### Future Intensity Frontier Initiatives in the US Particle Physics Program

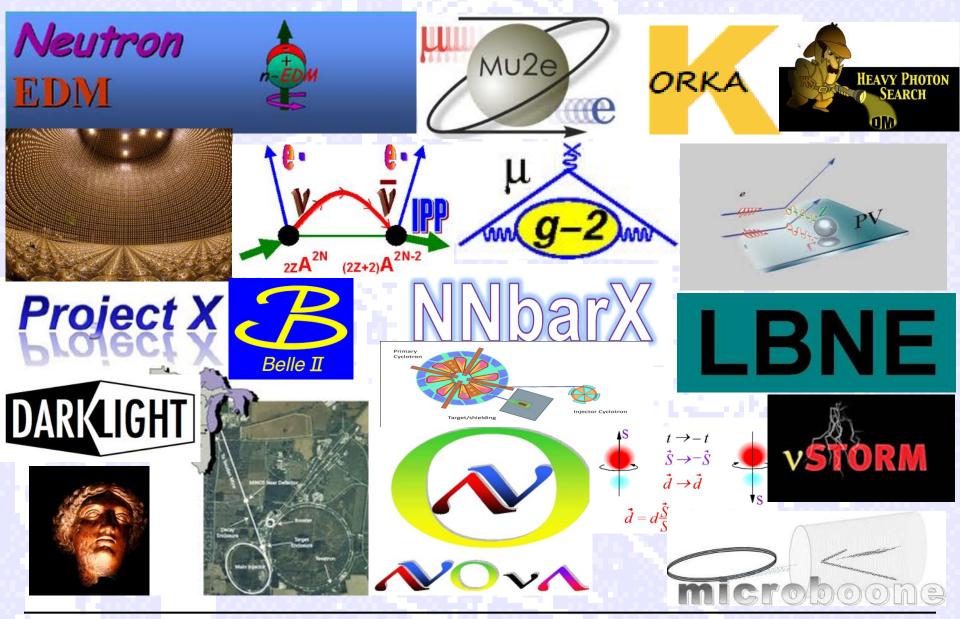
R. Tschirhart Fermilab KITP Snowmass on the Pacific May 29<sup>th</sup> 2013

### Material drawn from the Ongoing Snowmass Experience in 2013. A fraction of the IF activity...

- February:
- March:
- April:
- April:
- April:
- May:
- May:
- July:
- July:
- July-Aug:

Winter Workshop on Electric Dipole Moments, Fermilab. Neutrino Working group meeting, SLAC. Snowmass Workshop on Frontier Capability, BNL. Intensity Frontier All Hands Meeting, ANL. KAON-2013, Ann Arbor Michigan. First International Conference on Charged Lepton Flavor Violation, Lecce, Italy. International Symposium on Opportunities for Underground Physics in Snowmass, Asilomar, Ca. Journey through the Frontiers, SSI-2013, SLAC. Workshop on Neutrino Physics and Astrophysics, Lead/Deadwood, SD. SNOWMASS on the Mississippi.

### A Vibrant Community...



Snowmass on the Pacific @ KITP, May 29th 2013

### Facility subpanel recommendations accepted by the High Energy Physics Advisory Panel: HEPAP

- LBNE: Stage 1 begins a world leading program in neutrino physics +.... Science reach of Stage 1 is *important* and it lays the groundwork for an <u>absolutely central</u> facility. *Ready for construction*, planned start in 2016 and completed in 2023.
- MU2E: Will search for muon to electron conversion in the field of a nucleus with unparalleled sensitivity. It is *absolutely central. Ready for construction* starting in 2014, completed in 2018.
- **PROJECT X**: Unique world leading facility at Fermilab for intensity frontier physics. It is *absolutely central* and although it is pre CD0 it is *ready for construction*.
- nuSTORM: Muon storage ring that would provide neutrino beams with well defined flavor composition and spectrum. While the committee is not aware of major technical challenges in realizing nuSTORM, its performance requirements are not yet fully defined. While nuSTORM has great potential we don't know enough yet to assess nuSTORM's role in US world-leading science.

Mark Wise, HEPAP March 11th 2013

# In the absence of new facilities enabling new experiments...



From Hitoshi Murayama , ICFA October 2011

Snowmass on the Pacific @ KITP, May 29th 2013

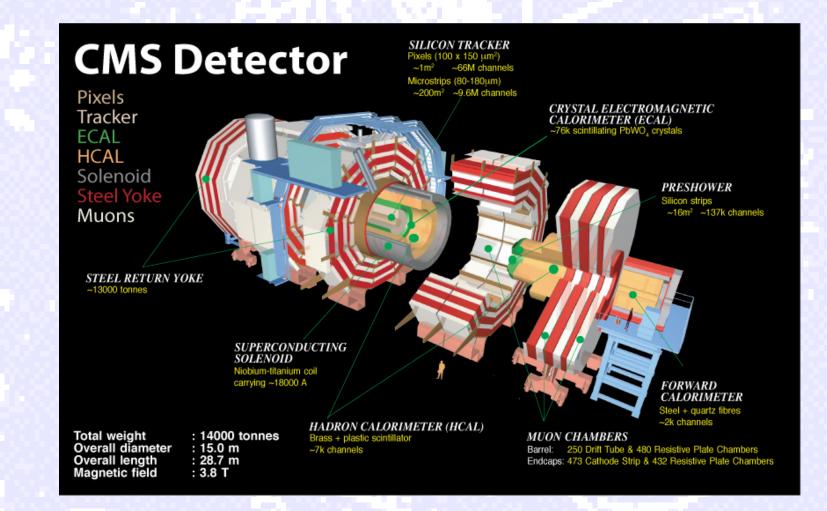
### Intensity Frontier Killer App? Not a single experiment! The science requires multiple probes.



Modified from Hitoshi Murayama , ICFA October 2011

Snowmass on the Pacific @ KITP, May 29th 2013

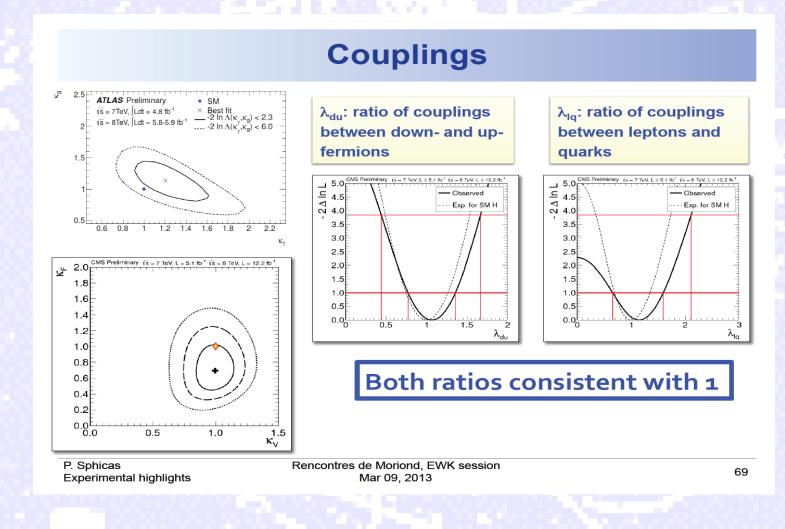
# What sub-detector in CMS or ATLAS is the Killer App?



Snowmass on the Pacific @ KITP, May 29th 2013

# What sub-detector in CMS or ATLAS is the Killer App?

For Discovery of the Standard Model Scalar??



### **CSS13 Working Groups**

Quark Flavor Physics: Joel Butler, Zoltan Ligeti, Jack Ritchie

Charged Lepton Processes Brendan Casey, Yuval Grossman, David Hitlin

Neutrinos Andre deGouvea, Kevin Pitts, Kate Scholberg, Sam Zeller

Baryon Number Violation Kaladi Babu, Ed Kearns

New Light, Weakly Coupled Particles Rouven Essig, John Jaros, William Wester

Nucleons, Nuclei & Atoms Krishna Kumar, Z.-T. Lu, Michael Ramsey-Musolf K, D & B Meson decays/properties

Precision measurements with muons, taus

SLAC

All experiments for properties of neutrinos. Accelerator & non-accel.

Proton decay, Neutron Oscillation

"Dark" photons, paraphotons, axions, WISPs

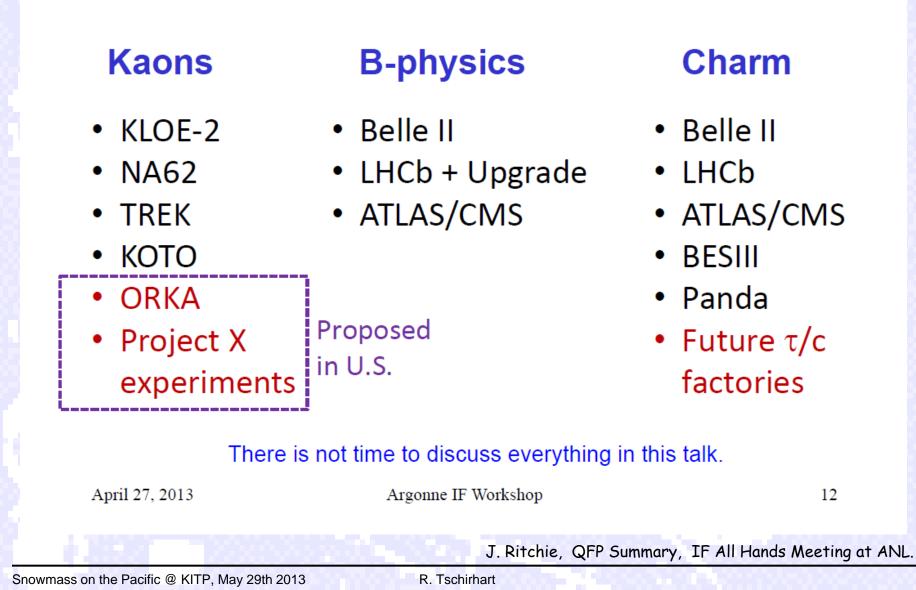
Properties of nucleons, nuclei or atoms (EDM), as related to HEP

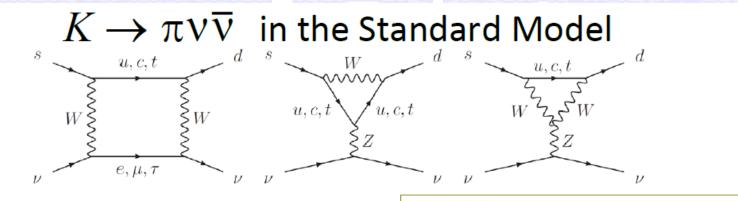
J. Hewett, IF All Hands Meeting at ANL.

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### **Quark Flavor Physics**

### The Experimental Quark Flavor Program



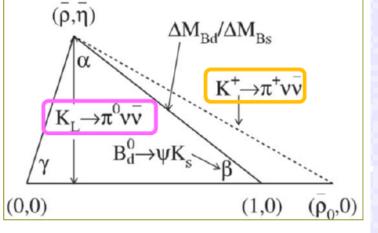


- A single effective operator  $(\overline{s}_L \gamma^{\mu} d_L) (\overline{v}_L \gamma_{\mu} v_L)$
- Dominated by top quark
- Hadronic matrix element shared with  $K \rightarrow \pi e v$
- Standard Model predictions precise

$$B(K^+_{I} \to \pi^+ \nu \overline{\nu})_{\rm SM} = (7.8 \pm 0.8) \times 10^{-11}$$
$$B(K^0_{I} \to \pi^0 \nu \overline{\nu})_{\rm SM} = (2.4 \pm 0.4) \times 10^{-11}$$

Brod, Gorbahn, and Stamou, PR D 83, 034030(2011)

- Largest uncertainty from CKM elements (which will improve)
- Remains clean in New Physics models
   April 27, 2013
   Argonne IF Workshop



$$\begin{array}{r} \pm 10\% \Rightarrow \pm 5\% \\ \pm 16\% \Rightarrow \pm 11\% \end{array}$$

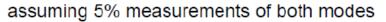


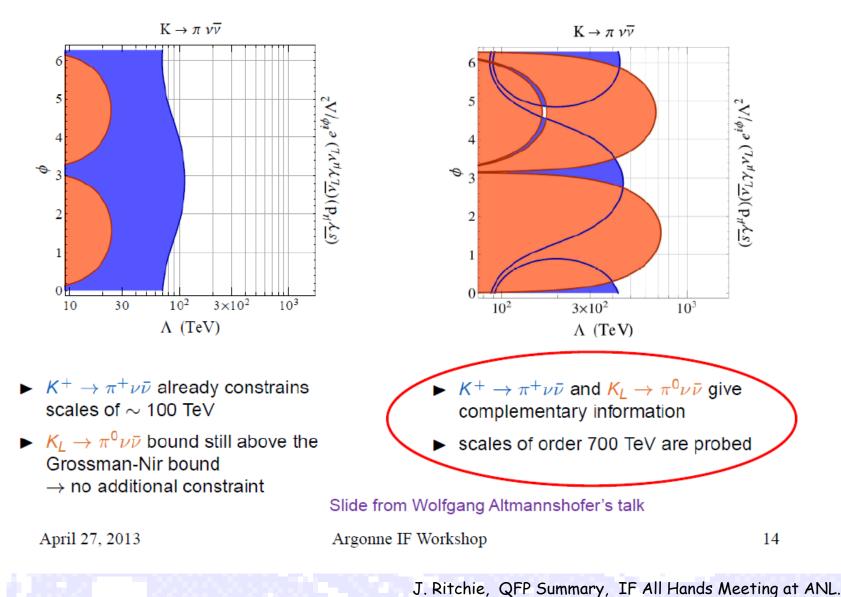
J. Ritchie, QFP Summary, IF All Hands Meeting at ANL.

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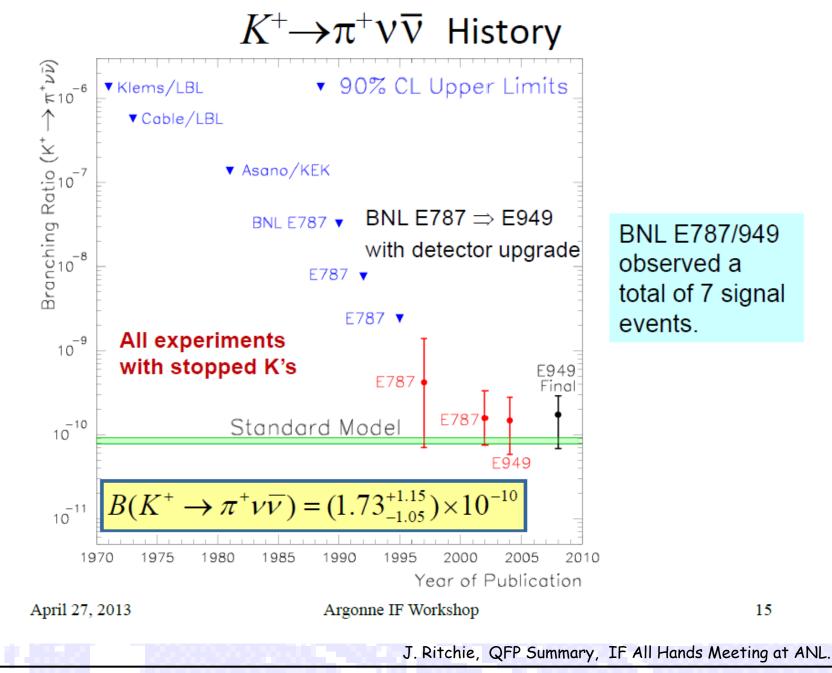
 $(\overline{s\gamma}^{\mu}d)(\overline{\nu_L}\gamma_{\mu}\nu_L) e^{i\phi}/\Lambda^2$ 

current situation

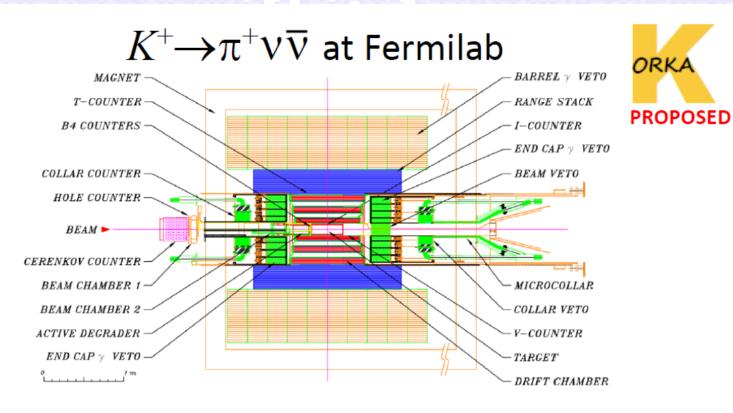




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Snowmass on the Pacific @ KITP, May 29th 2013



- Will employ the proven technique (stopped K) from BNL E787/949, with beam, detector, and data acquisition improvements
  - Goal to collect 1000 events, ~200 events per year (error  $\cong$  theory)
  - Does not require better background rejection than E949
- Will utilize existing facilities and infrastructure at FNAL (Main Injector protons, B0/CDF Hall, CDF superconducting solenoid)

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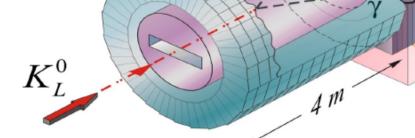
J. Ritchie, QFP Summary, IF All Hands Meeting at ANL.

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### $K_{l} \rightarrow \pi^{0} \nu \overline{\nu}$ at Project X

#### Based on the KOPIO concept

- Project-X beam energy is well-suited
- CW-linac time structure supports time-of-flight measurement
  - Provides kinematic info for background rejection



- High intensity from Project-X allows small beam (like KOTO)
  - 2-dimn constraint provides additional background rejection
- Reconstruct  $\pi^0 \rightarrow \gamma \gamma$  with a pointing calorimeter
- $4\pi$  photon and charged particle vetos
- Provides opportunity for a high statistics (~1000 event) measurement

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J. Ritchie, QFP Summary, IF All Hands Meeting at ANL.

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### **Kaon Projections**

#### a few K observables

Observable	SM Theory	Current Expt.	Future Experiments
$\mathcal{B}(K^+ \to \pi^+ \nu \overline{\nu})$	$7.8 \times 10^{-11}$	$1.73^{+1.15}_{-1.05} \times 10^{-10}$	${\sim}10\%$ measurement from NA62
			${\sim}5\%$ measurement from ORKA
			$\sim 2\%$ with Project X
$\mathcal{B}(K^0_L \to \pi^0 \nu \overline{\nu})$	$2.43 \times 10^{-11}$	$< 2.6 \times 10^{-8}$	1 <sup>st</sup> observation from KOTO
			${\sim}5\%$ measurement with Project X
$\mathcal{B}(K_L^0 \to \pi^0 e^+ e^-)_{SD}$	$1.4 \times 10^{-11}$	$< 2.8 \times 10^{-10}$	${\sim}10\%$ measurement with Project X
$\mathcal{B}(K^0_L \to \pi^0 \mu^+ \mu^-)_{SD}$	$3.5 \times 10^{-11}$	$< 3.8 \times 10^{-10}$	${\sim}10\%$ measurement with Project X
$ P_T $ in $K^+ \to \pi^0 \mu^+ \nu$	$\sim 10^{-7}$	< 0.0050	< 0.0003 from TREK
			< 0.0001 with Project X
$R_K = \Gamma(K_{e2}) / \Gamma(K_{\mu 2})$	$2.477\times 10^{-5}$	$(2.488 \pm 0.080) \times 10^{-5}$	$\pm 0.054 \times 10^{-5}$ from TREK
			$\pm 0.025 \times 10^{-5}$ with Project X
$\mathcal{B}(K^0_L \to \mu^{\pm} e^{\mp})$	$< 10^{-25}$	$<4.7\times10^{-12}$	$< 2 \times 10^{-13}$ with Project X

From the Report of the Heavy Quarks working group, Fundamental Physics at the Intensity Frontier (2012), arXiv:1205.2671

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J. Ritchie, QFP Summary, IF All Hands Meeting at ANL.

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### Message for Snowmass

- Flavor physics probes far above the TeV scale.
  - A necessary complement to LHC if new physics is found there.
  - Probes above the reach of LHC and other foreseeable machines.
- Existing facilities at Fermilab can support unparalleled rare K decay experiments (ORKA, and potentially others).
  - A cost effective way to mount quark-flavor experiments in this decade with significant potential to uncover new physics.
  - This opportunity is not open-ended (the world won't wait).
- Project X can open a new regime of sensitivity for rare K decay experiments in the next decade.
  - An order of magnitude beyond other kaon sources in the world.
- B-physics and charm physics will be led by non-U.S. programs for the foreseeable future.
  - These programs will do great physics! The U.S. should be actively involved in these experiments (Belle II and LHCb).

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### **Charged Lepton Processes**

#### Status of CLFV Searches MU2e 1 10<sup>-1</sup> Limit $\mu \rightarrow e\gamma$ 10<sup>-3</sup> • $\mu \rightarrow 3e$ 10<sup>-5</sup> $\mu N \to eN$ 10-7 **MEG 2013** arXiv:1303.0754 10<sup>-9</sup> 10-11 10<sup>-13</sup> 4 orders of magnitude MEG Upgrade improvement projected! 10<sup>-15</sup> PSI, MUSIC $\bigcirc$ Mu2e, COMET 10-17 Project X, PRIME 10<sup>-19</sup> 1940 1950 1960 1970 1980 1990 2000 2010 2020 2030 Year David Brown, Lawrence Berkeley National Lab н CLFV, Lecce, 2013 mu2e conversion at FNAL

Snowmass on the Pacific @ KITP, May 29th 2013

## Atomic Capture of µ-

Muzee

DIO

39%

- Stopped  $\mu^{-}$  is captured by an atom  $E_{bind} \sim 500 \text{ KeV}$ 
  - Falls to K-shell
  - Binding energy emitted as x-rays
- μ<sup>-</sup> can decay-in-orbit (DIO)
  - EM coupling to nucleus
- μ<sup>-</sup> can be captured by nucleus
  - resultant nucleus is unstable

440 fm

David Brown, Lawrence Berkeley National Lab

mu2e conversion at FNAL

27**A** 

27 A

CLFV, Lecce, 2013

Nuclear

Capture

61%

Snowmass on the Pacific @ KITP, May 29th 2013

0

R. Tschirhart

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### The Mu2e Experiment



- Goal: Discover µN→eN conversion
- Target sensitivity:  $R_{\mu e} = 6 \times 10^{-17}$  @ 90% C.L.
  - 4 orders of magnitude better than current limits
- Requires ~ 10<sup>18</sup> stopped muons
  - ~ 4×10<sup>20</sup> protons on target (3 year run @ 8 KW)
- Requires negligible (<1) background events</li>
- Many challenges for the beamline and detector design

David Brown, Lawrence Berkeley National Lab

12

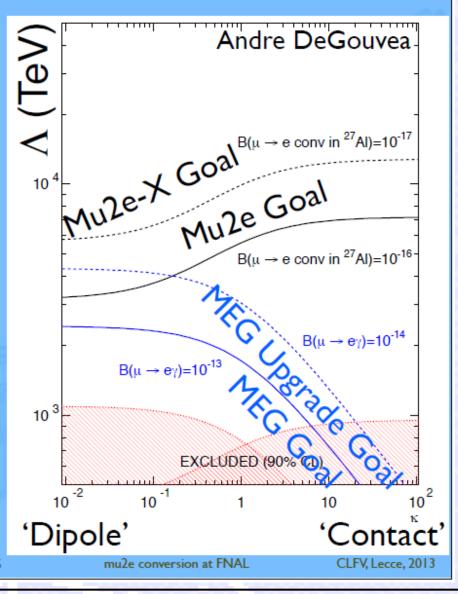
mu2e conversion at FNAL

CLFV, Lecce, 2013

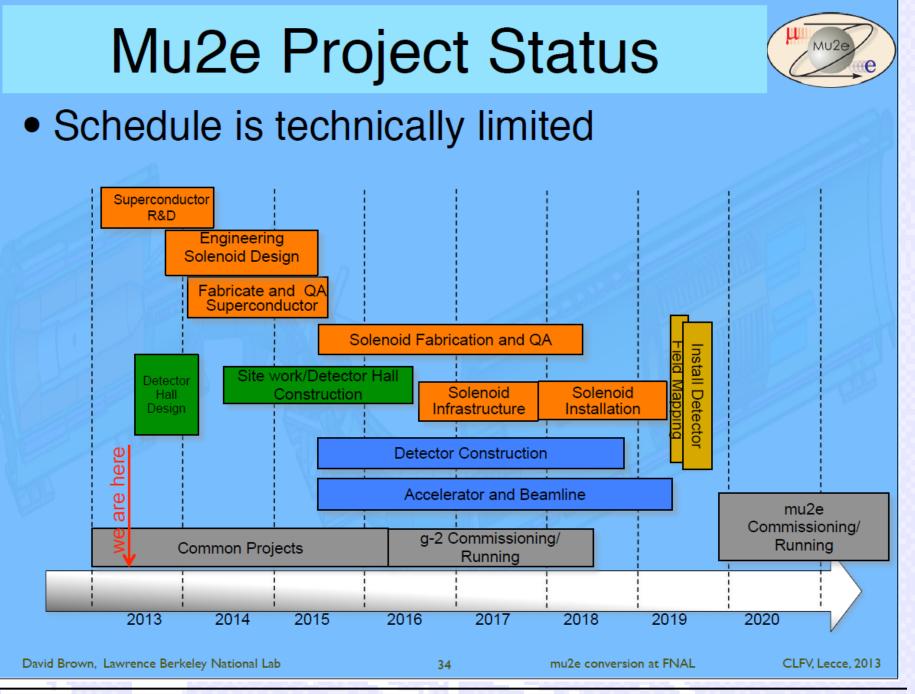
## **CLFV** Sensitivity



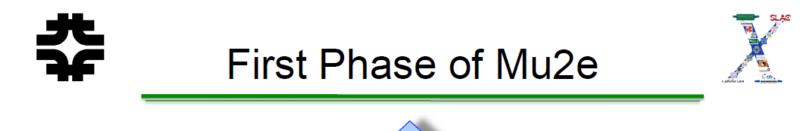
- Effective mass scales up to 10<sup>4</sup> TeV are accessible
- Mu2e is sensitive over the full  $\kappa$  range

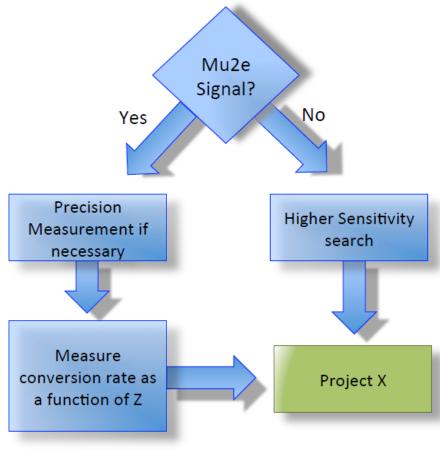


David Brown, Lawrence Berkeley National Lab



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R. Ray - Intensity Frontier Workshop

Snowmass on the Pacific @ KITP, May 29th 2013

April 26, 2013



### Mu2e Sees a Signal



V<sup>2</sup>

D

80

V. Cirigliano et al., phys. Rev. D80 013002 (2009)

3

2

1

20

670 ns

600

800

400

200

40

1695 n

Aluminum

POT pulse shape Muon Arrival Time Muon Decay Time

Search Window

1200

1400

1600

1800

Time (ns)

7

60

 $R_{\mu\nu}(Z)/R_{\mu\nu}(Al)$ 

0.1

0.08

0.06

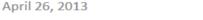
0.04

0.02

- If Mu2e sees a signal in its initial run, next step is to map out conversion rate for various target nuclei where model dependent effects vary by a factor of 3.
- However, muon lifetime varies with Z
  - Big impact on execution of measurement

Nucleus (Z)	Muon Lifetime (ns)
Al(13)	864
Ti(22)	329
Au(79)	73

Arrival time of  $\mu/\pi$  at stopping target Increase in muon decays and RPC as one looks earlier in time



R. Ray - Intensity Frontier Workshop



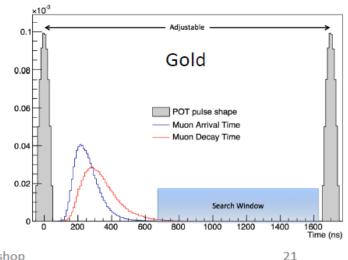


### Mu2e Sees a Signal



Stage 1 of Project X makes these measurements possible with the existing solenoid system.

- Narrower proton beam pulse, intrinsic extinction and beam power enables use of stopping targets with shorter muon lifetimes
  - Limit to how early we can search due to muon decay in-flight and RPC background, assuming same background level required.
  - Rely on beam power to wait more muon lifetimes
    - Instantaneous rates decrease with time – reduces backgrounds
  - Flexible time structure allows us to wait longer than 1695 ns if desirable
    - Optimization different for different target nuclei.



April 26, 2013

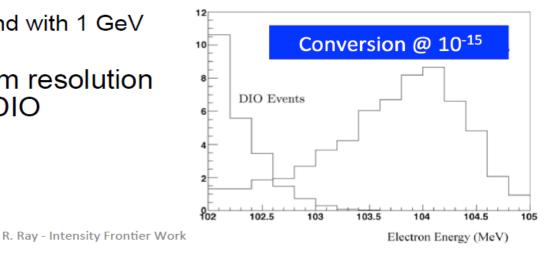
R. Ray - Intensity Frontier Workshop





Stage I of Project X makes it possible to push the sensitivity another order of magnitude.

- Reductions in background also required. (See Doug Glenzinski's talk)
  - Narrow proton beam pulse and intrinsic extinction provided by Project X reduces prompt backgrounds.
  - No pbar background with 1 GeV protons.
- Improved momentum resolution required to reduce DIO background.



April 26, 2013

### Neutrinos

### Final thoughts on the message for Snowmass

#### Yuval's talk on first day:

"Once you find an entrance, there will be an explosion in some direction that will carry on for decades"



That's happened for neutrinos!

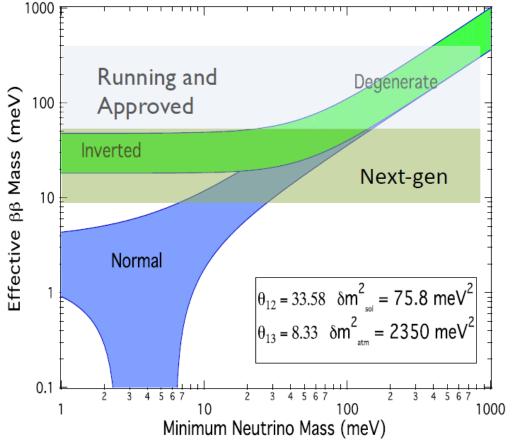
## We can build a world-class neutrino program along three lines:

- long-baseline oscillations
- neutrinoless double beta decay
- smaller experiments to search for new physics

#### Breadth, and connections between Frontiers, are important

CSS-2013 IF All Hands meeting April 25th-27th ANL

### Goals for Next-Generation $0\nu\beta\beta$



- Majorana vs Dirac
- Absolute Neutrino Mass
- Lepton Number Violation

• Next-generation  $\beta\beta$ experiments must cover the entire allowed region of the inverted hierarchy

• Ideas for probing the normal hierarchy exist

Elliot, Kauffman SLAC Snowmass neutrino meeting, March 2013

Snowmass on the Pacific @ KITP, May 29th 2013

# Several Experiments Running or Nearly Running to get down to 100 meV $m_{\beta\beta}$ -scale

- <sup>136</sup>Xe
  - EXO-200 and KamLAND-Zen currently running
  - Combined result:  $m_{\beta\beta} < 120 250 \text{ meV}$
  - NEXT to be running in 2014
- <sup>76</sup>Ge
  - GERDA running
  - MAJORANA DEMONSTRATOR coming online in the next few months
- Tellurium
  - CUORE0 online
  - CUORE online in 2015
- Selenium
  - SuperNEMO Demonstrator online in 2015
- Neodymium
  - SNO+ will come online in 2014

Elliot, Kauffman SLAC Snowmass neutrino meeting, March 2013

### Several Ideas to Get US to the Inverted Hierarchy

- Several isotopes and several experiments to exploit them
  - Xenon
    - nEXO (Liquid XeTPC)
    - NEXT (High pressure Xe Gas EL TPC)
    - KamLAND-Zen
  - Germanium
    - MAJORANA/GERDA
  - Tellurium
    - CUORE/Enriched CUORE
  - Selenium
    - SuperNEMO
  - Molybdenum
    - MOON
  - Neodymium
    - SNO+/Enriched SNO+

Elliot, Kauffman SLAC Snowmass neutrino meeting, March 2013

Snowmass on the Pacific @ KITP, May 29th 2013

# Theory Issues for 0νββ

- Support for nuclear theory effort on matrix elements
  - Auxiliary measurements to support understanding the matrix elements
- Support for particle physics theory efforts on exchange mechanism

Elliot, Kauffman SLAC Snowmass neutrino meeting, March 2013

From Michael Salamon at DURA Meeting this week.

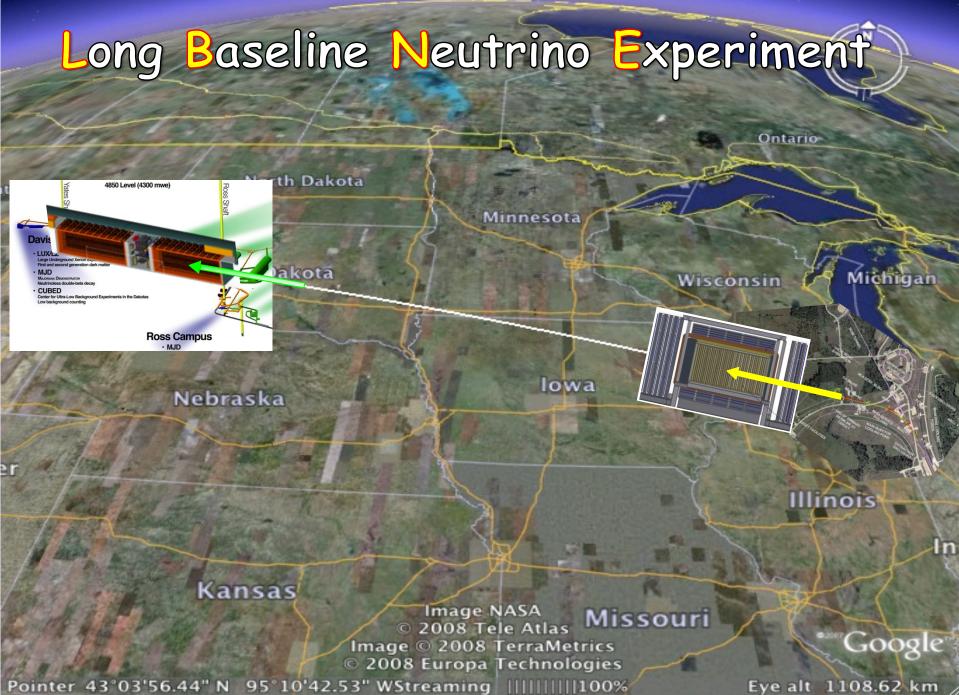
#### **DOE Double Beta Decay: Comments**

- DOE/Nuclear Physics is the steward for next-generation double beta decay experiments at DOE.
- DOE/HEP, however, is supporting EXO-200 for historical reasons, along with DOE/NP research and NSF support
- DOE/HEP (along with NSF) also is supporting all the R&D activities for the proposed 1-tonne scale next generation EXO, "nEXO."
- DOE/HEP and NP will establish a joint process to determine a selection process that involves both HEP and NP communities.
- After the time of selection, DOE/NP will become the sole DOE office supporting next-generation DBD projects.

A significant amount of the **technologies** and **facilities** used for double-beta decay overlap with the dark matter community, funded by DOE HEP.

Elliot, Kauffman SLAC Snowmass neutrino meeting, March 2013

LBNE-doc-7332

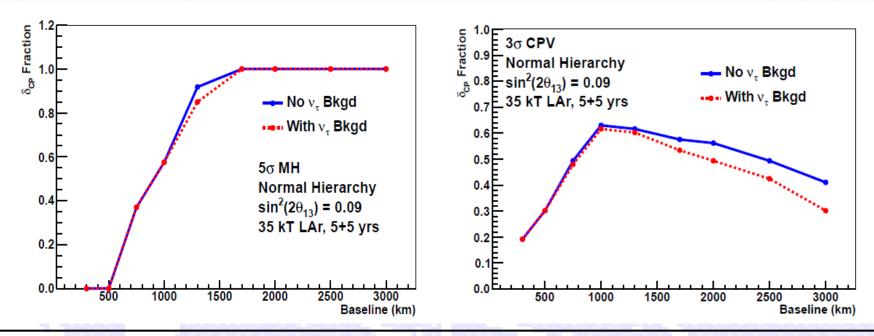


# LBNE is...

- A new neutrino beam at Fermilab
  - 700 kW proton beam, 2.3 MW capable
- A near neutrino detector
- An optimal 1300 km baseline: Fermilab-SURF
- A 34 kt Liquid Argon TPC with 4850' overburden
- This conceptual design...
  - Completed a successful CD-1 Director's Review (March 2012)
  - Updated cost estimate (July 2012): ~\$1.5B (incl. contingency + escalation)

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Snowmass on the Pacific @ KITP, May 29th 2013

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# Staging the Experiment

- March 2012 DOE asked us to stage LBNE construction
- An external review panel considered reconfiguration options including different far sites
  - We accepted the recommendation to proceed with emphasis on the most important aspects: 1300 km baseline and the full capability beam
- December 2012: CD-1 approval for \$867M first phase DOE funding
  - We have completed an extensive cost/schedule for 10 kt LAr far detector (LBNE10) on the surface but the design is **not** fixed
  - CD-1 approval explicitly allows for scope change enabled by new partners
- First phase goal: <u>greater than 10 kt</u> far detector <u>underground</u> and a <u>full</u> <u>capability near detector</u>
- In the past 3 months there has been considerable progress towards international partnerships (encouraged by European Strategy statement)

LBNE Leadership team, May 2013

# Strawman Plan

### DOE initial investment of \$867M

Additional Investment (TPC)	Capability Added	Science Gained	Science Priority
+\$140M	Underground placement	ATM nus, p-decay, SNB nus	Very High
+\$130-190	Near Detector	Enhanced LB physics, near detector physics	Very High
+\$200-350	Add FD mass	Precision CP and other 3-flavor paradigm measurements	Very High

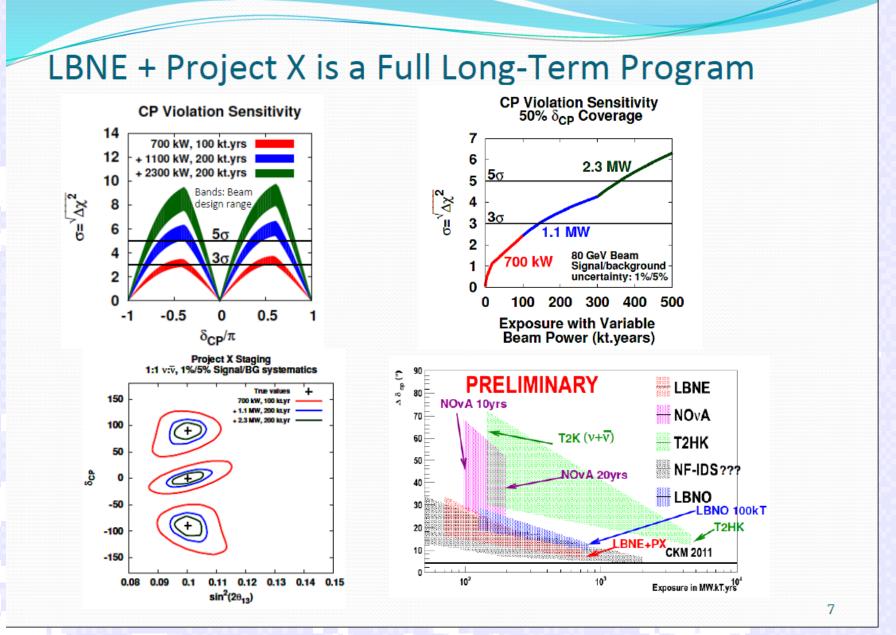
LBNE Leadership team, May 2013

Snowmass on the Pacific @ KITP, May 29th 2013

## LBNE10 Alone is Major Advance

δ<sub>CP</sub> Resolution **CP Violation Sensitivity** 1σ Resolution (degrees) 80 T2K+NOvA+LBNE10 5 T2K+NOvA T2K NOvA LBNE10 70 4 60 50  $\sigma = \sqrt{\Delta \chi^2}$ 30 3 40 LBNE10 30 2 20 1 10 2K+NOvA+LBNE10 0 0 -0.5 0.5 -1 0 1 -0.5 0.5 -1 0 1 δ<sub>CP</sub>/π Mass Hierarchy Sensitivity  $\delta_{CP}/\pi$ 10 T2K+NOvA+LBNE10 T2K+NOvA T2K NOvA LBNE10 8  $\sigma = \Delta \chi^2$ PRELIMINARY 6 5σ Bands: 1 $\sigma$  variations of  $\theta_{13}$ ,  $\theta_{23}$ ,  $\Delta m_{31}^2$ 4 (Fogli et al. arXiv:1205.5254v3) 2 0 -0.5 0.5 -1 0 1 6  $\delta_{CP}/\pi$ LBNE Leadership team, May 2013

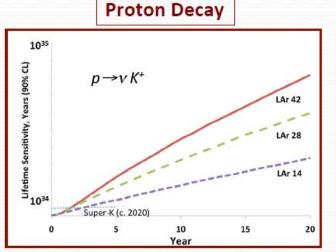
Snowmass on the Pacific @ KITP, May 29th 2013



#### LBNE Leadership team, May 2013

Snowmass on the Pacific @ KITP, May 29th 2013

## Underground Science

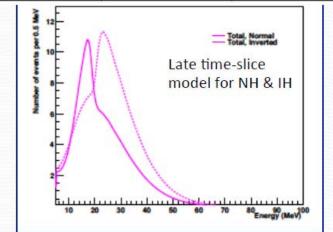


- LAr has high efficiency for SUSY-favored decay modes
- High spatial precision and energy resolution enable reconstruction of many potential decays modes

#### **Atmospheric Neutrinos**

#### Supernova Burst Neutrinos

Channel	Events, "Livermore" model	Events, "GKVM" mode
$\nu_e + {}^{40}\operatorname{Ar} \rightarrow e^- + {}^{40}\operatorname{K}^*$	2308	2848
$\nu_e + {}^{40}\operatorname{Ar} \to e^- + {}^{40}\operatorname{K*}$ $\bar{\nu}_e + {}^{40}\operatorname{Ar} \to e^+ + {}^{40}\operatorname{Cl*}$	194	134
$\nu_x + e^- \rightarrow \nu_x + e^-$	296	178
Total	2794	3160



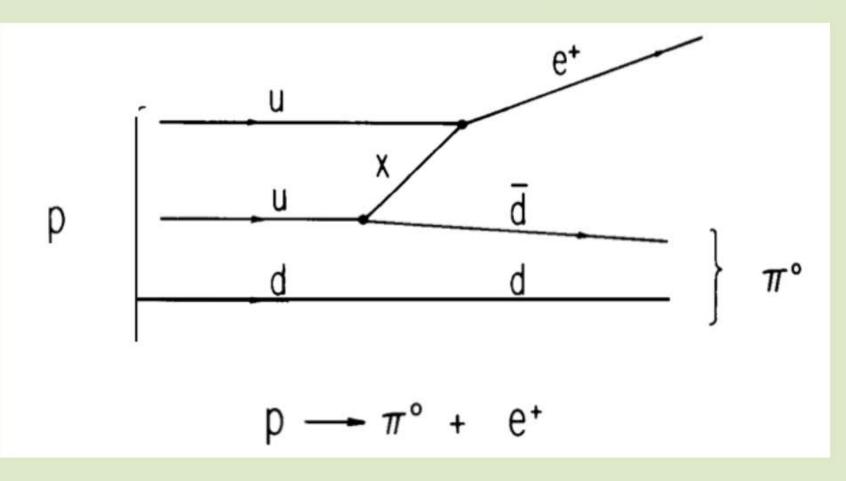
- SN at galactic core (10 kpc) 1000s interactions in 35 kt LAr in 10s of seconds
- Complementary to WCD
- Fantastic for particle physics and astrophysics (c.f. SN1987A ~dozen events significance)
- Independent determination of mass hierarchy (or add significance to beam measurement)
- $\theta_{23}$  octant sensitivity
- Nu-e sensitivity complementary to water Cherenkov detectors (anti-nu-e)

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LBNE Leadership team, May 2013

# **Baryon Number Violation**

# **Proton Decay**



Treasure hunt: from which Nobel Lecture is this diagram taken?

E. Kearns, Intensity Frontier All Hands Meeting, ANL

Snowmass on the Pacific @ KITP, May 29th 2013

## **Efficiency and Background Rates**

A. Bueno et al. hep-ph/0701101

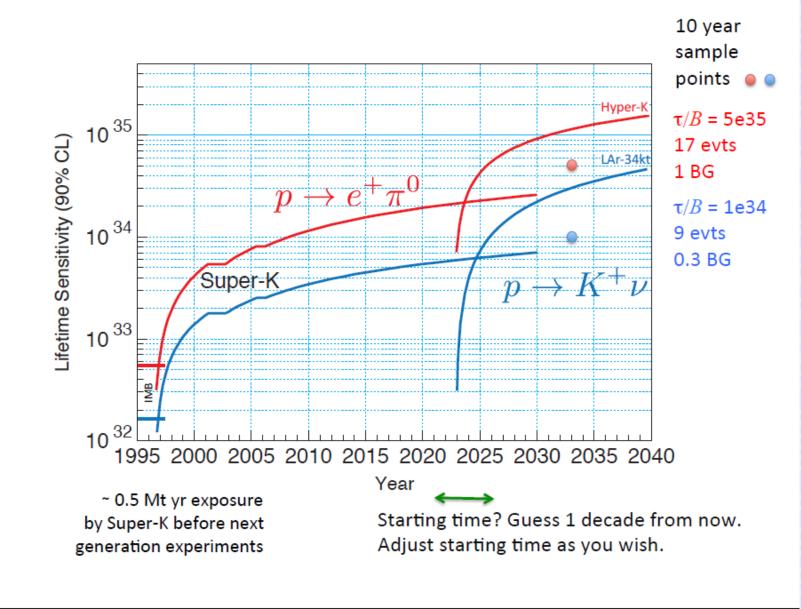
		Super-K Water Ch.		LAr (generic)		
	Mode	Efficiency	BG Rate (/Mt y)	Efficiency	BG Rate (/Mt y)	
	$e^+\pi^0$	45%	2	45% (?)	1	
B-L	ν Κ+	15%	2*	97%	1	
	μ+ Κ <sup>ο</sup>	8%	8	47%	<2	
B+L	μ⁻ π⁺ K⁺	?	?	97%	1	
	e⁻ K⁺	10%	3	96%	<2	
∆B=2	n nbar	12%	260	?	?	

For many modes, high efficiency and low BG rate makes up for smaller mass of Lar detectors

\* New analysis (Miura, BLV Heidelberg)

E. Kearns, Intensity Frontier All Hands Meeting, ANL

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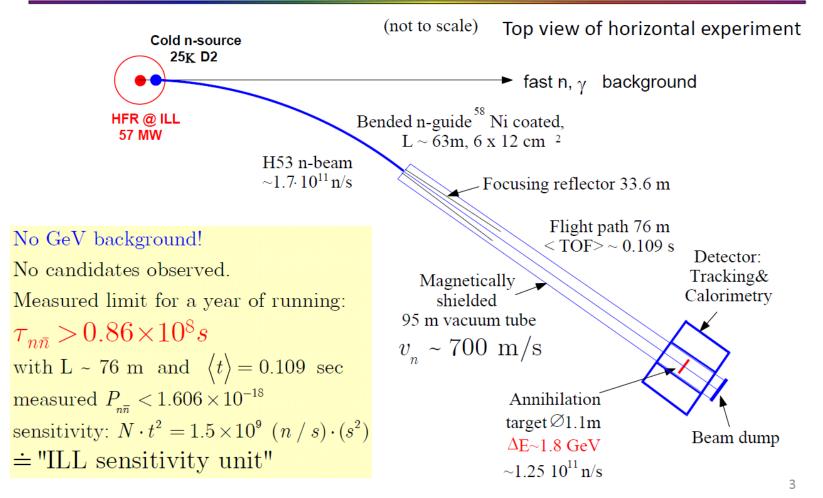


E. Kearns, Intensity Frontier All Hands Meeting, ANL

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#### Previous n-nbar search experiment with free neutrons

#### At ILL/Grenoble reactor in 89-91 by Heidelberg-ILL-Padova-Pavia Collaboration Z. Phys., C63 (1994) 409



Dubbers, Kamyshkov Project X Physics Study

## Free neutron antineutron oscillation

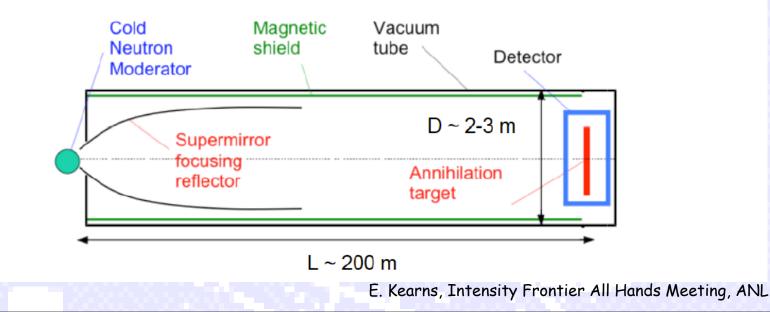
Expression of Interest

### Search for Neutron-Antineutron Transformation at Fermilab

#### The NNbarX Collaboration

need slow neutrons from high flux source, access of neutron focusing reflector to cold source, free flight path of ~200m

Improvement on ILL experiment by factor of ~1000 in transition probability is possible with horizontal experiment at Project X with existing n optics technology, sources, and moderators. Vertical experiment also possible

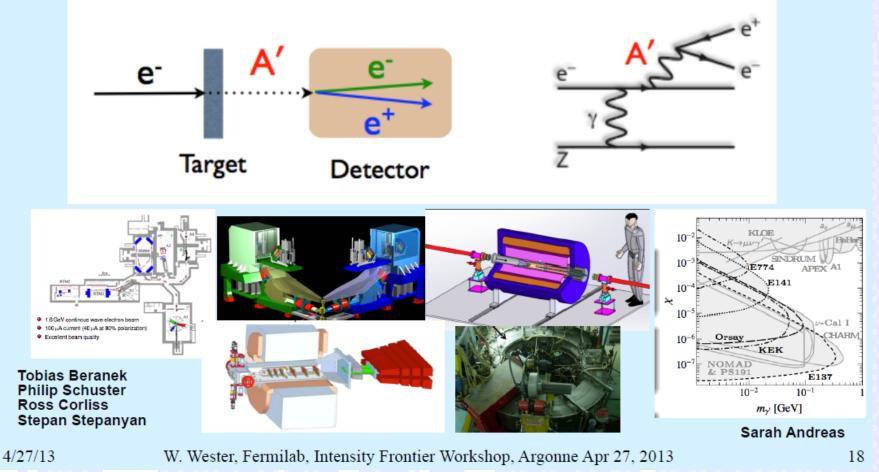


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# New Light, Weakly Coupled Particles

# e- fixed targets

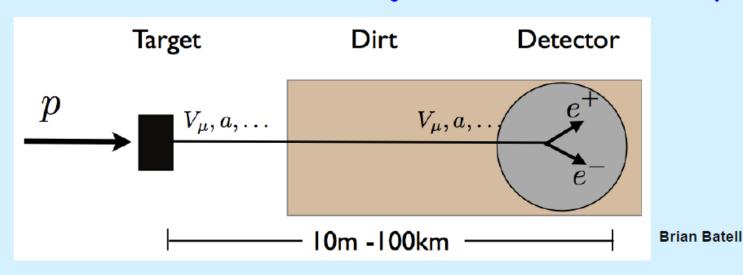
• Jefferson Lab, Mainz, VEPP-3, e- beam dump APEX, HPS, DarkLight



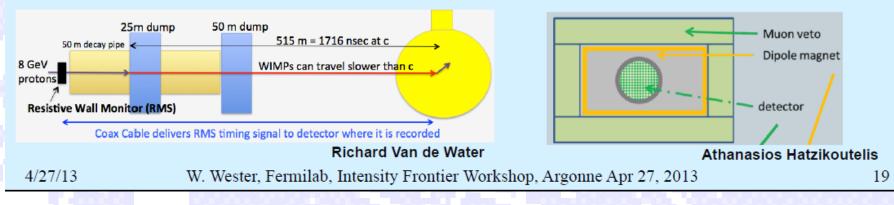
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# p fixed targets

Neutrino beamlines, Project X, beam dump

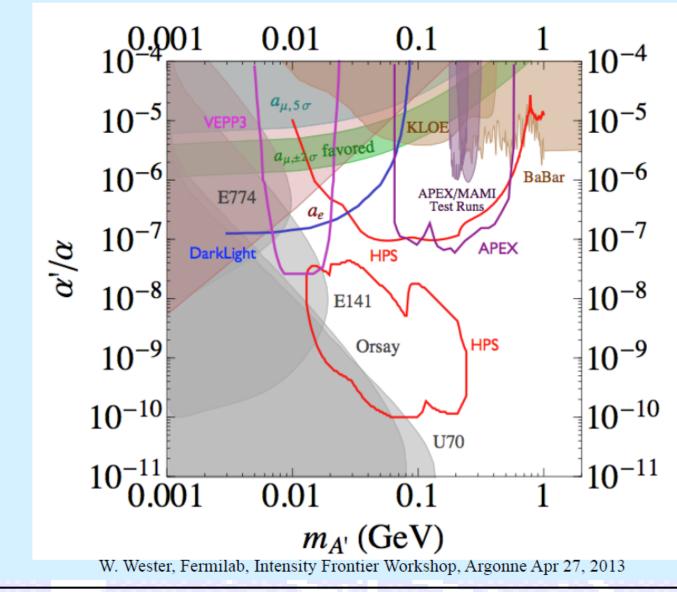


#### **MiniBooNE**



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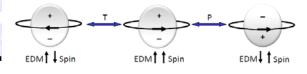
# Current and future status

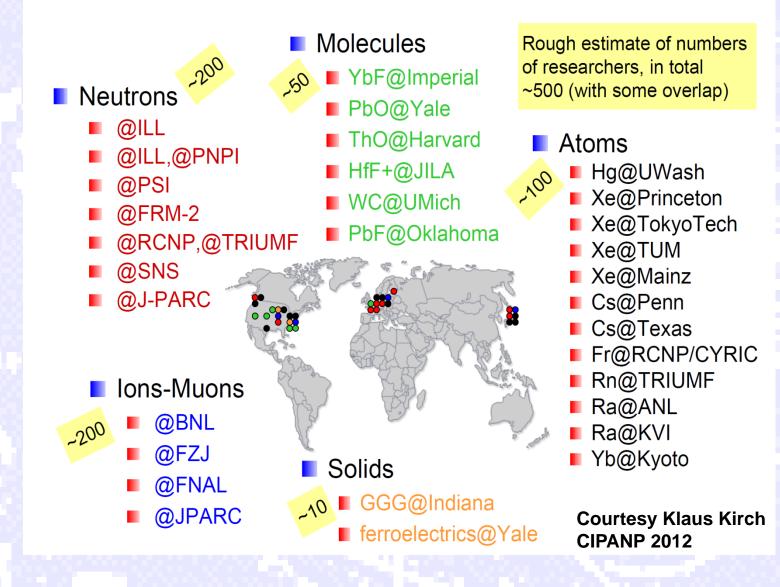


4/27/13

# Nucleons, Nuclei, Atoms

## **EDM Research Worldwide...**

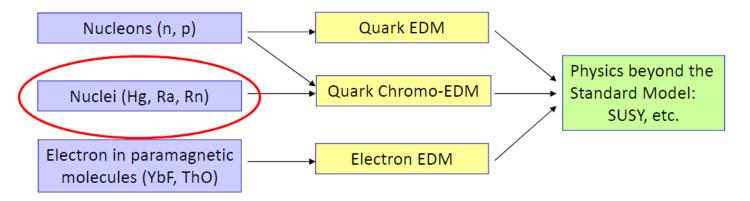




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#### **EDM Searches in Three Sectors**

Review article: *EDM of Nucleons, Nuclei, and Atoms* Engel, Ramsey-Musolf, van Kolck, arXiv:1303.2371 (2013)

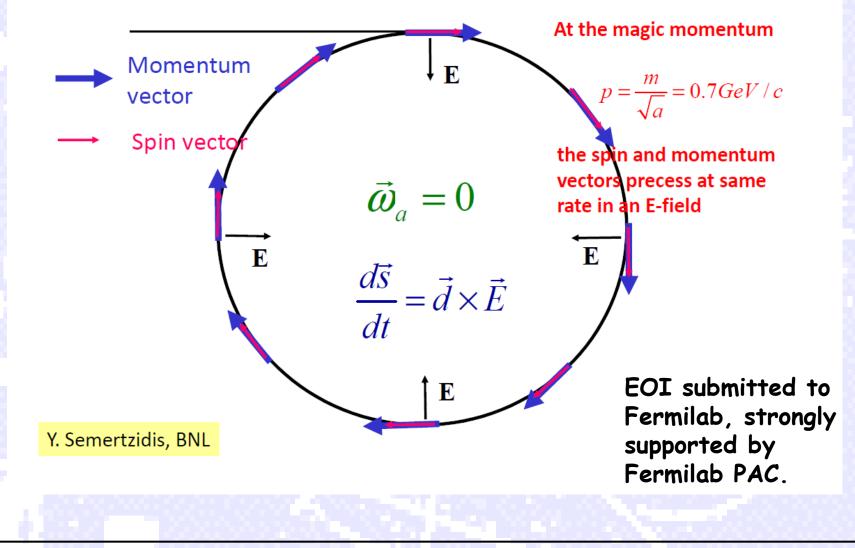


Sector	Exp Limit (e-cm)	Method	Standard Model
Electron	1 x 10 <sup>-27</sup>	YbF in a beam	10 <sup>-38</sup>
Neutron	3 x 10 <sup>-26</sup>	UCN in a bottle	10 <sup>-31</sup>
<sup>199</sup> Hg	3 x 10 <sup>-29</sup>	Hg atoms in a cell	10 <sup>-33</sup>
		M Bar	msev-Musolf (2009)

IVI. Ramsey-IVIUsolf (2009)

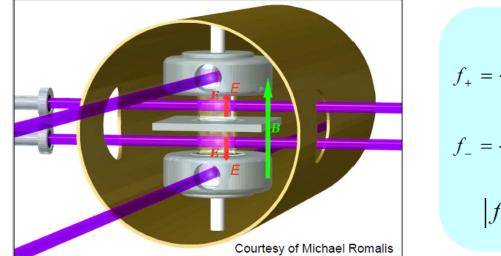
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## The proton EDM uses an ALL-ELECTRIC ring: spin is aligned with the momentum vector



#### The Seattle EDM Measurement (1980's - present)

<sup>199</sup>Hg) stable, high Z, groundstate  ${}^{1}S_{0}$ , I =  $\frac{1}{2}$ , high vapor pressure



$$f_{+} = \frac{2\mu B + 2dE}{h} \approx 15 \text{ Hz}$$
$$f_{-} = \frac{2\mu B - 2dE}{h} \approx 15 \text{ Hz}$$
$$\left| f_{+} - f_{-} \right| < 0.1 \text{ nHz}$$

#### **Limits and Sensitivities**

- Current: < 0.3 x 10<sup>-28</sup> e-cm Griffith *et al.*, Phys. Rev. Lett. (2009)
- Next 5 years: 0.03 x 10<sup>-28</sup> e-cm
- 2020 and beyond: 0.006 x 10<sup>-28</sup> e-cm





**Intensity Frontier** 

**Radioactive Isotope Facilities** 

- Sources of Enhancer Isotopes

R. Tschirhart

Radon (Rn)

• No stable isotopes

**Region of Enhancers** 

Francium (Fr)

• Favorable nuclear and atomic properties

Radium (Ra)

#### EDM of <sup>225</sup>Ra enhanced

Closely spaced parity doublet – Haxton & Henley (1983)

Large intrinsic Schiff moment due to octupole deformation

- Auerbach, Flambaum & Spevak (1996)

Relativistic atomic structure (<sup>225</sup>Ra / <sup>199</sup>Hg ~ 3)

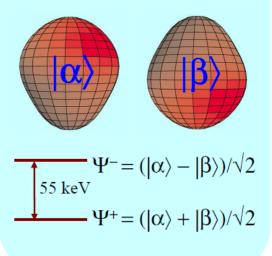
– Dzuba, Flambaum, Ginges, Kozlov (2002)

#### Parity doublet

225Ra:

 $| = \frac{1}{2}$ 

t<sub>1/2</sub> = 15 d



$$S \equiv \langle \psi_0 | \hat{S}_z | \psi_0 \rangle = \sum_{i \neq 0} \frac{\langle \psi_0 | \hat{S}_z | \psi_i \rangle \langle \psi_i | \hat{H}_{PT} | \psi_0 \rangle}{E_0 - E_i} + c.c.$$

#### Enhancement Factor: EDM (<sup>225</sup>Ra) / EDM (<sup>199</sup>Hg)

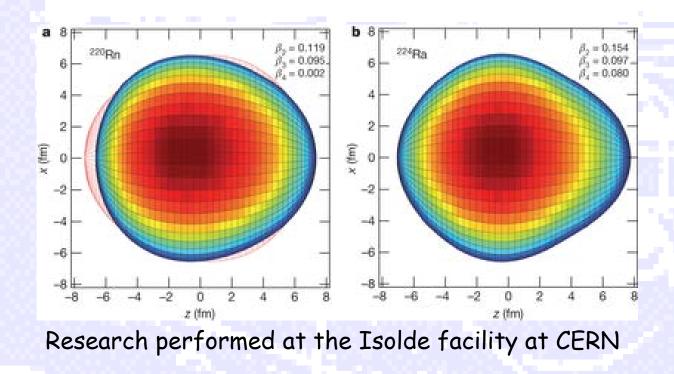
Skyrme Model	Isoscalar	Isovector	lsotensor
SIII	300	4000	700
SkM*	300	2000	500
SLy4	700	8000	1000

Schiff moment of <sup>225</sup>Ra, Dobaczewski, Engel (2005) Schiff moment of <sup>199</sup>Hg, Ban, Dobaczewski, Engel, Shukla (2010)

# Studies of pear-shaped nuclei using accelerated radioactive beams

L. P. Gaffney, P. A. Butler, M. Scheck, A. B. Hayes, F. Wenander, M. Albers, B. Bastin,
 C. Bauer, A. Blazhev, S. Bönig, N. Bree, J. Cederkäll, T. Chupp, D. Cline, T. E. Cocolios,
 T. Davinson, H. De Witte, J. Diriken, T. Grahn, A. Herzan, M. Huyse, D. G. Jenkins,
 D.T. Joss, N. Kesteloot, J. Konk

#### Nature 497,199-204(09 May 2013) doi:10.1038/nature12073



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# **Project-X:**

• Evolution of the existing Fermilab accelerator complex with the revolution in Super-Conducting RF Technology.



# **Project-X:**

• Evolution of the existing Fermilab accelerator complex with the revolution in Super-Conducting RF Technology.



# The Project-X Research Program

#### Neutrino experiments

A high-power proton source with proton energies between 1 and 120 GeV would produce intense neutrino sources and beams illuminating near detectors on the Fermilab site and massive detectors at distant underground laboratories.

#### Kaon, muon, nuclei & nucleon precision experiments

These could include world leading experiments searching for lepton flavor violation in muons, atomic, muon, nuclear and nucleon electron dipole moments (edms), precision measurement of neutron properties (e.g. n,nbar oscillations) and world-leading precision measurements of ultra-rare kaon decays.

### Platform for evolution to a Neutrino Factory and Muon Collider

Neutrino Factory and Muon-Collider concepts depend critically on developing high intensity proton source technologies.

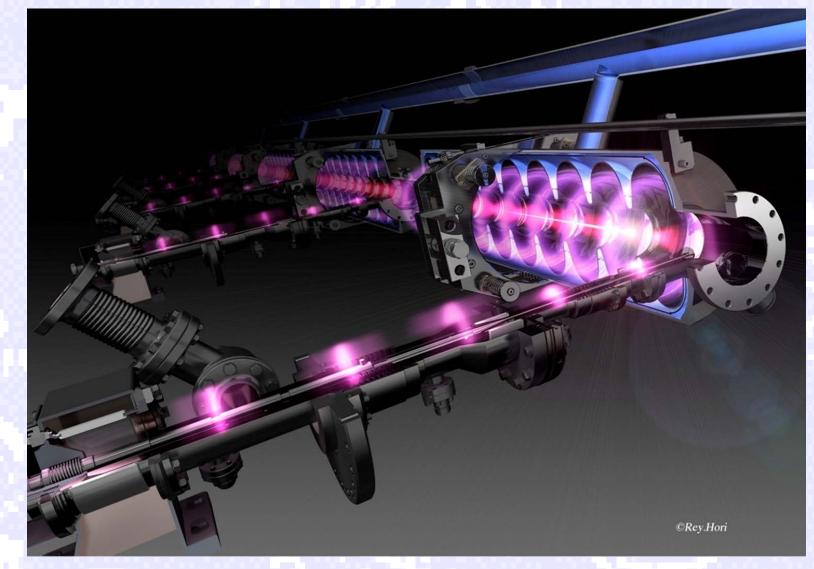
### Material Science and Nuclear Energy Applications

Accelerator, spallation, target and transmutation technology demonstrations which could investigate and develop accelerator technologies important to the design of future nuclear waste transmutation systems and future thorium fuel-cycle power systems. Possible applications of muon Spin Resonance techniques (muSR). as a sensitive probes of the magnetic structure of materials.

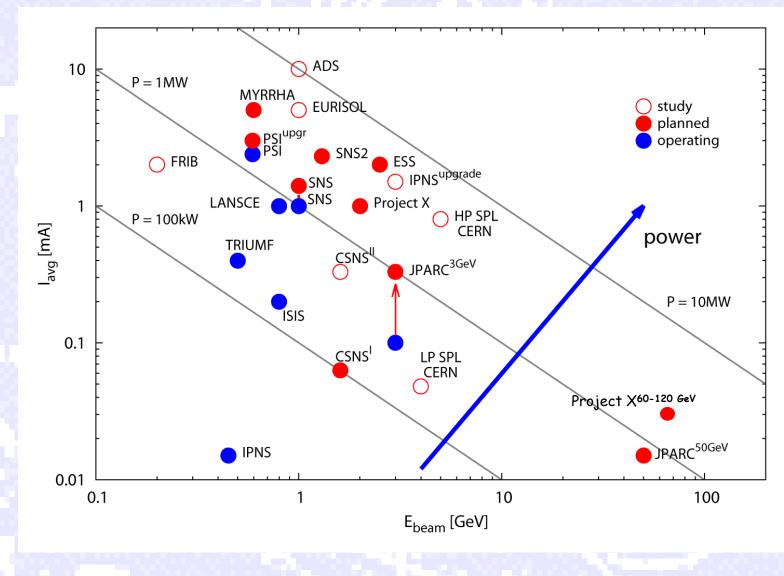
### Detailed discussion on Project X website

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## Beam Power is the Gateway to the Intensity Frontier...



## Beam Power is the Gateway to the Intensity Frontier...



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## Reference Design Linac Technology Map



LEBT RFQ MEBT	β <b>=0.11</b>	β=0.22 β	3=0.51	β <b>=0.61</b>	β <b>=0.9</b>	β <b>=1.0</b>
$\leftarrow$ RT $\rightarrow$	<		cw —			← Pulsed →
162.5 0.03-11		325 MH 10-177 N			MHz 3 GeV	1.3 GHz 3-8 GeV
Section	Freq	Energy (MeV)	Cav/ma	ig/CM	Тур	e
RFQ	162.5	0.03-2.1				
HWR (β <sub>G</sub> =0.1)	162.5	2.1-11	8/8/	/1	HWR, so	olenoid
SSR1 (β <sub>G</sub> =0.22)	325	11-38	16/8	/ 2	SSR, so	lenoid
SSR2 (β <sub>G</sub> =0.51)	325	38-177	35/2	1/7	SSR, so	lenoid
LB 650 (β <sub>G</sub> =0.61)	) 650	177-467	30/20	0/5	5-cell elliptic	al, doublet
HB 650 (β <sub>G</sub> =0.9)	650	467-1000	42/10	6/7	5-cell elliptic	al, doublet
HB 650 (β <sub>G</sub> =0.9)	650	1000-3000	120/30	0/15	5-cell elliptic	al, doublet
ILC 1.3 (β <sub>G</sub> =1.0)	1300	3000-8000	224 /28	8 /28	9-cell ellipti	cal, quad <sup>8</sup>

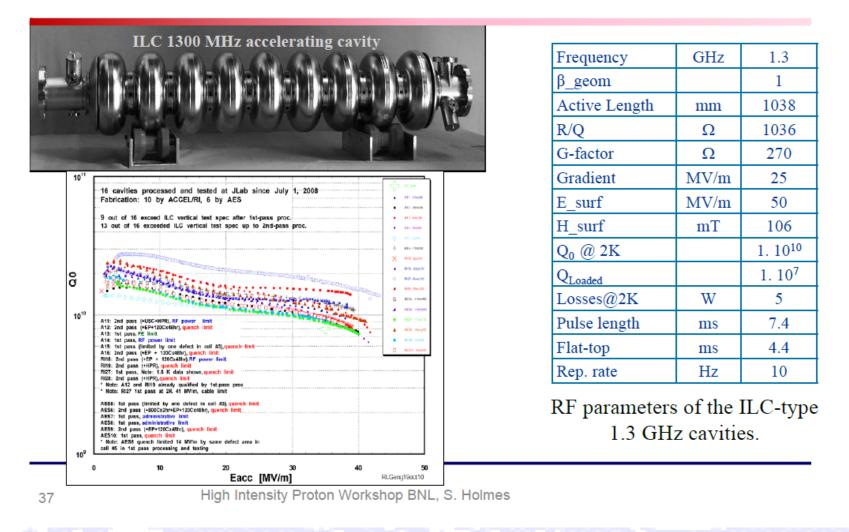
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Project X



## 1.3 GHz pulse linac





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# Evolution from the Energy Frontier to the Intensity Frontier at Fermilab...

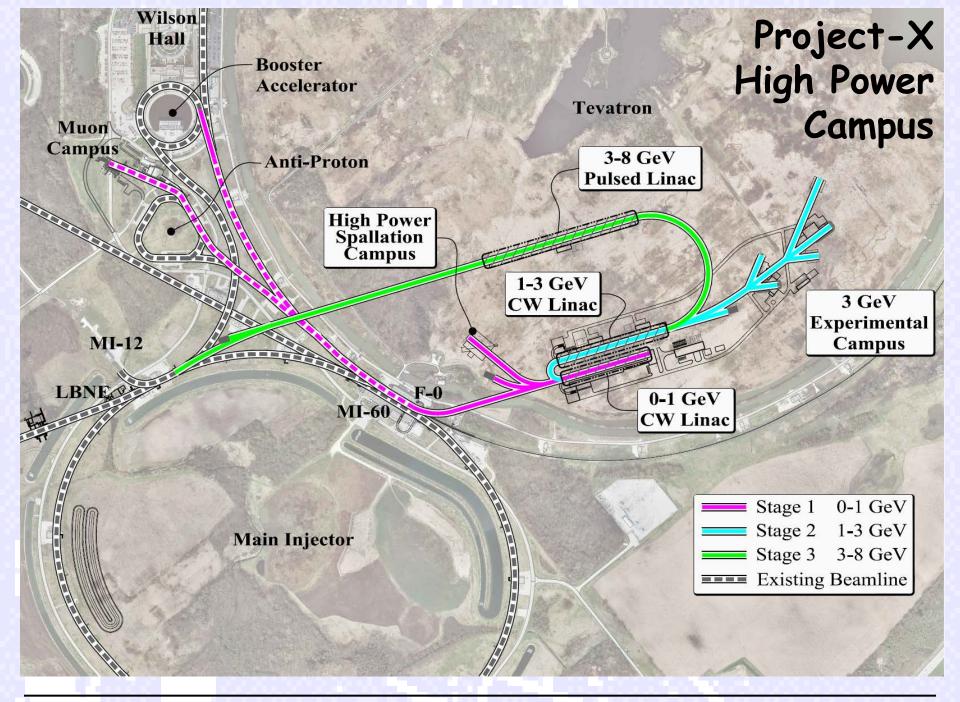
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# Evolution from the Energy Frontier to the Intensity Frontier at Fermilab...



# Evolution from the Energy Frontier to the Intensity Frontier at Fermilab...





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## Project X: Evolution of the Fermilab Accelerator Complex

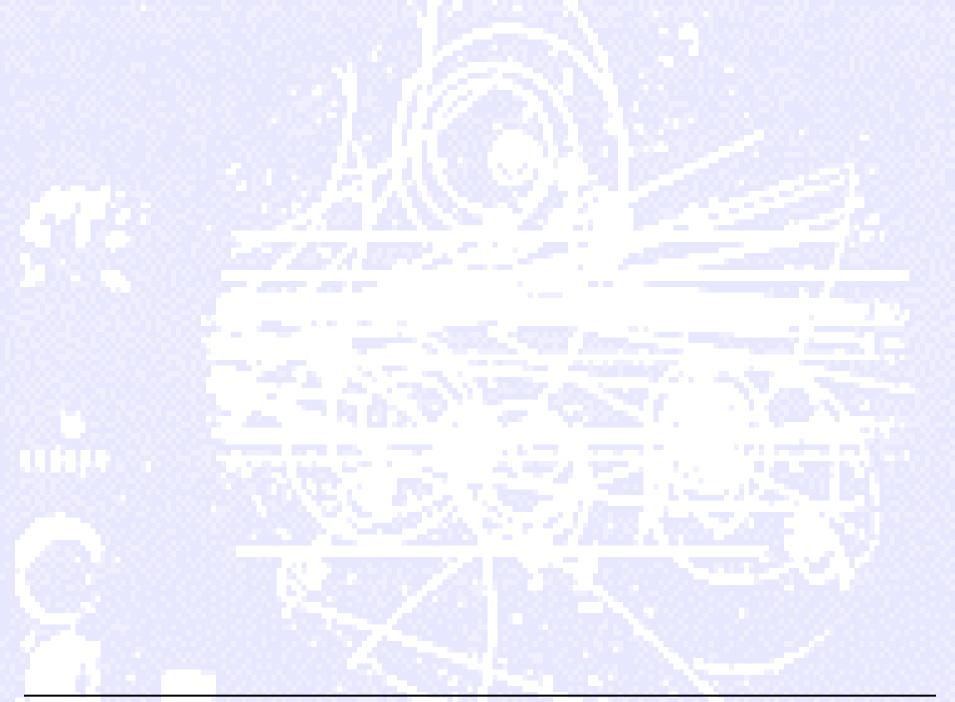
Program:	Onset of NOvA operations in 2013	Stage-1: 1 GeV CW Linac driving Booster & Muon, n/edm programs	Stage-2: Upgrade to 3 GeV CW Linac	Stage-3: Project X RDR	Stage-4: Beyond RDR: 8 GeV power upgrade to 4MW
MI neutrinos	470-700 kW**	515-1200 kW**	1200 kW	2450 kW	2450-4000 kW
8 GeV Neutrinos	15 kW +0-50kW**	0-42 kW* + 0-90 kW**	0-84 kW*	0-172 kW*	3000 kW
8 GeV Muon program e.g, (g-2), Mu2e-1	20 kW	0-20 kW*	0-20 kW*	0-172 kW*	1000 kW
1-3 GeV Muon program, e.g. Mu2e-2		80 kW	1000 kW	1000 kW	1000 kW
Kaon Program	0-30 kW** (<30% df from MI)	0-75 kW** (<45% df from MI)	1100 kW	1870 kW	1870 kW
Nuclear edm ISOL program	none	0-900 kW	0-900 kW	0-1000 kW	0-1000 kW
Ultra-cold neutron program	none	0-900 kW	0-900 kW	0-1000 kW	0-1000 kW
Nuclear technology applications	none	0-900 kW	0-900 kW	0-1000 kW	0-1000 kW
# Programs:	4	8	8	8	8
Total max power:	735 kW	2222 kW	4284 kW	6492 kW	11870kW
Total max power: * Operating point in range d ** Operating point in range d	epends on MI energy fo	or neutrinos.		6492 kW	11870kW

\*\* Operating point in range depends on MI injector slow-spill duty factor (df) for kaon program.

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# Summary

- Framework & Texture: The U.S. Intensity Frontier has a proposed framework and texture to deliver the science: LBNE & Project-X and a texture of experiments. Given the projected funding environment the framework must be strong enough to stand the test of time.
- Time: Large projects in Particle Physics will develop and evolve over decades. We know how to do/survive this, and the Tevatron/LHC is the most recent example of a robust framework in our field.
- Federation: Improving communication and ties among the texture of Intensity Frontier experiments and the theory community will strengthen the research program. Again, our field has demonstrated this in the evolution of the Energy Frontier program.
- **Resources:** Intensity Frontier researchers must reach out broadly to the funding agencies to communicate how particle physics spans agencies, and where synergies and leverage can be found. DOE/HEP, NP, NSF/NP, NIST, BES, etc.



### What do we not know about threeflavor oscillations?

			+ RSBL	Free Fluxes -	
			$3\sigma$ range	bfp $\pm 1\sigma$	
			$0.267 \rightarrow 0.344$	$0.302\substack{+0.013\\-0.012}$	$\sin^2  heta_{12}$
ibly	ls θ <sub>23</sub> non-negligil		$31.09 \rightarrow 35.89$	$33.36\substack{+0.81 \\ -0.78}$	$\theta_{12}/^{\circ}$
	greater or smaller	$ \models $	$0.342 \rightarrow 0.667$	$0.413^{+0.037}_{-0.025} \oplus 0.594^{+0.021}_{-0.022}$	$\sin^2  heta_{23}$
	than 45 deg		$35.8 \rightarrow 54.8$	$40.0^{+2.1}_{-1.5} \oplus 50.4^{+1.3}_{-1.3}$	$\theta_{23}/^{\circ}$
			$0.0156 \rightarrow 0.0299$	$0.0227\substack{+0.0023\\-0.0024}$	$\sin^2 heta_{13}$
	hasiaally		$7.19 \rightarrow 9.96$	$8.66\substack{+0.44\\-0.46}$	$\theta_{13}/^{\circ}$
	unknown		$0 \rightarrow 360$	$300^{+66}_{-138}$	$\delta_{ m CP}/^{\circ}$
			7.00  ightarrow 8.09	$7.50\substack{+0.18 \\ -0.19}$	$\left  {\Delta m^2_{21} \over 10^{-5} \ { m eV}^2}  ight $
	sign of ∆m unknown		$+2.276 \rightarrow +2.695$	$+2.473^{+0.070}_{-0.067}$	$\frac{\Delta m_{31}^2}{10^{-3}~{\rm eV}^2}({\rm N})$
)	(ordering of masses)		-2.649  ightarrow -2.242	$-2.427\substack{+0.042\\-0.065}$	$\frac{\Delta m_{32}^2}{10^{-3} \ {\rm eV}^2}  ({\rm I})$
า ท	sign of ∆n unknown (ordering		$7.19 \rightarrow 9.96$ $0 \rightarrow 360$ $7.00 \rightarrow 8.09$ $+2.276 \rightarrow +2.695$	$8.66^{+0.44}_{-0.46}$ $300^{+66}_{-138}$ $7.50^{+0.18}_{-0.19}$ $+2.473^{+0.070}_{-0.067}$	$ \begin{array}{c} \theta_{13}/^{\circ} \\ \delta_{\rm CP}/^{\circ} \\ \hline \frac{\Delta m_{21}^2}{10^{-5} \ {\rm eV}^2} \\ \frac{\Delta m_{31}^2}{10^{-3} \ {\rm eV}^2} \left( {\rm N} \right) \\ \Delta m_{21}^2 \end{array} $

CSS-2013 IF All Hands summary April 25th-27th ANL

Snowmass on the Pacific @ KITP, May 29th 2013

## **Outstanding 'anomalies'**

LSND @ LANL (~30 MeV, 30 m) Excess of  $\overline{\nu}_{e}$  interpreted as  $\overline{\nu}_{\mu} \rightarrow \overline{\nu}_{e}$ 

## $\Rightarrow \Delta m^2 \sim 1 \text{ eV}^2$ : inconsistent with 3 v masses

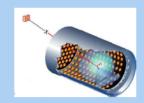
## MiniBooNE @ FNAL (v, v ~1 GeV, 0.5 km)

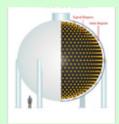
- unexplained >3 σ excess for E < 475 MeV in neutrinos (inconsistent w/ LSND oscillation)
- no excess for E > 475 MeV in neutrinos (inconsistent w/ LSND oscillation)
- small excess for E < 475 MeV in antineutrinos (~consistent with neutrinos)
- small excess for E > 475 MeV in antineutrinos (consistent w/ LSND)
- for E>200 MeV, both nu and nubar consistent with LSND

Also: possible deficits of reactor  $\overline{\nu}_e$  ('reactor anomaly') and source  $\nu_e$  ('gallium anomaly')

**Sterile neutrinos??** (i.e. no normal weak interactions) Some theoretical motivations for this, both from particle physics & astrophysics. Or some other new physics??

CSS-2013 IF All Hands meeting April 25th-27th ANL







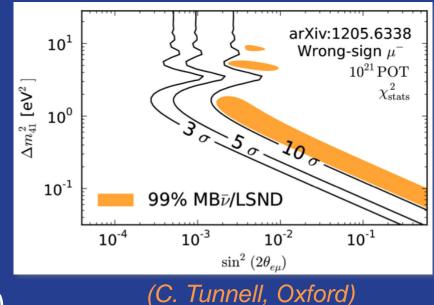
## vSTORM



•  $v_e \rightarrow v_\mu$  appearance (CPT conjugate to MiniBooNE)

• testbed for future  $\mu$  storage rings

 v's from a few-GeV muon storage ring aimed at near & far magnetized iron detectors



(110 collaborators, 37 institutions) https://indico.fnal.gov/conferenceDisplay.py?confld=6794



## Stage-1 Accelerator Resources:

- Promotes the Main Injector (MI) to a Mega-Watt class machine for neutrinos, and increases the potential beam power for other medium power MI experiments (e.g. ORKA, nu-STORM).
- Unshackles the µ→e (Mu2e) experiment from the Booster complex: Potentially increases sensitivity of Mu2e by x10 - x100 with 1-GeV CW drive beam.
- High power spallation target optimized for ultra-cold neutron and atomic-edm particle physics experiments and neutron⇔anti-neutron oscillation experiments.
- Capability to drive polarized protons to a proton-edm experiment.
- Increases the available integrated 8 GeV power for other experiments (e.g. short-baseline neutrinos) from the Booster complex by liberating Mu2e.

