

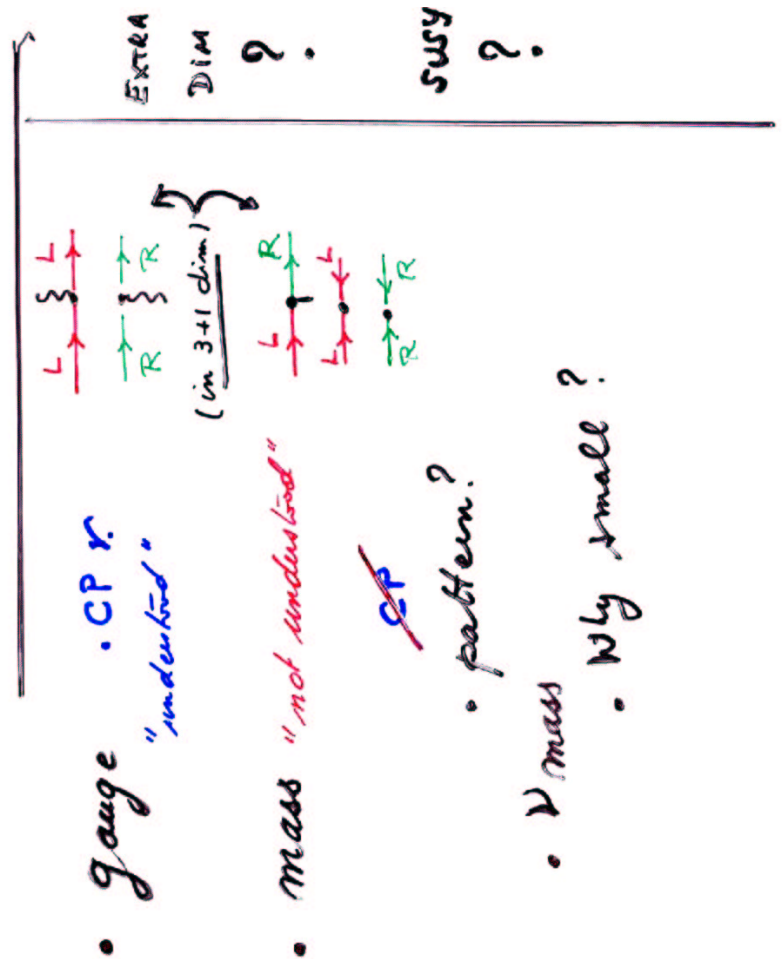
- INTRO

. is mass low-energy? (2)

- $\nu$  mass from susy (3a)
- $\nu$  and extra dim (3b)
- $\nu_R$  and  $SU(2)_R$  chiralities (3c)
  
- mass 4a
- CP estadi: 4b
  
- recreation: true  $\nu$  telescope 5
- current work 6

JMF  
KITP 1 mai 03

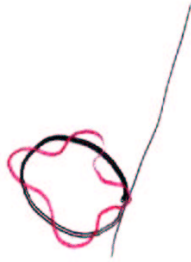
Gathering ideas for a talk  
to <sup>other</sup> neutrino fans.



Possible <sup>views</sup> sources of MASS.

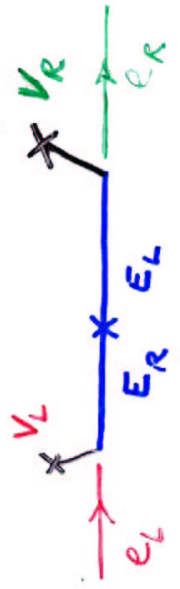
*mostly (fermions)*

- Bare mass term  $\rightarrow$  vectorlike th ~~degenerate multiples.~~  $\times$
- Min. energy of a particle  $\dots$  but only after choice of physical vacuum  $\times$
- $L-R$  transition (in  $3+1$  dim...)
- $\dots$  a low-energy phenomenon? *not necessarily* (see e.g.  $M_2$ )  $\rightarrow$  2.a
- Majorana (see below)
- Kaluza-Klein energy (extra. dim.)



2

a 2-scale model for all masses (cf. 5.Ma)



$\hookrightarrow$  chiral sym. breaking at high energy (e.g. fundamental mass for right E)

+ same for all other particles.

$$m \sim \frac{\nu_L \nu_R}{M}$$

$\rightarrow$  separate

$\hookrightarrow$  group ~~chirality~~

2a

More questions with Neutrino mass.

Why so small?

- $\nu_R$ . see-saw  $\rightarrow$  see later (3c)
- MSSM . nat. small...  $\nu_L$  Major. mass  $\sim 10^{-3} \text{ eV}$   $\xrightarrow{\text{Tanaka et al.}} (3a)$   
 • no  $\nu_R$
- SM:  $\rightarrow \left| \langle \chi^0 \rangle \right|$   $\langle \chi^0 \rangle \ll \nu$   
 (e.g. in LR models  $\nu_L \nu_R$ )  
 $\langle \chi_L^0 \rangle = \nu_L \sim \frac{\nu_L \nu_R}{\nu_R}$
- Coupling to large extra-dim LED  $\rightarrow$  funny oscillations patterns (3b)

3

$\nu_L$  mass from MSSM

S. Troitsky  
M. LeComanor  
prof 12/99

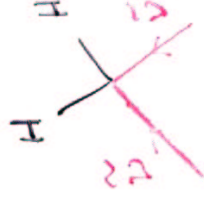
totally  $\nu_R$  - free!  
 No particle added\*

- $F = \text{Ausy breaking}$  } gravity mediated.
- $M \sim 10^{19} \text{ GeV}$  usually keep  $\frac{F}{M} \dots$  here go to  $\frac{F}{M^2} \sim h$

$$h(\tilde{L}_i H_j \epsilon_{ij})^2$$

R-parity conserving

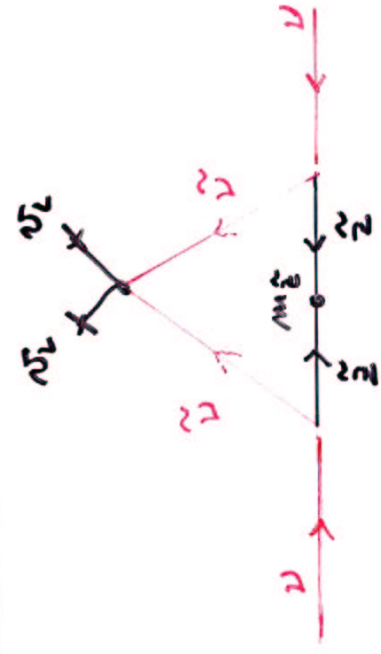
SU(2) singlet



(\* beyond MSSM years!) 3.a.1

→ generate

$$\frac{F}{M} \sim 1 \text{ TeV}$$



$$\frac{\hbar}{32\pi^2} \frac{g^2 < H^2}{m_{\nu}^2 \cdot \cos^2 \theta_w} f(\tilde{m}_Z/\tilde{m}_\nu) \quad \cdot 2 \rightarrow 0.6$$

$$\sim \hbar g^2 \frac{(F/M)^2}{(F/M)} \sim \boxed{\frac{F^2}{M^3} \sim 10^{-3} \text{ eV}}$$

checks:

- no chiral combination
- ...

unrelated

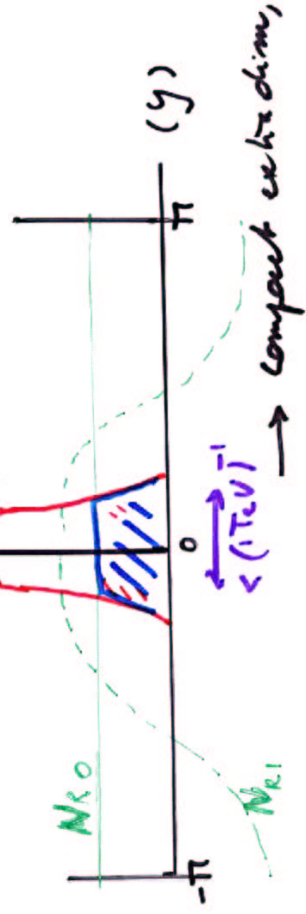
to  $m_e$  !

3.4.2

1) and extra-dim: strange oscillations and disappearance.

(CP and extra-dim: later)

$$m_{\text{min}} = (10^{-3} \text{ eV})^{-1}$$



(R modes also = bulk)

← Yukawa

← KK mass

$$m \sim 2.5 \times \frac{1/R}{1 \text{ TeV}}$$

Small

$$\xi = \frac{m}{1/R}$$

strength of mixing...

3.6.1



Intuitive picture:

(1 flavour).

few modes, ~ harmonic  
 $\xi \ll 1$



many modes  $\rightarrow$  chaotic, spikes.  
 $\rightarrow$  disappearance



"spread in extra-dim"

36.2

Survival

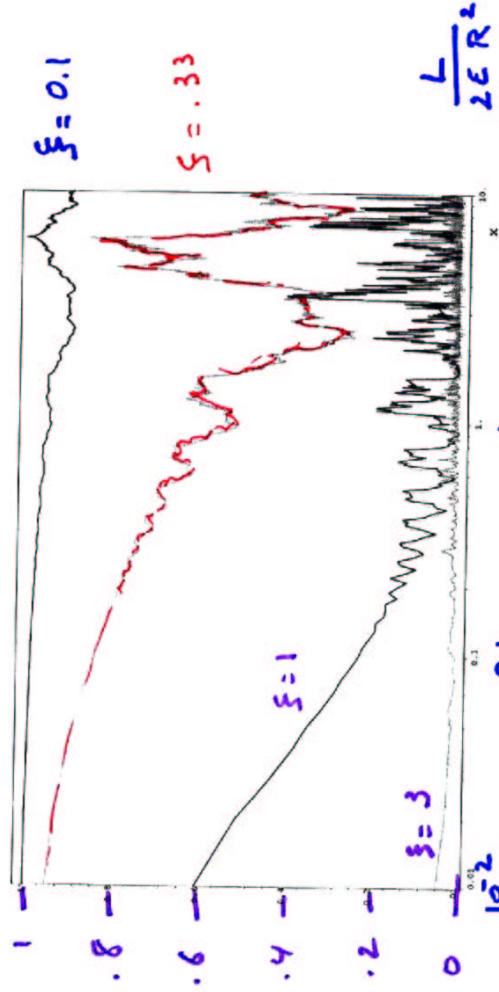


FIG. 1. Survival probability for different values of  $\xi$  ( $\xi = 1/10, 1/3, 1, 3$  from top to bottom), as functions of  $x = \frac{L}{2\epsilon R^2}$ .

Survival probability  $\sim L$

363

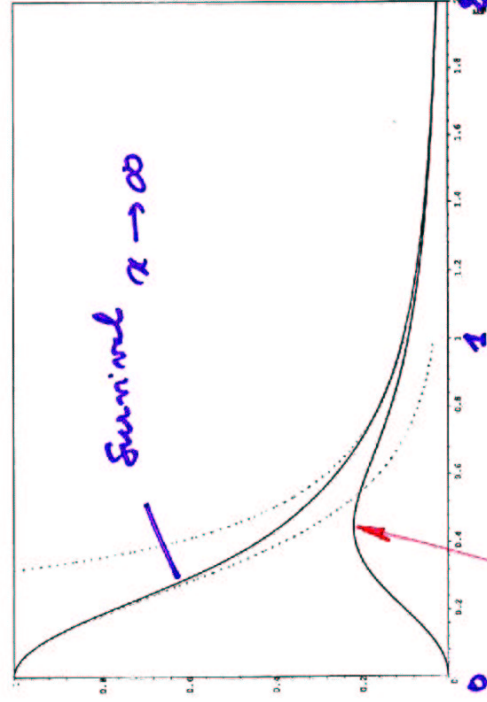


FIG. 3. Mean survival probability  $\langle P_{\nu_e \nu_e} \rangle$  and fluctuations  $\sigma(P)$  as functions of  $\xi$ . The dashed lines give the small  $\xi$  ( $\langle P_{\nu_e \nu_e} \rangle \approx (U_{00})^4$ ), and the large  $\xi$  ( $\langle P_{\nu_e \nu_e} \rangle \approx \frac{1}{2} \frac{U_{00}^2}{U_{10}^2}$ ) approximations.

$\sigma(P_{\text{survival}})$  fluctuations.

Come, see, Gouvea, Liu, Menden, V. Cheryck. 2001

3.5.4

Possibility for mixing:

$$\nu_e > \nu_{\text{back}}$$

$$\nu_\mu$$

$$\nu_e + \nu_\mu \text{ disappears}$$

$$\nu_e - \nu_\mu \text{ stays}$$

The  $\left\{ \begin{array}{l} \text{HEK} \\ \text{SNO} \end{array} \right\} \rightarrow \text{fine! (disappearance ..)}$

Supernova limit  $\rightarrow$  Removed (Remaining  $\nu_{\mu e}$ )

Now:

x need std oscillations for SNO.

x allow max amount of disappearance compatible with  $\delta B$  uncertainty.

but might be the only way to explore LED ... 3.5.5.

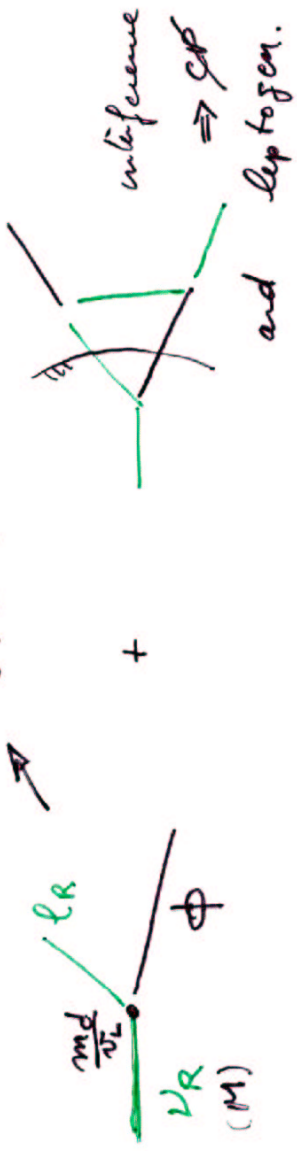
$\nu_R$  See-saw.  $\nu_R$ ? Only particle with "Majorana" (self-) mass?  
 $\Rightarrow$  Who ordered  $\nu_R$ ?  $\rightarrow$  arbitrary singlet  
 SM  $SU(3) \times SU(2) \times U(1)$   $\left\{ \begin{array}{l} 5 \\ 10 \end{array} \right.$   
 $SU(10) \supset SU(3) \times SU(2) \times SU(2) \times U(1)$   $\left\{ \begin{array}{l} 1 = \nu_R \\ 16 \end{array} \right.$   
 $\hookrightarrow$  anomaly-free

$\nu_R$  mass likely related to  $W_R$  mass  
 $\hookrightarrow \lambda_{Ai} \bar{\nu}_R$   
  

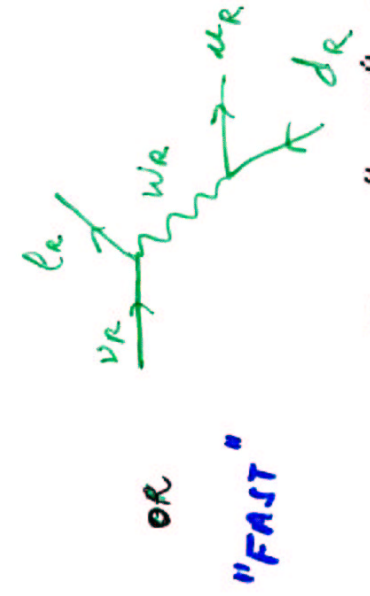
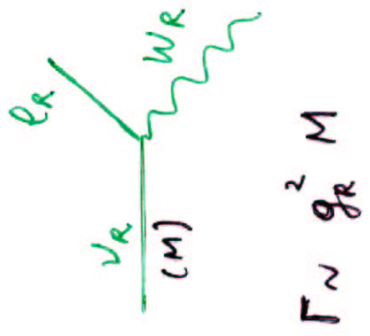

3.C.1

Dilution!

$$\Gamma \sim \sum_j \left| \frac{m_{dij}}{v} \right|^2 \cdot M.$$



$\frac{m_d}{v}$  small  $\Rightarrow$  out-of-equil. decay.



$$\Gamma \sim g_R^4 \frac{M^4}{M_{WR}^4} \cdot M.$$

3.C.2

Dilution:

usual result /  $(1 + \frac{g_R^4 \frac{M^4}{M_{WR}^4}}{\sum (\frac{m_{Dij}}{M_{WL}})^2 \cdot g_L^2})$

$\Rightarrow \frac{M_{WR}}{M_{WL}} < \sqrt[4]{\text{Max dilution acceptable}} \cdot \sqrt{\frac{\sum (m_{Dij})}{M_{WL}}}$

"Worst case": lightest  $M_1$  mostly related to  $\frac{\sqrt{\frac{m_{Dij}}{M_{WL}}}}$  by  $\sim M_e \sim \frac{1}{400}$

on the other hand, helps in meeting out-of-equil. condition...

$(\Gamma_R)^{-1} > \tau_{\text{expansion}} \approx \sqrt{g_*} \cdot \frac{T^2}{M_{\text{plck}}}$

3.C.3

With "typical" mix  $m_d \sim \begin{pmatrix} 0 \sim \sqrt{m_e m_\mu} & 0 \\ \sqrt{m_e m_\mu} \sim m_\mu & \sqrt{m_e m_\tau} \\ 0 & \sqrt{m_e m_\tau} \sim m_\tau \end{pmatrix}$   
 and  $\begin{pmatrix} M_1 & & \\ & M_2 & \\ & & M_3 \end{pmatrix}$  (max phase).

we get:

if lightest is	} $M_1$	$M_i >$
		$M_2$
		$M_3$
		$M_i >$
		$10^7 \text{ GeV}$
		$0.17$
		$3 \cdot 10^{11} \text{ GeV}$
		$0.11$
		$4 \cdot 10^{11} \text{ GeV}$
		$7 \cdot 10^{-3}$
		$M/M_{WR}$

(Carlier prof. Louis)

3.C.4



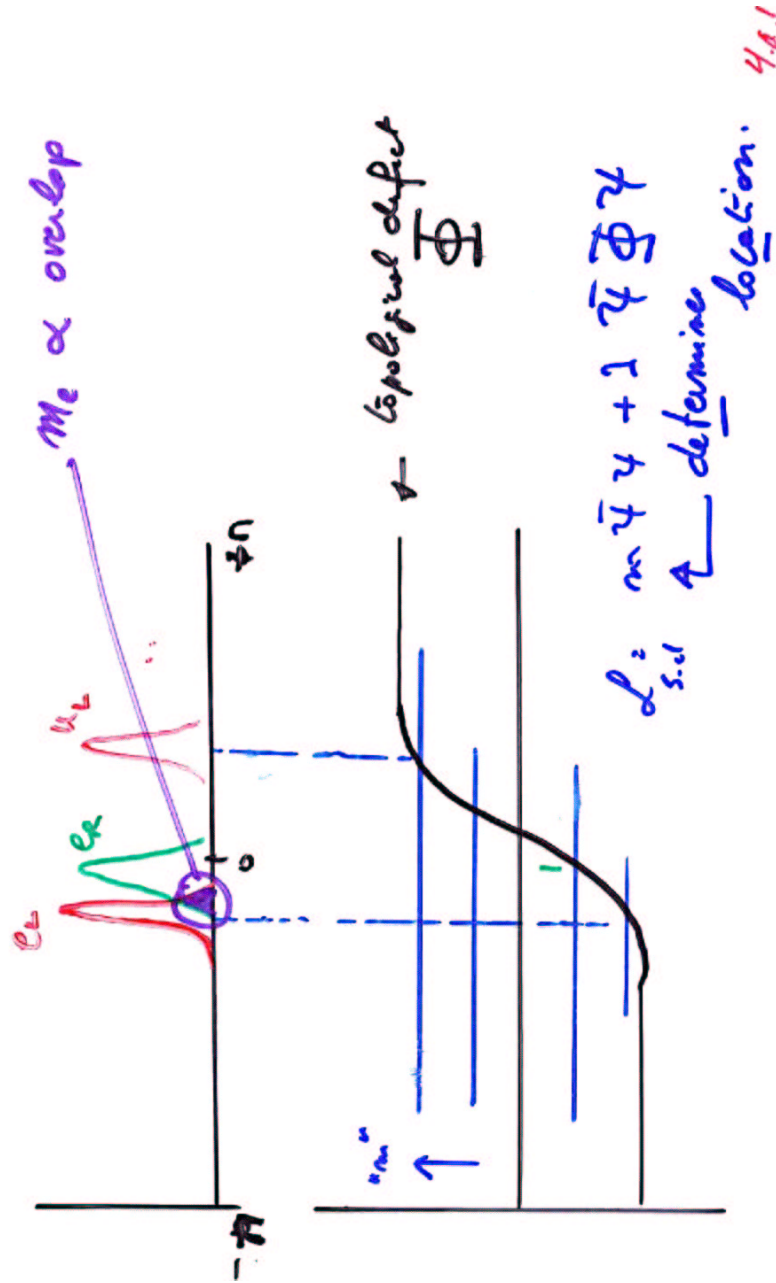
# Mass and ~~CP~~ from Extradim.?

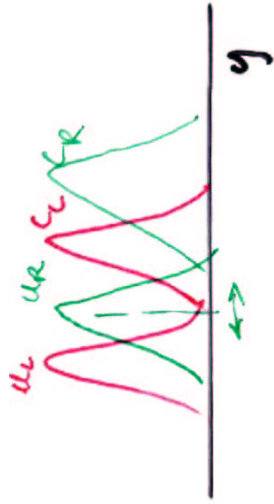
- how to generate mass pattern?

- horiz. sym, textures (refer. independt.?)
- overlaps in Extra-dimensions → 4.a

- how to generate ~~CP~~ in fundamental,  
 UNIFIED theory gauge is CP-conserving ✓  
 $\mathcal{L} \begin{matrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{matrix}$  Spontaneous ~~CP~~ breaking? → 4.b  
 extradim.? → 4.

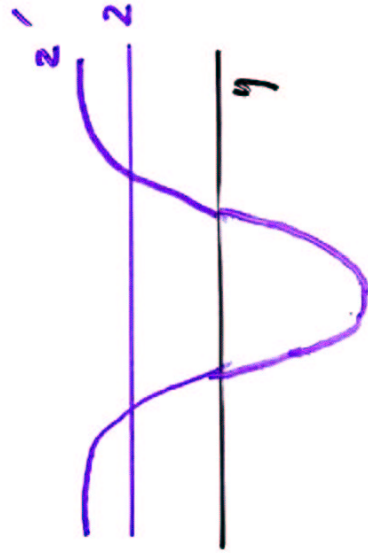
## 5th. dimension cartography? (a curvity popular idea)





Small  $\gamma$  shift  $\leftrightarrow$  large  $m_{4-d}$  change!

NOT PREDICTIVE FOR MASSES (FIT).



$z$ : flavor diagonal

$z'$ : F.C.N.C. + non-universal  $\sqrt{g}$

$\rightarrow$  LHC?   
 (first possible prediction)

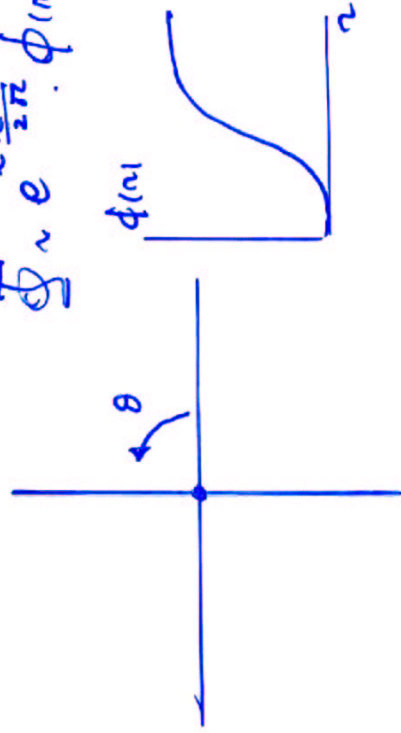
CP? by hand (Branco, Mohapatra... ) 4.a.2

A more constrained scheme:

(Trnicky-kebanov, IMF)

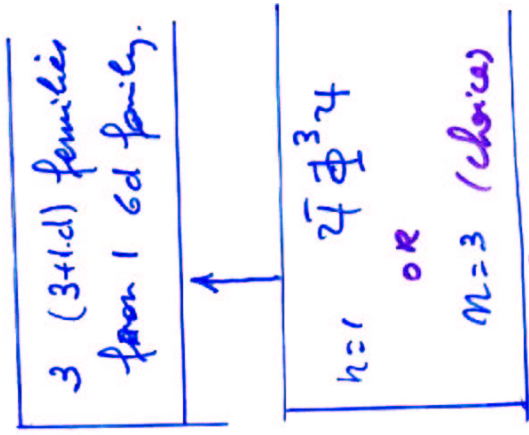
$(3+1) + 2$  dim.

$$\Phi \sim e^{i \frac{n\theta}{2\pi}} \phi(n)$$



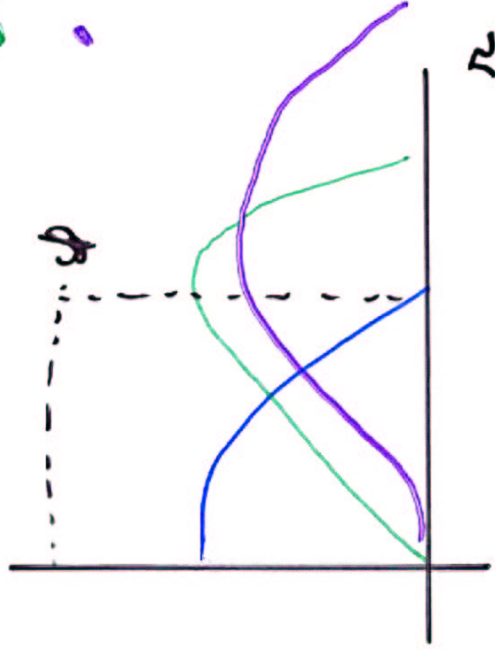
"Vortex solution"

Index Theorem:  $N$  massless chiral fermions bound to vortex.



4.a.3

$$\begin{aligned}
 \phi_1 e^{i\theta} &\sim \pi^c \\
 \phi_2 e^{i\frac{\theta}{2n}} &\sim \pi^c \\
 \phi_3 e^{i\frac{\theta}{2n}} &\sim \pi^c
 \end{aligned}$$



overlap with  $\phi$

$$\rightarrow \begin{pmatrix} m_1 & \times \\ \times & m_2 \\ \times & \times & m_3 \end{pmatrix}$$

$R$  hierarchy of masses.

A selection rule by winding nb.  
 add 1 Major field  $\rightarrow$   $\times$  terms.

4.4.4

• scheme is constrained, but  $\exists$  solutions!  $\nabla$

•  $Z'$ : dominant  $\sim e^{i\theta/2n}$



just FCNC  $\rightarrow K_L \rightarrow \bar{\mu} e$

$$m_{Z'} > 100 \text{ TeV}$$

•  $D$  can be embedded.

e.g.: leave  $U_R$  in bulk

• Constraints: form factor on 2d sphere  
 to have compact space.

4.4.5

~~CP~~ in fundamental theories?

- Gauge  $\leftrightarrow$  CP
- if fund. th. rests on gauge (N=2 sup, .. extra dim...)
- $\Rightarrow$  Why is CP broken?
- $\rightarrow$  Several "condensates"  $\rightarrow$  Spontaneous CP
- $\rightarrow$  geometrical  $\rightarrow$  toy model SD



4.6.1

(N. Grosse, double Lopen-H. prof)

5 dim = (3+1)+1

$$\mathcal{L} = \bar{\psi} (i \partial \not{\!} + g A \not{\!} \gamma_5 - \bar{\psi} M \psi)$$

CP-cons. real

(P odd.  $\therefore$  can explain)

$\gamma_4 = i \gamma_5$

5  $\rightarrow$  4 dim.

$A^{0,1,2,3} \Rightarrow A^\mu$   
 $A^4 \Rightarrow \phi$

$\Rightarrow$  can have  $\bar{\psi} \gamma_4 \psi = i \bar{\psi} \gamma_5 \psi$

4.6.2



⇒ mass term becomes:

$$\bar{\psi} \psi (M) + i \bar{\psi} \gamma_5 \psi \langle \phi \rangle$$

... complex

but  $\langle \phi \rangle = \langle A^4 \rangle$  not gauge invt.!!

-- in fact,  $\int dy A^4(y) \Rightarrow$  gauge invariant

if  $y$  compact,  $- \pi \rightarrow + \pi$

Hosotani  
(Wilson)



$$\oint A^4(y) dy = (\text{virtual flux of } \vec{B} \text{ in loop}) \Rightarrow \text{gauge invt.} \quad \text{4.6.3}$$

⇒ gauge  $M \in \mathbb{C}$  for 1 fermion

⇒ CP phase may be gauged away!

⇒ must involve internal group

$$\bar{\psi} (\alpha \partial^a + g \tau_i \gamma^B A^i_B) \psi$$

⇒ Hosotani loop  $\int \rightarrow$  breaks internal group  $\sim$  CP

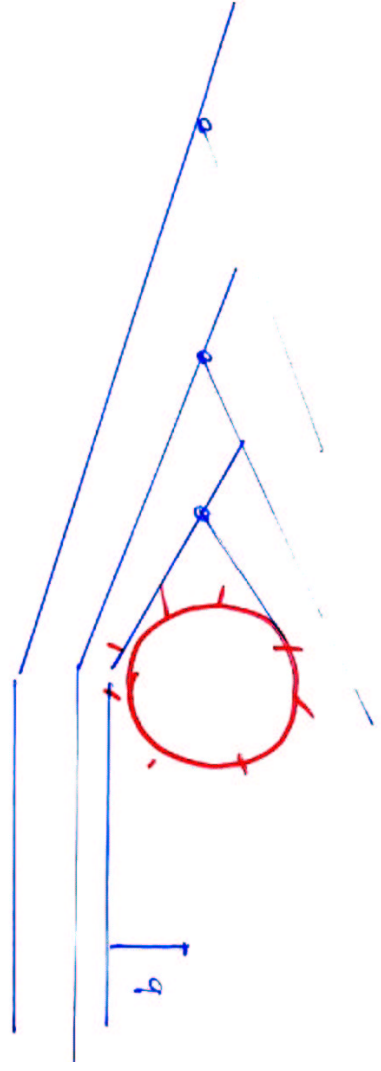
⇒ Toy examp with  $S(U(2)) \rightarrow$  edm / radiative?  $SO(11)$   
 can be extended to  $D \dots$  4.6.4

For fun: a TRUE  $\nu$  telescope

↳ (refractor)

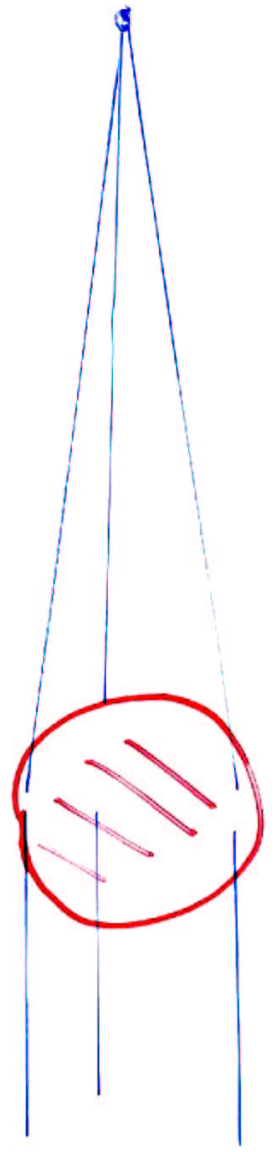
gravitational lensing

of  $\nu$ :  
 defl.  $\downarrow$  with  $b$   
 $\Rightarrow$  no focusing!



5.1

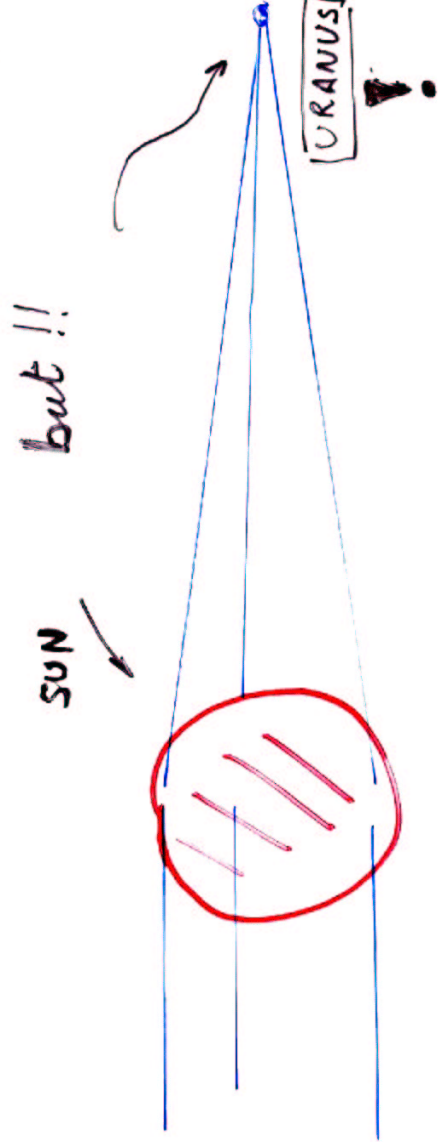
$\nu$  can cross through matter ( $E \ll 100 \mu\text{eV}$ )  
 deflection  $\nearrow$  with  $b$  for  $b < R$



$\rightarrow$  good focusing possible!

5.2

↳ can cross through matter  
 deflection ↗ with  $b$  for  $b < R$  ( $E \ll 100 \mu\text{eV}$ )

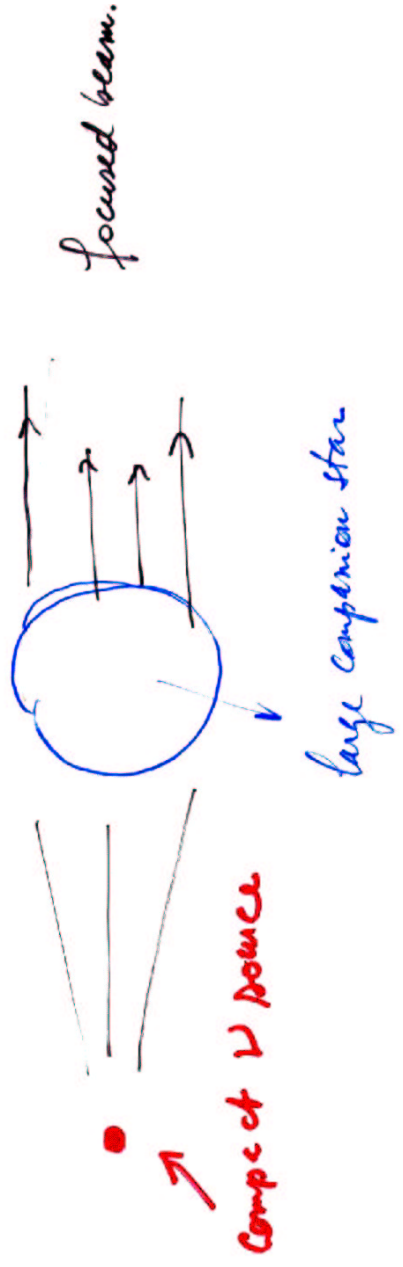


→ good focusing possible!  
 (too bad!)

5.3.2a

only possibility:

↳ light lense:



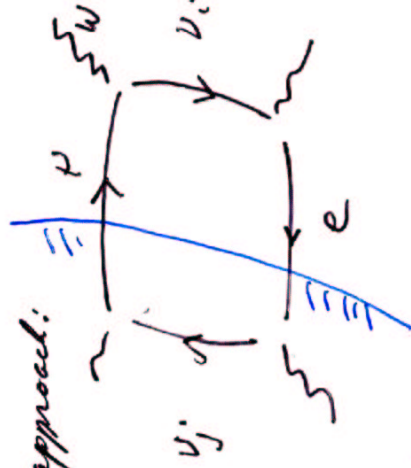
5.3

Current work:

How to make Jarlskog determinants appear explicitly in CP?   
 *Cecilia*

e.g:  $\nu_\mu \rightarrow \nu_e \neq \bar{\nu}_\mu \rightarrow \bar{\nu}_e$

try "unitarity graph" approach:



similar to  $p \rightarrow e$  in  $W$  background.

6.1

Expand in mass insertions:



... but why does the same process not  $\Rightarrow \mu \rightarrow e \gamma$  ?   
  $\neq \bar{\mu} \rightarrow \bar{e} \gamma$

would violate TCP: no compensating channel...

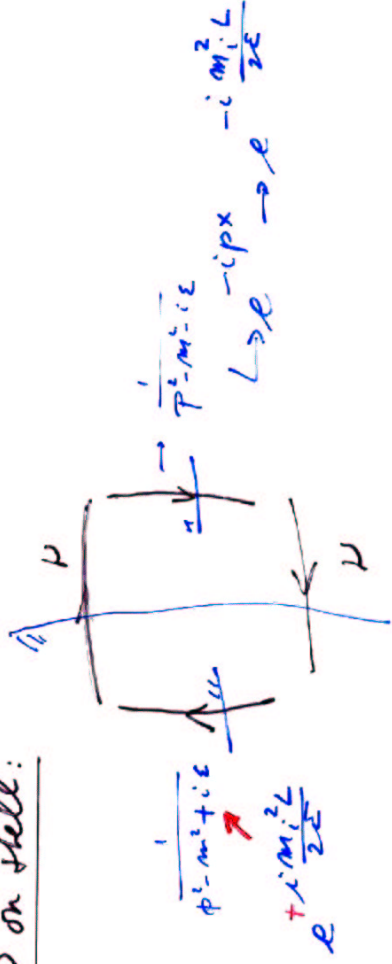
check indeed: Mass expansion in cut graphs  $\rightarrow 0$    
 at order  $(M/M^2)^3$ :  $[ (MM^\dagger)(MM^\dagger) ]_{2,1} [ (MM^\dagger) ]_{1,2}$    
  $+ [ (MM^\dagger) ]_{2,1} [ (MM^\dagger)(MM^\dagger) ]_{1,2}$    
  $\Rightarrow$  real

$\rightarrow$  need to put  $\nu$  on-shell   
 (impossible in  $\mu \rightarrow e \gamma$ )

6.2



With  $\nu$  on shell:



$\Rightarrow$  get at same order:  $\propto [(MM^\dagger)(MM^\dagger)]_{21} [MM^\dagger]_{12}$

$$- (MM^\dagger)_{21} [(MM^\dagger)(MM^\dagger)]_{12}$$

$$\rightarrow (m_1^2 - m_3^2)(m_2^2 - m_3^2)(m_3^2 - m_i^2) \cdot \text{Im}(U_{ii}^\dagger U_{2i}^\dagger U_{12} U_{21})$$



6.3

(exact result  $\propto \text{Im}(U_{12}) \cdot \sin \frac{m_{12}^2 L}{4E} \cdot \sin \frac{m_{23}^2 L}{4E} \sin \frac{m_{31}^2 L}{4E}$ )

(in a way, compensation is in time space, - or L-dependence - not visible in momentum space).

6.4