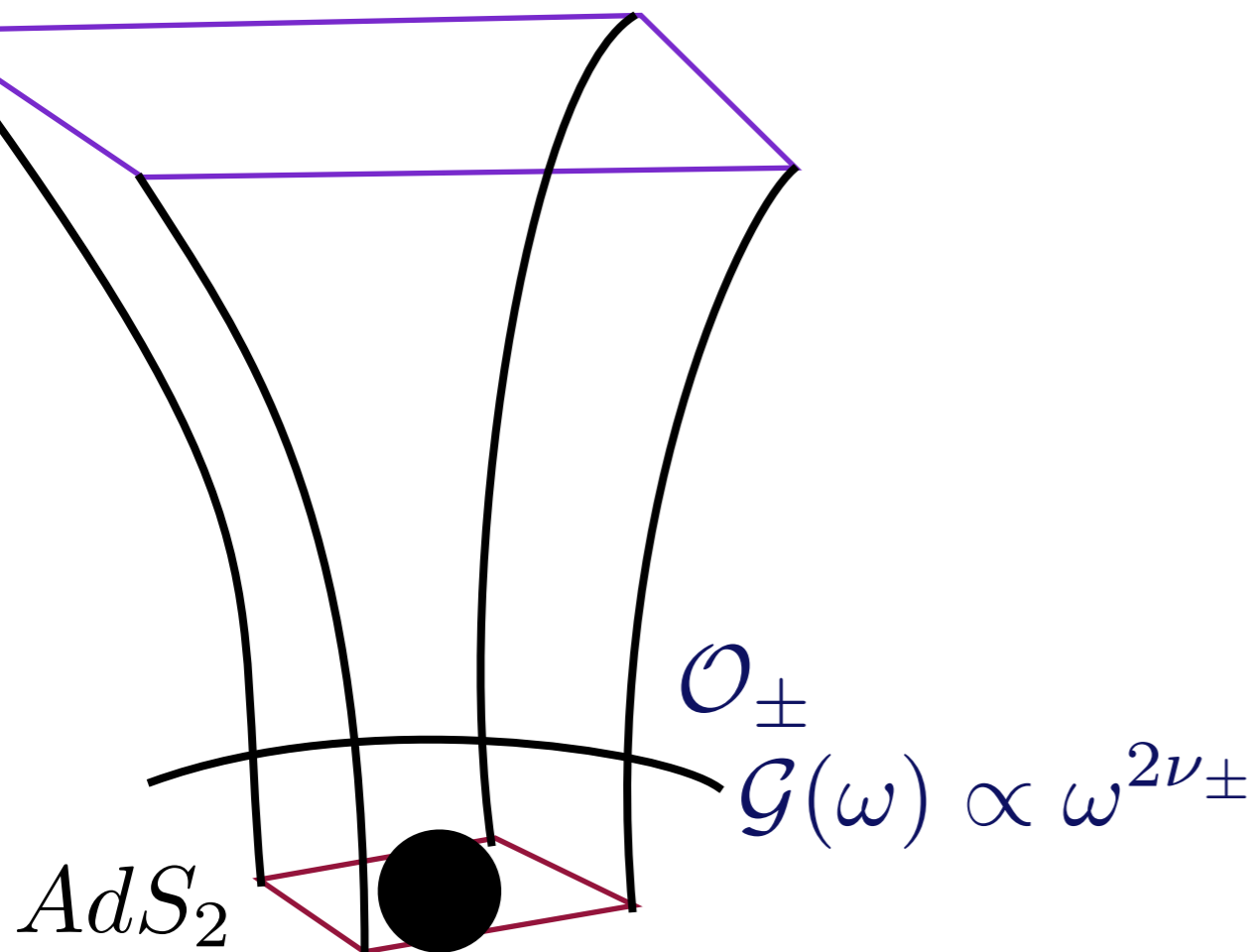
finite $T=0$ entropy

$$S \propto T^{2/z} \neq 0 \quad T \rightarrow 0$$

$$\vec{x} \rightarrow \lambda^{(1/z=\infty)} \vec{x} = \vec{x}$$

$$\{\vec{x}, \tau, \zeta\} \rightarrow \{\vec{x}, \lambda\tau, \lambda\zeta\}$$

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$$ds^2 = \frac{L_2^2}{\zeta^2} (-d\tau^2 + d\zeta^2) + \frac{r_0^2}{R^2} d^2\vec{x}$$

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boundary physics:
 $AdS_2 \times R^2$

$$\sqrt{-g}i\bar{\psi}(D - m)\psi$$

AdS-RN
MIT, Leiden group



$$G(\omega, k) = \frac{Z}{v_F(k - k_F) - \omega - h\omega \ln \omega}$$

marginal Fermi
liquid

??



Mott gap

How to Destroy the Fermi surface?

decoherence

$$\sqrt{-g}i\bar{\psi}(D - m)\psi$$

AdS-RN
MIT, Leiden group



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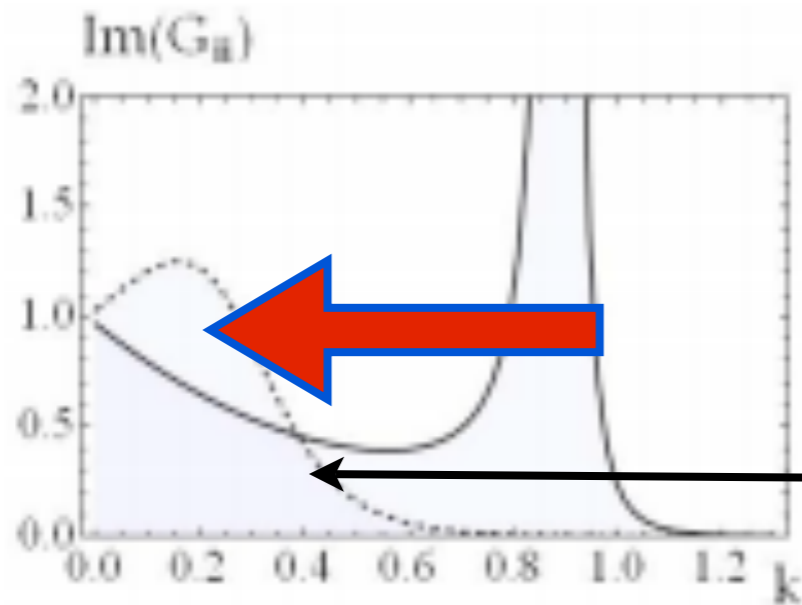
??



Mott gap

How to Destroy the Fermi surface?


decoherence



log-oscillatory regime

$$S_{\text{probe}}(\psi, \bar{\psi}) = \int d^d x \sqrt{-g} i \bar{\psi} (\Gamma^M D_M - m + \dots) \psi$$

$$S_{\text{probe}}(\psi, \bar{\psi}) = \int d^d x \sqrt{-g} i \bar{\psi} (\Gamma^M D_M - m + \dots) \psi$$



what is hidden here?

$$S_{\text{probe}}(\psi, \bar{\psi}) = \int d^d x \sqrt{-g} i \bar{\psi} (\Gamma^M D_M - m + \dots) \psi$$

what is hidden here?

consider

$$\sqrt{-g} i \bar{\psi} (\not{D} - m - ip \not{F}) \psi$$

$F_{\mu\nu} \Gamma^{\mu\nu}$

$$S_{\text{probe}}(\psi, \bar{\psi}) = \int d^d x \sqrt{-g} i \bar{\psi} (\Gamma^M D_M - m + \dots) \psi$$

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QED anomalous magnetic moment of an electron
(Schwinger 1949)

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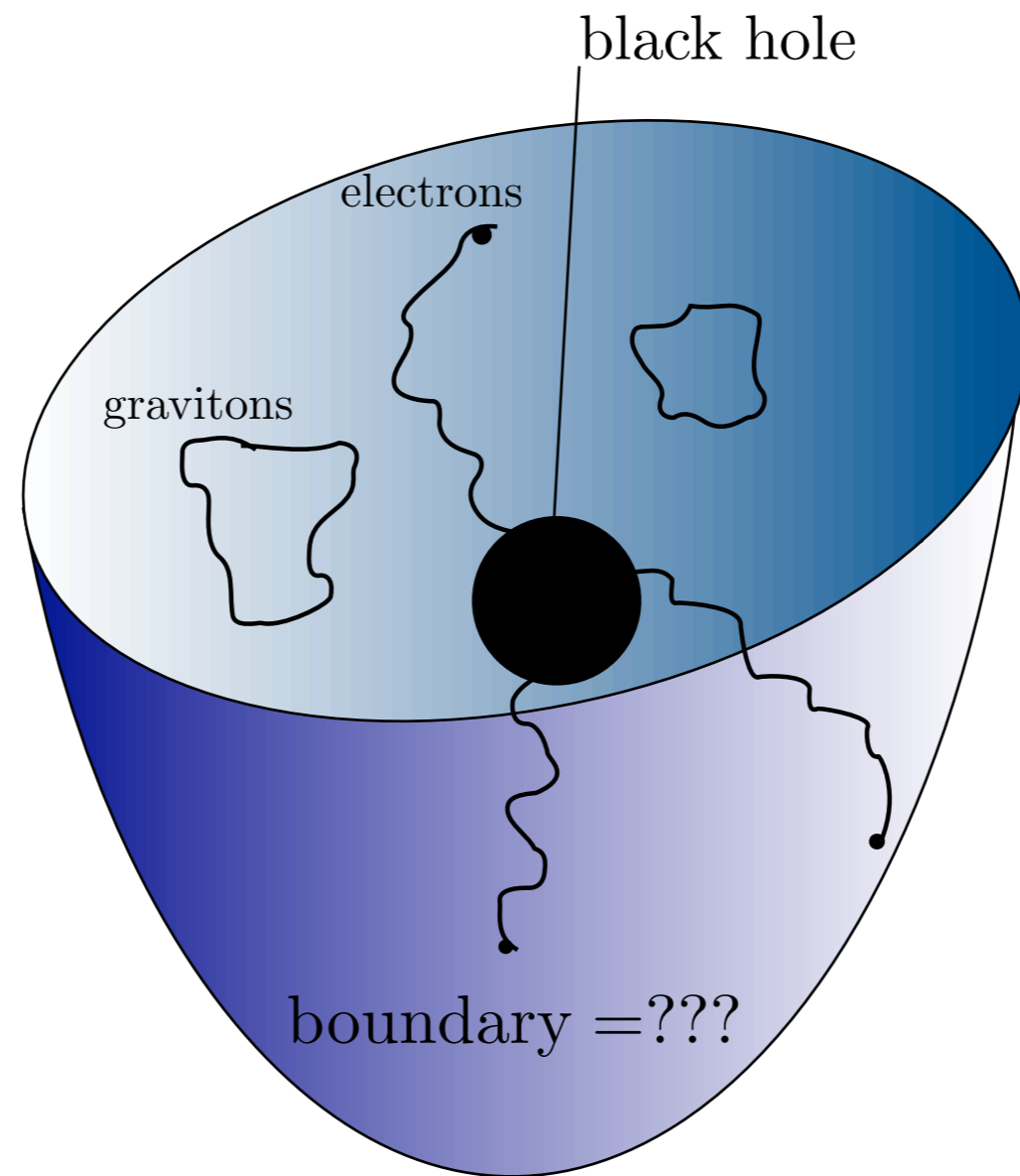
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$F_{\mu\nu} \Gamma^{\mu\nu}$

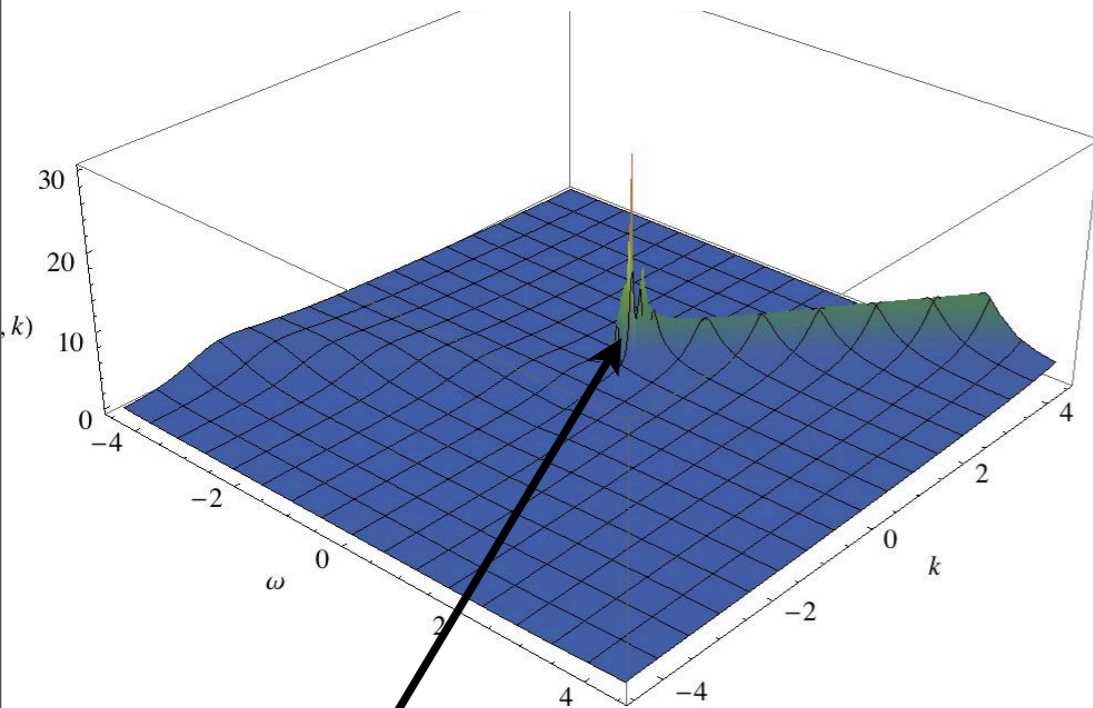
QED anomalous magnetic moment of an electron
(Schwinger 1949)

fermions in RN $\text{Ads}_{\{d+1\}}$ coupled to a gauge field
through a dipole interaction



How is the spectrum modified?

$P=0$

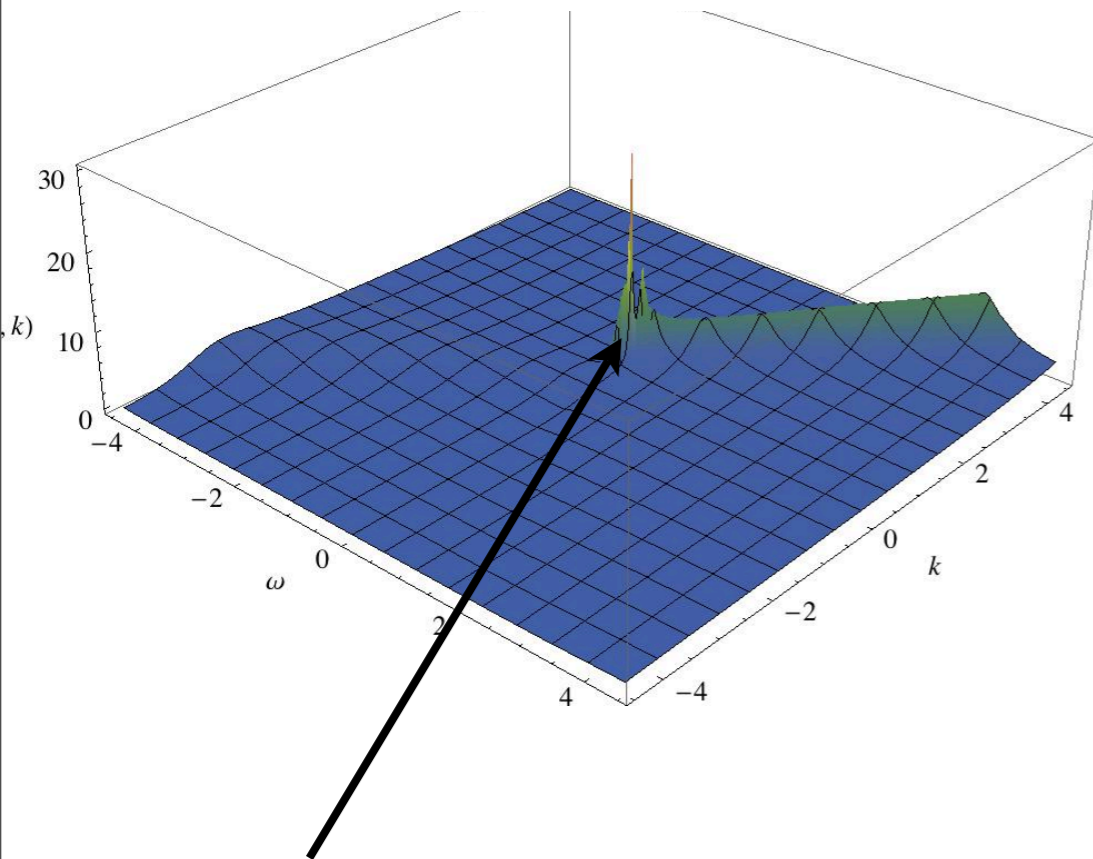


Fermi
surface
peak

How is the spectrum modified?

P

$P=0$

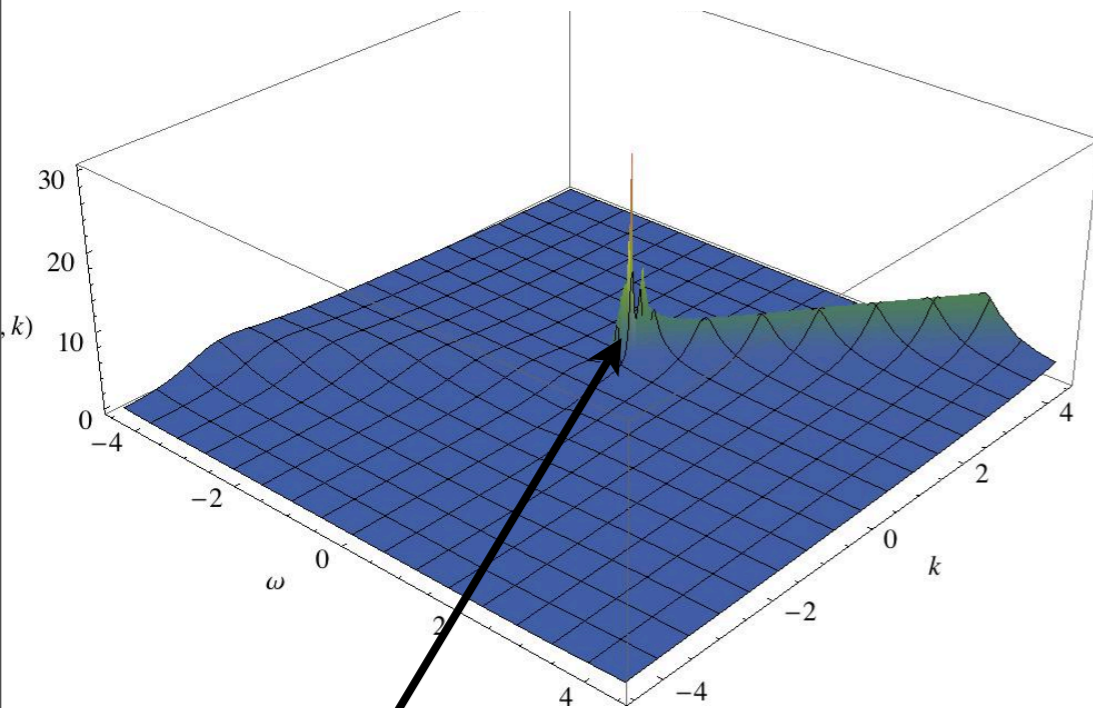


Fermi
surface
peak

How is the spectrum modified?

P

$P=0$



Fermi
surface
peak

$$-1.54 < p < -0.53$$

$$1 > \nu_{k_F} > 1/2$$

$$\Re \omega \propto k - k_F$$

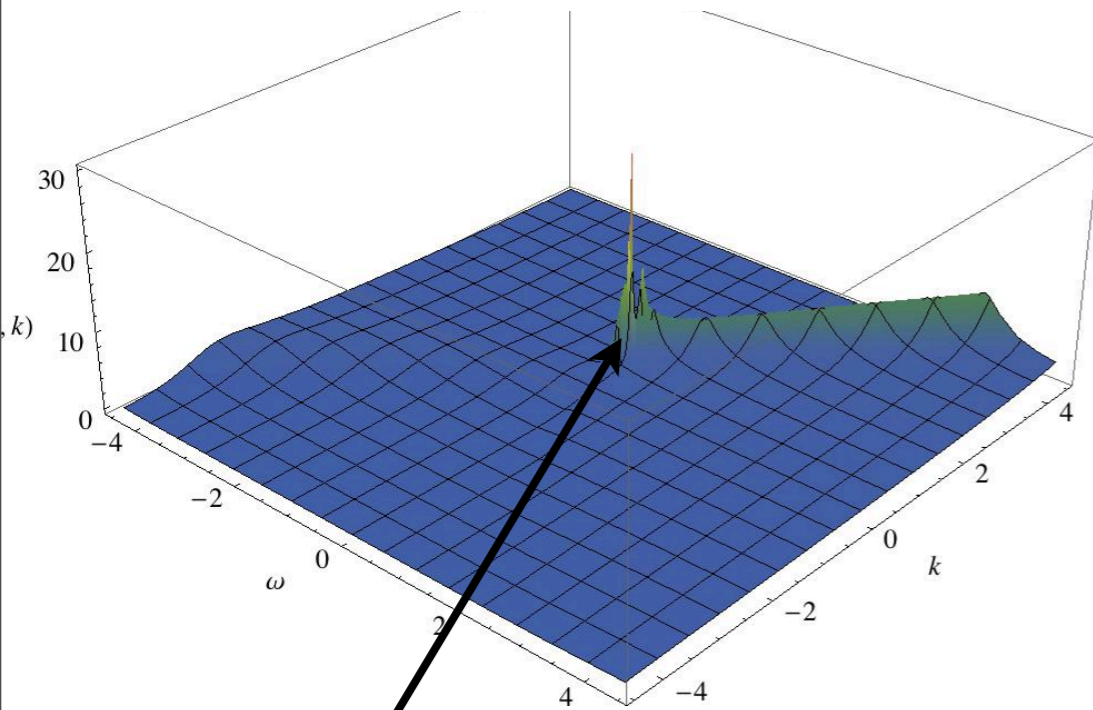
$$\Im \omega \propto (k - k_F)^{2\nu_{k_F}}$$

'Fermi Liquid'

How is the spectrum modified?

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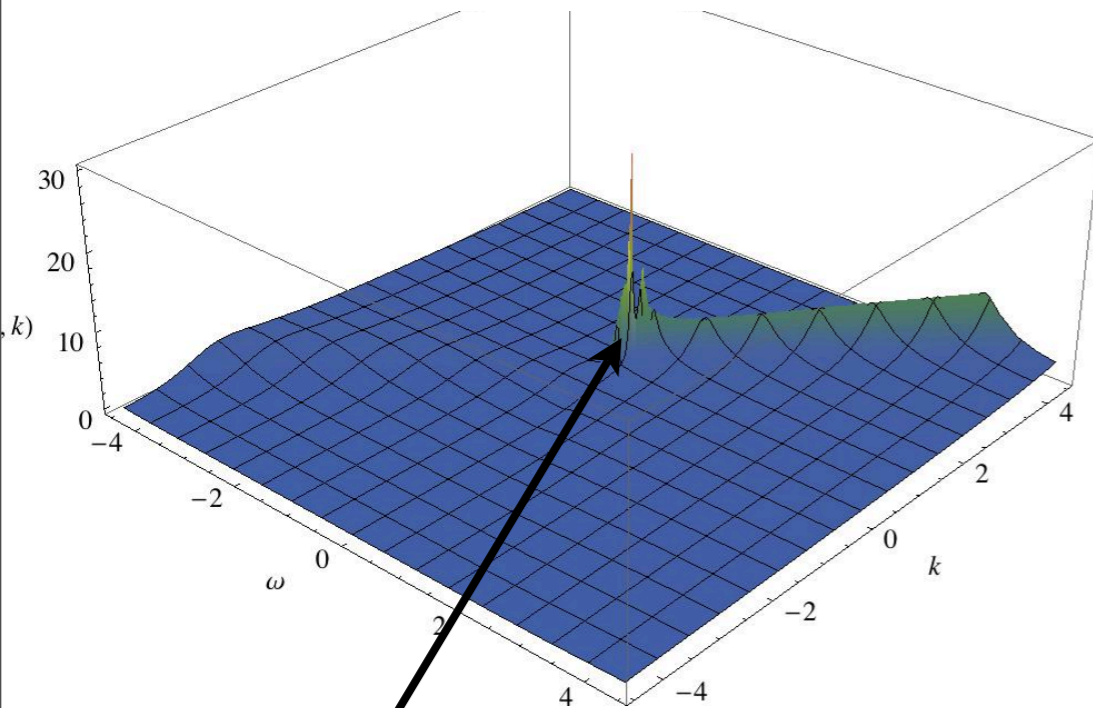


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How is the spectrum modified?

P

$P=0$



$$p = -0.53$$

$$\nu_{k_F} = 1/2$$

MFL

$$-0.53 < p < 1/\sqrt{6}$$

$$1/2 > \nu_{k_F} > 0$$

$$\Re\omega = \Im\omega \propto (k - k_F)^{1/(2\nu_{k_F})}$$

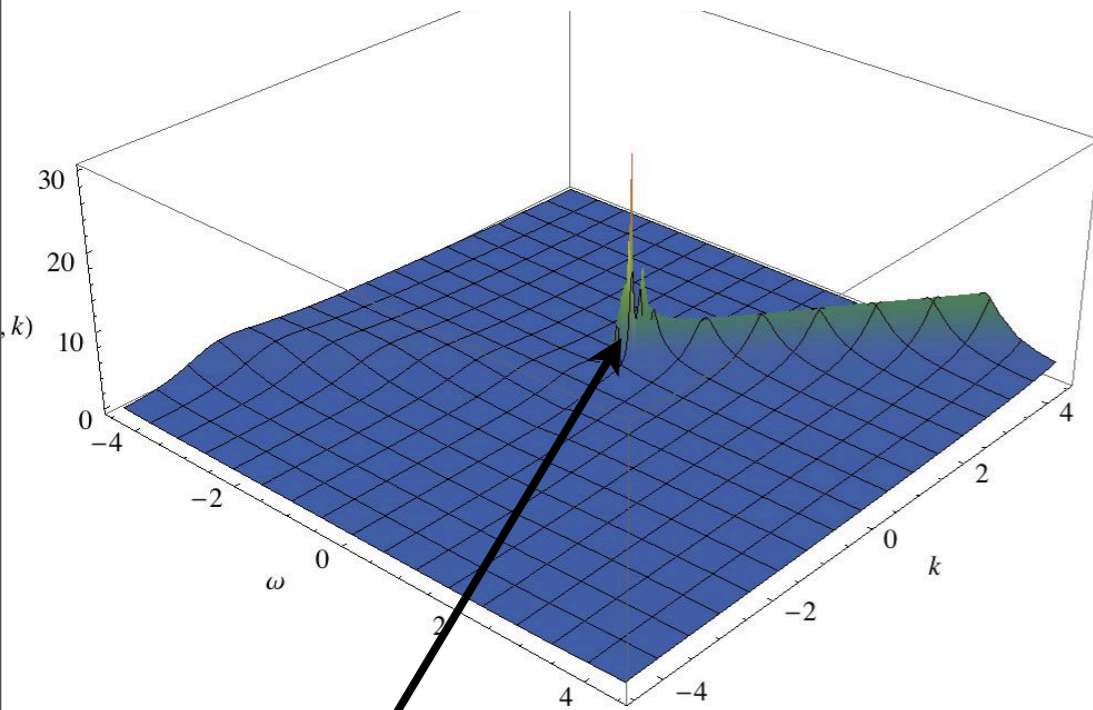
NFL

Fermi surface peak

How is the spectrum modified?

P

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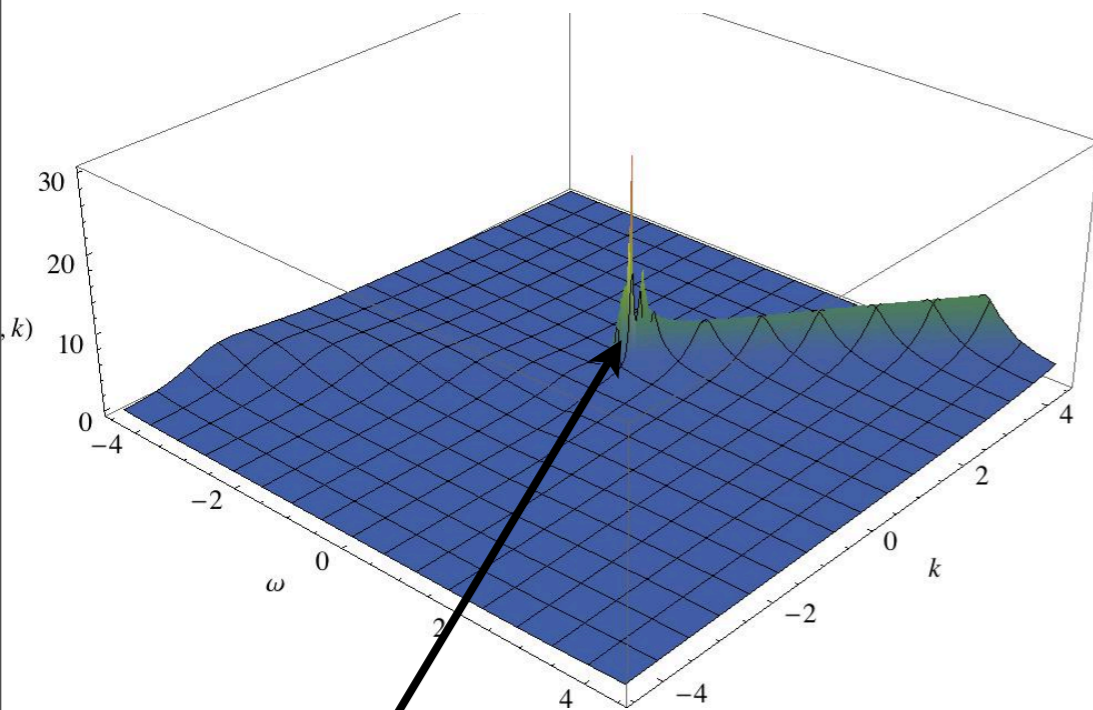


Fermi
surface
peak

How is the spectrum modified?

$P > 4.2$ P

$P=0$

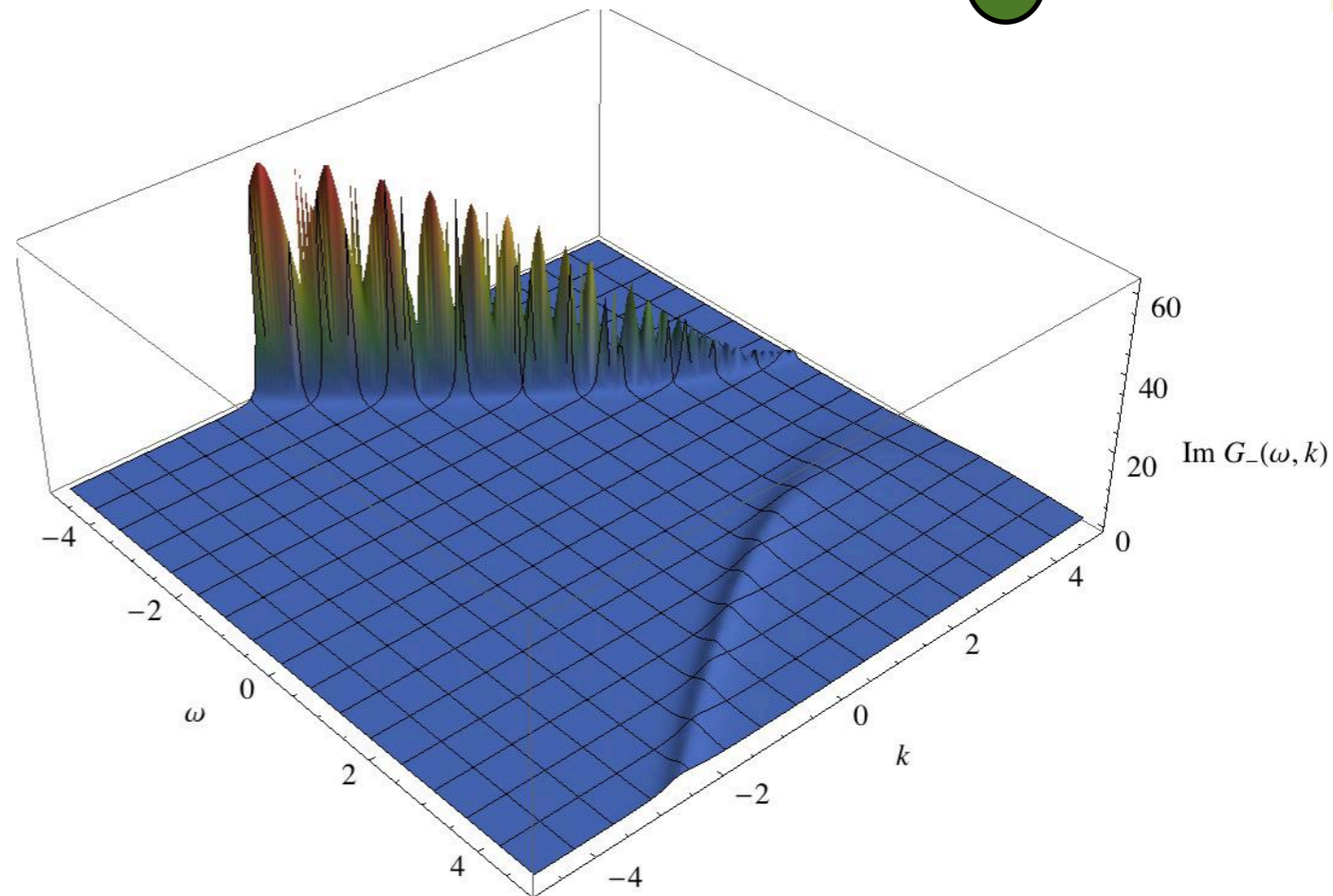
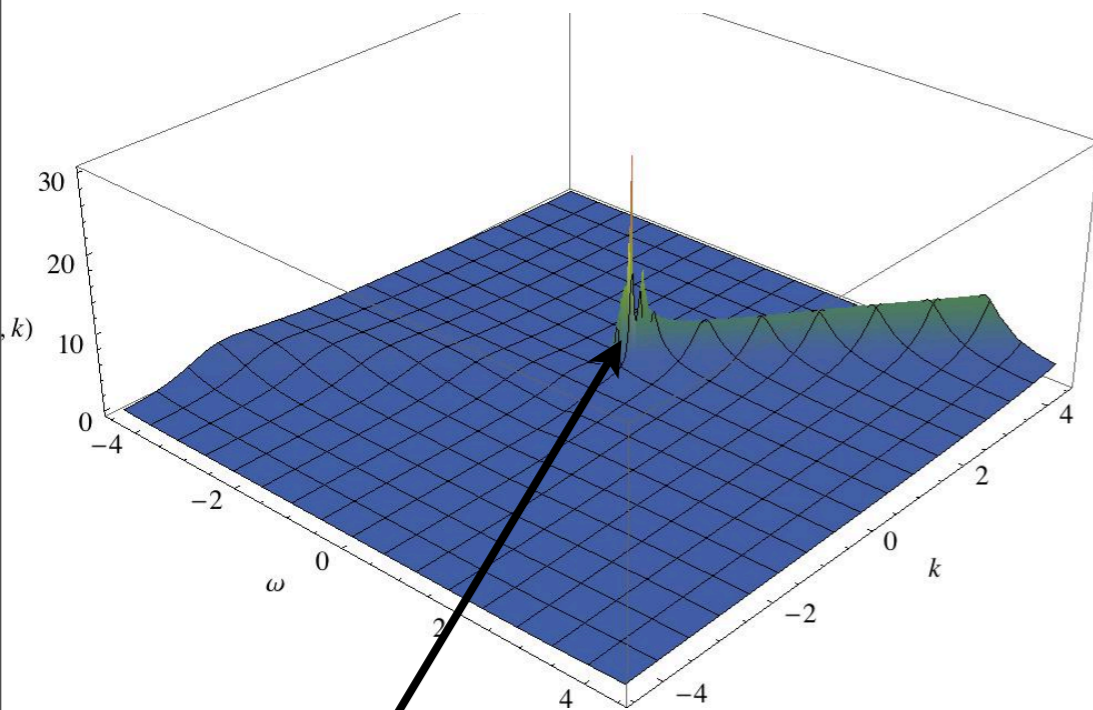


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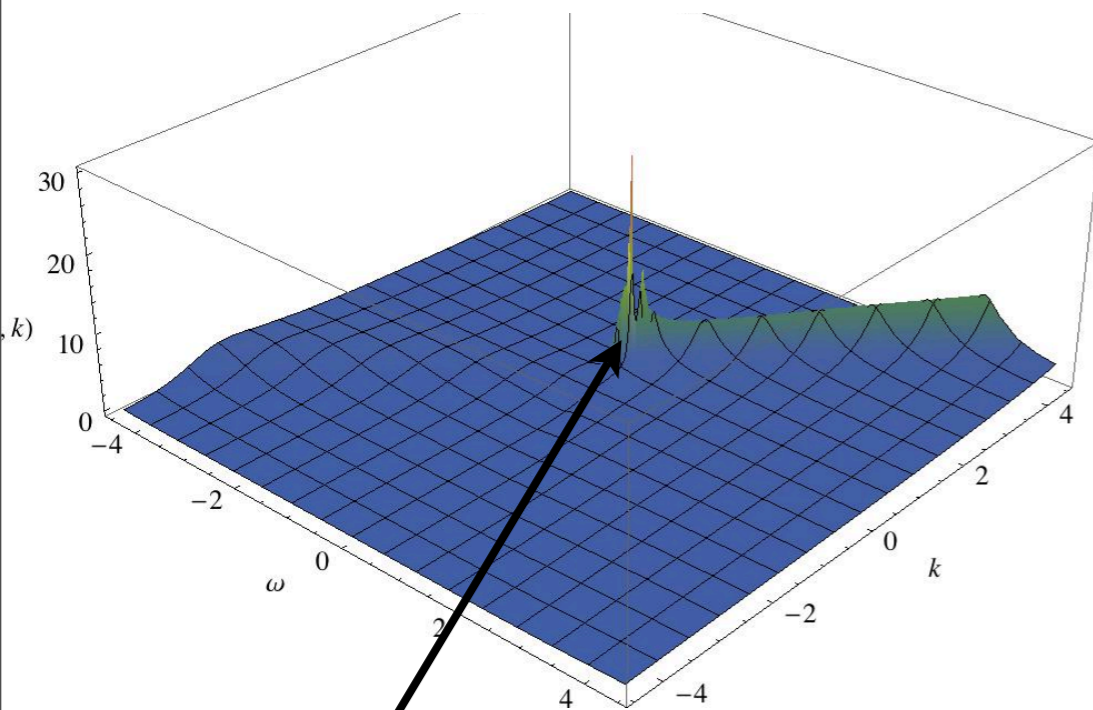


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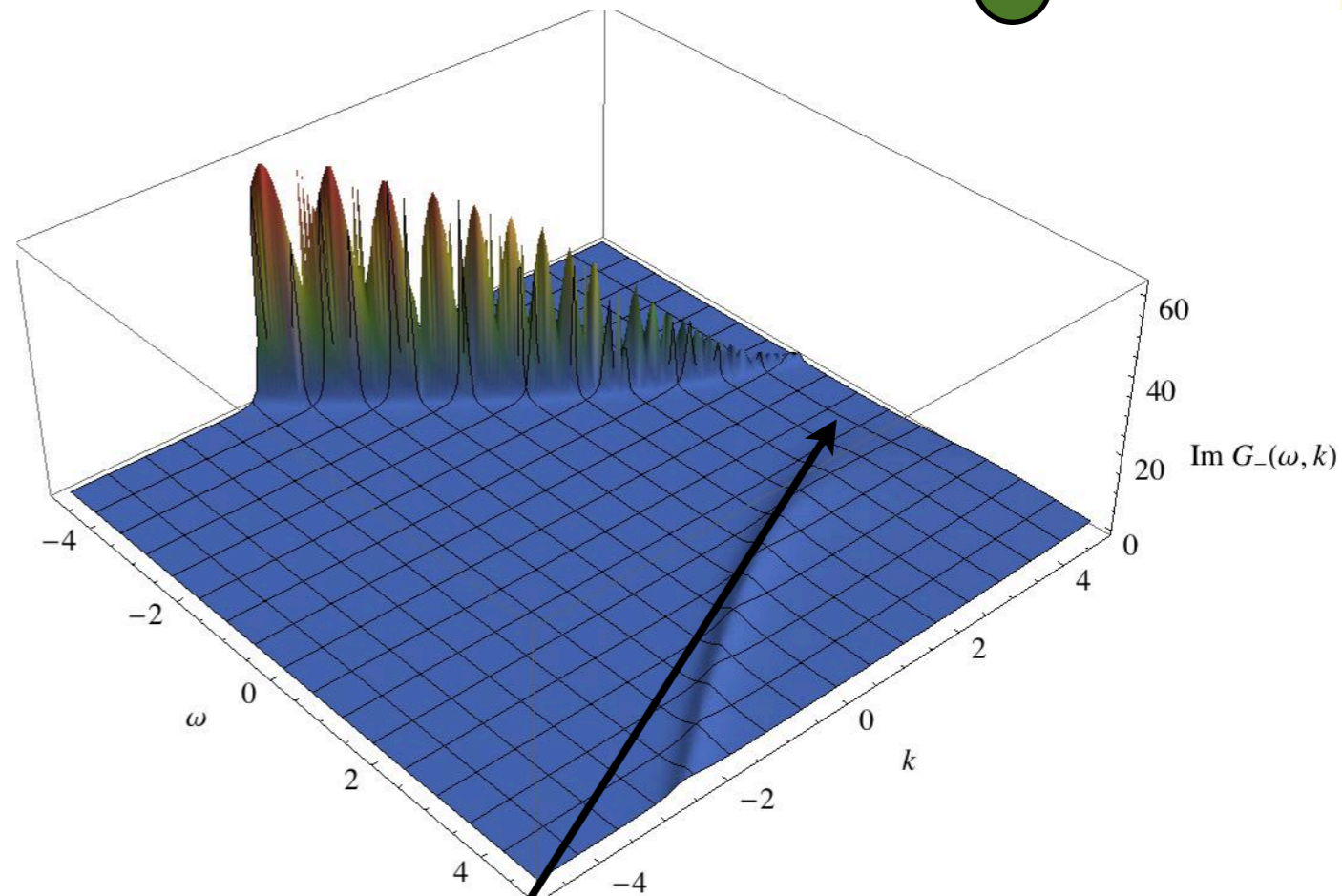
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surface
peak

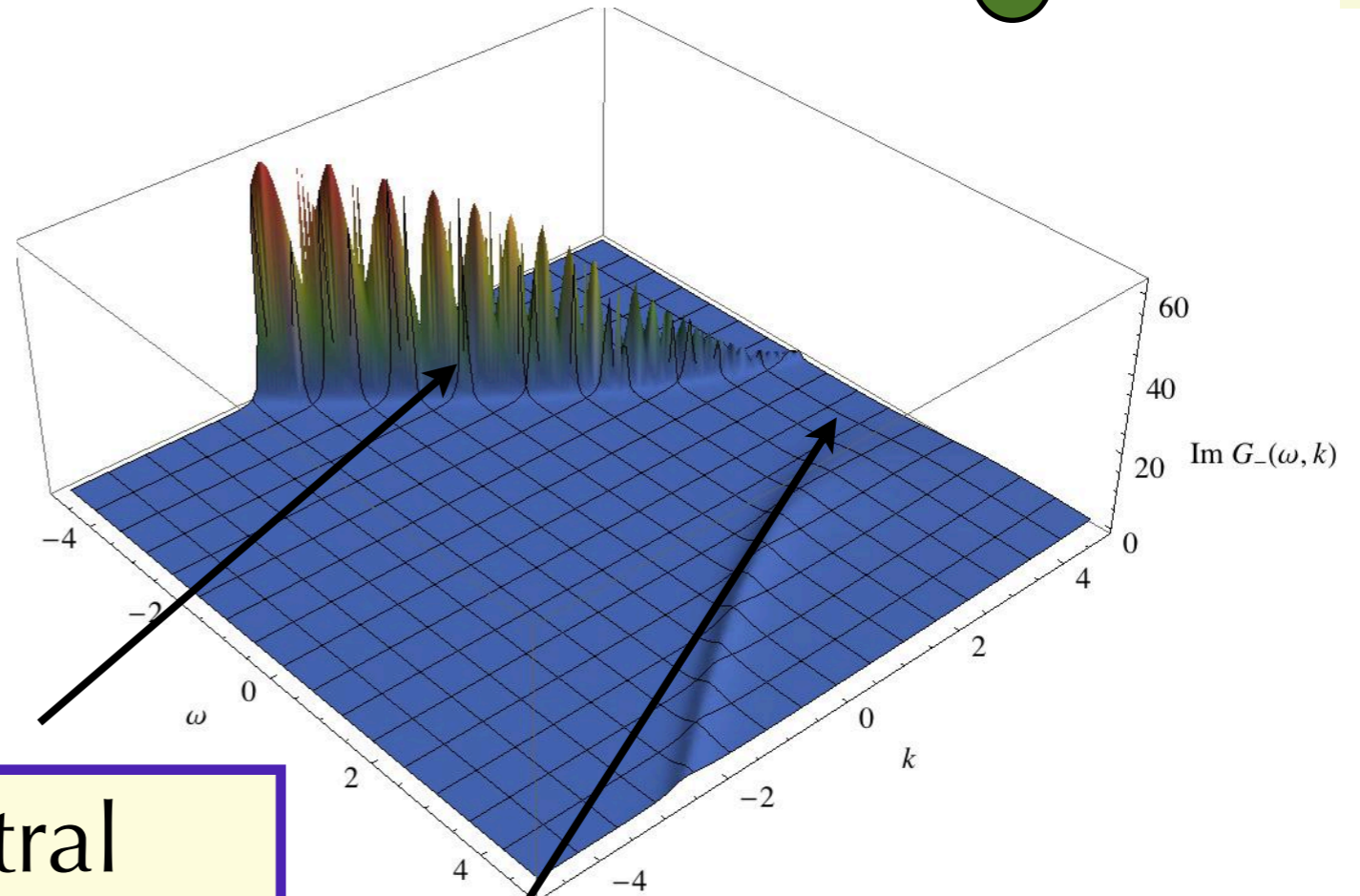
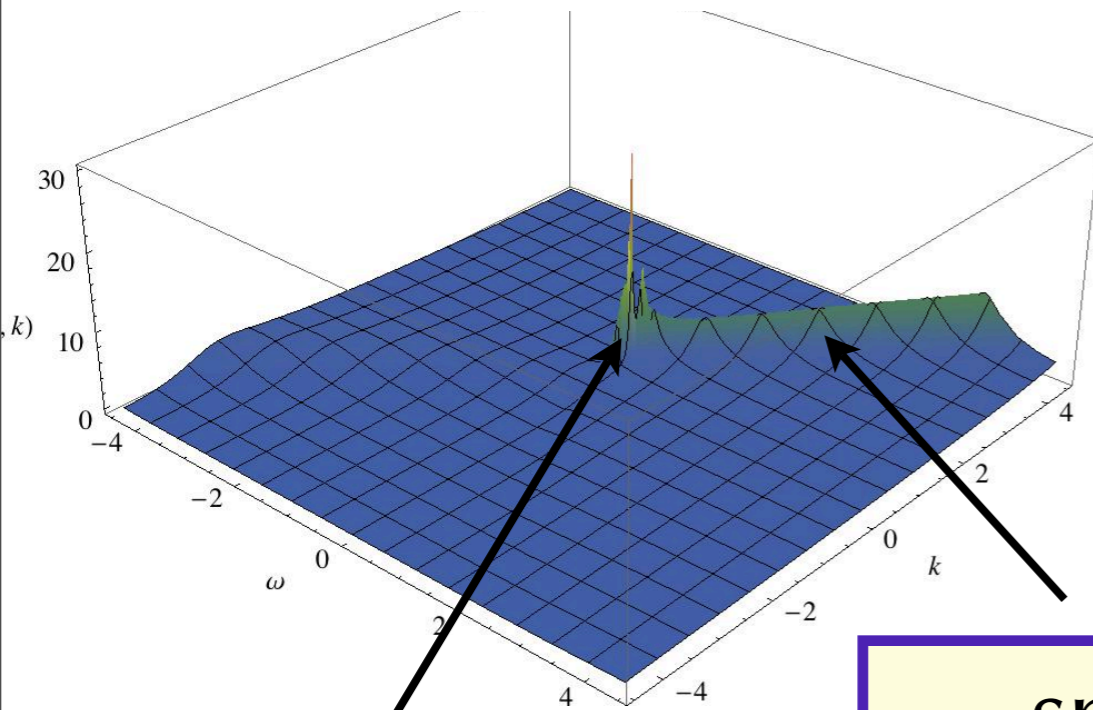


dynamically generated gap:

How is the spectrum modified?

$P > 4.2$ **P**

$P=0$



spectral weight transfer

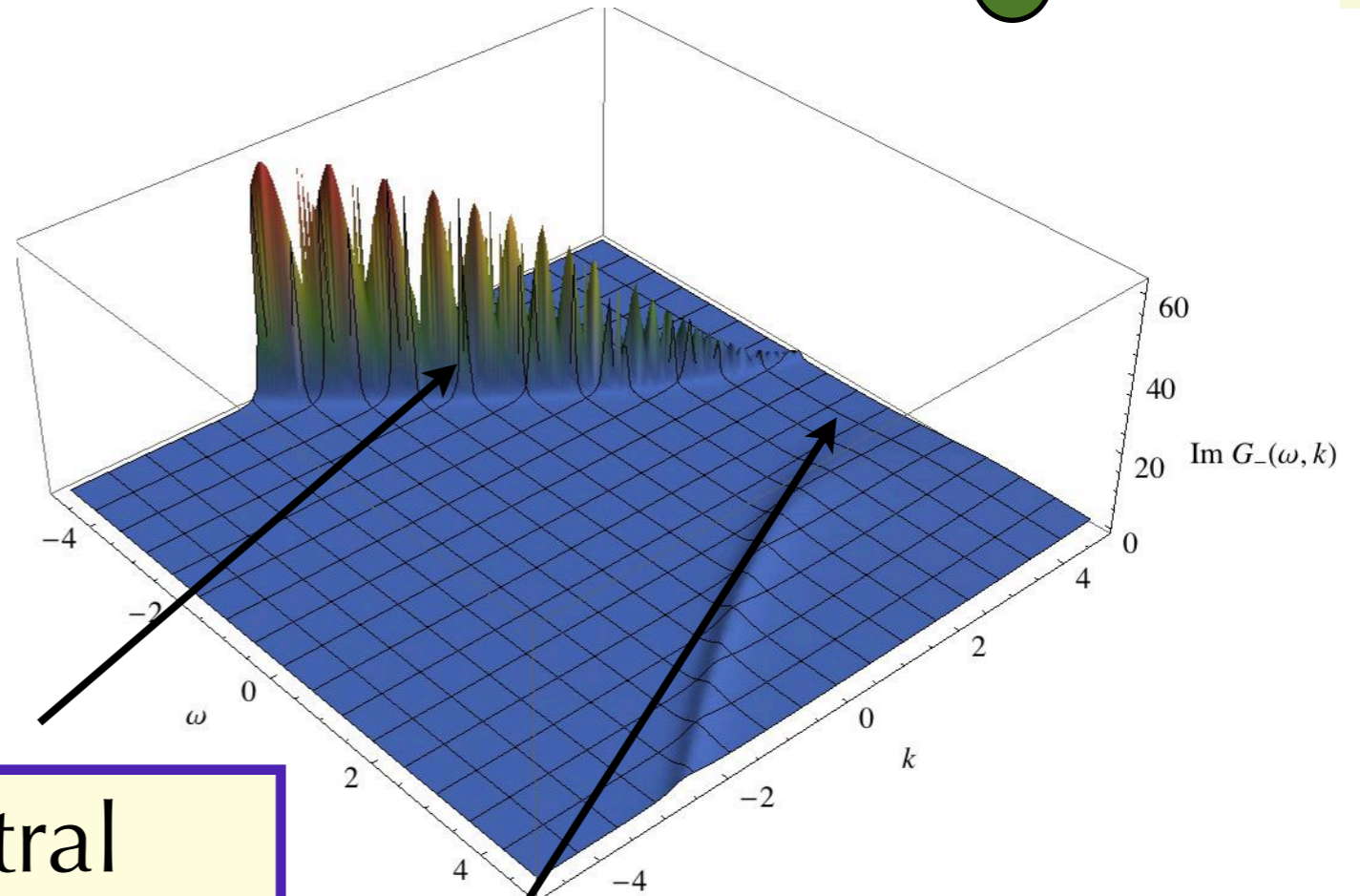
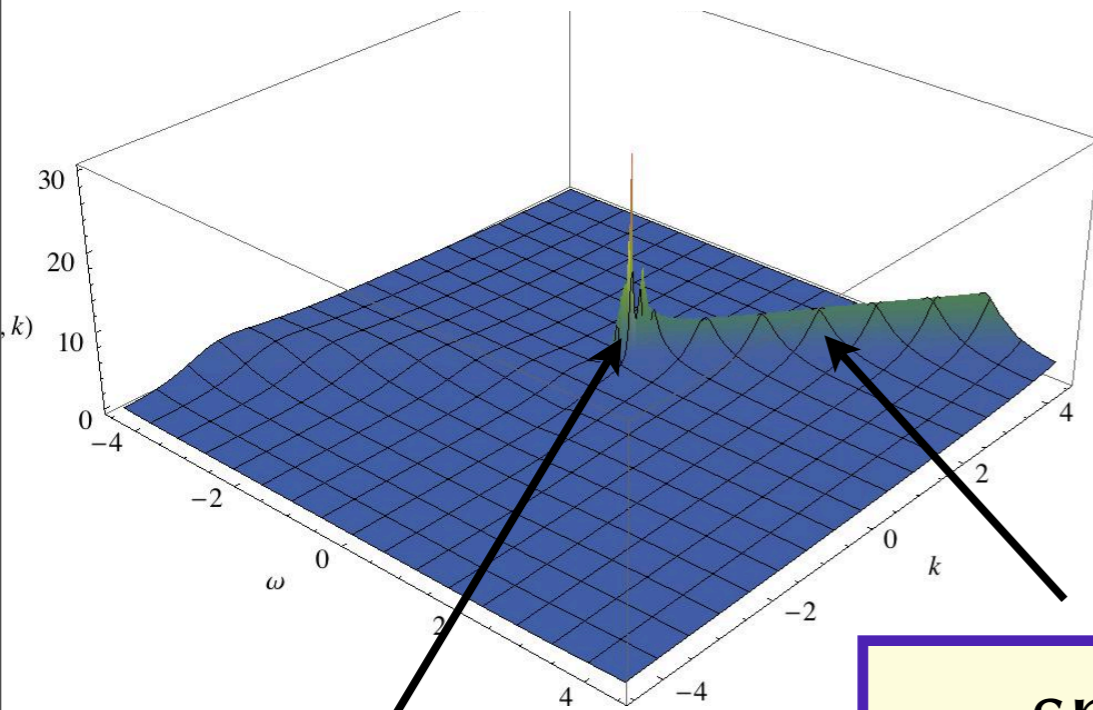
Fermi surface peak

dynamically generated gap:

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$P > 4.2$ **P**

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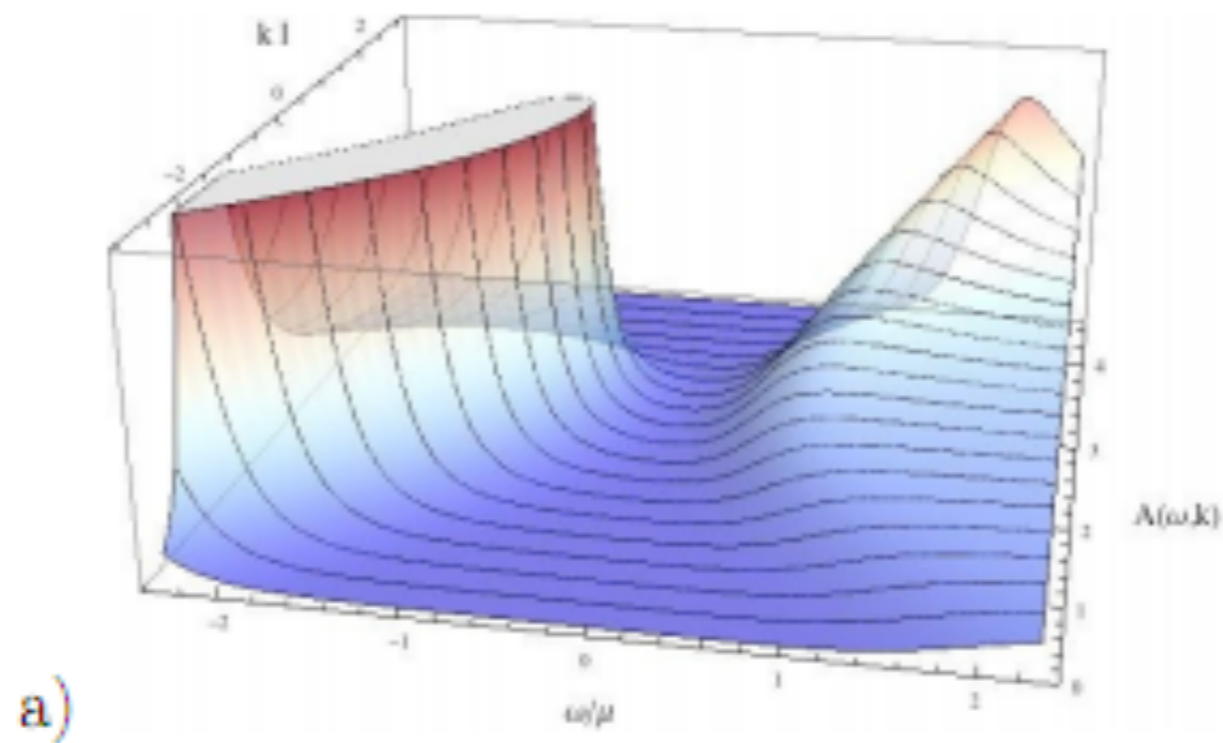


spectral weight transfer

Fermi surface peak

dynamically generated gap:

Gubser, Gauntlett, 2011 similar results



Sonner, 2011 (top-down gravitino model)

Mechanism?

$$\psi \propto ar^\Delta + br^{-\Delta}$$

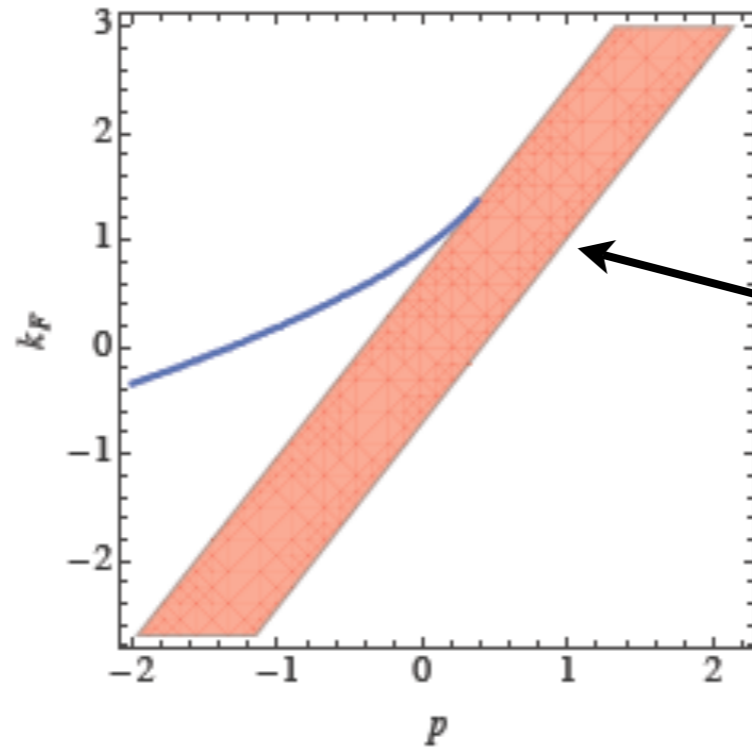
IR

UV

QFT

operators \mathcal{O}

where is k_F



log-oscillatory

k_F moves into log-oscillatory region: IR \mathcal{O}_\pm acquires a complex dimension

near horizon

radial Dirac Equation

$$\psi_{I\pm}(\zeta) = \psi_{I\pm}^{(0)}(\zeta) + \omega \psi_{I\pm}^{(1)}(\zeta) + \omega^2 \psi_{I\pm}^{(2)}(\zeta) + \dots$$

$$-\psi_{I\pm}^{(0)''}(\zeta) = i\sigma_2 \left(1 + \frac{qe_d}{\zeta}\right) - \frac{L_2}{\zeta} \left[m\sigma_3 + \left(pe_d \pm \frac{kL}{r_0} \right) \sigma_1 \right] \psi_{I\pm}^{(0)}(\zeta),$$

$$e_d = 1/\sqrt{2d(d-1)}$$

$$m_k^2 = m^2 + \left(pe_d \pm \frac{kL}{r_0} \right)^2$$

p: time-reversal breaking mass term (in bulk)

$$\nu_k^\pm = \sqrt{m_{k\pm}^2 L_2^2 - q^2 e_d^2 - i\epsilon, \delta} \quad \pm = \nu_k^\pm + 1/2$$

near horizon

radial Dirac Equation

$$\psi_{I\pm}(\zeta) = \psi_{I\pm}^{(0)}(\zeta) + \omega \psi_{I\pm}^{(1)}(\zeta) + \omega^2 \psi_{I\pm}^{(2)}(\zeta) + \dots$$

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$$\nu_k^\pm = \sqrt{m_{k\pm}^2 L_2^2 - q^2 e_d^2 - i\epsilon, \delta} \quad \pm = \nu_k^\pm + 1/2$$

scaling dimension is complex!

What does a complex scaling dimension mean?

continuous
scale invariance

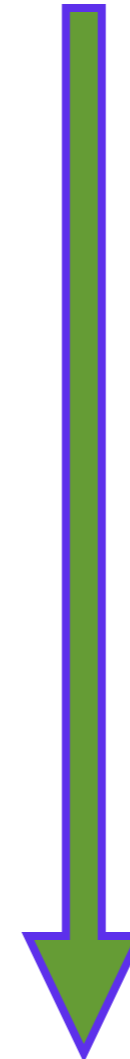


$$\mathcal{O} = \mu(\lambda)\mathcal{O}(\lambda r)$$



$$1 = \mu(\lambda)\lambda^{\Delta}$$

$$\Delta = -\frac{\ln \mu}{\ln \lambda}$$



Δ is real,
independent of
scale

what about complex Δ

$$1 = \mu \lambda^{\Delta}$$

$$e^{2\pi in} = 1 = \mu\lambda^\Delta$$

Discrete scale invariance (DSI)

$$e^{2\pi i n} = 1 = \mu \lambda^{\Delta}$$

n=0:CSI

$$\Delta = -\frac{\ln \mu}{\ln \lambda} + \frac{2\pi i n}{\ln \lambda}$$

Discrete scale invariance (DSI)

$$e^{2\pi i n} = 1 = \mu \lambda^{\Delta}$$

n=0:CSI

$$\Delta = -\frac{\ln \mu}{\ln \lambda} + \frac{2\pi i n}{\ln \lambda}$$

scaling dimension depends on scale

$$\lambda_n = \lambda^n$$

magnification

example



example

example



example



n iterations

length

$$3^{-n}$$

number of segments

$$2^n$$

example



n iterations

length

$$3^{-n}$$

number of segments

$$2^n$$

scale invariance only
for $\lambda_p = 3^p$

example



n iterations

length

$$3^{-n}$$

number of segments

$$2^n$$

scale invariance only
for $\lambda_p = 3^p$

discrete scale invariance:

$$D = -\frac{\ln 2}{\ln 3} + \frac{2\pi i n}{\ln 3}$$

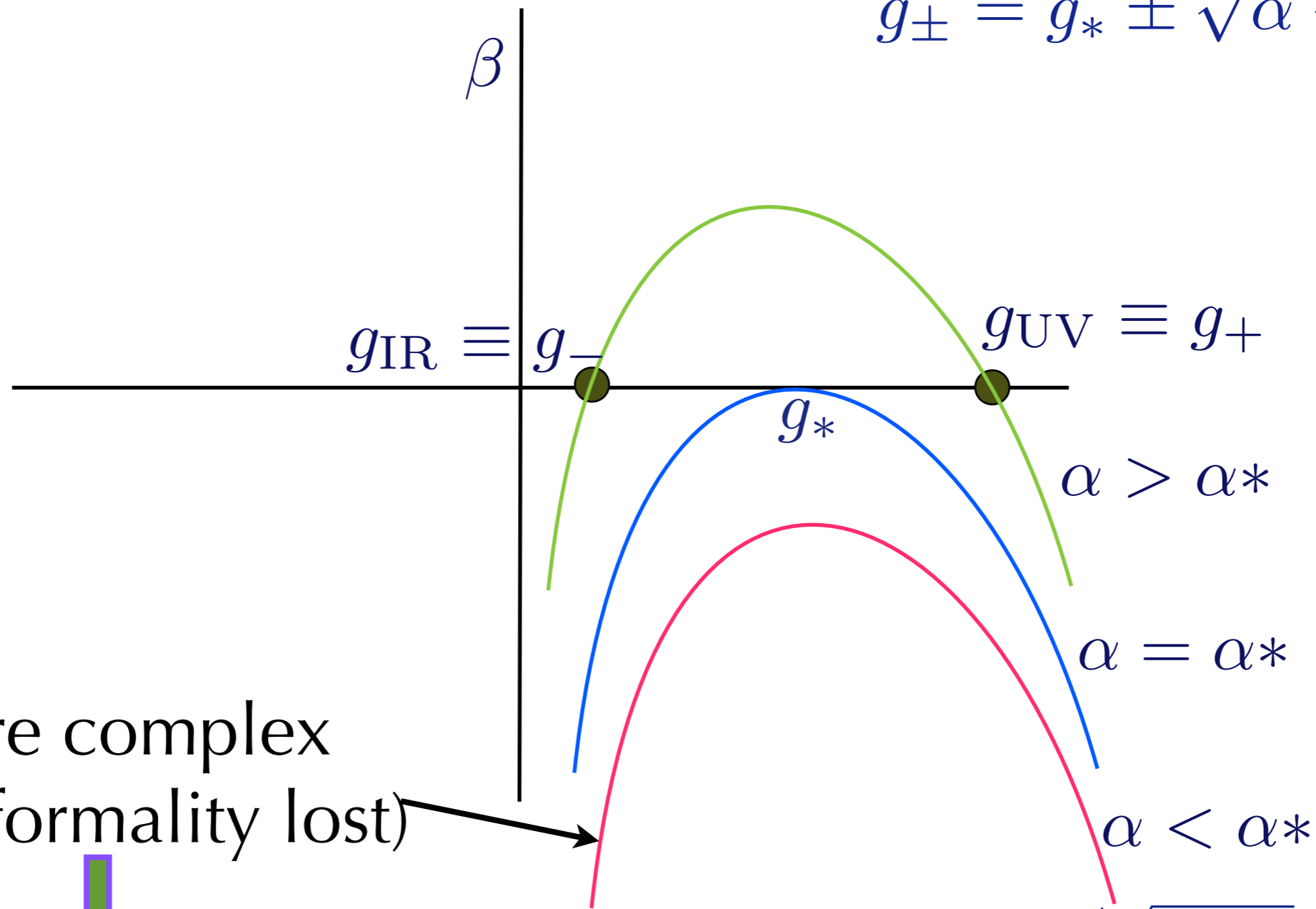
discrete scale invariance

hidden scale
(length, energy,...)

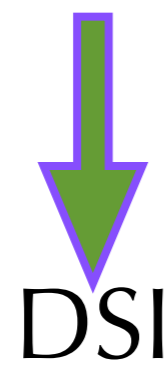
toy model: merging of UV and IR fixed points

$$\beta = (\alpha - \alpha_*) - (g - g_*)^2$$

$$g_{\pm} = g_* \pm \sqrt{\alpha - \alpha_*}$$



g_{\pm} are complex
(conformality lost)



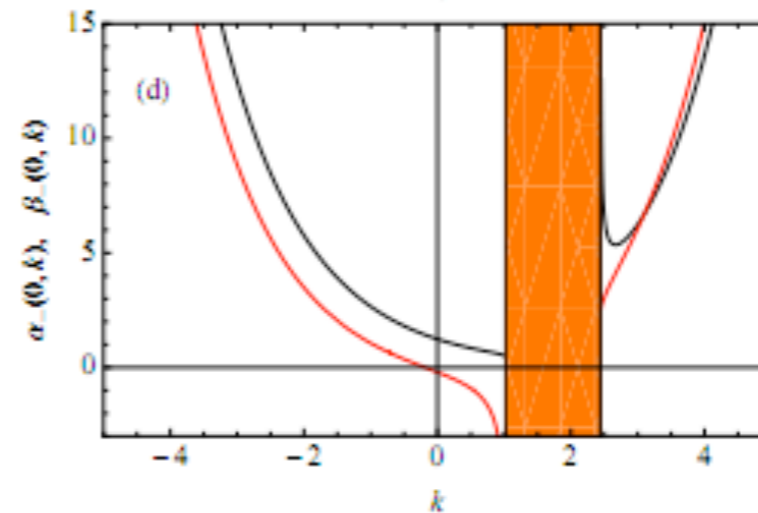
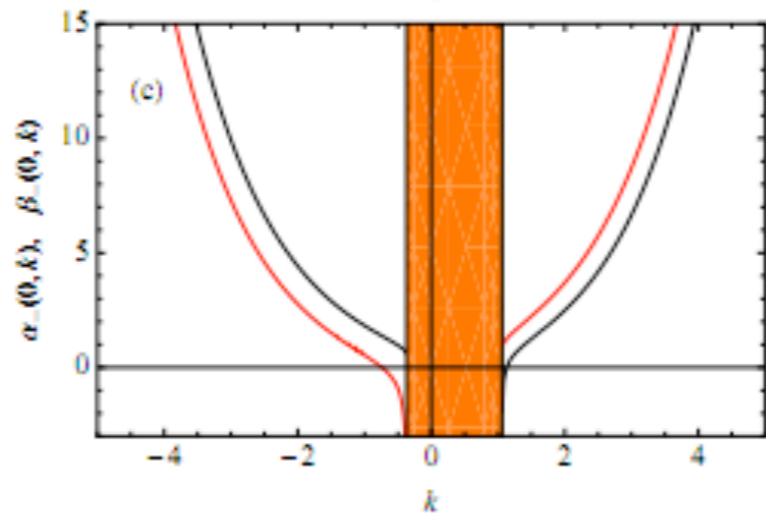
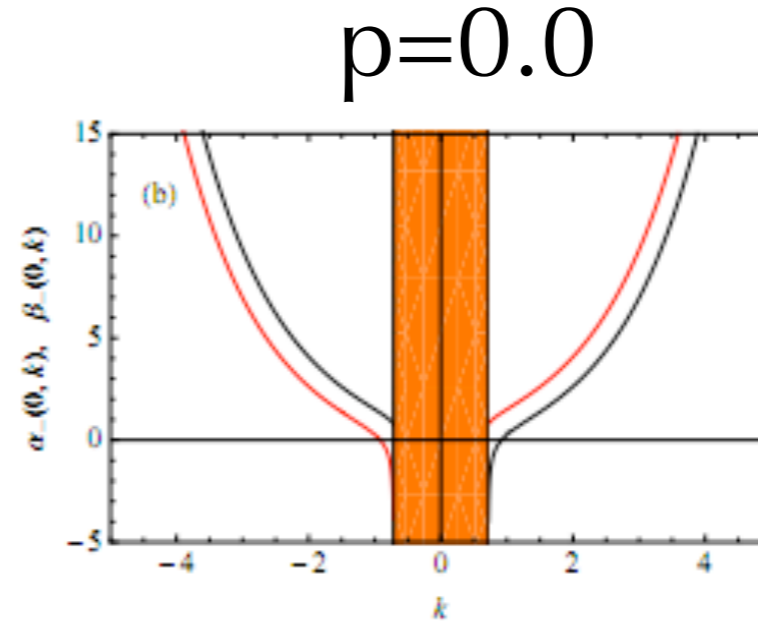
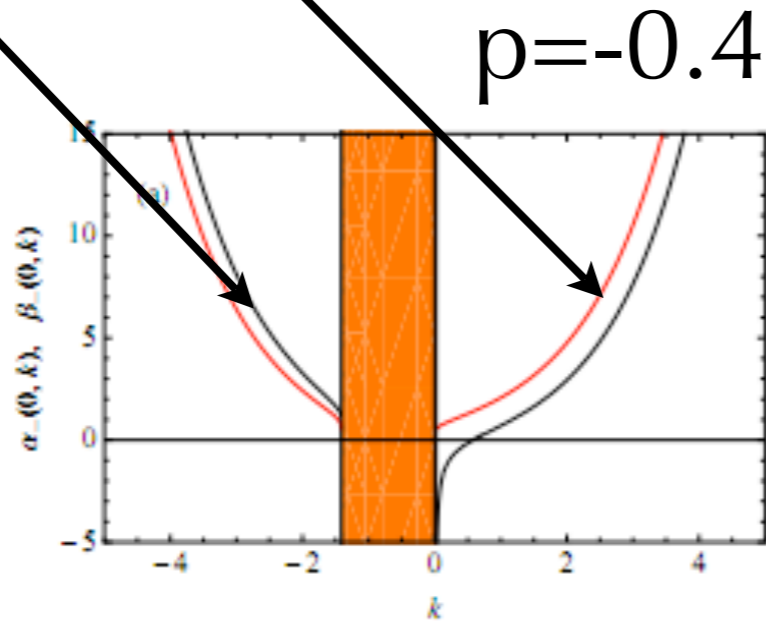
DSI

$$\Lambda_{\text{IR}} = \Lambda_{\text{UV}} e^{-\pi/\sqrt{\alpha_* - \alpha}}$$

BKT transition

Kaplan, arxiv:0905.475

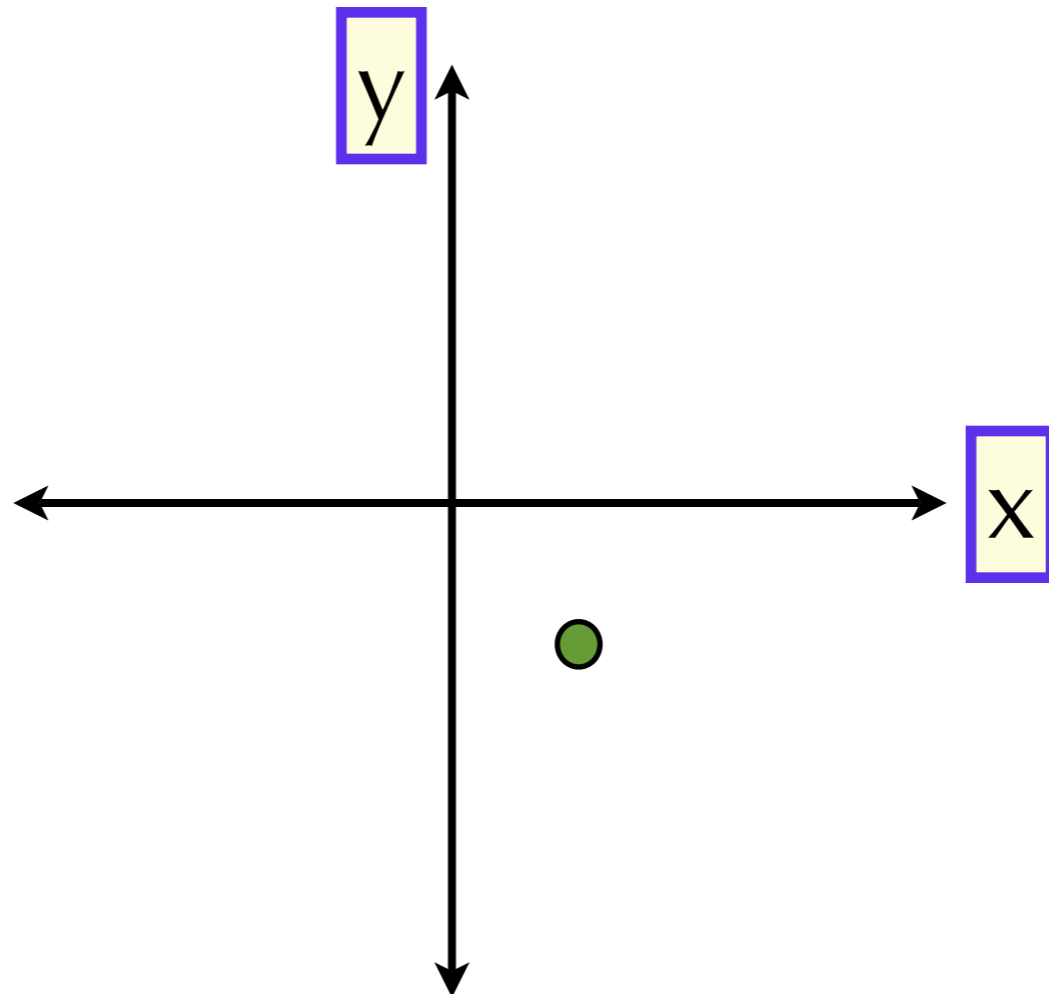
$$G = \frac{\beta_-(0, k)}{\alpha_-(0, k)}$$



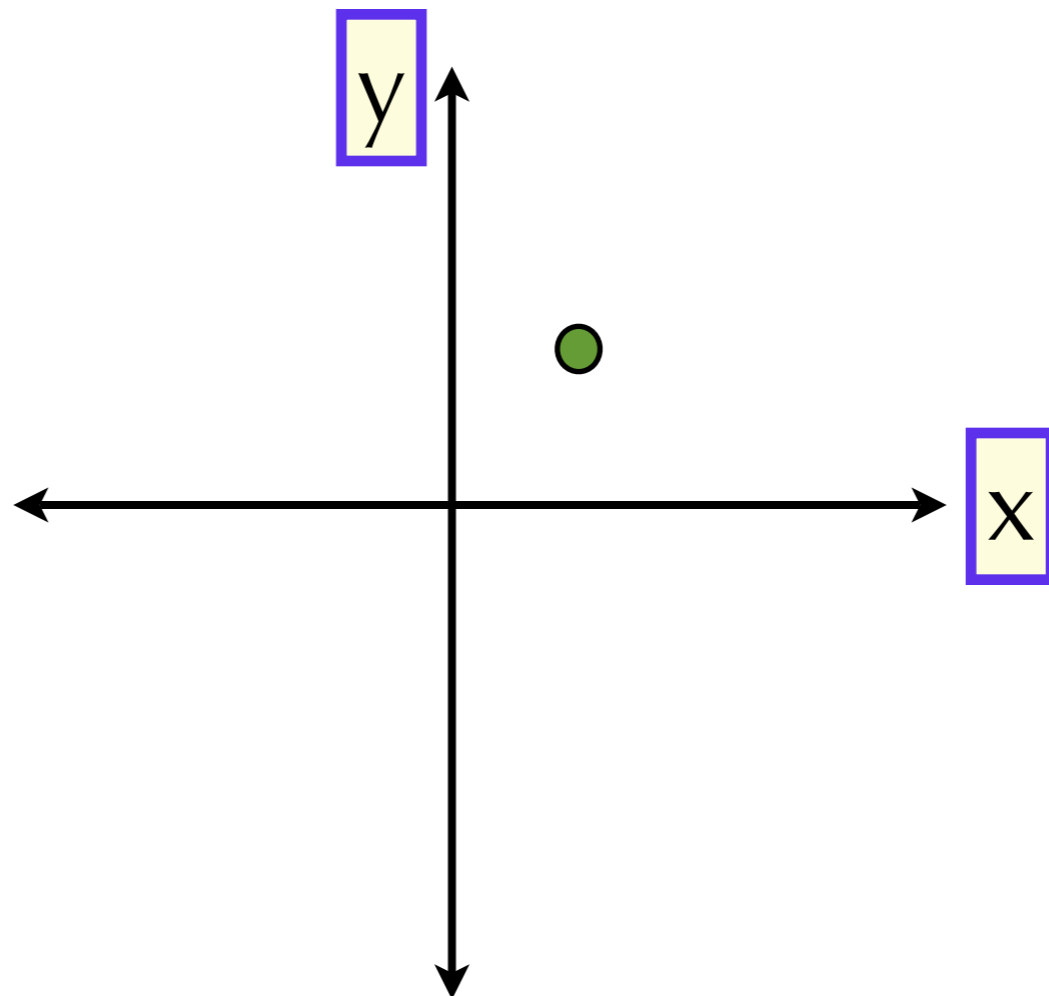
no poles outside log-oscillatory region for $p > 1/\sqrt{6}$

Mott physics and DSI are linked!

is there an instability? (violation of BF bound)

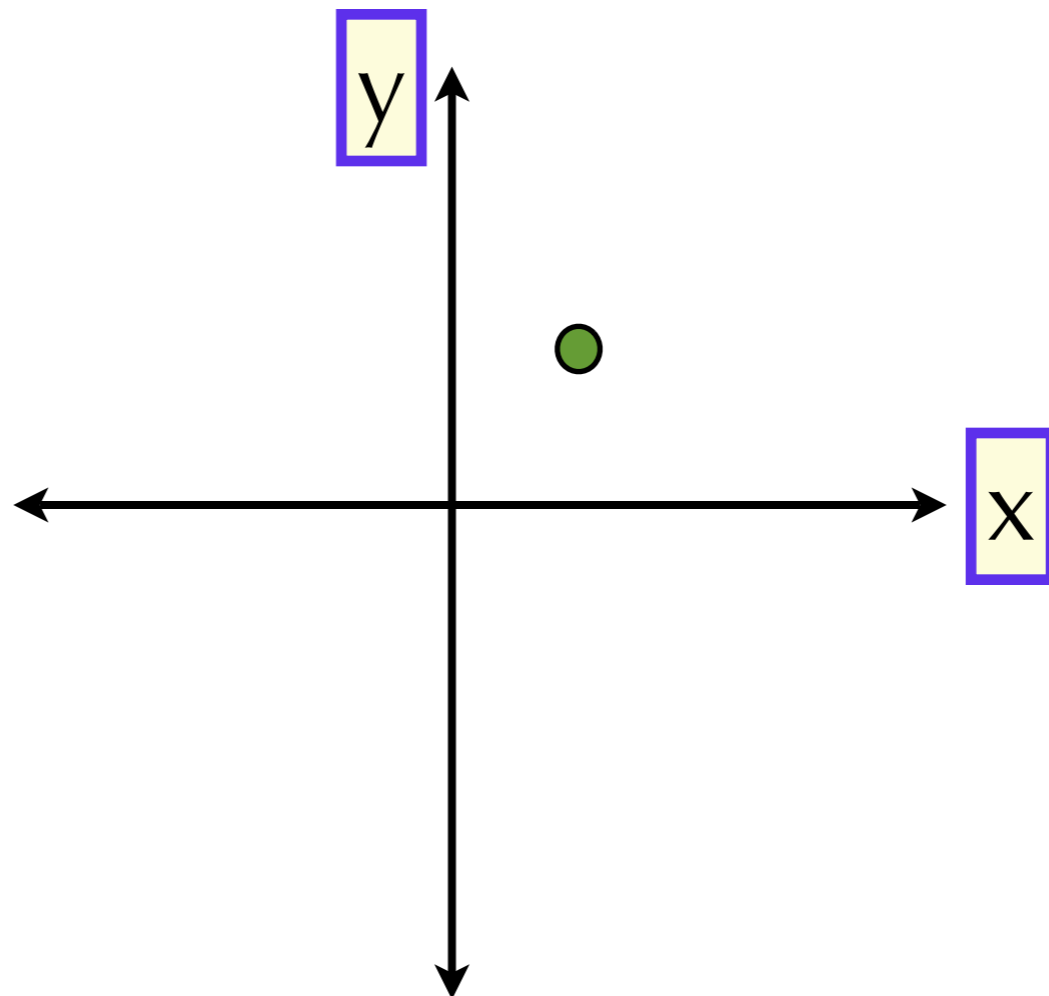


is there an instability? (violation of BF bound)



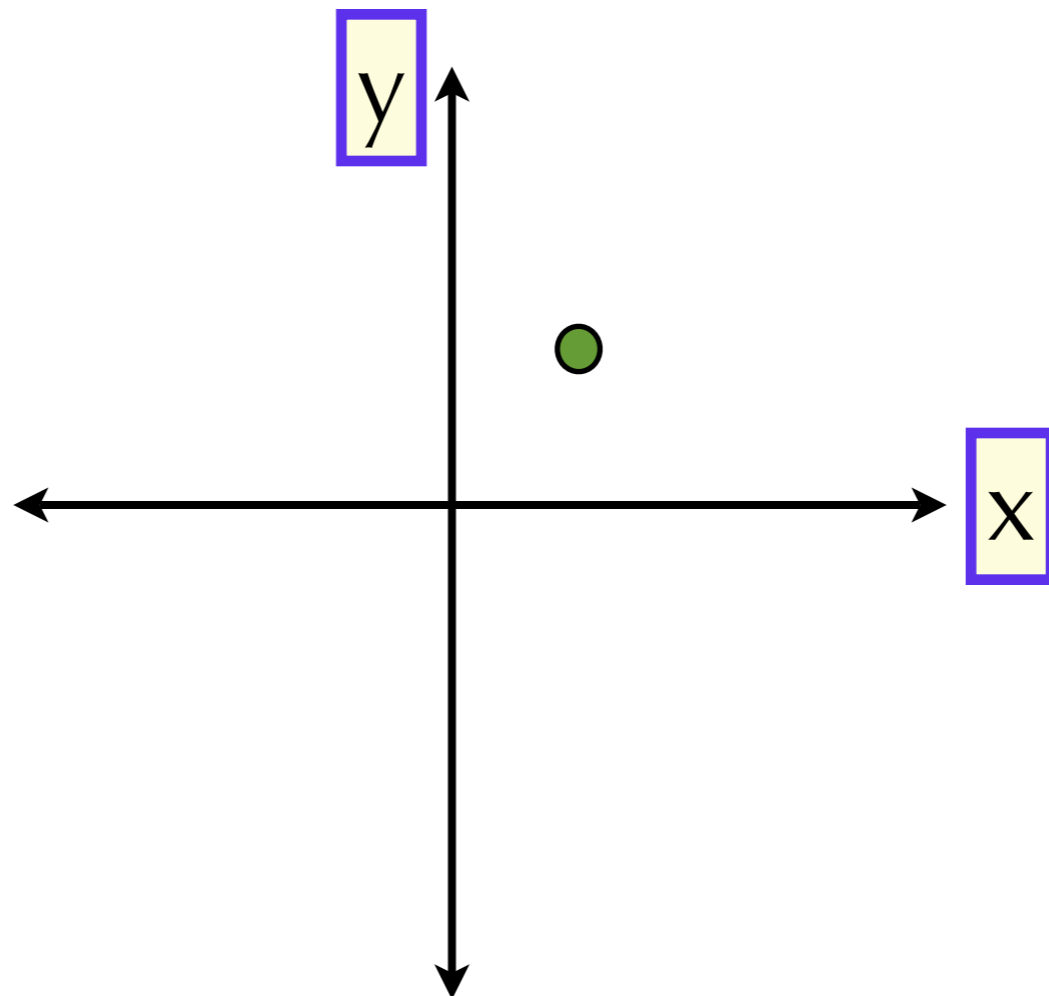
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$$E \rightarrow E + i\gamma \quad \gamma > 0$$



is there an instability? (violation of BF bound)

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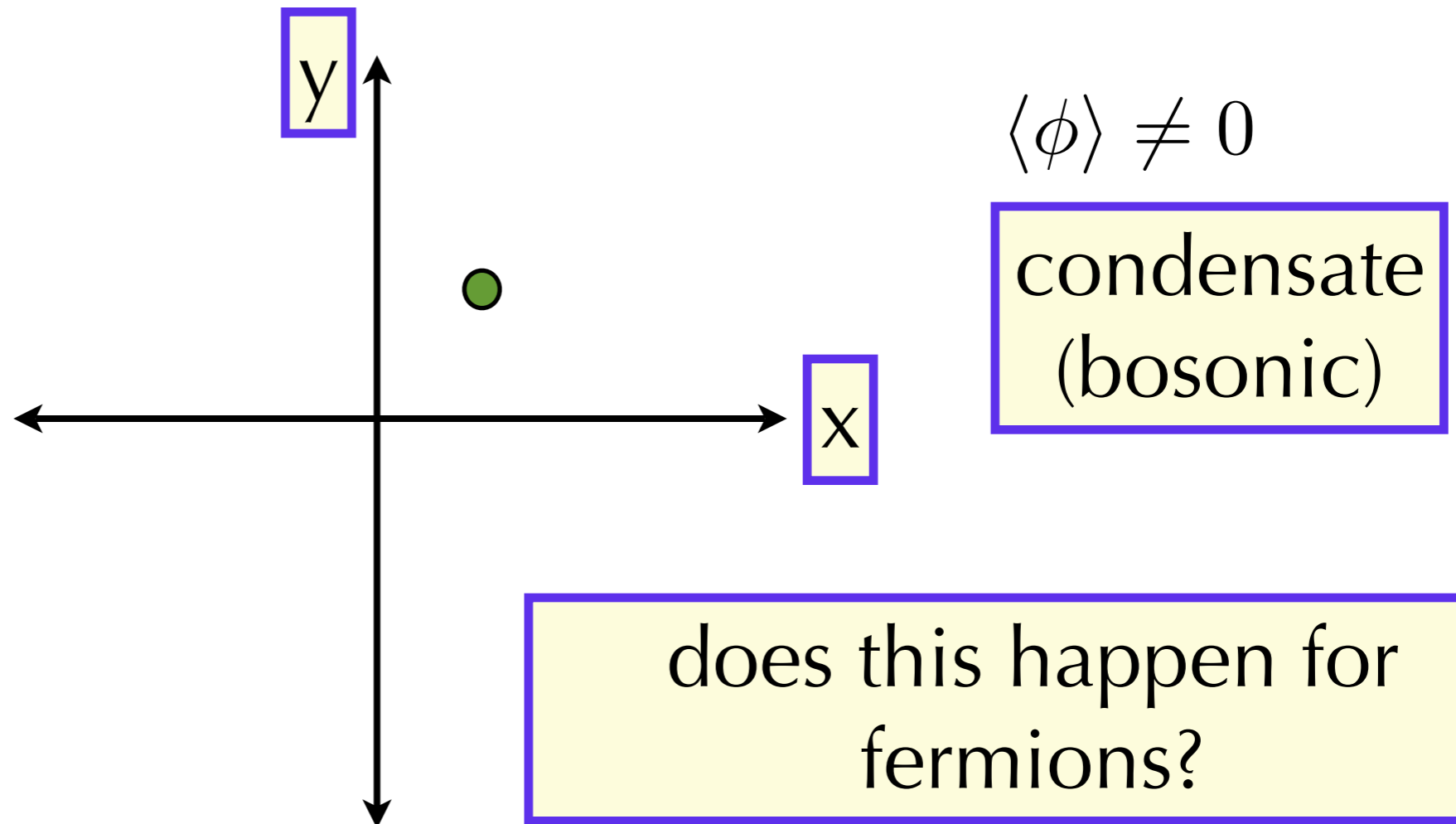


$$\langle \phi \rangle \neq 0$$

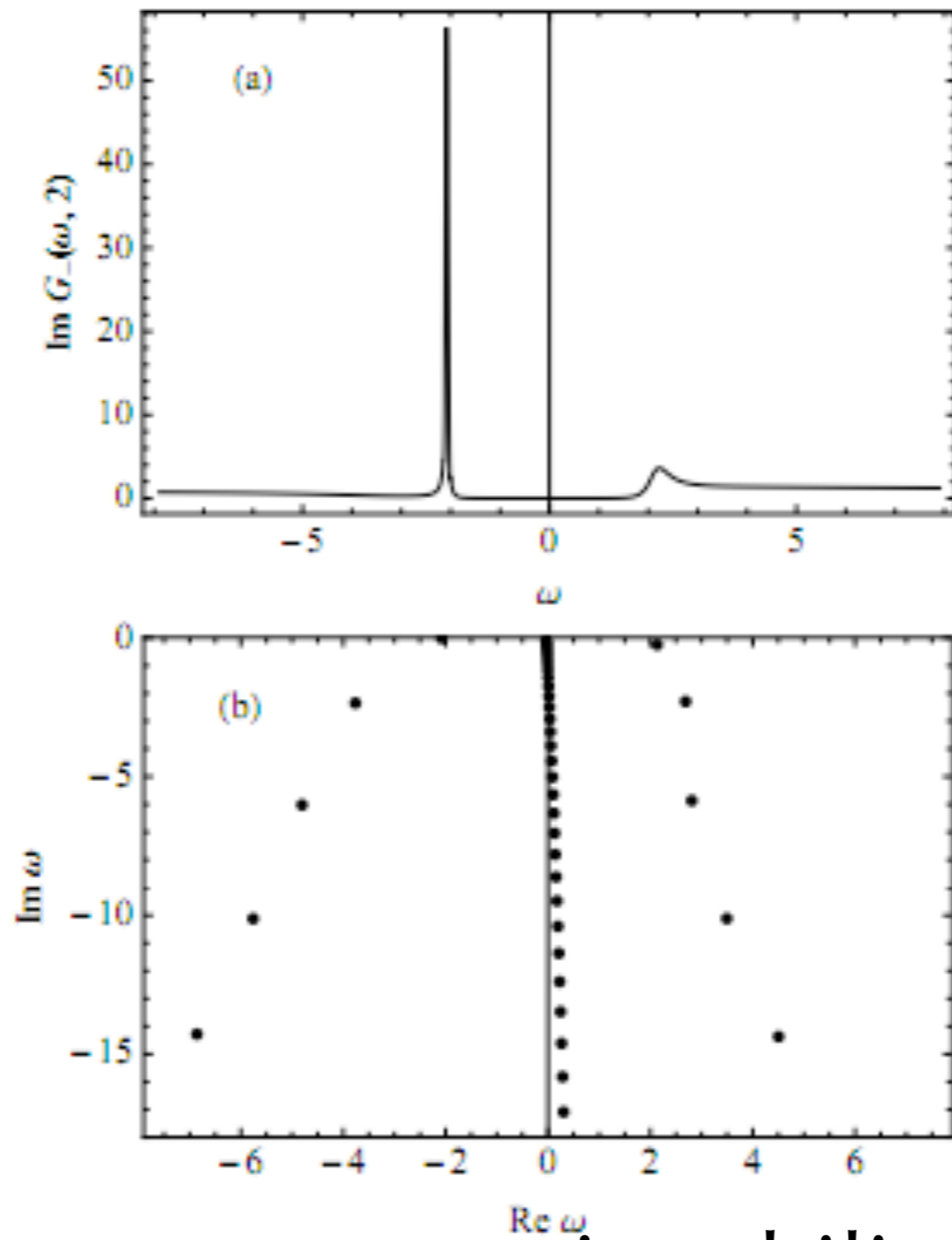
condensate
(bosonic)

is there an instability? (violation of BF bound)

$$E \rightarrow E + i\gamma \quad \gamma > 0$$

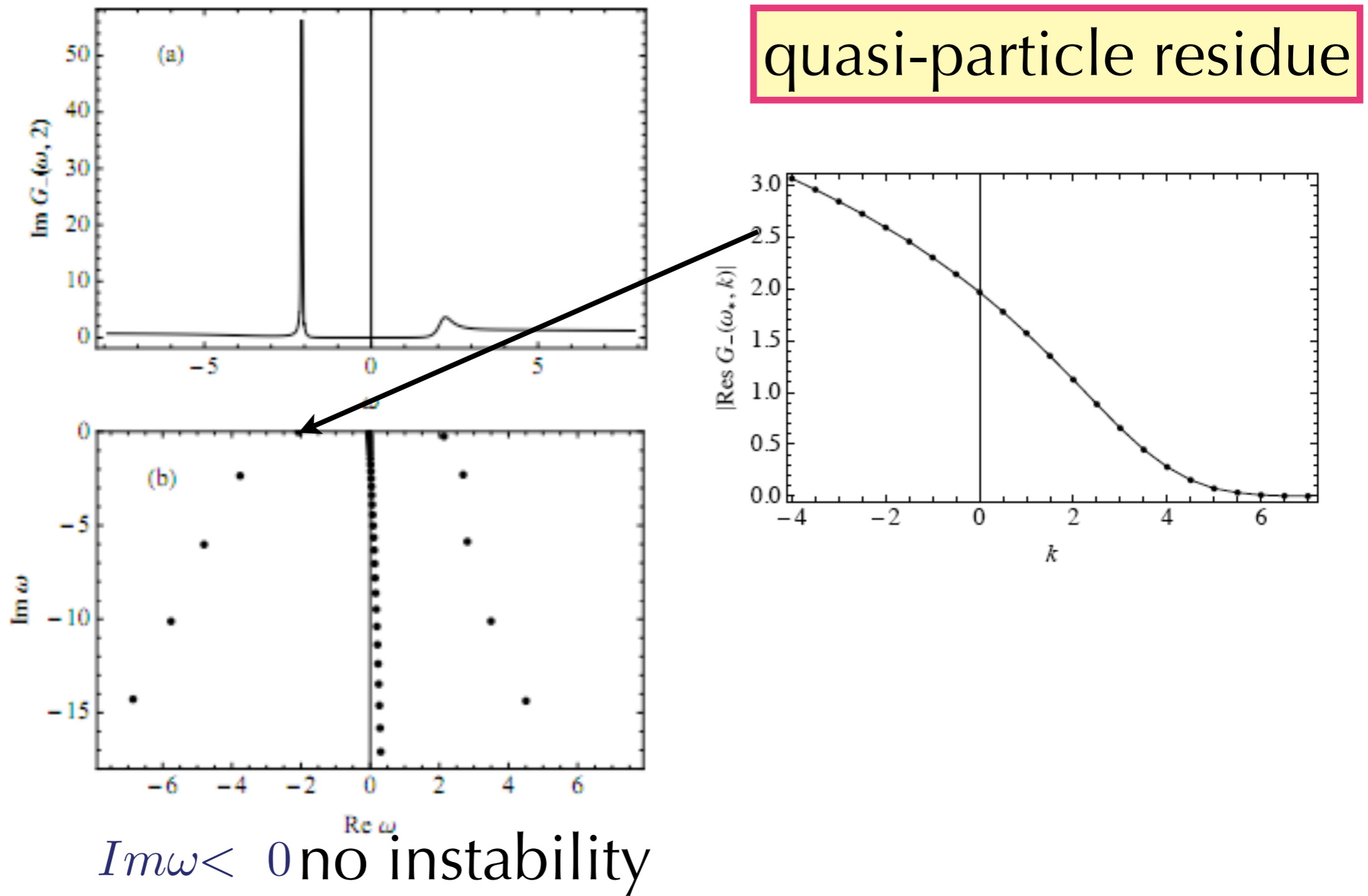


quasi-normal modes



$\text{Im } \omega < 0$ no instability

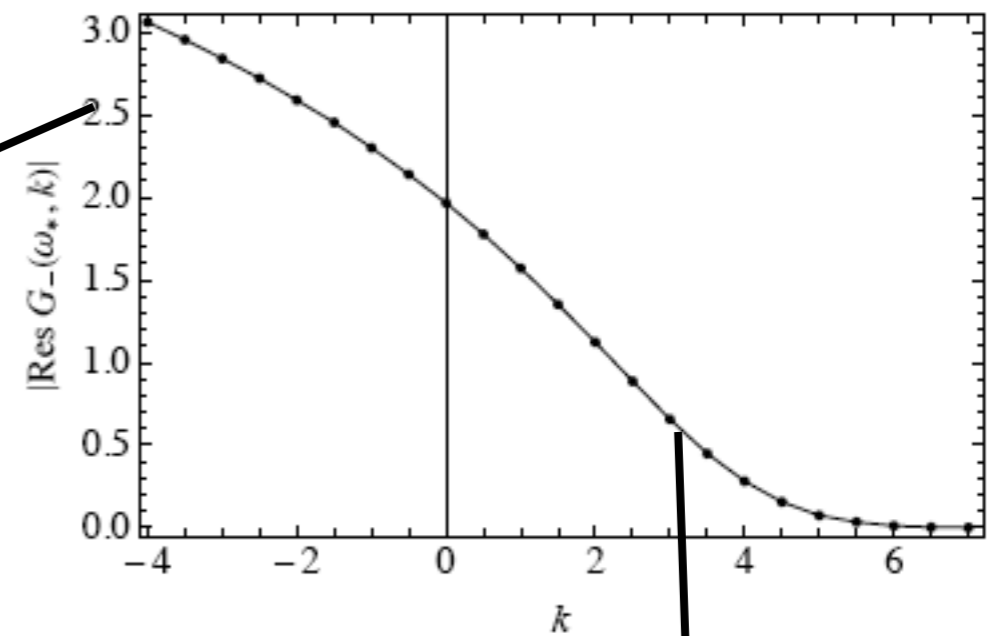
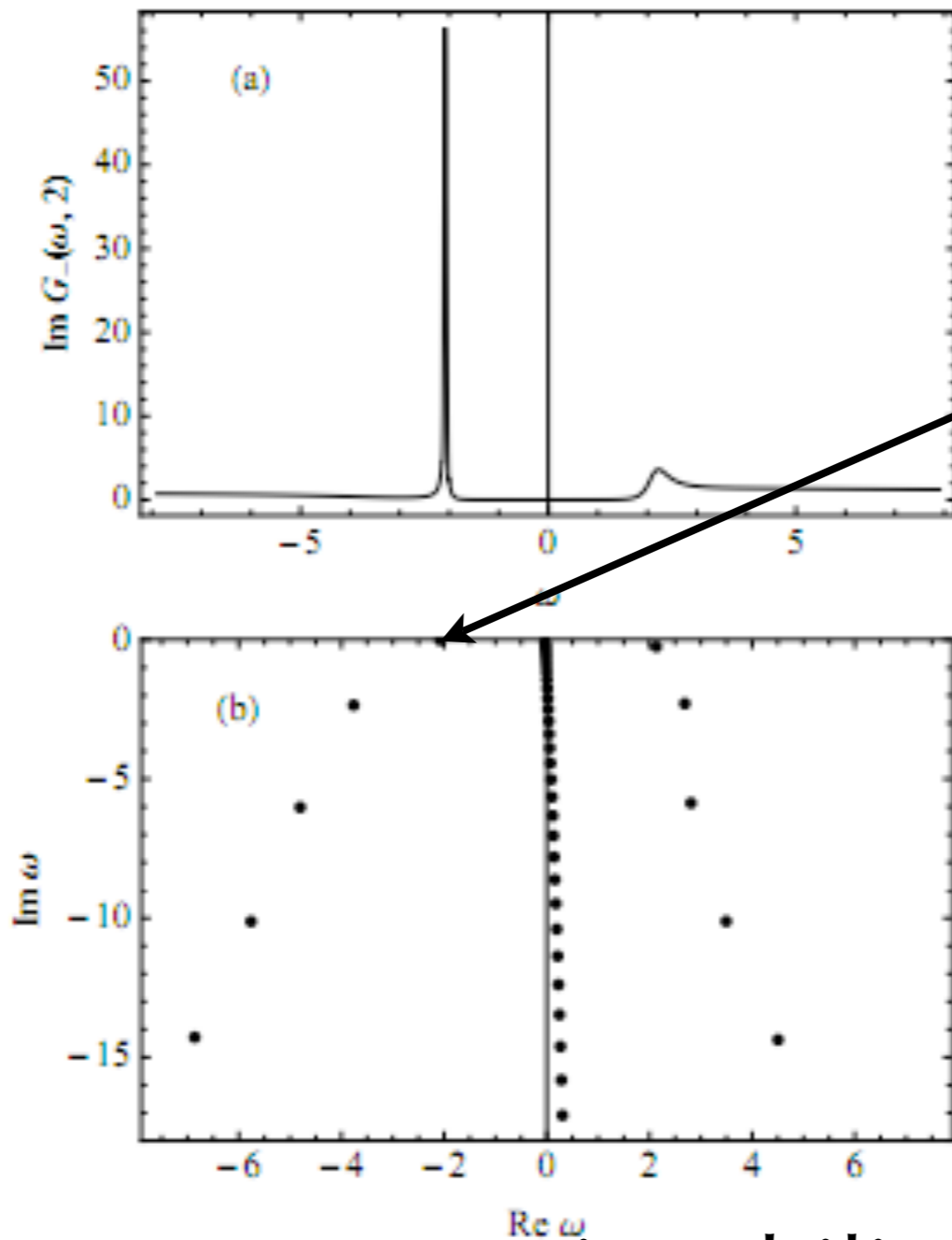
quasi-normal modes



quasi-normal modes

but the residue
drops to zero:
opening of a gap

quasi-particle residue

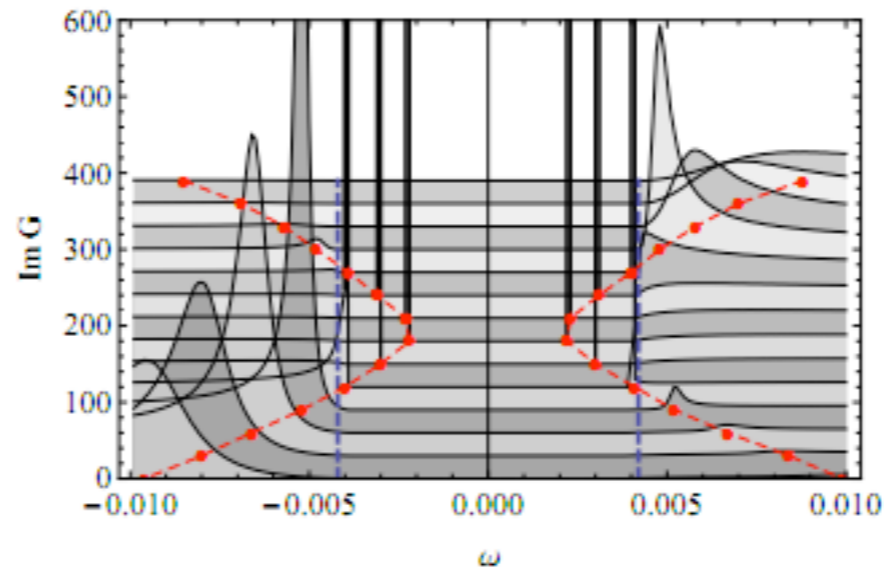


$$G_R(\omega, \mathbf{k}) = \frac{\mathbf{b}^{(0)} + \omega \mathbf{b}^{(1)} \mathcal{G}_{\mathbf{k}}(\omega)}{\mathbf{a}^{(0)} + \omega \mathbf{a}^{(1)} \mathcal{G}_{\mathbf{k}}(\omega)}$$

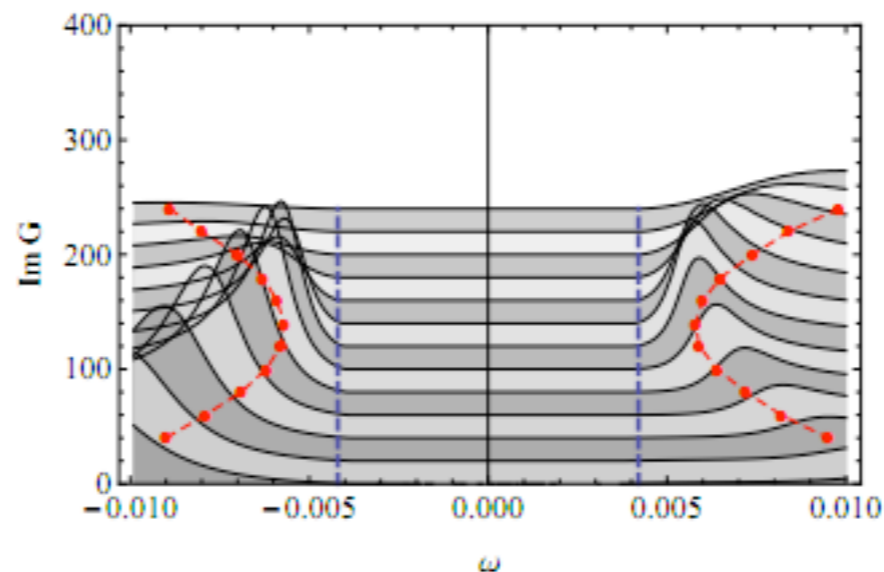
$Im \omega < 0$ no instability

$$S(\varphi) + S_D(\zeta) + \eta_5 \int d^{d+1}x \sqrt{-g} \varphi \bar{\zeta} C \Gamma^5 \zeta^T + cc$$

$$\eta_5 = 0.025$$

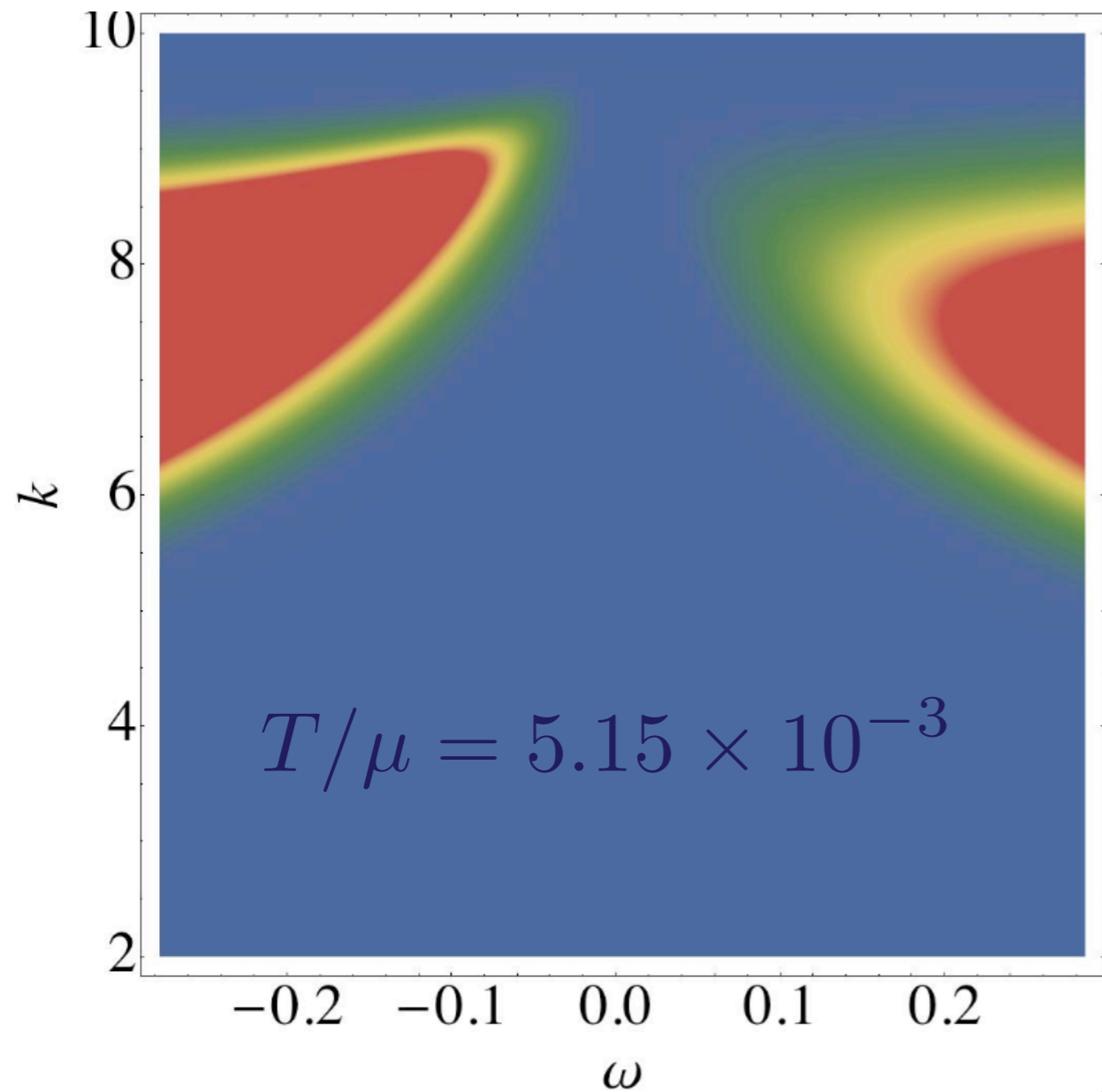


$$\eta_5 = 0.075$$

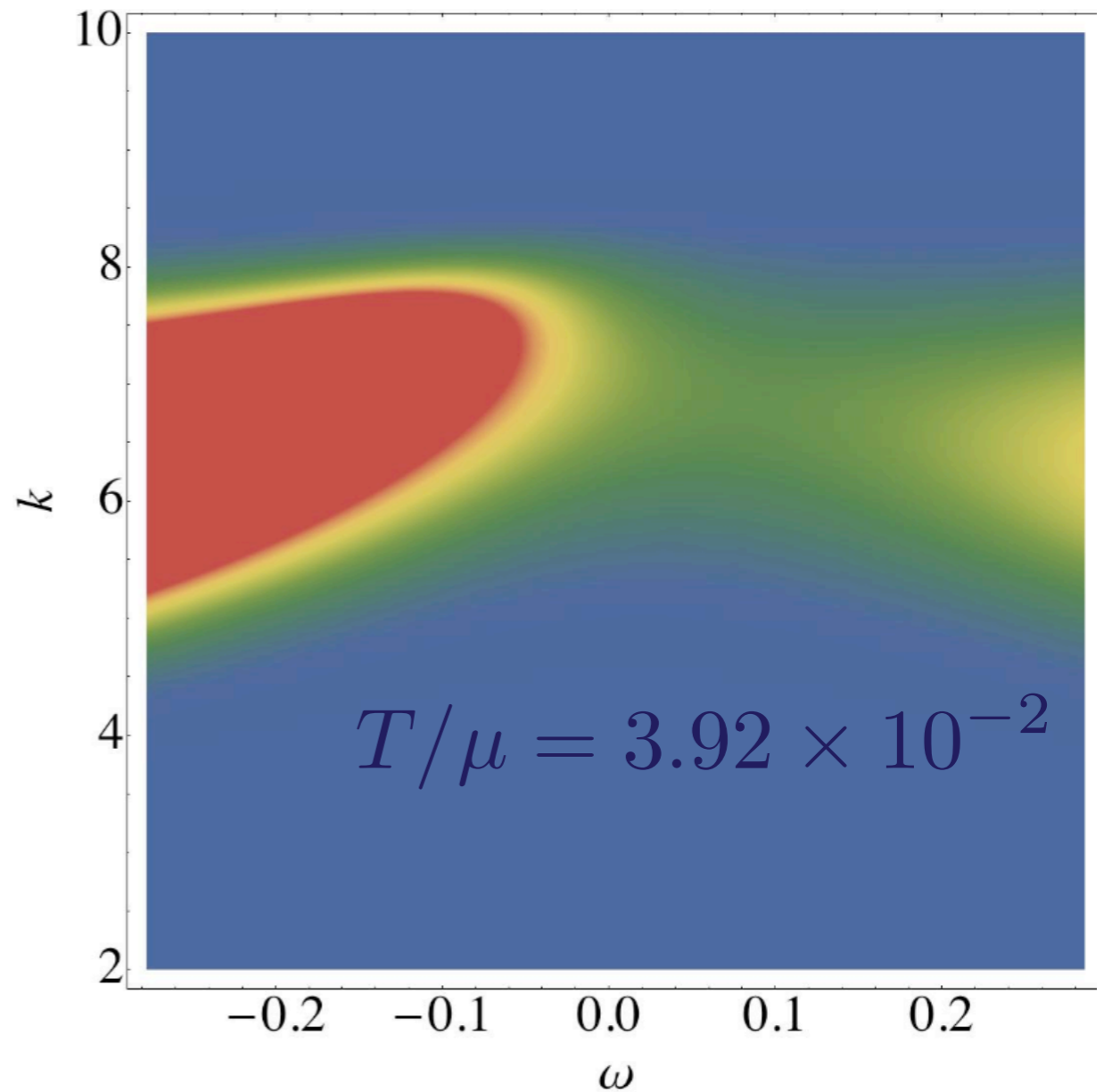


majorana coupling kills qp peaks
through coupling to scaling dimension of spinor field

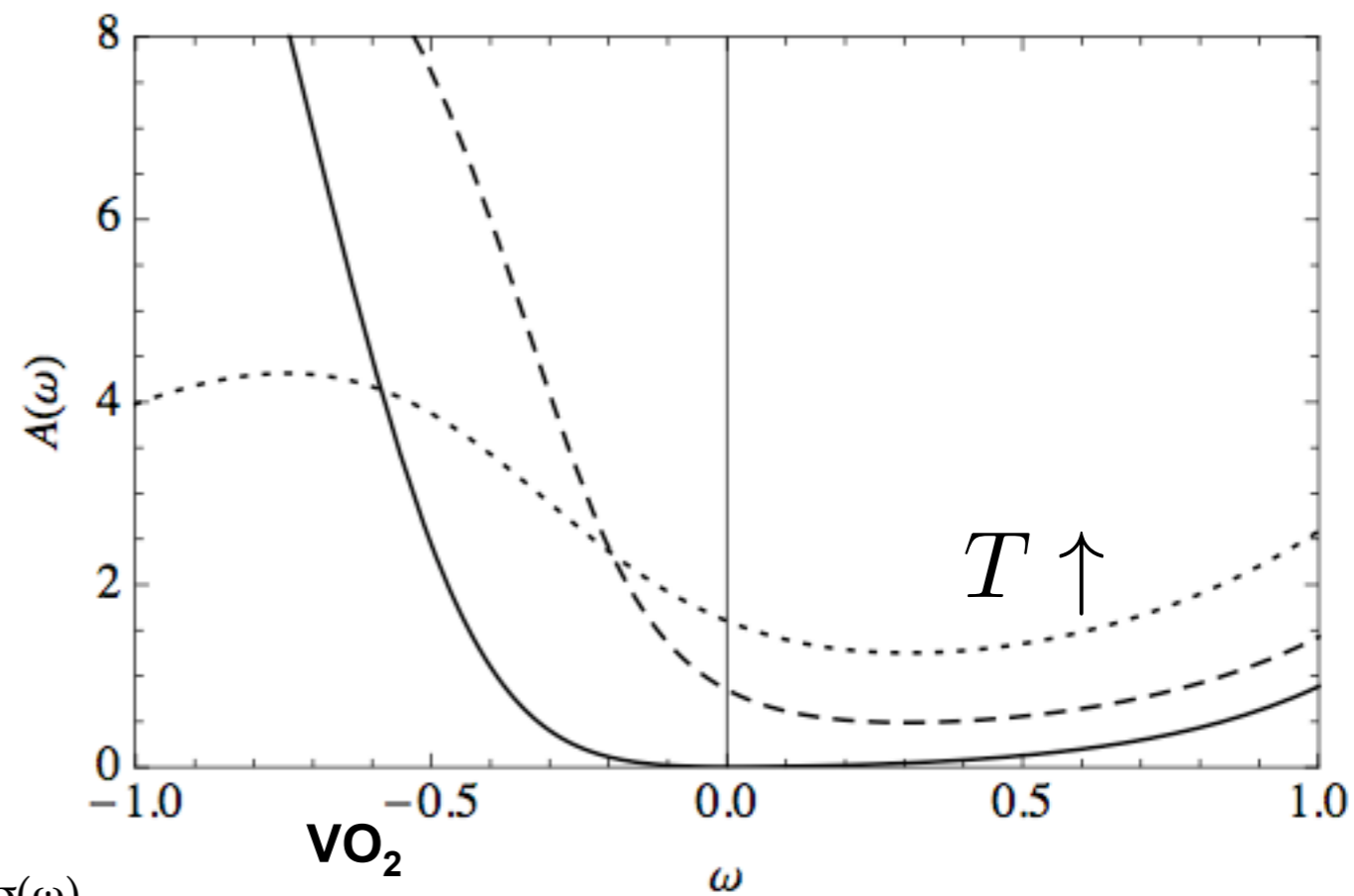
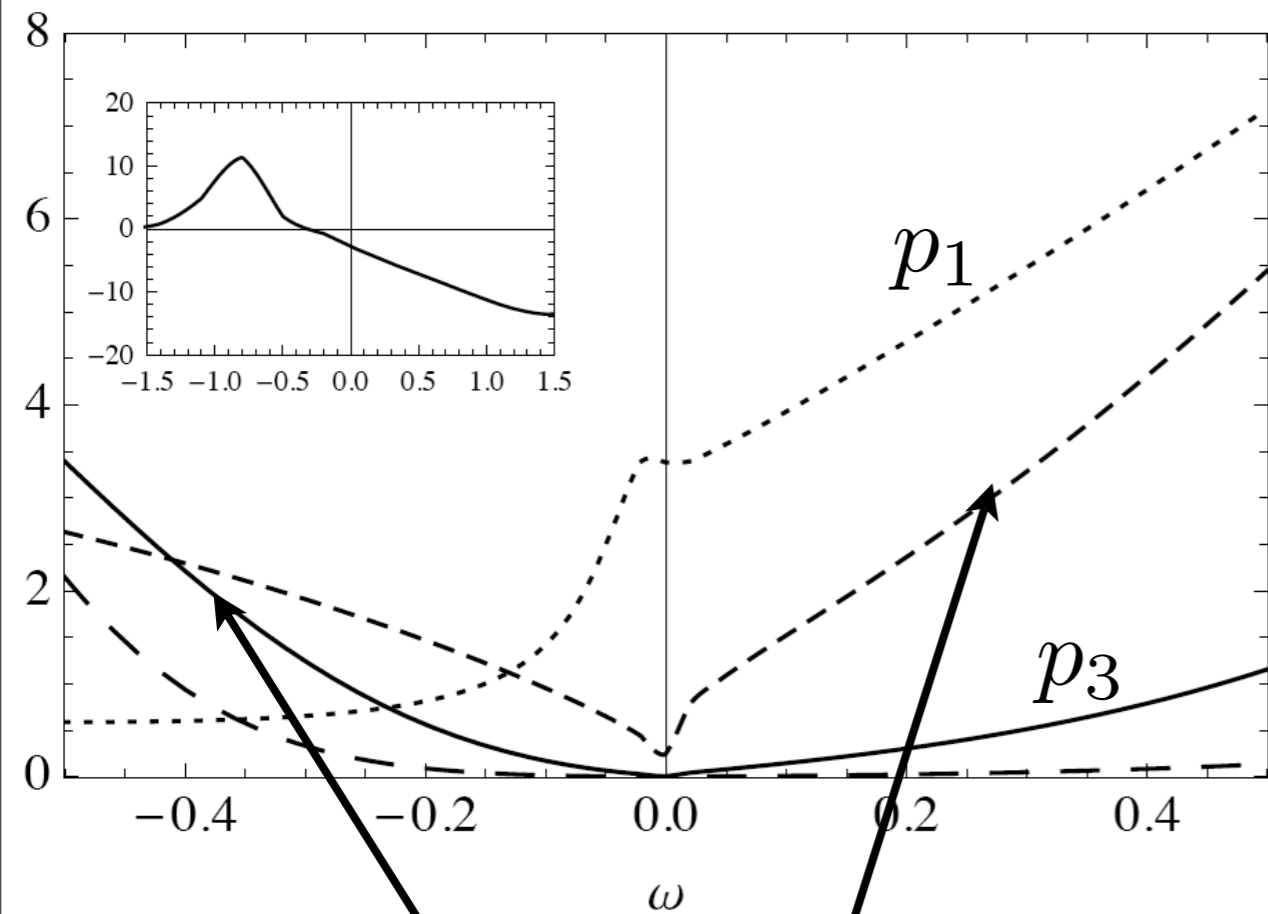
Finite Temperature Mott transition from Holography



$$\frac{\Delta}{T_{\text{crit}}} \approx 10$$

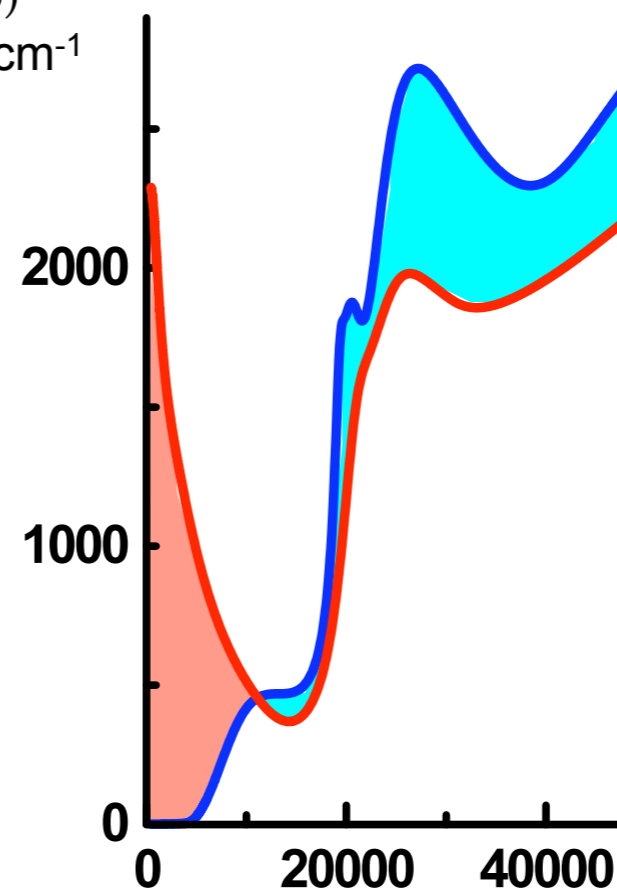


$$\frac{\Delta}{T_{\text{crit}}} \approx 20$$
 vanadium oxide



spectral weight
transfer
UV-IR mixing

$\sigma(\omega)$
 $\Omega^{-1}\text{cm}^{-1}$



$$G(\omega, k) = G(\omega\lambda^n, k)$$



discrete scale invariance
in energy



emergent IR scale



energy gap: Mott gap (Mottness)

$$G(\omega, k) = G(\omega\lambda^n, k)$$



discrete scale invariance
in energy



emergent IR scale

no condensate



energy gap: Mott gap (Mottness)

continuous scale invariance



discrete scale invariance
in energy

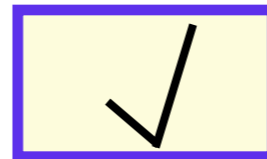
is this the symmetry that is ultimately
broken in the Mott
problem?

a.) yes

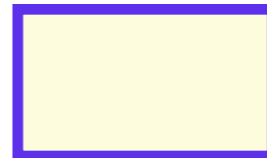
b.) no

holography

a.) yes



b.) no



holography

VO₂,
cuprates,...

a.) yes



b.) no

holography

VO₂,
cuprates,...

a.) yes



b.) no

Does scaling in VO₂ obey:

$$\Lambda_{\text{IR}} = \Lambda_{\text{UV}} e^{-\pi/\sqrt{U_c - U}}$$

holography

VO₂,
cuprates,...

a.) yes



b.) no

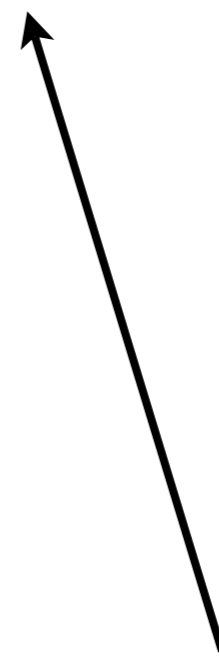
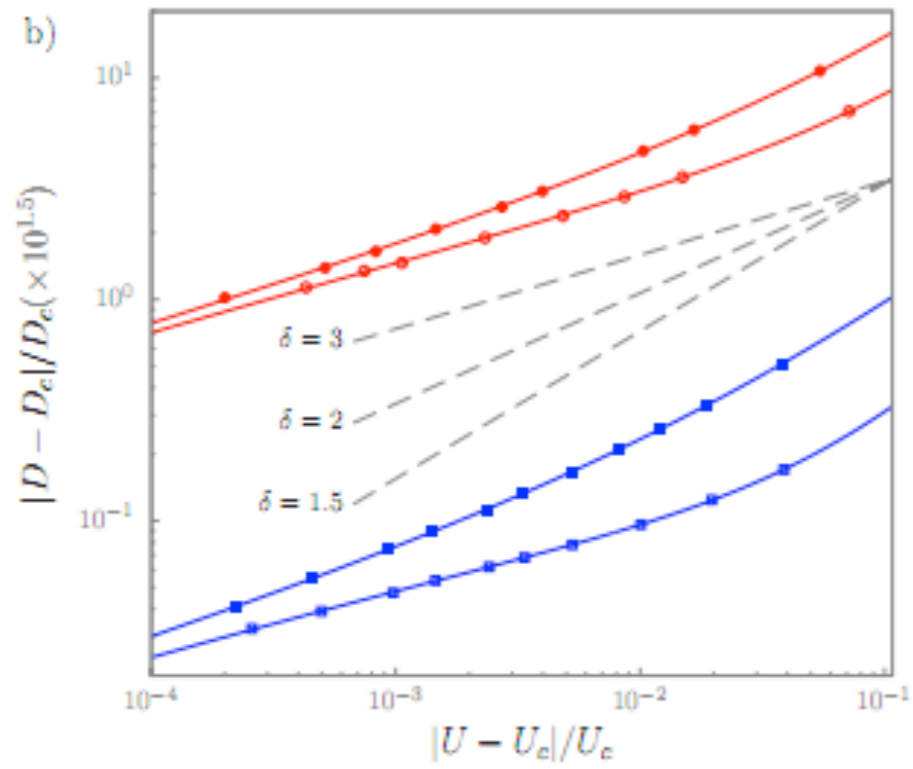
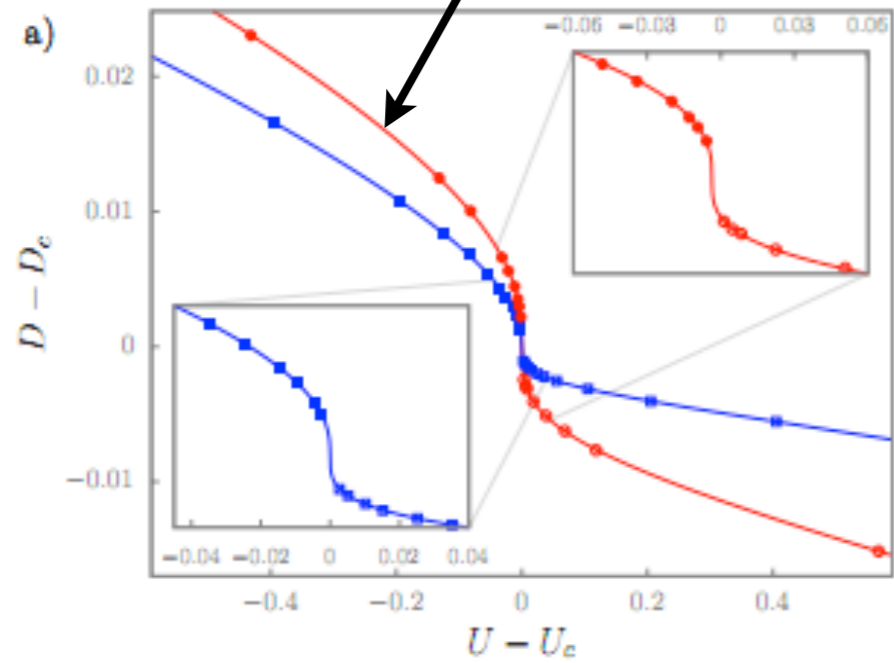
Does scaling in VO₂ obey:

$$\Lambda_{\text{IR}} = \Lambda_{\text{UV}} e^{-\pi/\sqrt{U_c - U}}$$

if yes: holography has solved the Mott
problem

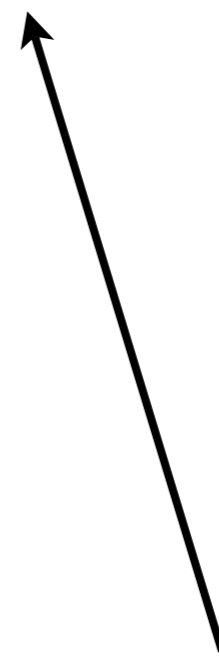
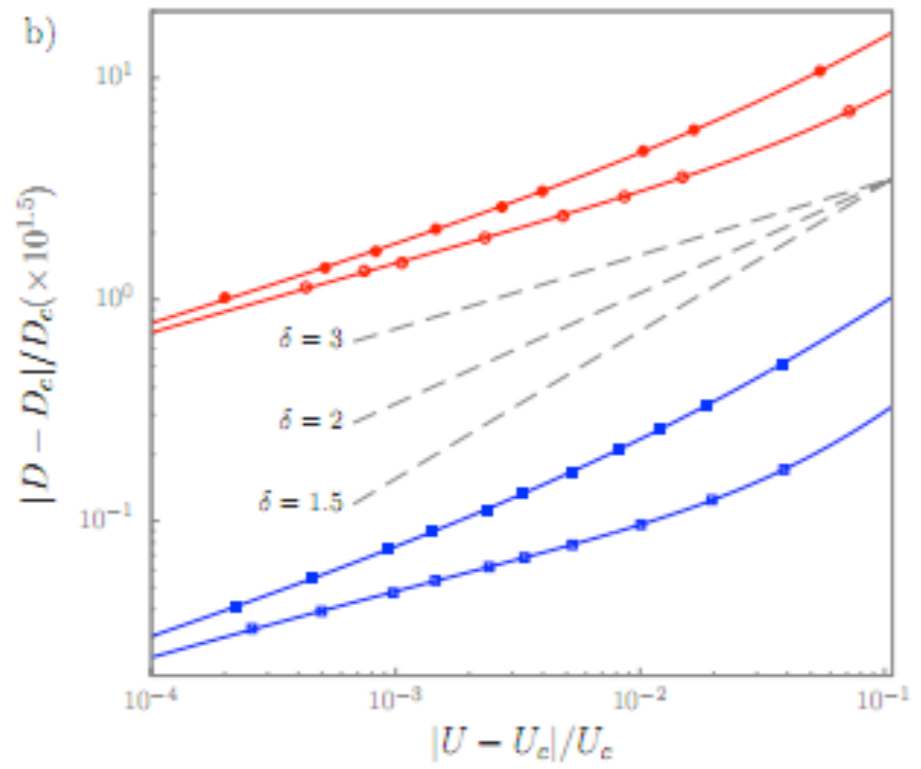
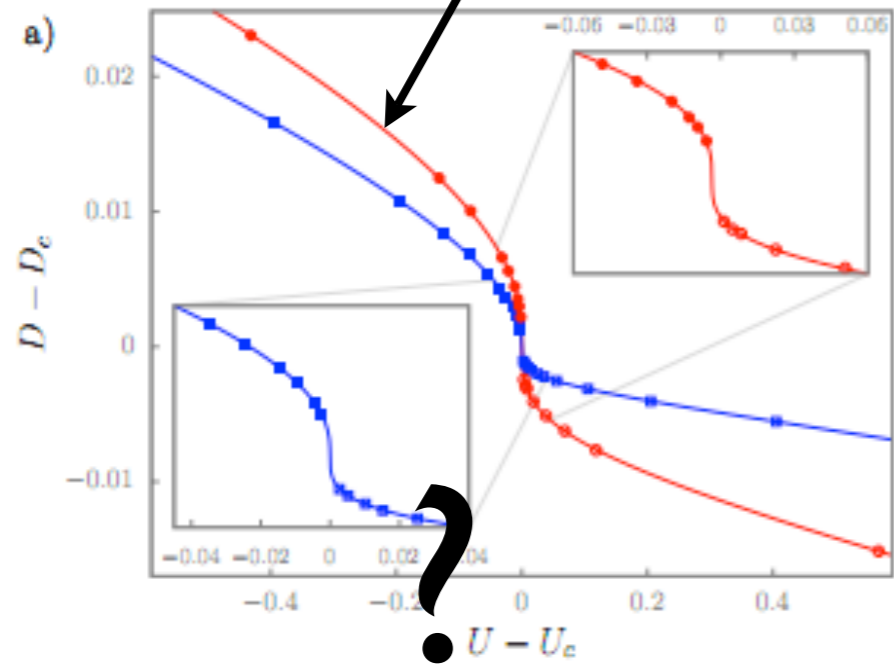
$$f(U) = c_1 \text{sgn}(\delta U) |\delta U|^{1/\delta} + c_2 |\delta|^{2/\delta} + c_3 \delta U + D_c$$

4 parameters



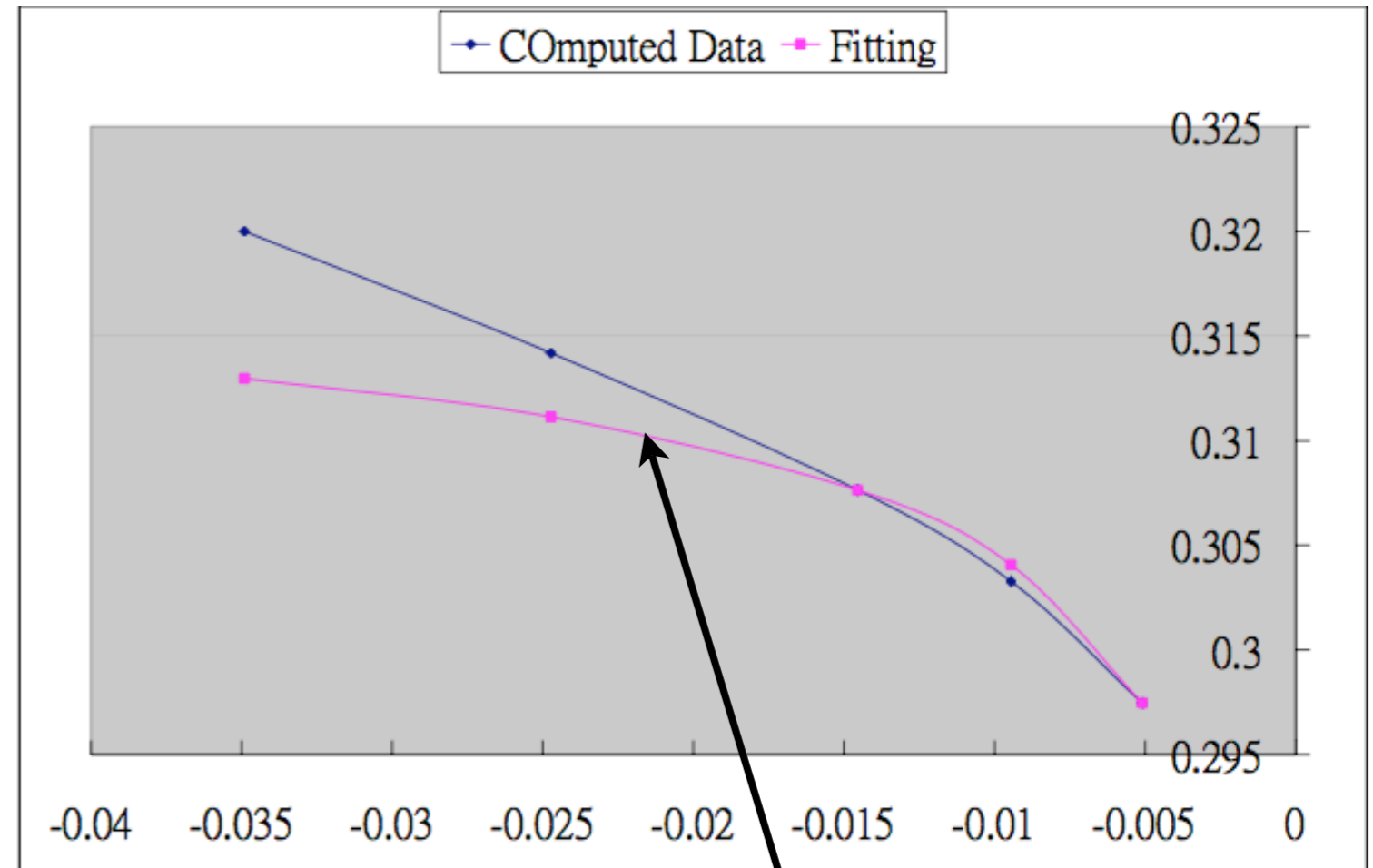
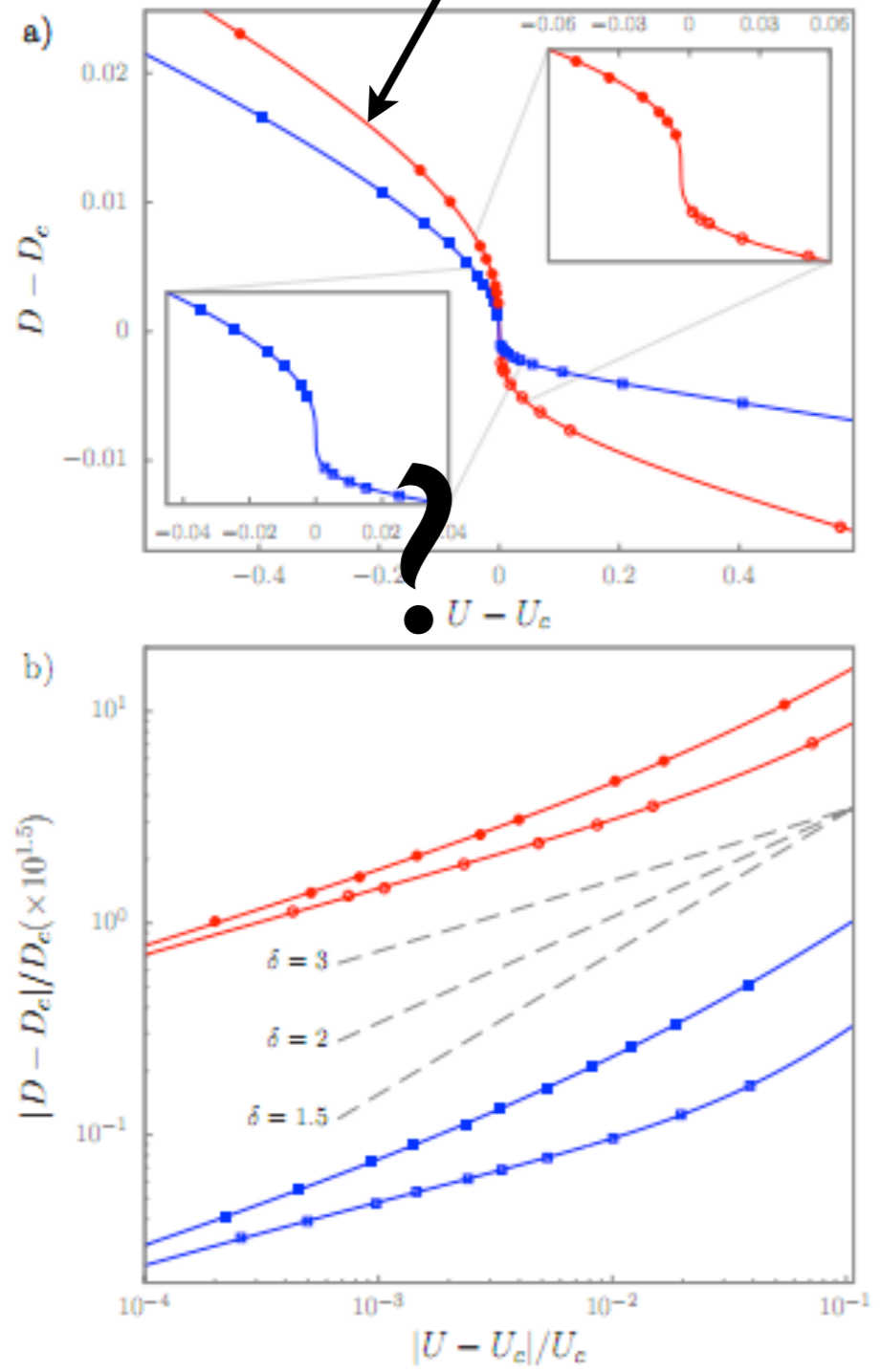
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4 parameters



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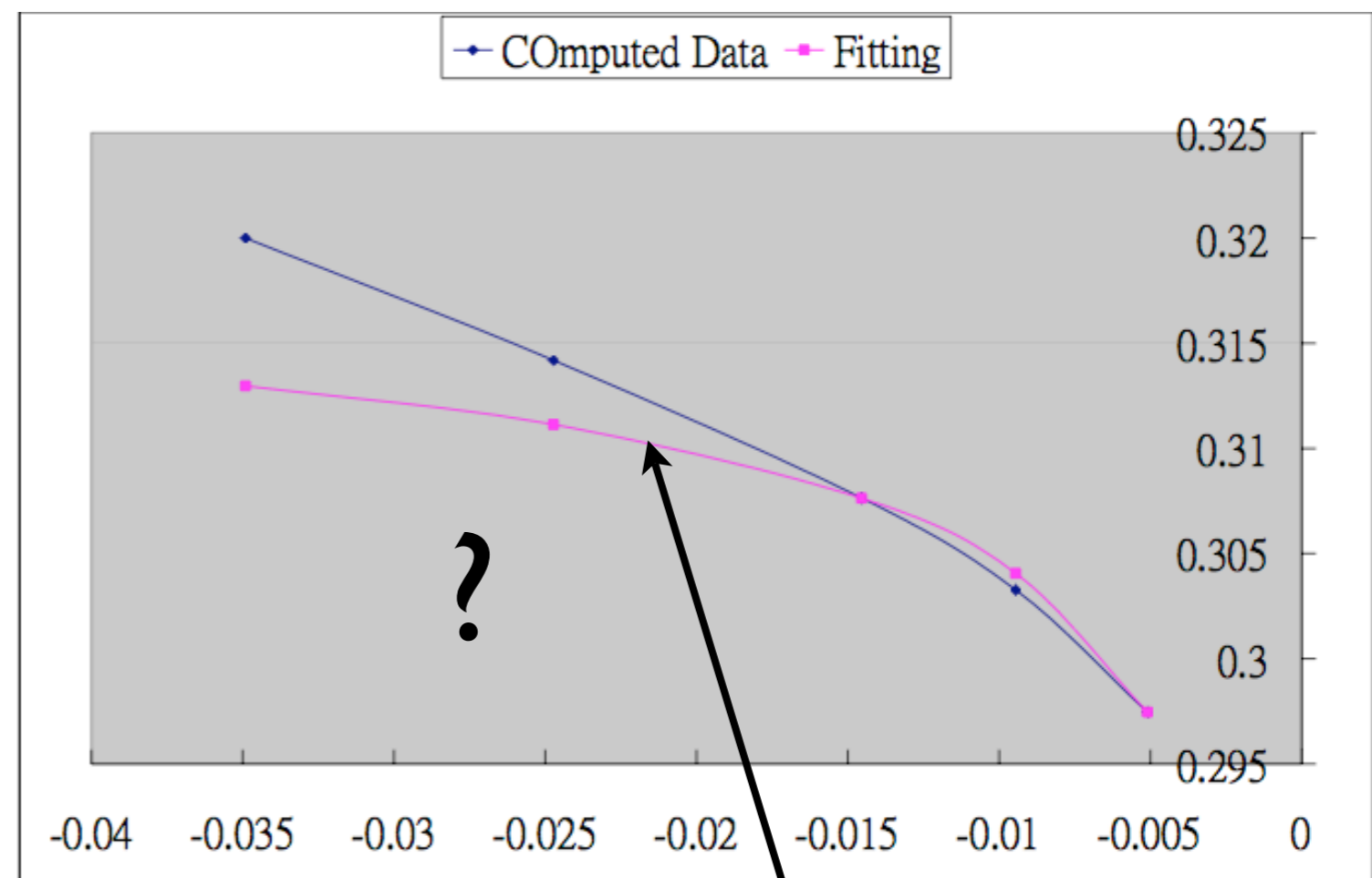
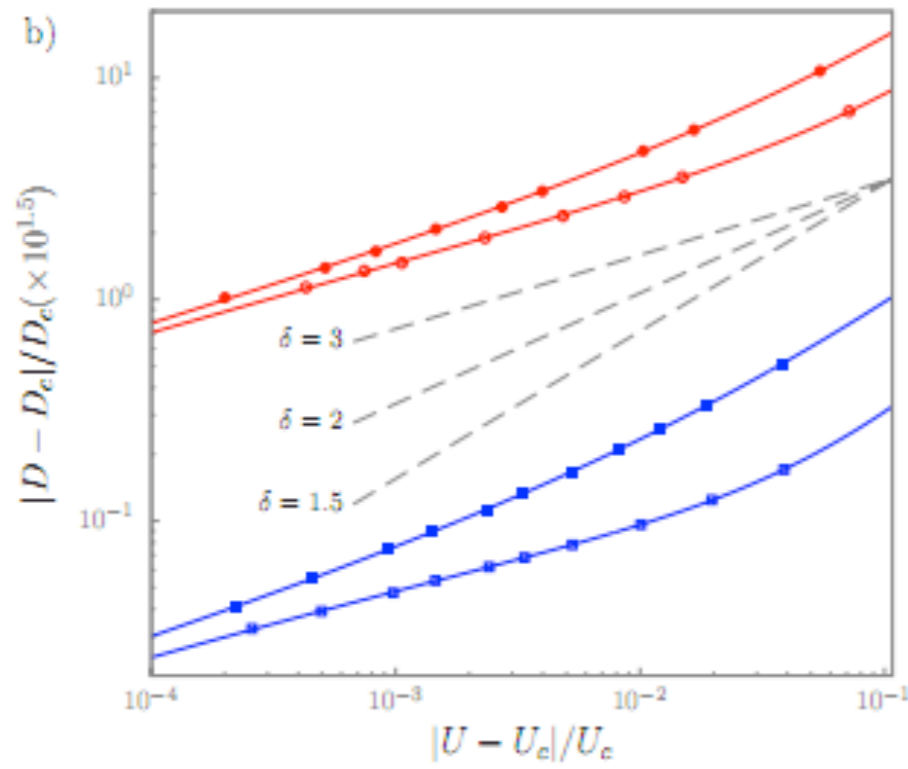
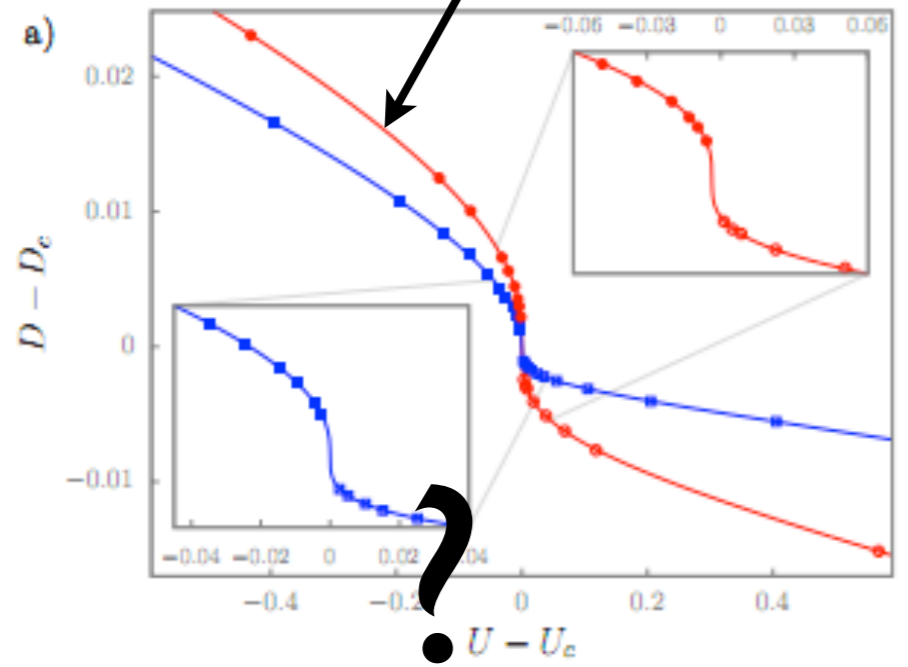
4 parameters



$$\Lambda_{\text{IR}} = \Lambda_{\text{UV}} e^{-\gamma/\sqrt{U_c - U}}$$

$$f(U) = c_1 \text{sgn}(\delta U) |\delta U|^{1/\delta} + c_2 |\delta|^{2/\delta} + c_3 \delta U + D_c$$

4 parameters



$$\Lambda_{\text{IR}} = \Lambda_{\text{UV}} e^{-\gamma/\sqrt{U_c - U}}$$