

TREATMENT OF BOUNDARIES IN NUMERICAL RELATIVITY

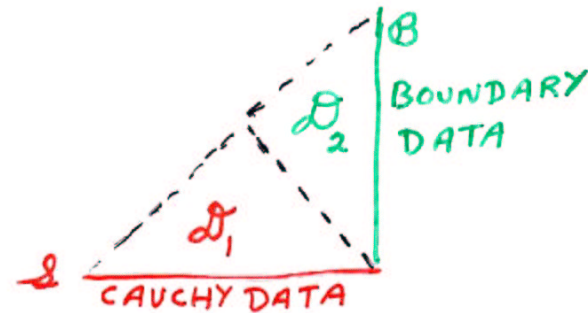
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based upon collaborative work with

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THE INITIAL-BOUNDARY VALUE PROBLEM FOR
GENERAL RELATIVITY

$$\text{EVOLUTION-CONSTRAINT SYSTEM: } \left. \begin{array}{l} \mathcal{E} = 0 \\ \mathcal{C} = 0 \end{array} \right\} \Rightarrow G_{\mu\nu} = 0$$



ASSUME CAUCHY IS OK.
WHAT CAN GO WRONG IN \mathcal{D}_2 ?

MATHEMATICALLY:

- NOT WELL-POSED
- CONSTRAINTS NOT SATISFIED

NUMERICALLY:

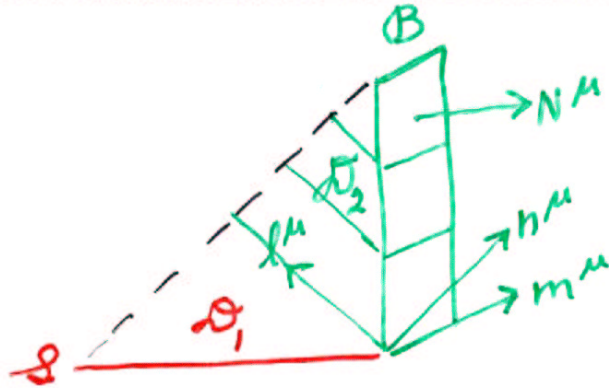
- UNSTABLE

PHYSICALLY:

- BOUNDARY DATA WRONG
- CAN'T EXTRACT WAVEFORM

THE DETAILS DEPEND UPON THE SYSTEM

SIMPLE EXAMPLE: NULL EVOLUTION IN \mathcal{D}_2



INTRODUCE NULL TETRAD ASSOCIATED WITH FOLIATION OF \mathcal{B}

$$g_{\mu\nu} = -\ell_{(\mu}n_{\nu)} + m_{(\mu}\bar{m}_{\nu)}$$

ASSUME \mathcal{B} EITHER EXPANDS OR CONTRACTS IN ℓ^μ DIRECTION

EVOLUTION SYSTEM \mathcal{E} : $G_{\mu\nu}\ell^\nu = 0$ $G_{\mu\nu}m^\mu m^\nu = 0$

INTRODUCE PROJECTOR ASSOCIATED WITH UNIT SPATIAL NORMAL N^μ TO \mathcal{B}

$$h^\mu_\nu = \delta^\mu_\nu - N^\mu N_\nu$$

CONSTRAINTS (3) \mathcal{C} : $h^\rho_\nu G_{\rho\sigma} N^\sigma = 0$

CONSTRAINT PROPAGATION

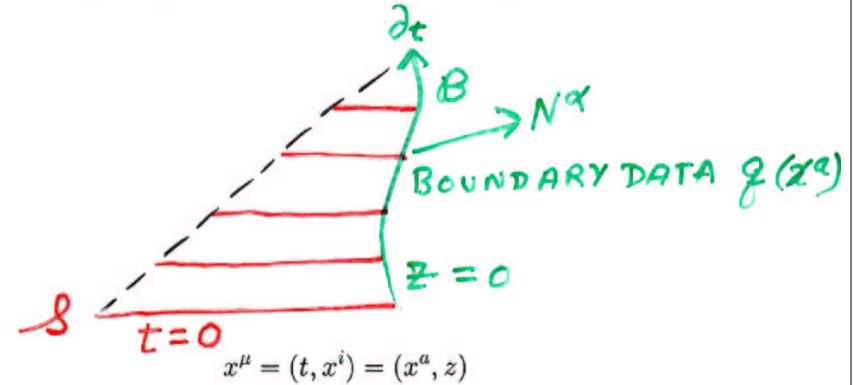
$$\mathcal{C} = \frac{1}{r^2} [r^2 \mathcal{C}]_{\mathcal{B}}$$

WELL-POSED ???

ROBUST \checkmark

CAUCHY BOUNDARY FOR A SCALAR FIELD ON A CURVED BACKGROUND

(SYMMETRIC) HYPERBOLIC SYSTEM: $g^{\mu\nu}\nabla_\mu\nabla_\nu\Phi = 0$



BOUNDARY FLUX: $\mathcal{F} = N^\mu T_{t\mu} = -(\partial_t\Phi)N^\mu\partial_\mu\Phi$

WELL-POSED BOUNDARY CONDITIONS: $\mathcal{F} \geq 0$

HOMOGENEOUS DIRICHLET BOUNDARY CONDITION:

$$\partial_t\Phi = 0$$

HOMOGENEOUS SOMMERFELD BOUNDARY CONDITION:

$$\partial_t\Phi + \sqrt{-g_{tt}}N^\mu\partial_\mu\Phi = 0$$

HOMOGENEOUS NEUMANN BOUNDARY CONDITION:

$$N^\mu\partial_\mu\Phi = 0$$

INHOMOGENEOUS BOUNDARY DATA $q(x^a)$

$$\partial_t\Phi = q(x^a), \quad \partial_t\Phi + \sqrt{-g_{tt}}N^\mu\partial_\mu\Phi = q(x^a), \quad N^\mu\partial_\mu\Phi = q(x^a)$$

WELL-POSED INITIAL-BOUNDARY VALUE PROBLEM FOR THE LINEARIZED EINSTEIN EQUATIONS IN THE HARMONIC GAUGE

$$g_{\mu\nu} = \eta_{\mu\nu} + \delta g_{\mu\nu}$$

$$\gamma^{\mu\nu} = \delta(\sqrt{-g}g^{\mu\nu})$$

EVOLUTION SYSTEM \mathcal{E}

$$\square\gamma^{ij} = 0$$

$$\partial_\mu\gamma^{\mu\nu} = \partial_t\gamma^{t\nu} + \partial_i\gamma^{i\nu} = 0$$

CONSTRAINT SYSTEM \mathcal{C}

$$\square\gamma^{t\nu} = 0$$

THE INITIAL-BOUNDARY VALUE PROBLEM FOR THIS SYSTEM IS WELL POSED FOR FREE DIRICHLET, SOMMERFELD OR NEUMANN BOUNDARY DATA FOR THE COMPONENTS γ^{ij} .

THE CORRESPONDING EVOLUTION CODE IS ROBUST

THE BAD NEWS:

- IT IS NOT NOT KNOWN HOW TO GENERALIZE THIS SYSTEM TO THE NONLINEAR CASE.
- NO KNOWN NONLINEAR METRIC-CONNECTION SYSTEM HAS THIS FLEXIBILITY OF BOUNDARY DATA.

BOUNDARIES IN LINEARIZED APPROXIMATION

- J. Stewart, *Class. Quantum Grav.*, **15**, 2865 (1998)
- B. Szilágyi, B. Schmidt and J. Winicour, *Phys. Rev.*, **D65**, 064015 (2002).
- G. Calabrese, L. Lehner and M. Tiglio G. Calabrese, L. Lehner and M. Tiglio, *Phys. Rev. D*, **65**, 104031 (2002).
- G. Calabrese, L. Lehner, D. Neilsen, J. Pullin, O. Reula, O. Sarbach, M. Tiglio, gr-qc
- G. Calabrese, J. Pullin, O. Reula, O. Sarbach, M. Tiglio, gr-qc
- S. Frittelli and R. Gómez, gr-qc

BOUNDARIES FOR NONLINEAR EINSTEIN EQUATIONS

- H. Friedrich and G. Nagy, *Commun. Math. Phys.*, **201**, 619 (1999).

EVOLUTION VARIABLES: $e_\mu, \Gamma_{\mu\nu}^\rho, C_{\mu\nu}^\rho$

THIS IS THE ONLY NONLINEAR SYSTEM WHICH IS KNOWN TO ADMIT PHYSICALLY GENERAL BOUNDARY CONDITIONS: ONLY THE WEYL CURVATURE REQUIRES BOUNDARY CONDITION AND ANALOGUES OF DIRICHLET, NEUMANN AND SOMMERFELD ARE ALLOWED.

POOR MAN'S VERSION

- B. Szilágyi and J. Winicour gr-qc/0205044

BASED ON
CAUCHY PROBLEM IN HARMONIC COORDINATES:
Y. Fours-Bruhat, *Acta. Math.*, **88**, 141 (1955).
A. E. Fisher and J. E. Marsden, *Comm. Math. Phys.*, **28**, 1 (1972).

WELLPOSED INITIAL-BOUNDARY VALUE PROBLEM FOR
NONLINEAR SYSTEMS WITH CHARACTERISTIC BOUNDARIES:
P. Secchi, *Arch. Rational Mech. Anal.*, **134**, 155 (1996).

HARMONIC INITIAL-BOUNDARY VALUE PROBLEM

REDUCED EVOLUTION SYSTEM: $\gamma^{\mu\nu} = \sqrt{-g}g^{\mu\nu}$

$$\gamma^{\alpha\beta}\partial_\alpha\partial_\beta\gamma^{\mu\nu} + S^{\mu\nu}(\gamma, \partial\gamma) = 0$$

WELL-POSED FOR ANY DISSIPATIVE BOUNDARY CONDITIONS, e.g. DIRICHLET, SOMMERFELD, NEUMANN

CONSTRAINTS: $H^\mu = \partial_\nu\gamma^{\mu\nu} = \hat{H}^\mu(x^\rho, \gamma)$
FOR BREVITY SET $\hat{H}^\mu(x^\rho, \gamma) = 0$

REDUCED EQUATIONS IMPLY

$$\gamma^{\alpha\beta}\partial_\alpha\partial_\beta H^\mu + C^{\mu\alpha}_\beta\partial_\alpha H^\beta + D^\mu_\beta H^\beta = 0$$

UNIQUENESS IMPLIES $H^\mu = 0$ IF IT SATISFIES A DISSIPATIVE HOMOGENEOUS BOUNDARY CONDITION.

THIS IS NOT EASY TO ARRANGE.

EXAMPLE: DIRICHLET CONDITION $H^\mu|_B = 0$.

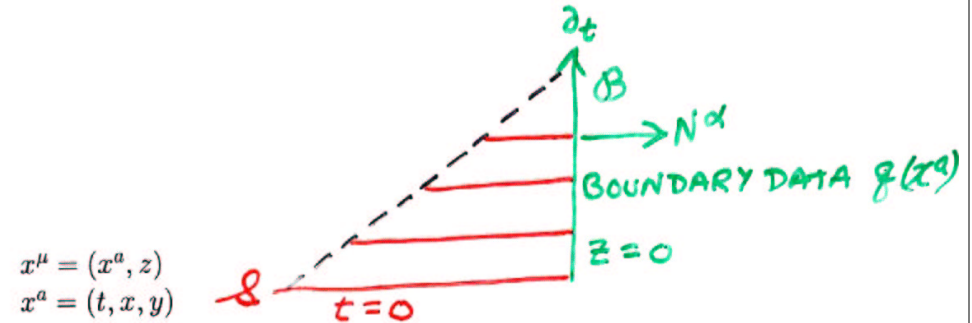
LET BOUNDARY BE AT $z = 0$ WITH $x^\mu = (x^a, z)$. THEN

$$H^z = \underline{\partial_a\gamma^{za}} + \partial_z\gamma^{zz} = 0$$

$$H^a = \partial_b\gamma^{ab} + \underline{\partial_z\gamma^{az}} = 0$$

NAIVE BOUNDARY DATA FOR $\gamma^{\mu\nu}$ IMPROPERLY POSES BOTH DIRICHLET AND NEUMANN CONDITIONS ON γ^{az}

WELL-POSED HOMOGENEOUS BOUNDARY DATA



ONE CHOICE THAT WORKS:

$$\left. \begin{aligned} \gamma^{za}|_B = 0 \\ \partial_z\gamma^{zz}|_B = 0 \\ \partial_z\gamma^{ab}|_B = 0 \end{aligned} \right\} \Rightarrow \left\{ \begin{aligned} H^z|_B = 0 \\ \partial_z H^a|_B = 0 \end{aligned} \right.$$

⇒ CONSTRAINTS SATISFIED

INHOMOGENEOUS BOUNDARY DATA $q(x^a)$

BOUNDARY HARMONIC GAUGE FREEDOM (SHIFT)

$$\gamma^{za}|_B = q^a(x^b)\gamma^{zz}|_B$$

NOTE: BOUNDARY DATA FOR $\gamma^{za}|_B$ DEPENDS ON DATA FOR $\gamma^{zz}|_B$ WHICH CAN ONLY BE DETERMINED BY CARRYING OUT THE EVOLUTION

BOUNDARY NORMAL: $\partial^n = \frac{1}{N^z}N^\mu\partial_\mu = \partial_z + q^a\partial_a$
NEUMANN DATA: $q^{zz} = \partial^n\gamma^{zz}|_B$

$$H^z|_B = 0 \Rightarrow q^{zz} = -\partial_a q^a \gamma^{zz}|_B$$

REMAINING NEUMANN BOUNDARY DATA

$$q^{ab} = \partial^n \gamma^{ab}|_{\mathcal{B}}$$

RELATED TO EXTRINSIC CURVATURE K^{ab} OF BOUNDARY.

$$\partial^n H_a|_{\mathcal{B}} = 0 \implies$$

$$\sqrt{-h} D_b (K_a^b - \delta_a^b K) + \sqrt{g^{zz}} K_{ab} H^b - \frac{g^{zz}}{2} H_b \partial_a q^b = 0$$

HERE h_{ab} AND D_a ARE THE METRIC AND CONNECTION INTRINSIC TO \mathcal{B}

This forms a symmetric hyperbolic system which determines the 6 pieces of Neumann data q^{ab} in terms of 3 free functions, as well as the free (boundary gauge) data q^a and boundary values of the variables γ^{zz} , γ^{ab} and $\partial_z \gamma^{za}$ which must be determined by the evolution.

ANY SOLUTION OF THE REDUCED EQUATIONS WITH THIS BOUNDARY DATA SATISFIES THE CONSTRAINTS.

IS THE INITIAL-BOUNDARY PROBLEM WELL-POSED???

NUMERICAL IMPLEMENTATION

SOME DIFFICULT CHOICES

- FIRST DIFFERENTIAL ORDER OR SECOND SECOND ORDER IN TIME FIRST ORDER IN SPACE

- CUBIC BOUNDARY OR SPHERICAL CUBIC

- GENERAL BOUNDARY GAUGE OR $\gamma^{za}|_{\mathcal{B}} = 0$
 $\gamma^{za}|_{\mathcal{B}} = q(x^b) \gamma^{zz}|_{\mathcal{B}}$

- HARMONIC FORCING TERMS OR $\partial_\mu \gamma^{\mu\nu} = 0$
 $\partial_\mu \gamma^{\mu\nu} = \hat{H}^\mu(x^\rho, \gamma)$

- BOUNDARY ACCURACY
1ST ORDER IN NORMAL DIRECTION, 2ND ORDER TANGENTIALLY

- NUMERICAL STENCILS, DISSIPATION, ...
BAG OF "TRICKS"

TESTS OF NAIVE ALGORITHM

- ROBUST STABILITY
- LINEARIZED WAVE CONVERGENCE TESTS
- NONLINEAR GAUGE WAVE CONVERGENCE TESTS

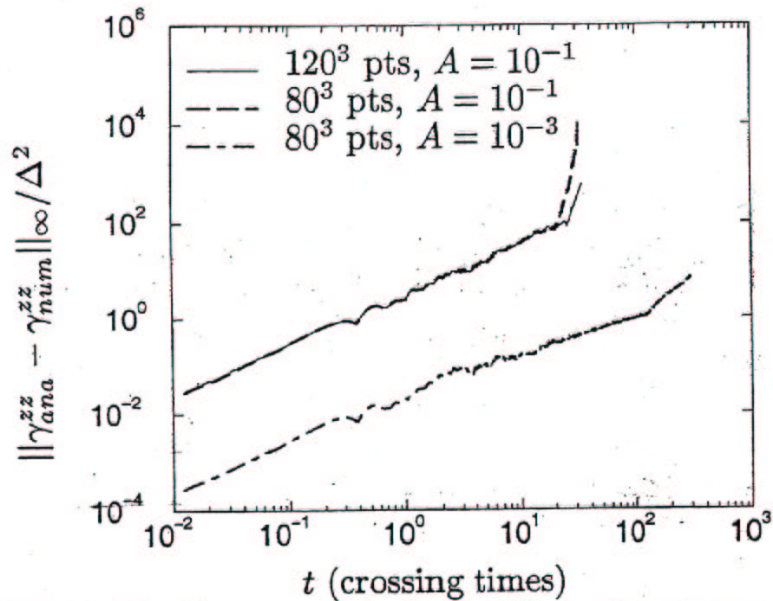
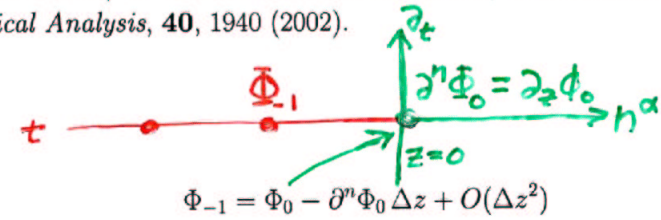


FIG. 2. The L_∞ norm of the finite-difference error, rescaled by a factor of $1/\Delta^2$, for a gauge-wave. The upper two (mostly overlapping) curves demonstrate convergence to the analytic solution for a wave with amplitude $A = 10^{-1}$ evolved for 30 crossing times with gridsizes 80^3 and 120^3 . The lower curve represents evolution of the same gauge-wave with $A = 10^{-3}$ for 300 crossing times with gridsize 80^3 .

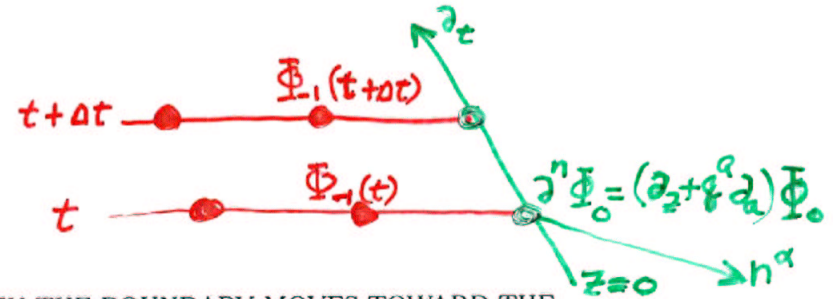
NEUMANN BOUNDARY CONDITIONS FOR THE SCALAR WAVE EQUATION IN SECOND DIFFERENTIAL ORDER FORM

CURVED BOUNDARY AT REST IN A MINKOWSKI METRIC BACKGROUND

H.-O. Kreiss, N. A. Petersson and J. Yström, *SIAM Journal on Numerical Analysis*, 40, 1940 (2002).

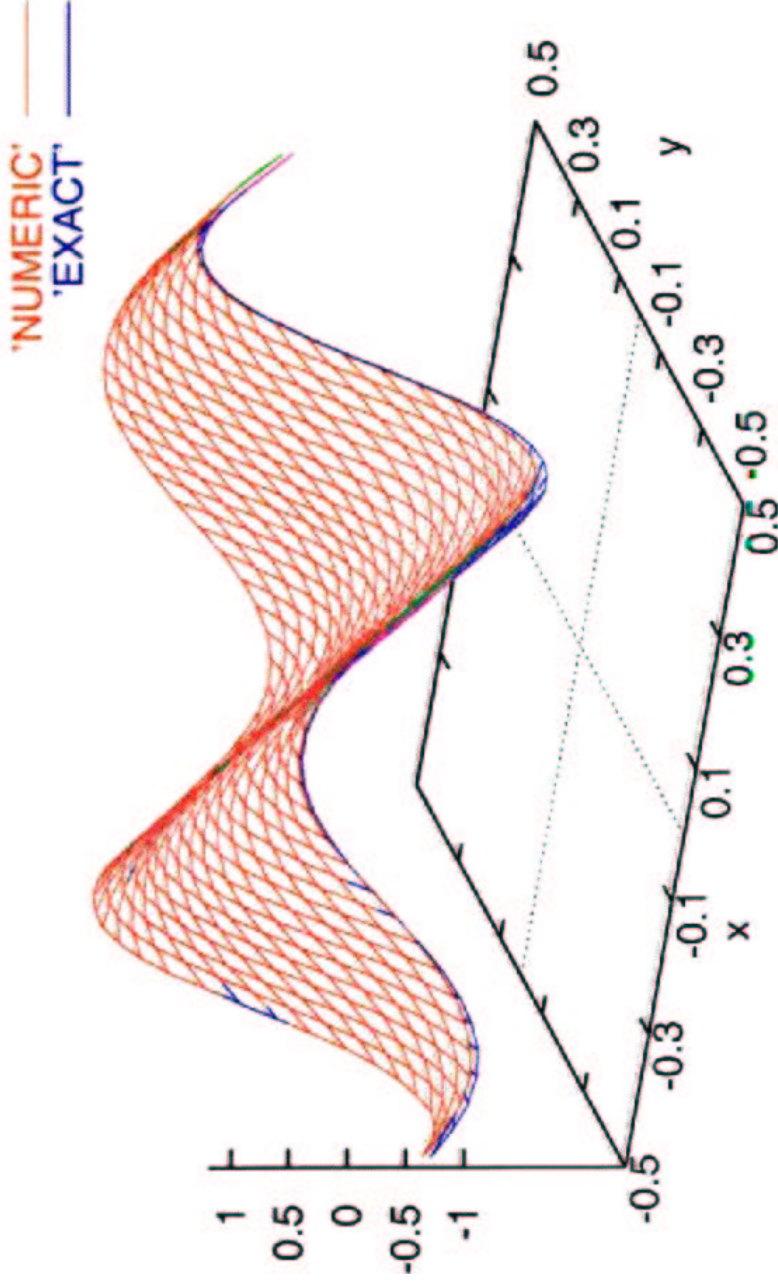


MOVING BOUNDARY ON A DYNAMIC BACKGROUND SPACETIME



WHEN THE BOUNDARY MOVES TOWARD THE CAUCHY INTERIOR THE INTERPOLATION STENCIL FOR Φ_0 INVOLVES FUTURE TIME LEVELS.

time = 20



**THE BOTTOM LINE
IS TO COMPUTE WAVEFORMS
FROM BINARY BLACK HOLES ...**

**BUT YOUR CODE
HAS TO PASS BASIC TESTS
IF THE WAVEFORMS CAN BE TRUSTED.**