



### Mean flow and fluctuations in a modified Taylor-state MHD flow: recent results of the DTS experiment

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*DTS* set-up
Super-rotation !
Modified Taylor state (mean flow)
Waves (fluctuations)
Conclusions



### +DTS set-up

### Super-rotation ! Modified Taylor state (mean flow) Waves (fluctuations) Conclusions



### DTS experiment

- 50 litres of liquid sodium
- Outer sphere radius
   a=21cm
- Global rotation
  - (*f* < 15 Hz)
- → Differential rotation
   (|∆f/ < 45 Hz)</li>

*DTS* = Rotating spherical Couette flow in a dipolar magnetic field

8 mT Elsasser number  $\Lambda = \sigma B^2 / \rho 2\pi f \approx 1$ Magnetostrophic regime

> *Cardin et al, MHD, 2002 Nataf et al, GAFD, 2006*

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aimant

sodiun

liquide



### The DTS experiment is NOT a dynamo BUT:

### The flow modifies the magnetic field (Rm up to 40)

### The magnetic field modifies the flow (N of order 10)

### Global rotation can be present Numerical modelling is tractable



Brito et al, Exp. Fluids, 2001 Nataf et al, 2006, 2008

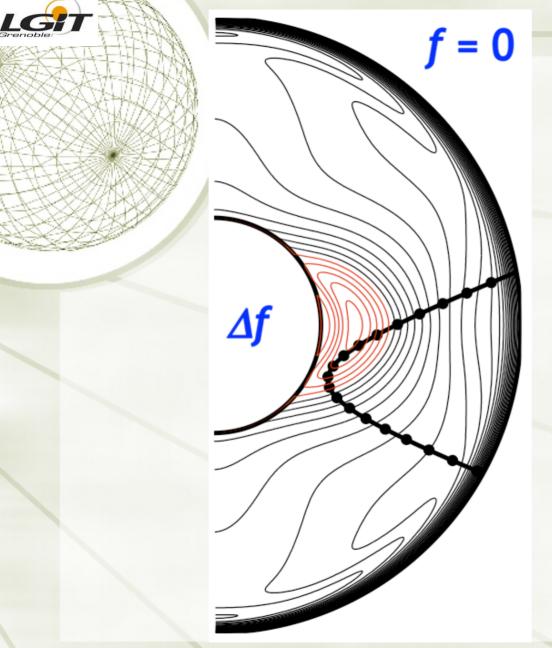


# *DTS* set-up Super-rotation ! Modified Taylor state (mean flow) Waves (fluctuations) Conclusions



### Outer sphere at rest

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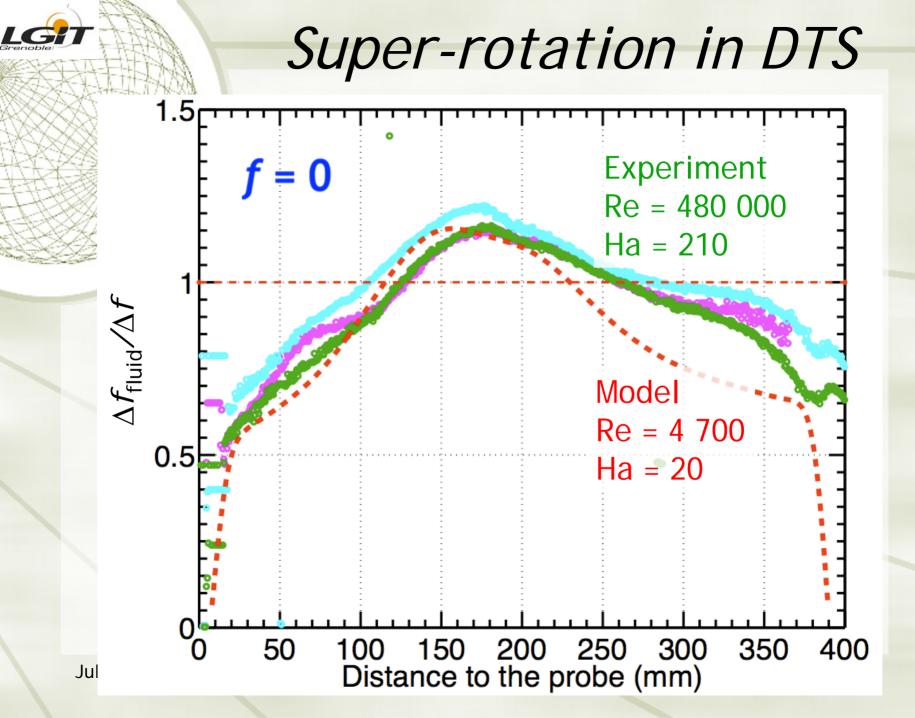
Axisymmetric non-linear model (D. Jault)

Re = 4700 Ha = 20 Pm = 7.5 10<sup>-6</sup>

Angular velocity isocontours

Super-rotation in red

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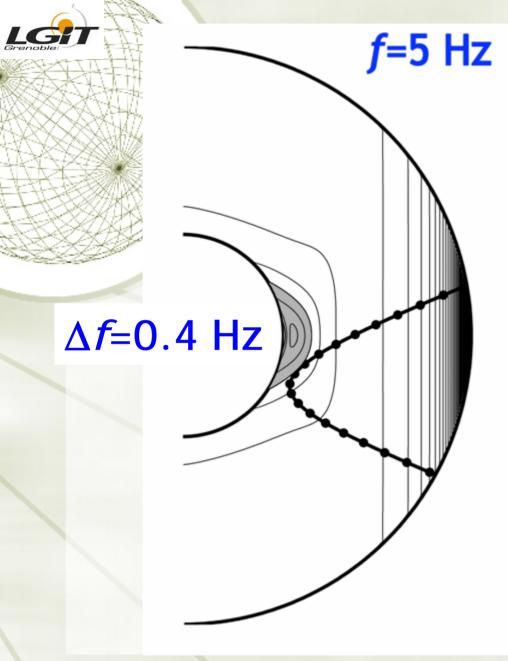




# *DTS* set-up Super-rotation ! Modified Taylor state (mean flow) Waves (fluctuations) Conclusions



### Rotating outer sphere



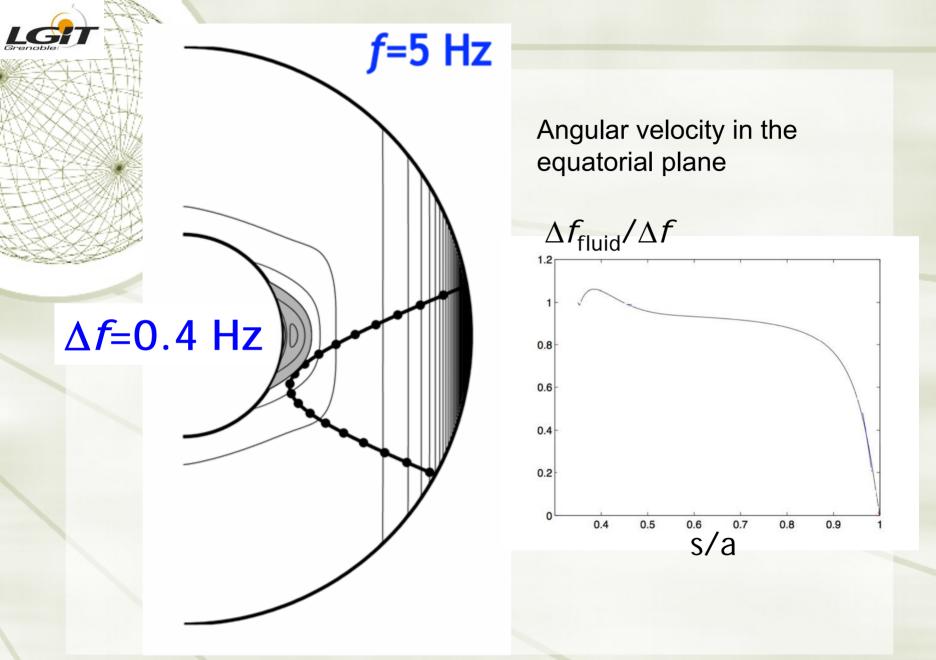
Axisymmetric non-linear model (D. Jault)

Re = 60 000 Ha = 210 Pm = 7.5  $10^{-6}$ E = 4.7  $10^{-7}$ 

Angular velocity isocontours

Super-rotation shaded

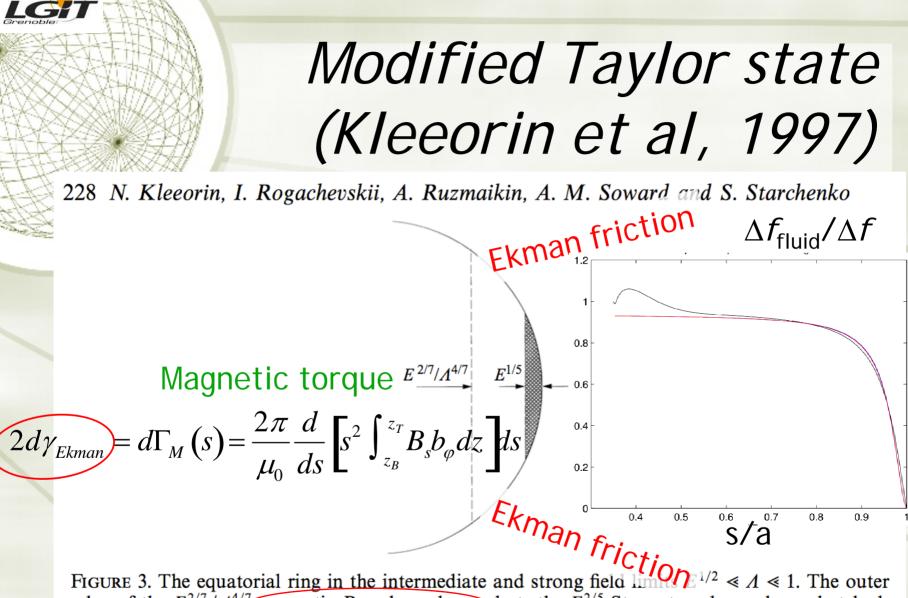
Nataf et al, PEPI, 2008





### Taylor state (Taylor, 1963)

 $\Delta f_{\rm fluid} / \Delta f$ Ζ<sub>T</sub> 1.2  $0 = \eta \Delta_H b_{\varphi} + s B_s \frac{d\Delta f(s)}{ds}$ 0.8 S Magnetic torque 0.6  $0 = d\Gamma_M(s) = \frac{2\pi}{\mu_0} \frac{d}{ds} \left[ s^2 \int_{z_B}^{z_T} B_s b_{\varphi} dz \right] ds$ 0.4 0.2 0 0.4 0.6 s/a 0.5 0.8 0.9  $Z_{R}$ 

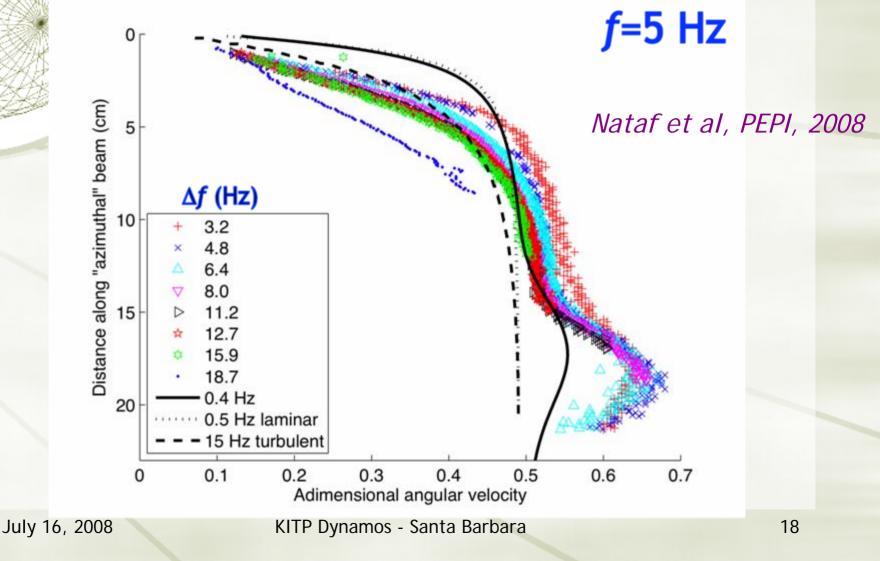


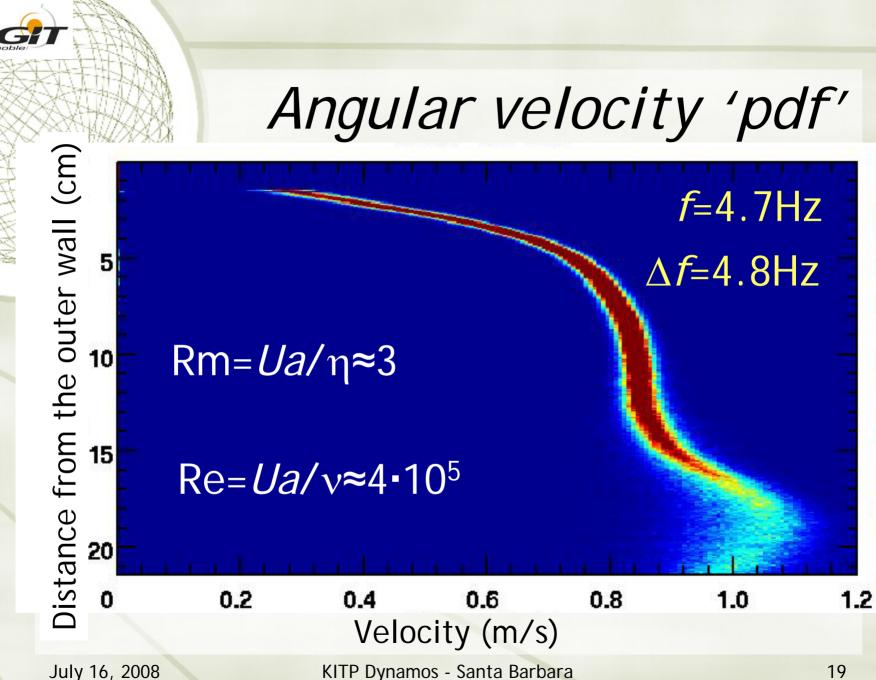
edge of the  $E^{2/7}/\Lambda^{4/7}$  magnetic-Proudman layer abuts the  $E^{2/5}$  Stewartson layer shown hatched.

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### Experimental, numerical and theoretical velocity profiles

LGI





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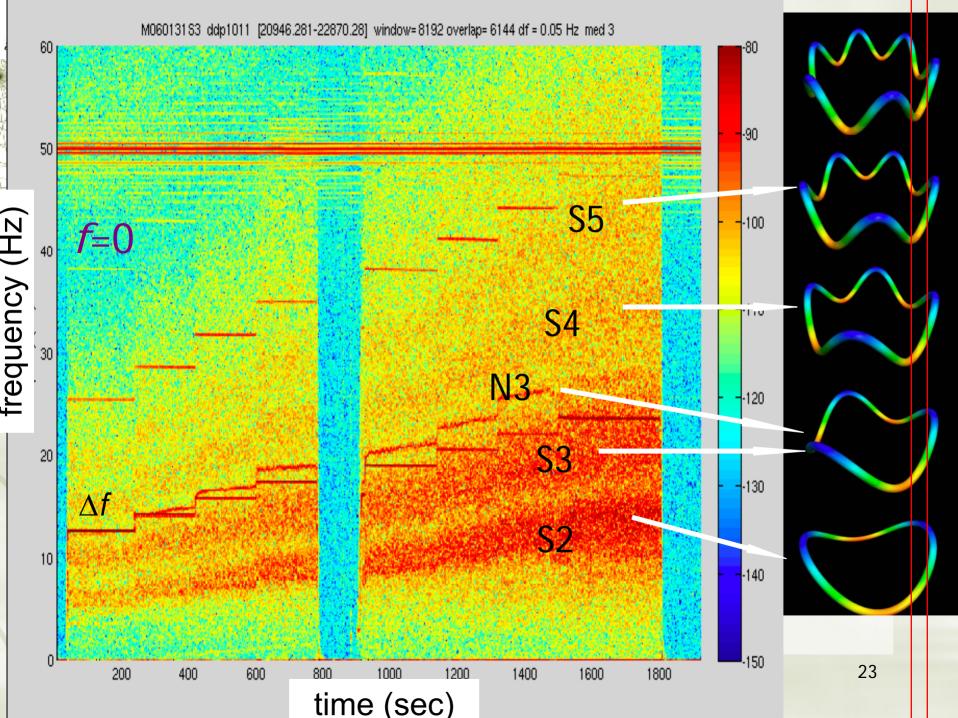
Electric potentials and meridional velocities

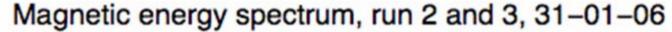
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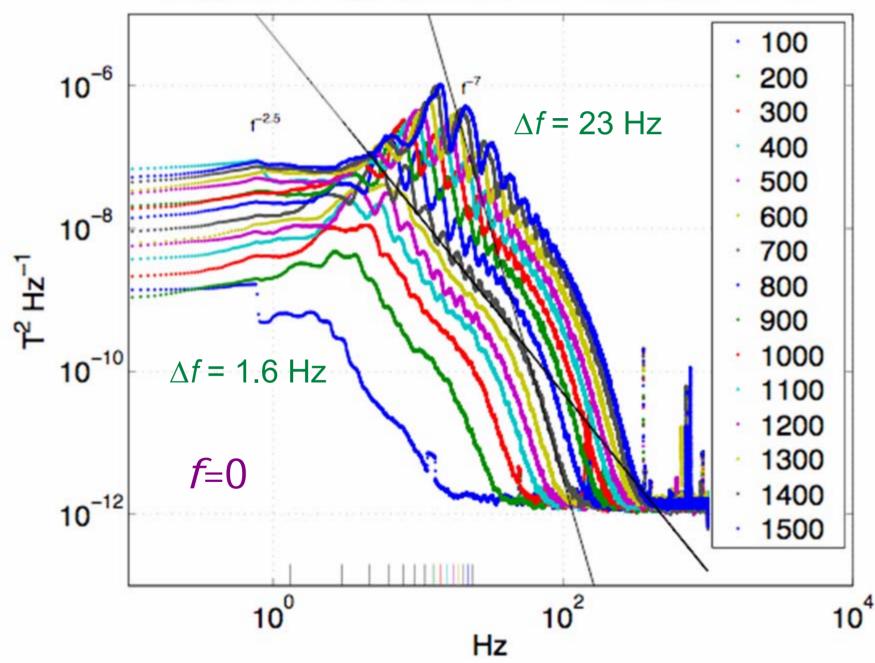


### Outer sphere at rest

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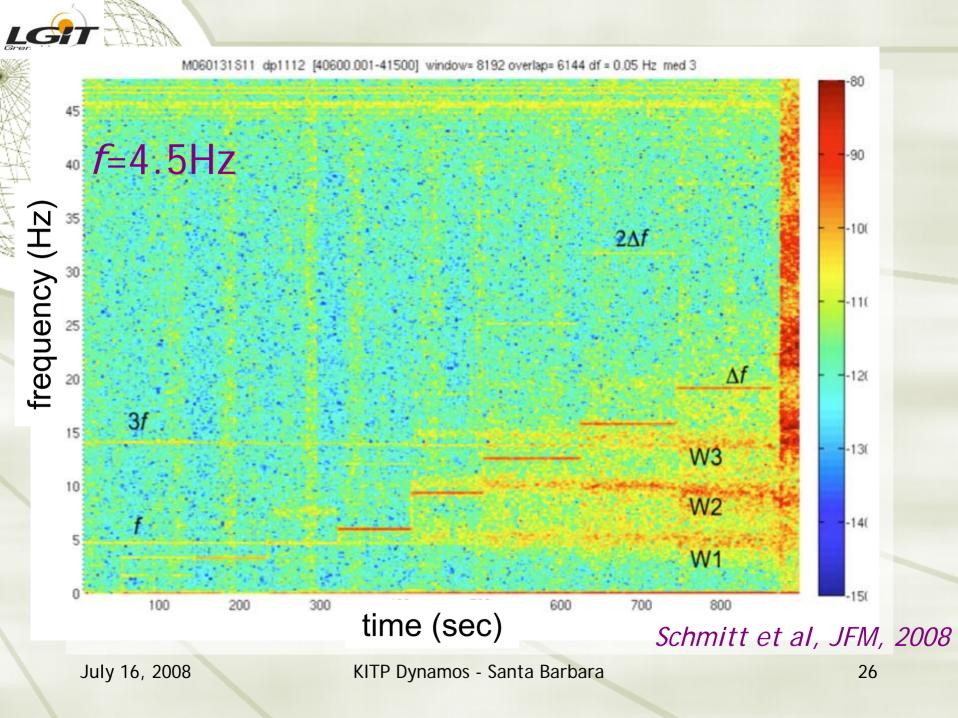


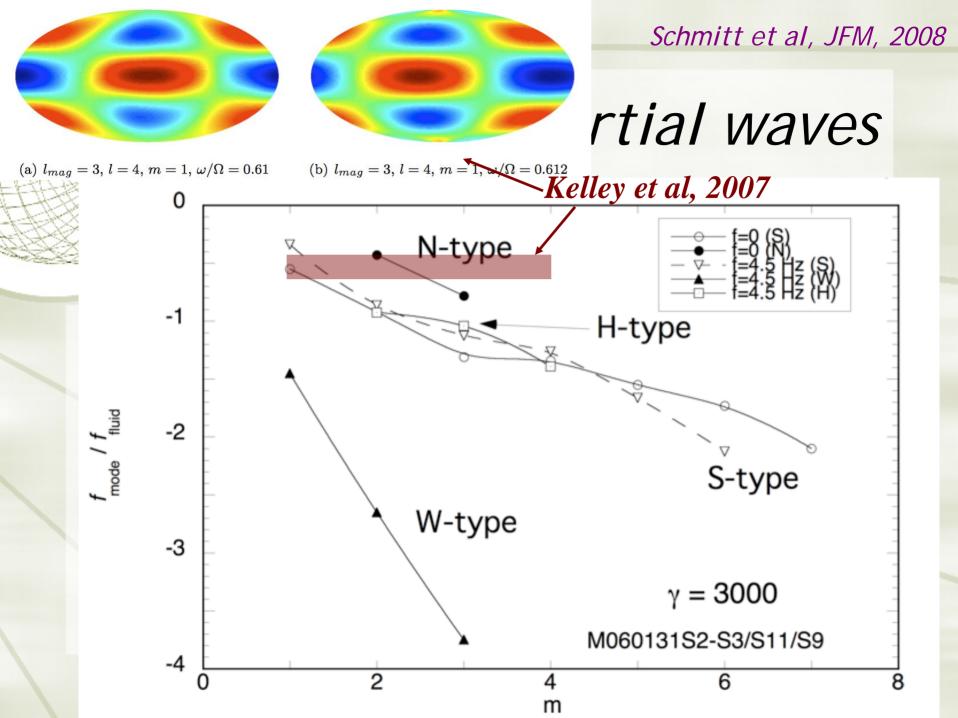






### Rotating outer sphere







## *DTS* set-up Super-rotation ! Modified Taylor state (mean flow) Waves (fluctuations) Conclusions



### Conclusions (1/3)

Large (25%) super-rotation observed for *f*=0
Mean flow for *f*≠0 :

- Quasi-geostrophic (QG) domain where Λ<1 : Coriolis dictates the geometry, Lorentz provides the driving.
- Modified Taylor state : magnetic torque balances friction in the Ekman layers.
- Experimental data well explained when turbulent layers are considered.

Nataf et al, PEPI, 2008



### Conclusions (2/3)

### Fluctuations :

- Dominated by non-axisymmetric waves/modes
- Dispersion relations not compatible with inertial modes
- ✦ Fluctuations for f≠0 :
  - Almost no fluctuation/wave when Ro>0
  - The combined constraints of rotation and magnetic field kill turbulence (except in the boundary layers)

Schmitt et al, JFM, 2008



### Conclusions (3/3)

Speculations on turbulence in the Earth's core : a very special kind :

- For a prescribed magnetic field, and because of rotation, the flow has no freedom left to fluctuate.
- The magnetic field can fluctuate, but only under the action of the flow.
- There is no dynamical constraint on the density field: it can fluctuate, but only under the action of the flow.

Nataf & Gagnière, CRAS, 2008



### To foster the discussion...

#### Flow at Re=10<sup>6</sup> is highly turbulent not in DTS

#### One can use subgrid models of turbulence not in DTS

 The induced magnetic field is large where the imposed field is large

#### not in DTS

- Joule dissipation dominates when B is large perhaps not in DTS
- It is difficult to measure flow velocity in liquid sodium expts yes, even in DTS, but we did it...

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Géodynamo



### A few articles

- Nataf & Gagnière, On the peculiar nature of turbulence in planetary dynamos, CRAS, accepted, 2008.
- Schmitt et al., Rotating spherical Couette flow in a dipolar magnetic field: experimental study of magneto-inertial waves, *J. Fluid Mech., in* press, 2008.
- Jault D., Axial invariance of rapidly rotating diffusionless motions in the Earth's core interior, *Phys. Earth Planet. Inter.*, **166**, 67-76, 2008.
- Nataf et al., Rapidly rotating spherical Couette flow in a dipolar magnetic field: an experimental study of the mean axisymmetric flow, *Phys. Earth Planet. Inter.*, *accepted*, 2008.
- Nataf et al., Experimental study of super-rotation in a magnetostrophic spherical Couette flow, *Geophysical and Astrophysical Fluid Dynamics*, 100, 281-298, 2006.
- Schaeffer & Cardin, Quasi-geostrophic kinematic dynamos at low magnetic Prandtl numbers, *Earth Planet. Sci. Lett.*, 245, 595-604, 2006.
- Schaeffer & Cardin, Rossby-wave turbulence in a rapidly rotating sphere, Nonlinear Processes in Geophysics, 12, 947-953, 2005.
- Cardin et al., Towards a rapidly rotating liquid sodium dynamo experiment, *Magnetohydrodynamics*, 38, 177-189, 2002.
- Dormy et al., A super-rotating shear layer in magnetohydrodynamic spherical Couette flow, J. Fluid Mech., 452, 263-291, 2002.
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