

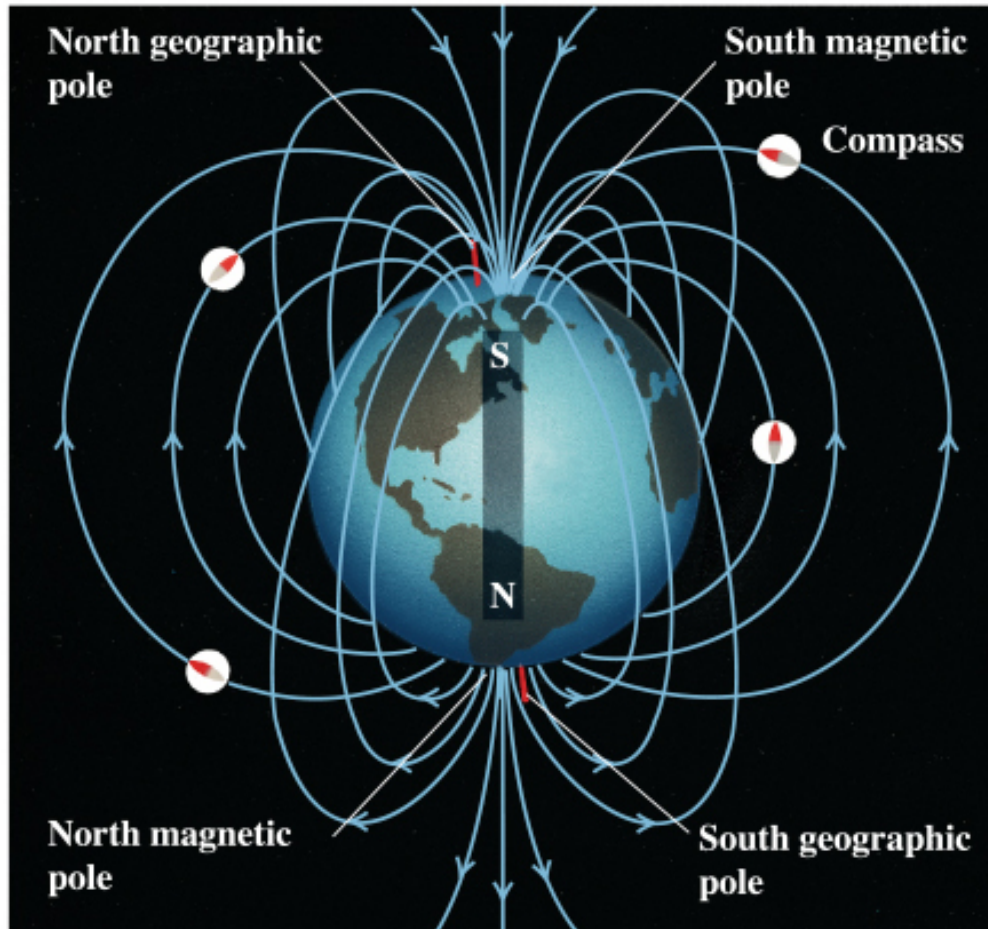
"the nature of turbulence"
KITP, may 2011

turbulence and the dynamo

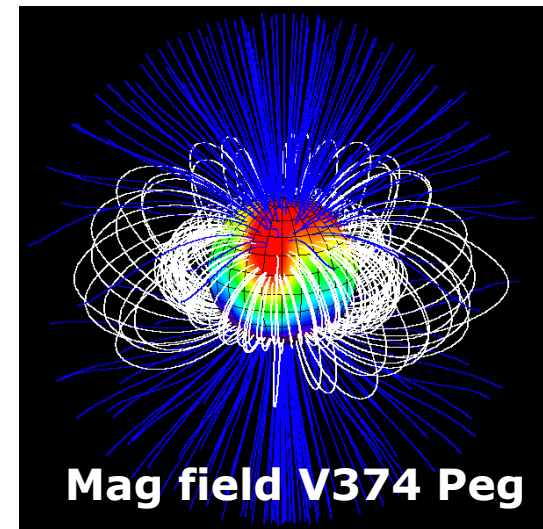
Jean-François Pinton
pinton@ens-lyon.fr



Focus on : Earth (liquid metals) & experiments



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OUTLINE

1. **Dynamos**

Scalar, scalar gradient and B-equations
Is turbulence helpful or detrimental?
Experiments

2. **Mean field MHD**

easy and hard
Examples
Measurements

3. **Turbulence in VKS**

the α - ω “Duddley & James” failed dynamo
the α - ω ferroDynamo.
mechanism and low dimensional behavior
turbulence and noise.

4. **Turbulent MHD**

Saturation: large B-field experiment.

5. **Next**

Advection / diffusion / stretching equations

- passive scalar

$$\partial_t C + \mathbf{u} \cdot \nabla C = \kappa \Delta C$$

- gradient of a passive scalar

$$\mathbf{G} = \nabla C$$

$$\partial_t \mathbf{G} = -\nabla \cdot (\mathbf{u} \cdot \mathbf{G}) + \kappa \Delta G$$

- vorticity

$$\partial_t \omega = \nabla \times (\mathbf{u} \times \omega) + \nu \Delta \omega$$

- magnetic field

$$\partial_t \mathbf{B} = \nabla \times (\mathbf{u} \times \mathbf{B}) + \eta \Delta \mathbf{B}$$

Question : $\frac{d}{dt} \int_V d^3r x^2$

anti-dynamo theorems

Antidynamo theorem : Cowling 1934

Axisymmetric magnetic field:

$$\mathbf{B} = \nabla A_\phi \times \mathbf{e}_\phi + B_\phi \mathbf{e}_\phi$$

Ohm's law:

$$\lambda \nabla \times \mathbf{B} = \mathbf{u} \times \mathbf{B} - \nabla U - \partial_t \mathbf{A}$$

For a steady process, the $C = r A_\phi$ component reads

$$u_r \partial_r C + u_z \partial_z C = \lambda \left(\partial_r^2 C + \partial_z^2 C - (1/r) \partial_r C \right)$$

However, dynamos with **small** deviations from axisymmetry are possible -- Bragiinsky 1976. Or small anisotropy in the diffusivity – Lortz 1989.

Dynamo : stretch > Joule

- Induction eqn. $\partial_t \mathbf{B} = \nabla \times (\mathbf{u} \times \mathbf{B}) + \eta \Delta \mathbf{B}$

- Magnetic Reynolds number: $Rm = \frac{UL}{\eta}$

- Turbulence $Re = \frac{UL}{\nu} = \frac{\eta}{\nu} \frac{UL}{\eta} = \frac{Rm}{Pm}$

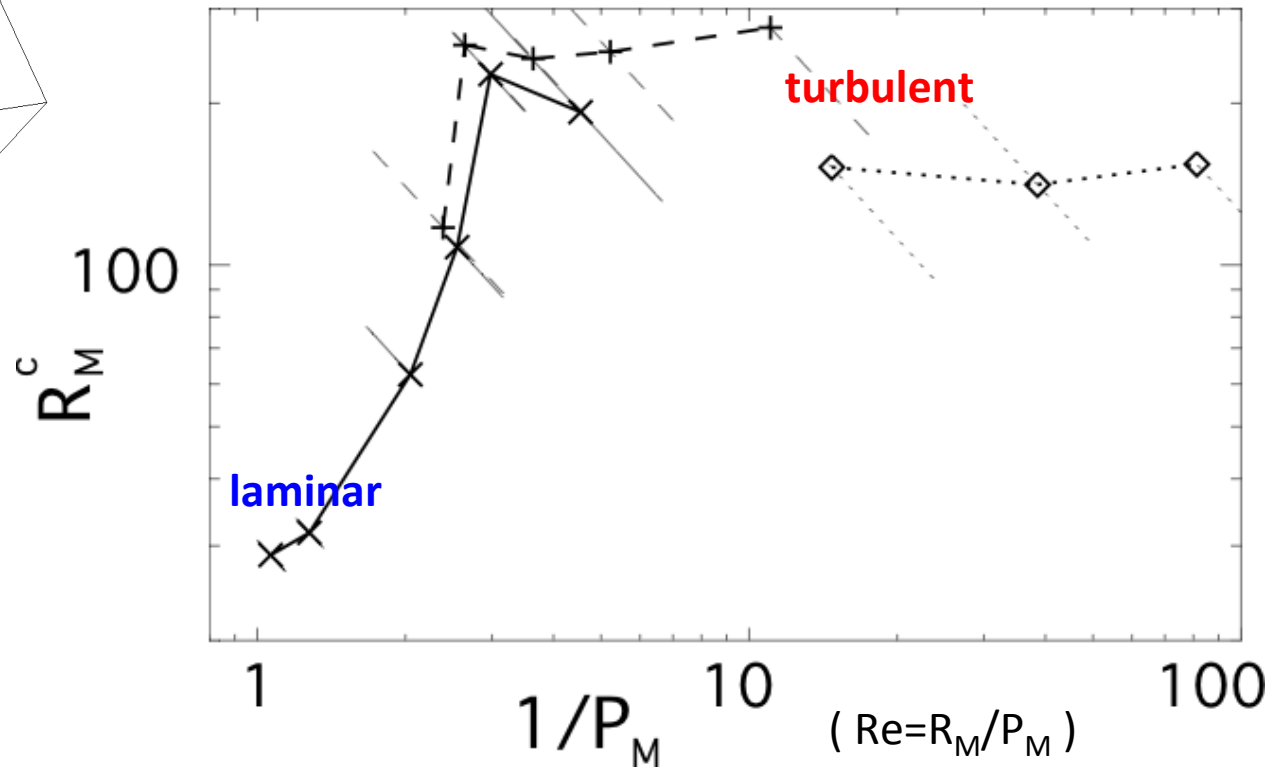
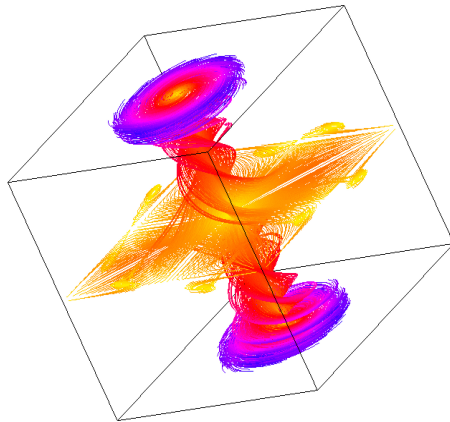
- Power $Rm = \mu\sigma \left(\frac{PL}{\rho} \right)^{1/3}$

$Rm^c < Rm^{MAX}$, not $Rm^{MAX} > Rm^c$!!!!

Numerical study of the Taylor Green flow

Y. Ponty, P. Mininni et al., *PRL* **94** (2005)

$$\mathbf{F}_{\text{TG}}(k_0) = 2F \begin{bmatrix} \sin(k_0 x) \cos(k_0 y) \cos(k_0 z) \\ -\cos(k_0 x) \sin(k_0 y) \cos(k_0 z) \\ 0 \end{bmatrix}$$



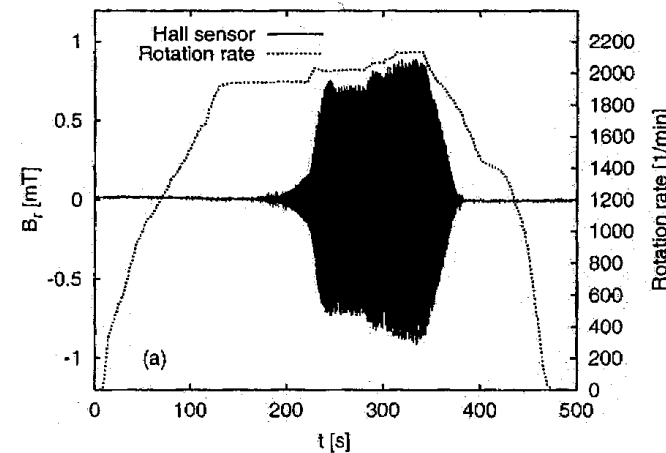
Exact solutions of the induction eqn

$$\partial_t \mathbf{B} = \nabla \times (\mathbf{u} \times \mathbf{B}) + \eta \Delta \mathbf{B}$$

Ponomarenko – Riga dynamo

A. Gailitis et al., PRL (2001)

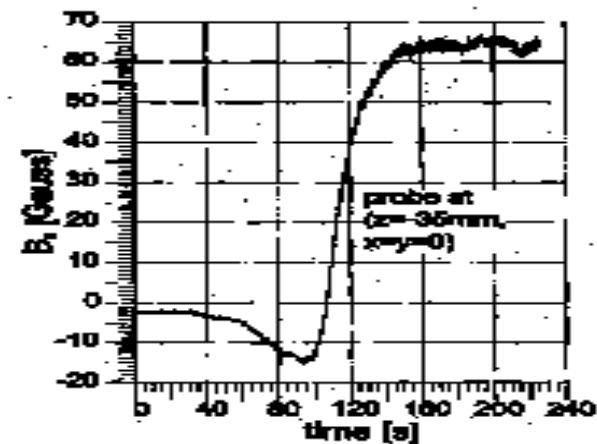
- Helical wave
- Threshold \approx laminar prediction
- α - ω dynamo
- no further dynamics.

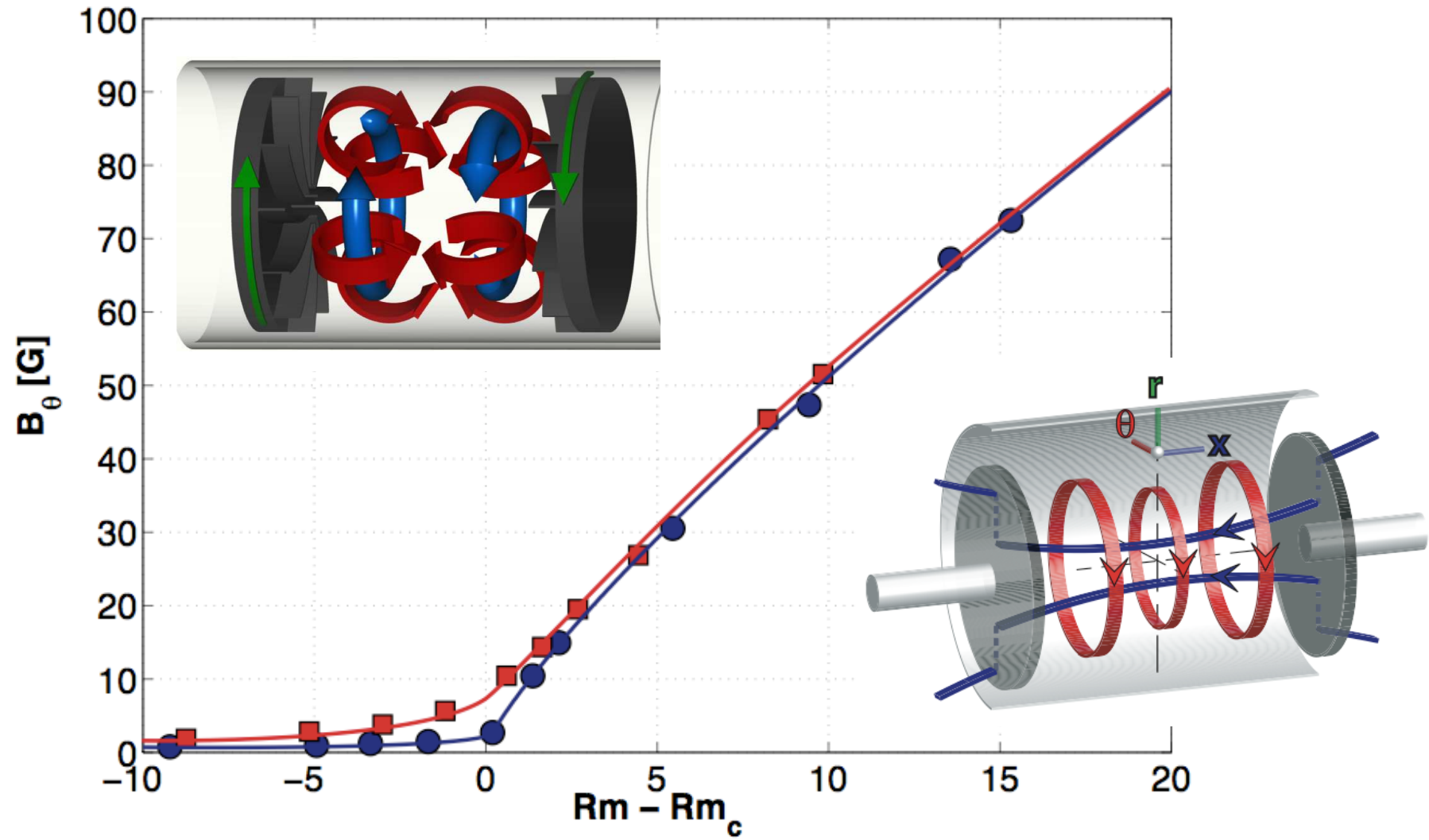


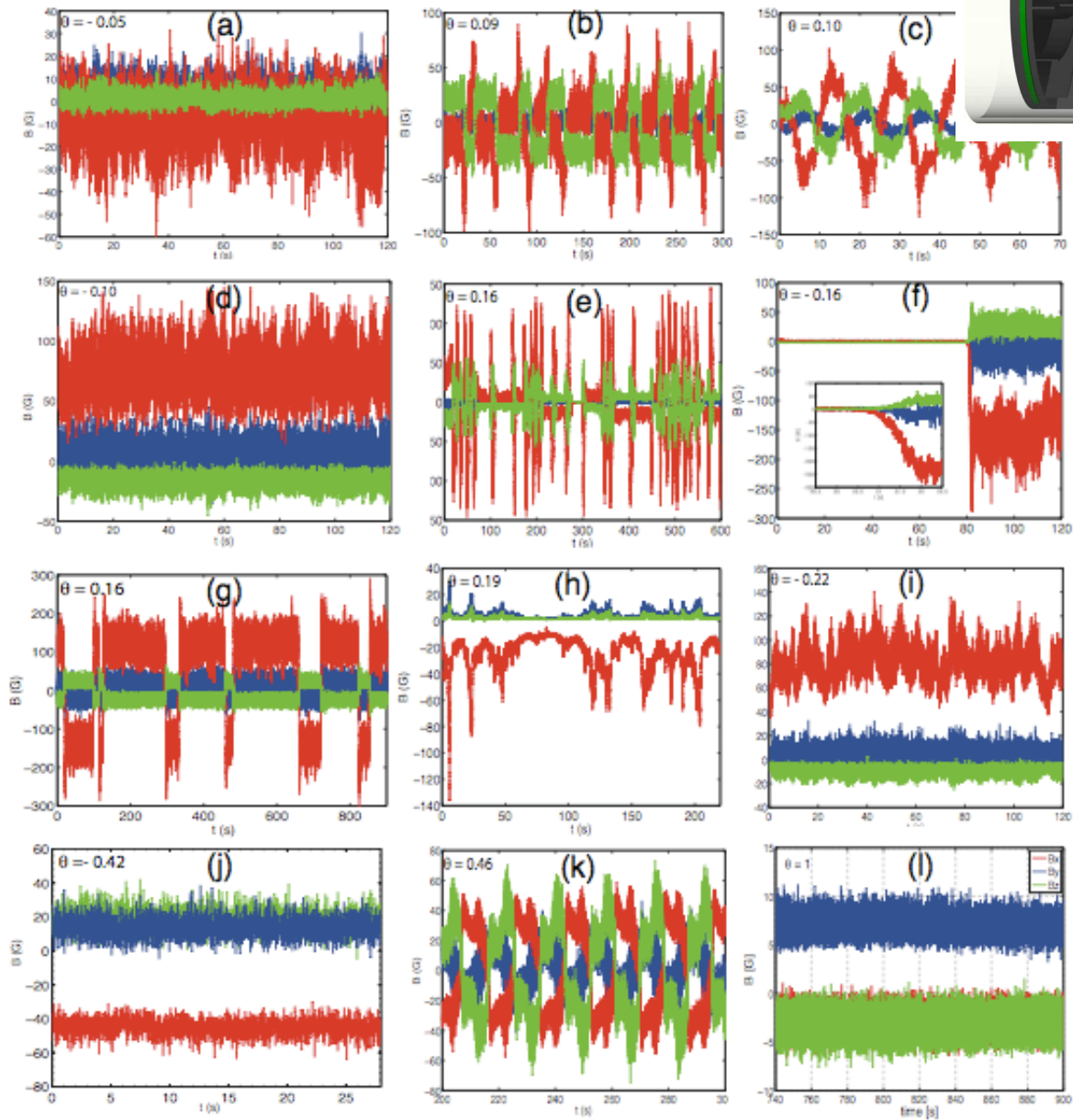
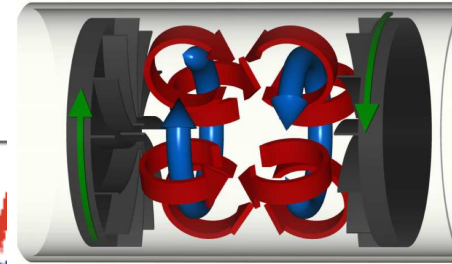
G.O. Roberts – Karlsruhe dynamo

Müller & Stieglitz, PoF (2001)

- Stationary dynamo
- Threshold \approx laminar prediction
- α^2 dynamo
- no further dynamics.







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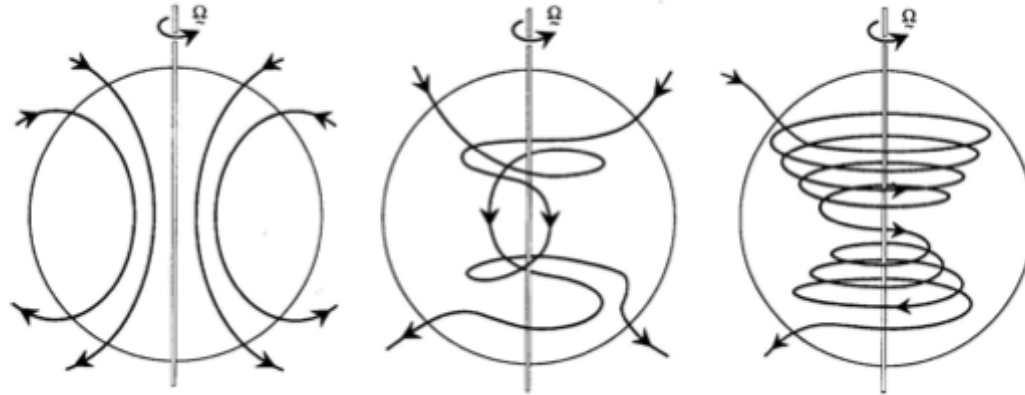
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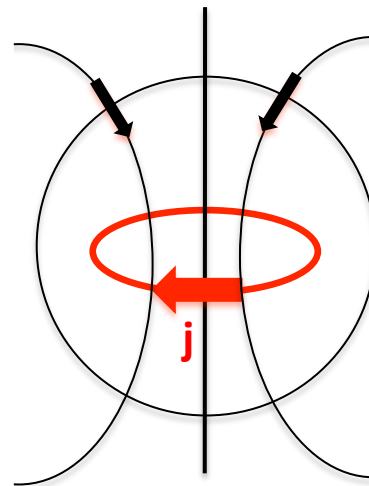
Saturation: large B-field experiment.

5. **Next**

POL to TOR : the ω effect



TOR to POL : the α effect



$$\langle \mathbf{u} \times \mathbf{b} \rangle = \alpha \mathbf{B}$$

$$\mathbf{j} // \mathbf{B}$$

Mean-field MHD 101

- Mean field MHD versus Prandtl mixing length

$$\tilde{B} = B + b$$

$$\tilde{u} = U + u = 0 + u$$

$$\langle \text{induction eqn} \rangle \Rightarrow \partial_t B = \nabla \times \langle u \times b \rangle + \lambda \Delta B$$

$$\partial_t b = \nabla \times (u \times B + u \times b - \langle u \times b \rangle) + \lambda \Delta b$$

$$\text{first order smoothing} \Rightarrow u \times b \approx \langle u \times b \rangle$$

$$\partial_t b - \lambda \Delta b = \nabla \times (u \times B) \Rightarrow b = F_u(B)$$

$$\langle u \times b \rangle = \alpha B + \beta \nabla \times B$$

compute α and β from $u(x)$

$$\text{enjoy } \partial_t B = \alpha \nabla \times B + (\lambda + \beta) \Delta B$$

Mean-field MHD 102

- calculation for homegenous isotropic (non mirror symmetric) turbulence

$$\bar{B} = B + b$$

$$\bar{u} = U + u = 0 + u$$

$$\langle \text{induction eqn} \rangle \Rightarrow \partial_t B = \nabla \times \langle u \times b \rangle + \lambda \Delta B$$

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compute α and β from $u(x)$

$$\langle u_i(\mathbf{x}, t) u_j(\mathbf{x}', t') \rangle = R_{ij}(\mathbf{r}, \tau)$$

$$R_{ij}(\mathbf{r}, \tau) = \int \int \Phi_{ij}(\mathbf{k}, \omega) e^{i(\mathbf{k} \cdot \mathbf{r} - \omega \tau)} d\mathbf{k} d\omega$$

$$E(k, \omega) = \frac{1}{2} \int_{S_k} \Phi_{ii}(\mathbf{k}, \omega) dS$$

$$F(k, \omega) = i \int_{S_k} \varepsilon_{ikl} k_k \Phi_{il}(\mathbf{k}, \omega) dS$$

$$\alpha \approx -\frac{1}{3\lambda} \int k^{-2} F(k) dk$$

$$\beta \approx \frac{2}{3} \lambda^{-1} \int_0^\infty k^{-2} E(k) dk$$

$$\alpha \sim u_0 R_M$$

$$\beta \sim u_0 \ell_0 R_M$$

The run-away problem

- Diffusivity and stretching increase at the same rate ?

$$\partial_t \mathbf{B} = \alpha \nabla \times \mathbf{B} + \beta \Delta \mathbf{B} \quad \alpha \text{ and } \beta \text{ from } \langle u_i u'_j \rangle$$

$$Rm_{\text{eff}} = \frac{\alpha l}{\beta} \stackrel{?}{=} \text{Const.} < Rm^c$$

- YES: Schekochihin et al., *APJ* **629** (2005)
- NO: Ponty et al., *PRL* **94** (2005)
- NO: Karlsruhe and Riga exp. : $Rm^c(\text{exp}) = Rm^c(\text{laminar})$
Denisov et al., *JETP Lett.*, **88** 2008
- MAYBE:?? Frick et al., *PRL* 2010

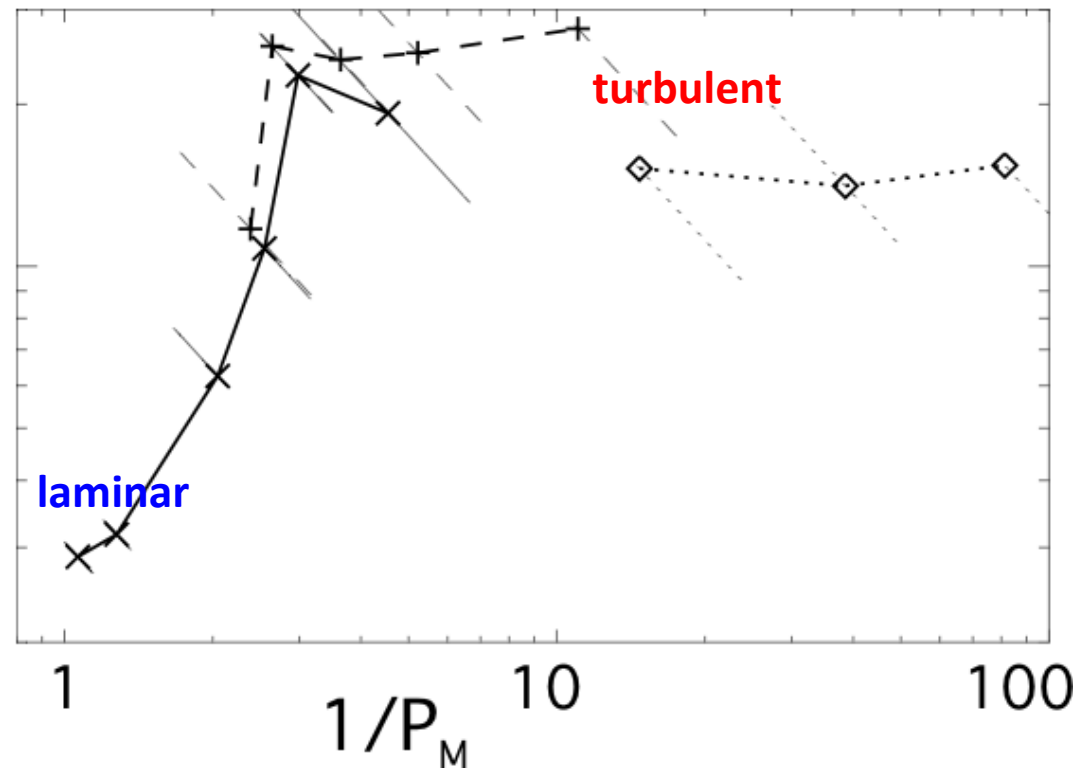
The run-away problem

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$$Rm_{\text{eff}} = \frac{\alpha l}{\beta} \stackrel{?}{=} \text{Const.} < Rm^c$$

- YES: Schekochihin
- NO: Ponty et al., $F \approx 100$
- NO: Karlsruhe and Denisov et al.
- MAYBE:?? Frick et al



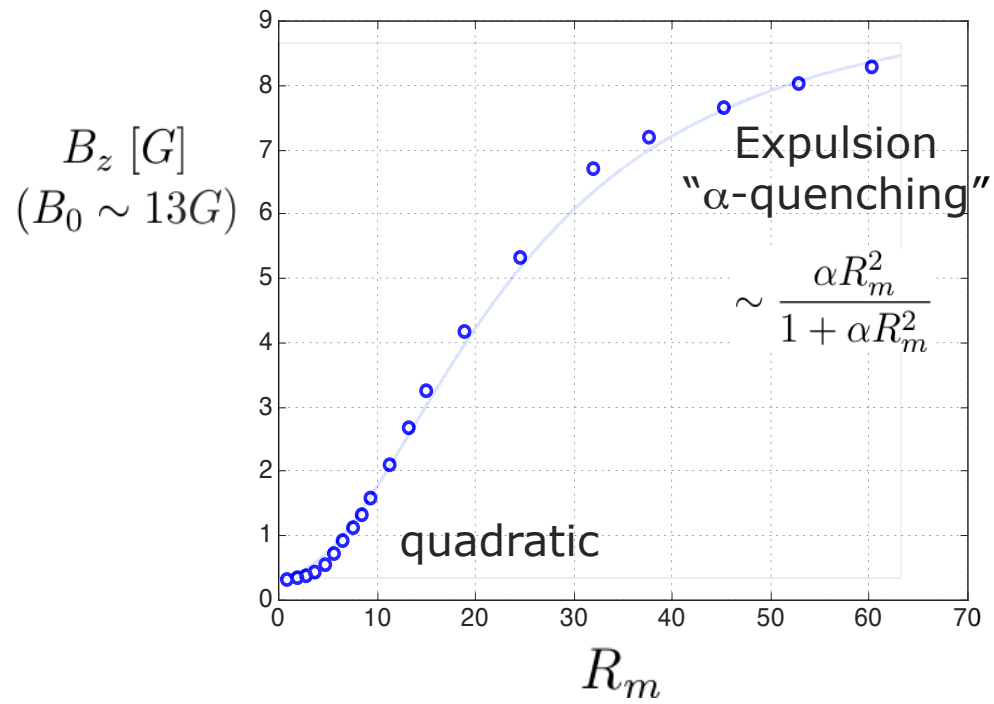
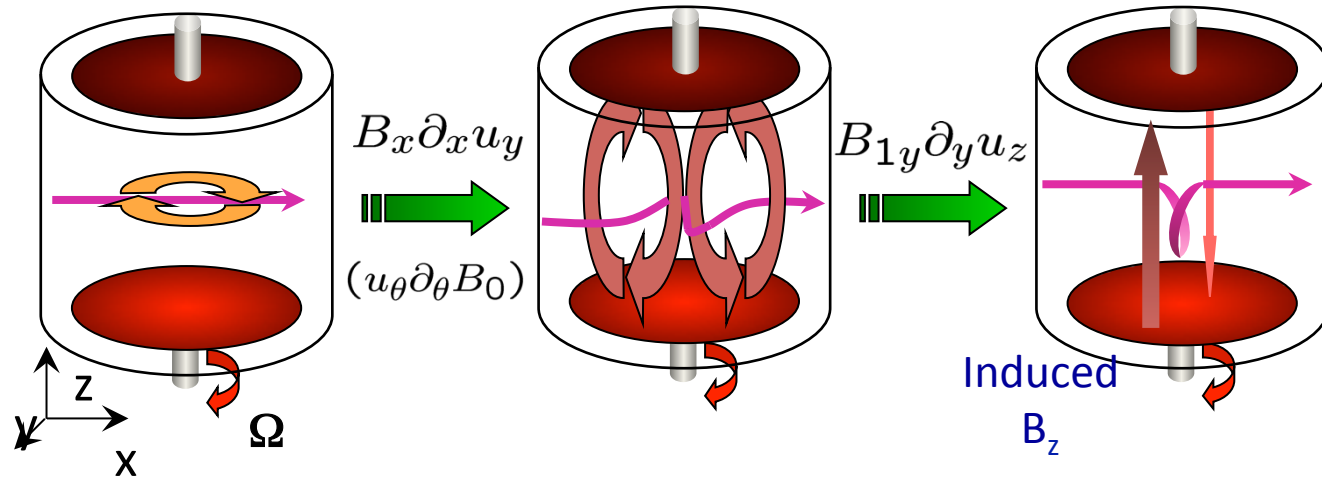
Mean-field MHD 715

- Now, do it right ... Rädler and Stepanov

$$\begin{aligned}
 \mathcal{E} = & -\alpha_1^{(\Omega)}(\mathbf{g} \cdot \boldsymbol{\Omega})\bar{\mathbf{B}} - \alpha_2^{(\Omega)}((\boldsymbol{\Omega} \cdot \bar{\mathbf{B}})\mathbf{g} + (\mathbf{g} \cdot \bar{\mathbf{B}})\boldsymbol{\Omega}) \\
 & -\alpha_1^{(W)}(\mathbf{g} \cdot \mathbf{W})\bar{\mathbf{B}} - \alpha_2^{(W)}((\mathbf{W} \cdot \bar{\mathbf{B}})\mathbf{g} + (\mathbf{g} \cdot \bar{\mathbf{B}})\mathbf{W}) \\
 & -\alpha^{(D)}\hat{\alpha}(\mathbf{g}, \mathbf{D}) \circ \bar{\mathbf{B}} \\
 & -(\gamma^{(0)}\mathbf{g} + \gamma^{(\Omega)}\mathbf{g} \times \boldsymbol{\Omega} + \gamma^{(W)}\mathbf{g} \times \mathbf{W} + \gamma^{(D)}\mathbf{g} \circ \mathbf{D}) \\
 & \times \bar{\mathbf{B}} - \beta^{(0)}\boldsymbol{\nabla} \times \bar{\mathbf{B}} - \beta^{(D)}\mathbf{D} \circ (\boldsymbol{\nabla} \times \bar{\mathbf{B}}) \\
 & -(\delta^{(\Omega)}\boldsymbol{\Omega} + \delta^{(W)}\mathbf{W}) \times (\boldsymbol{\nabla} \times \bar{\mathbf{B}}) \\
 & -(\kappa^{(\Omega)}\boldsymbol{\Omega} + \kappa^{(W)}\mathbf{W}) \circ (\boldsymbol{\nabla}\bar{\mathbf{B}})^{(s)} - \kappa^{(D)}\hat{\kappa}(\mathbf{D}) \circ (\boldsymbol{\nabla}\bar{\mathbf{B}})^{(s)}
 \end{aligned}$$

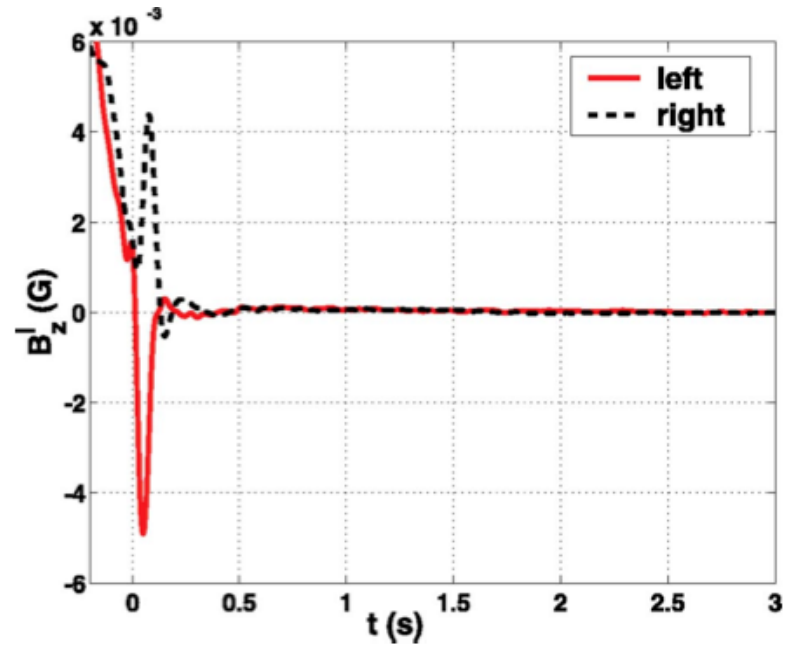
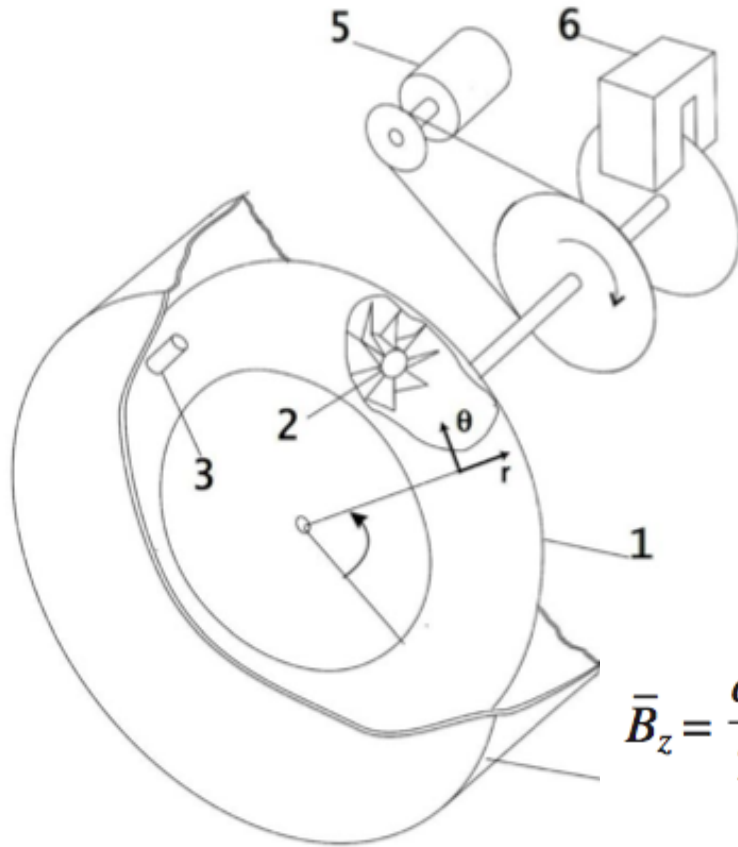
K.-H. Rädler and R. Stepanov, Phys. Rev. E 73, 056311 (2006)

Induction : helicity and alpha effect



Perm Torus experiment : alpha

Stepanov et al., *PRE* 73 (2006)



$$\bar{B}_z = \frac{\alpha \pi r_0^2}{2\lambda R} B_0 \Rightarrow [\alpha] \approx \frac{2\lambda R \bar{B}_z}{\pi r_0^2 B_0} \sim (3 \pm 0.3) \times 10^{-3} \text{ m/s.}$$

$\alpha_{\text{max}} \sim 10 \text{ cm/s}$

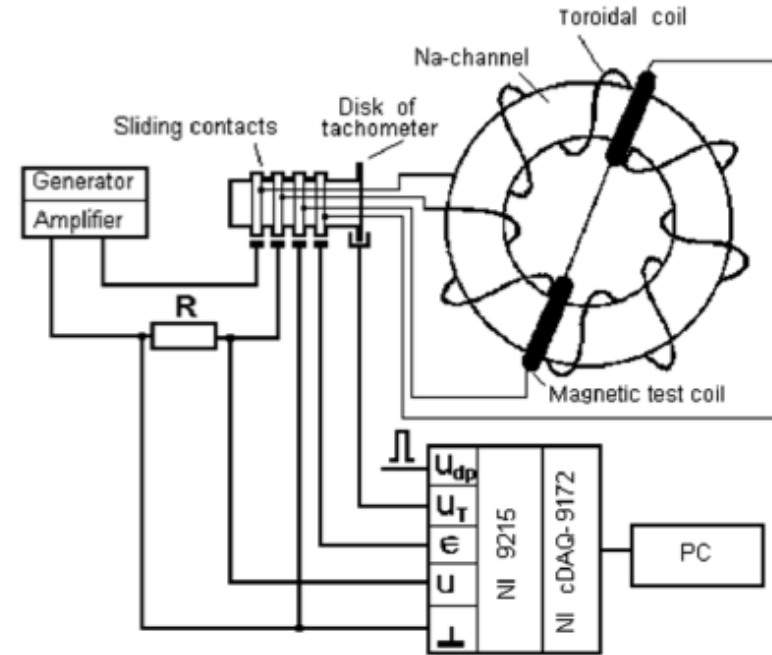
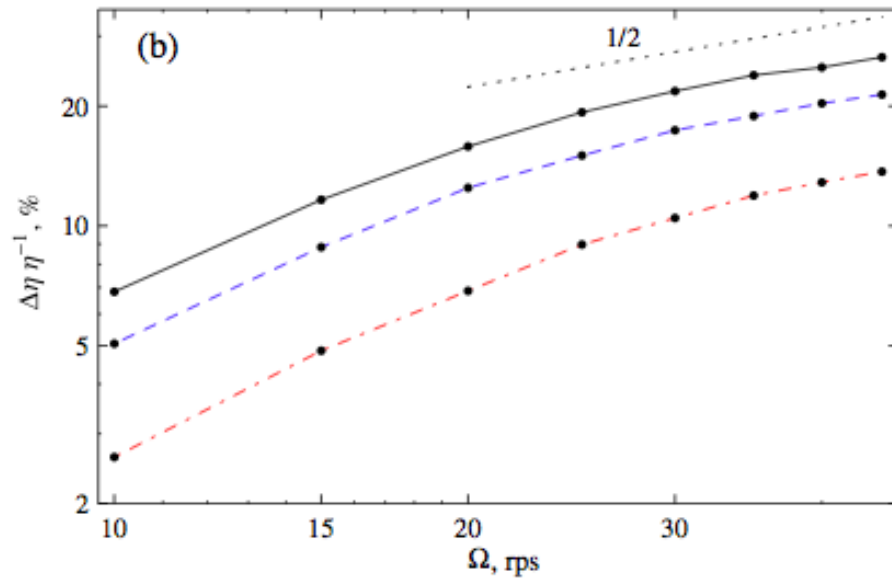
$$\alpha/u \sim 10^{-2} - 3$$



$$\frac{\alpha(\text{Na})}{\alpha(\text{Ga})} \sim \frac{L(\text{Na}) R_m(\text{Na})}{L(\text{Ga}) R_m(\text{Ga})}$$

Perm Torus experiment : beta

Frick et al., *PRL* 105 (2010)



$$\lambda \sim \lambda_{\text{mol}}(1 + 0.25) \text{ at } R_M \sim 50$$

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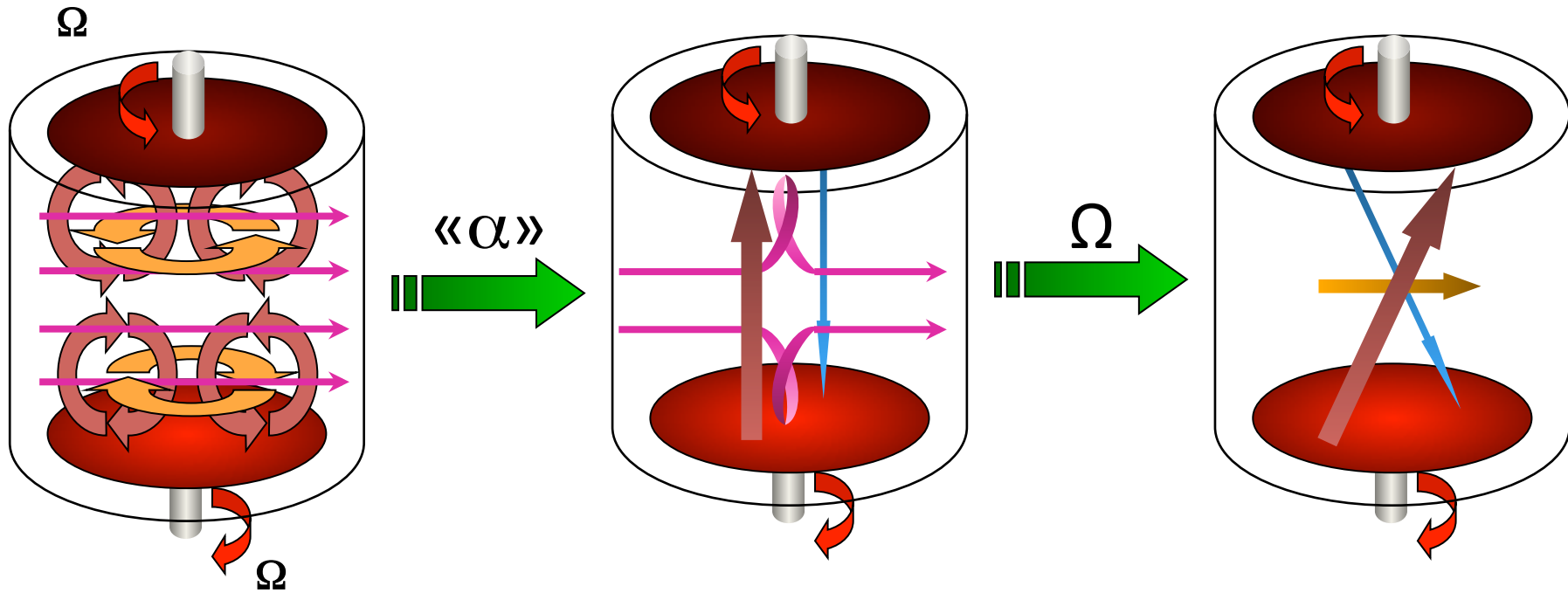
turbulence and noise.

4. **Turbulent MHD**

Saturation: large B-field experiment.

5. **Next**

VKS, as a Duddley and James s2t2 dynamo ?

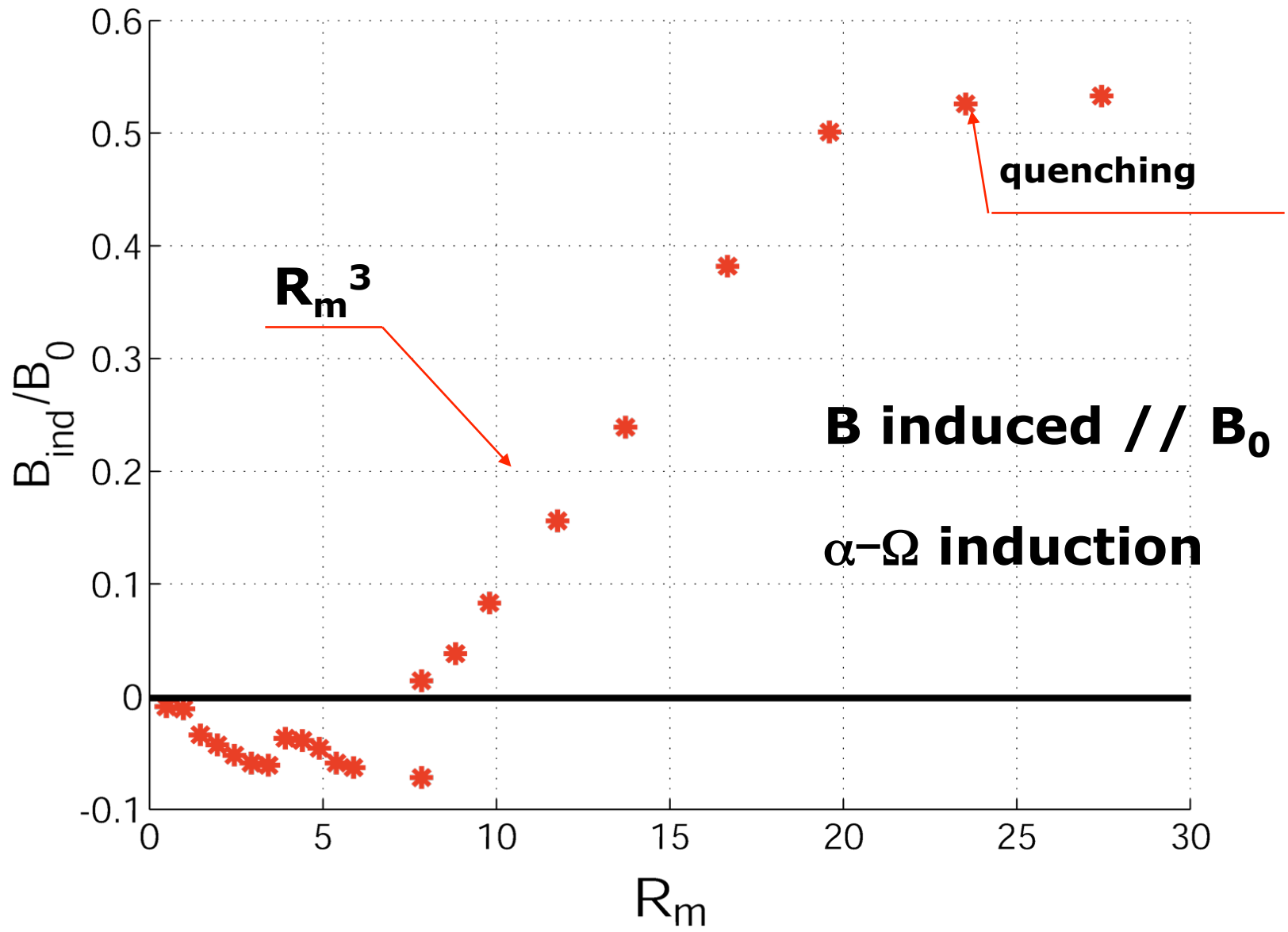


Induced
 $B // B_0$



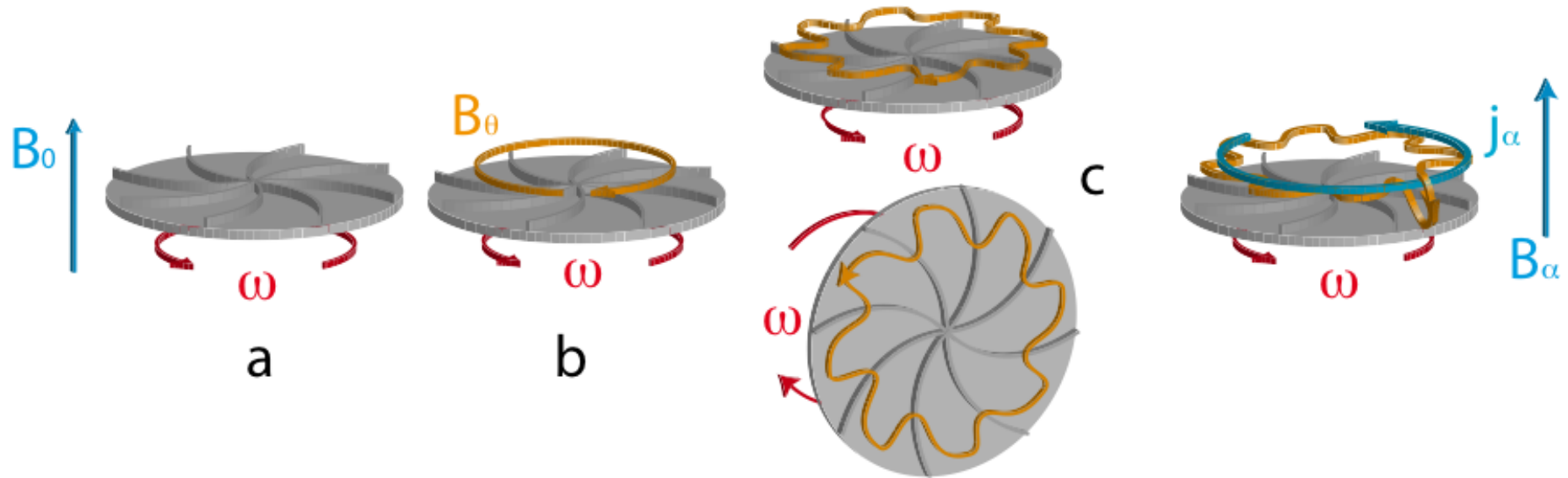
α - ω loop-back mechanism

VKS1 measurement
(M. Bourgoïn PhD thesis)



VKS dynamo generation mechanism

G. Verhille et al., *NJP* 12 (2010)



- **LARGE ω -effect enhanced by iron**
- **WEAK α -effect, several possibilities**

- $m=0$ is promoted by ferro impellers
- the neutral mode is NOT the one predicted for $\mu_r=1$

Generation mechanism and ferromagnetic impellers

A. Giesecke et al., *PRL* 104 (2010)

Kinematic approach,
MND flow
 μ_r variations implemented in rotating disk
uniform α

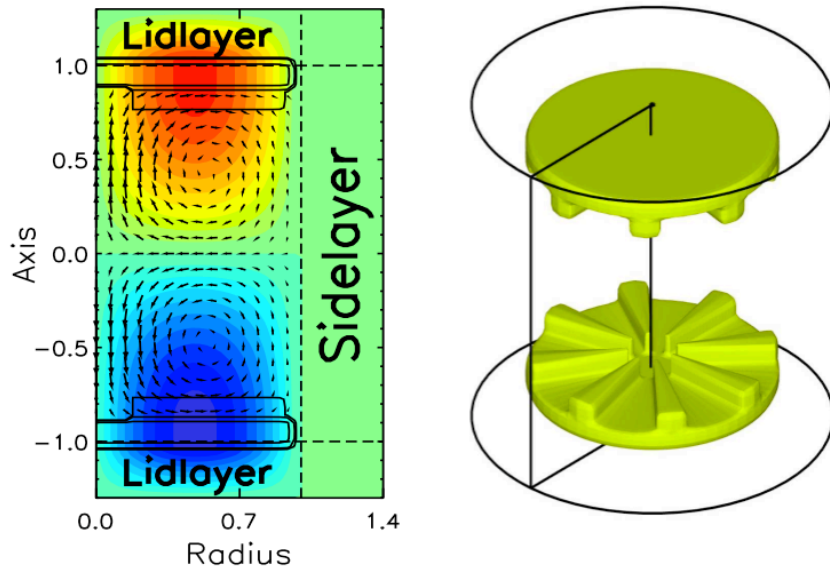
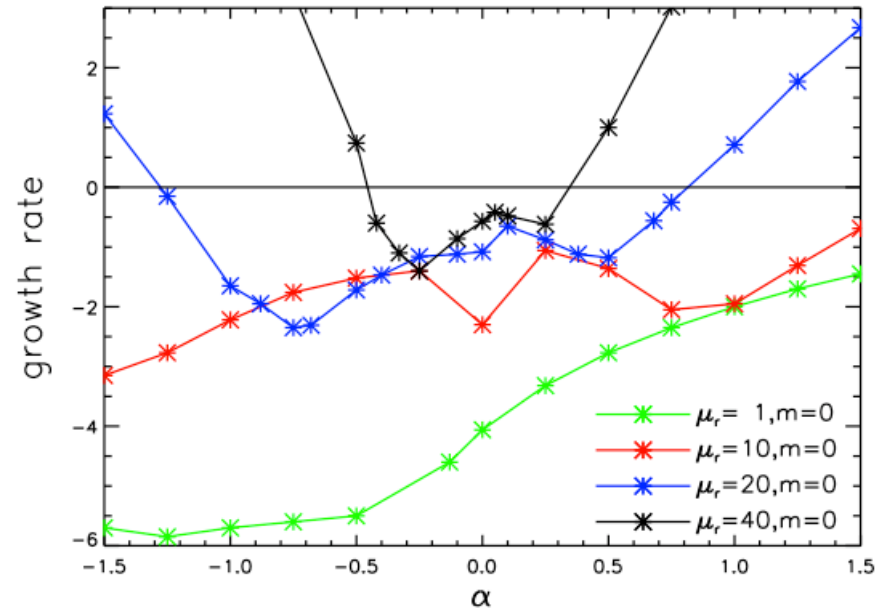
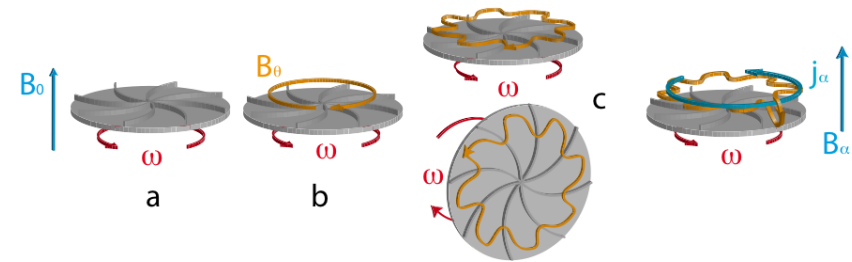


FIG. 1: Left panel: Structure of the prescribed axisymmetric velocity field. The color coded pattern represents the azimuthal velocity and the arrows show the poloidal velocity field. The black solid lines represent the shape of the impeller system (disk and blades). Right panel: assumed permeability distribution; isolevel at $\mu_r = 35$ where the peak value is given by $\mu_r = 40$. In the fluid region μ_r is equal to unity.



- $\alpha \approx 0.15$ m/s for $\mu_r = 40$
- $\alpha \approx 15$ m/s for $\mu_r = 1$

LARGE ω , WEAK α



α - ω dynamo

Threshold corresponds to $Rm^\alpha \cdot Rm^\omega \approx 1$

$$Rm^\alpha = \frac{\alpha h}{\eta} \sim 0.2\alpha$$

$$Rm^\omega = K \frac{Uh}{\eta} \sim 0.4K$$

$$Rm^\alpha \cdot Rm^\omega \sim 0.1 \alpha K$$

$\alpha \approx 0.1$, Ga measurements

Stepanov et al., *PRE* **73** (2006)

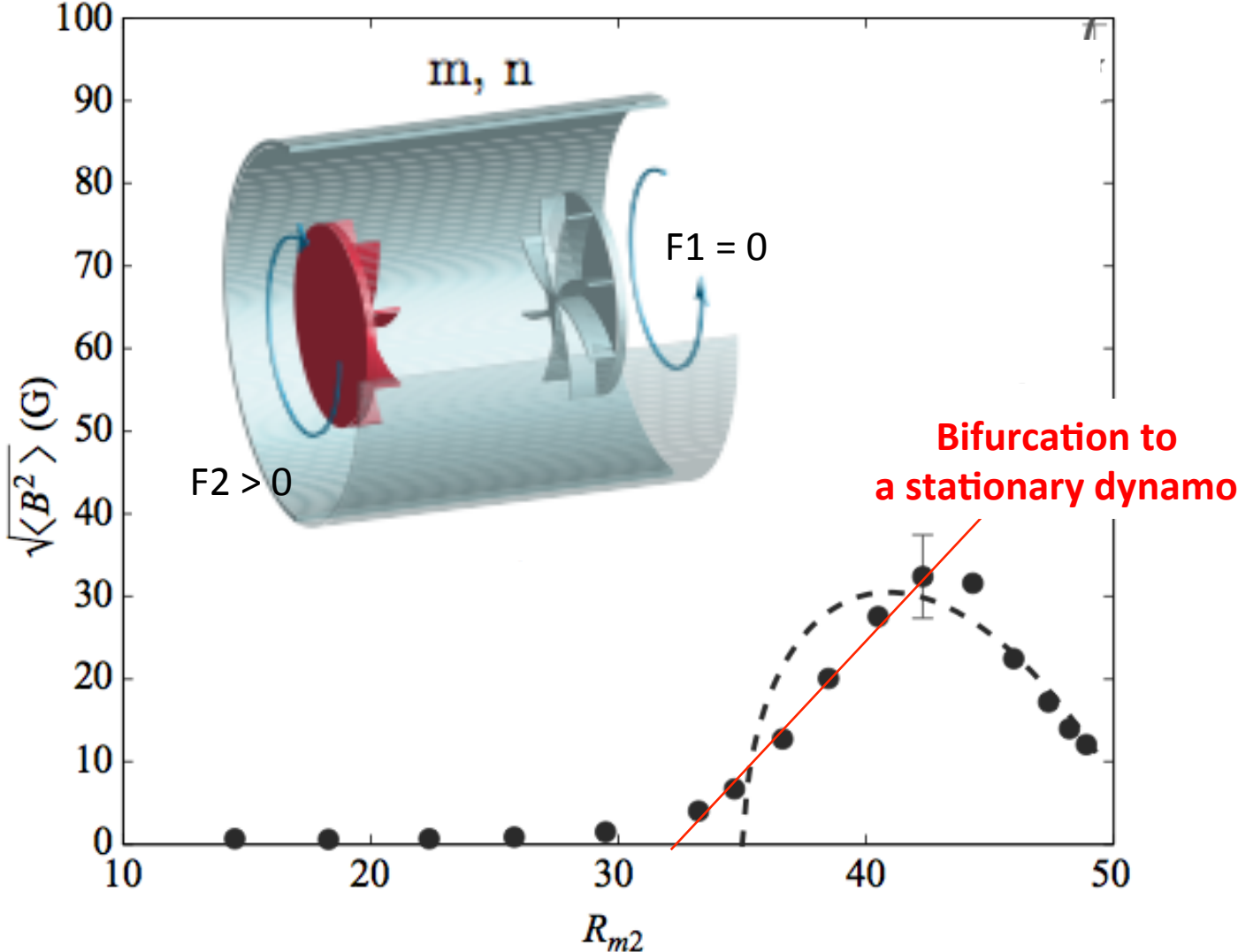
K: iron enhancement,

$K \approx 100$, Ga measurements.

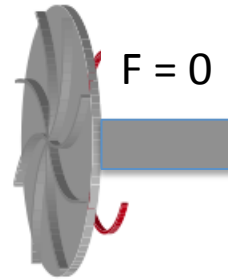
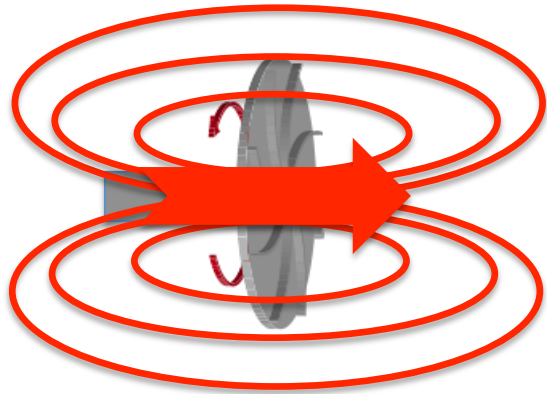
G. Verhille et al., *NJP* **12** (2010)

1Disk – dynamo generation

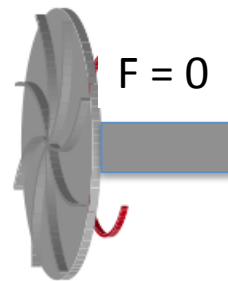
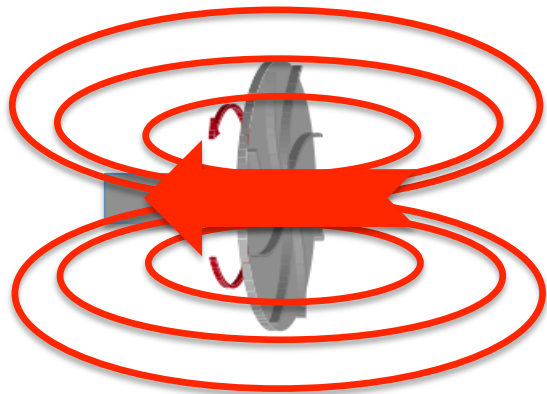
M. Berhanu et al., *JFM* 641 (2009)



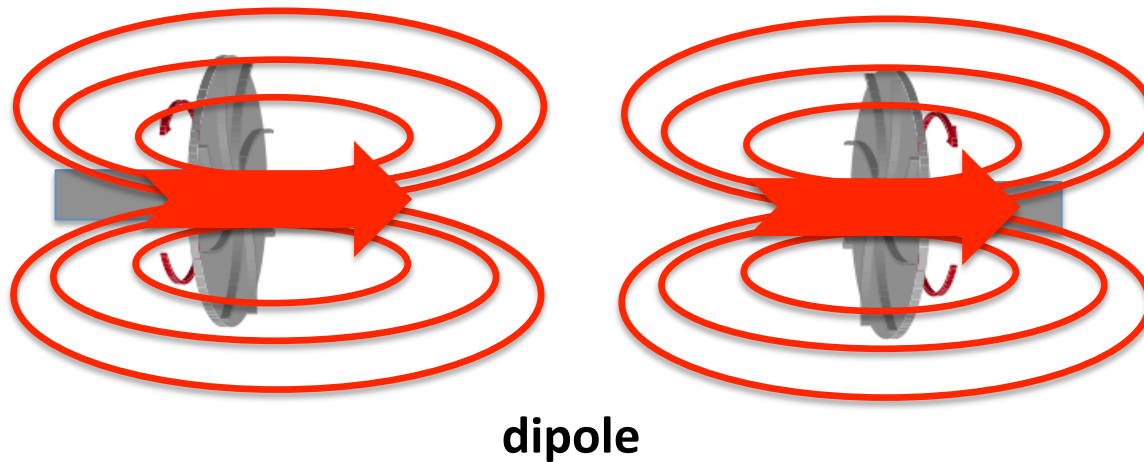
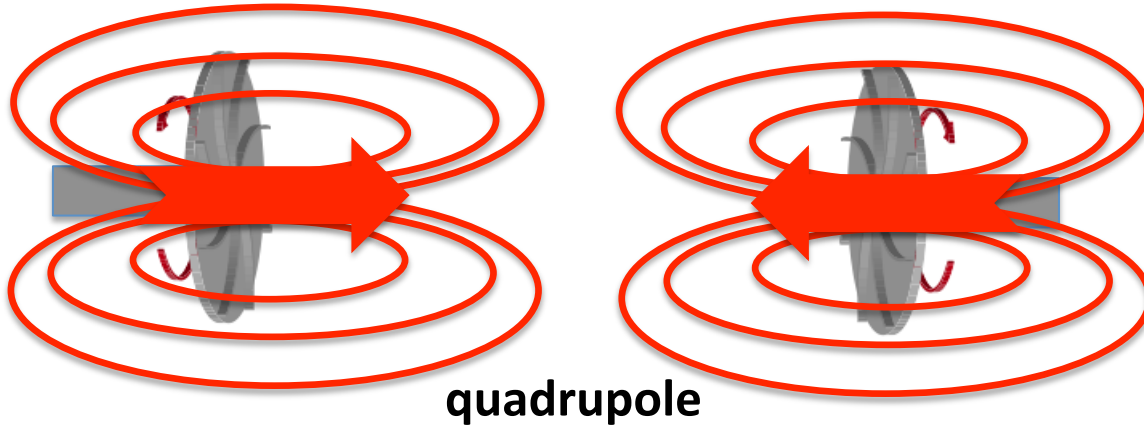
VKS / 1 Disk dynamo



B and $-B$ are
equally possible



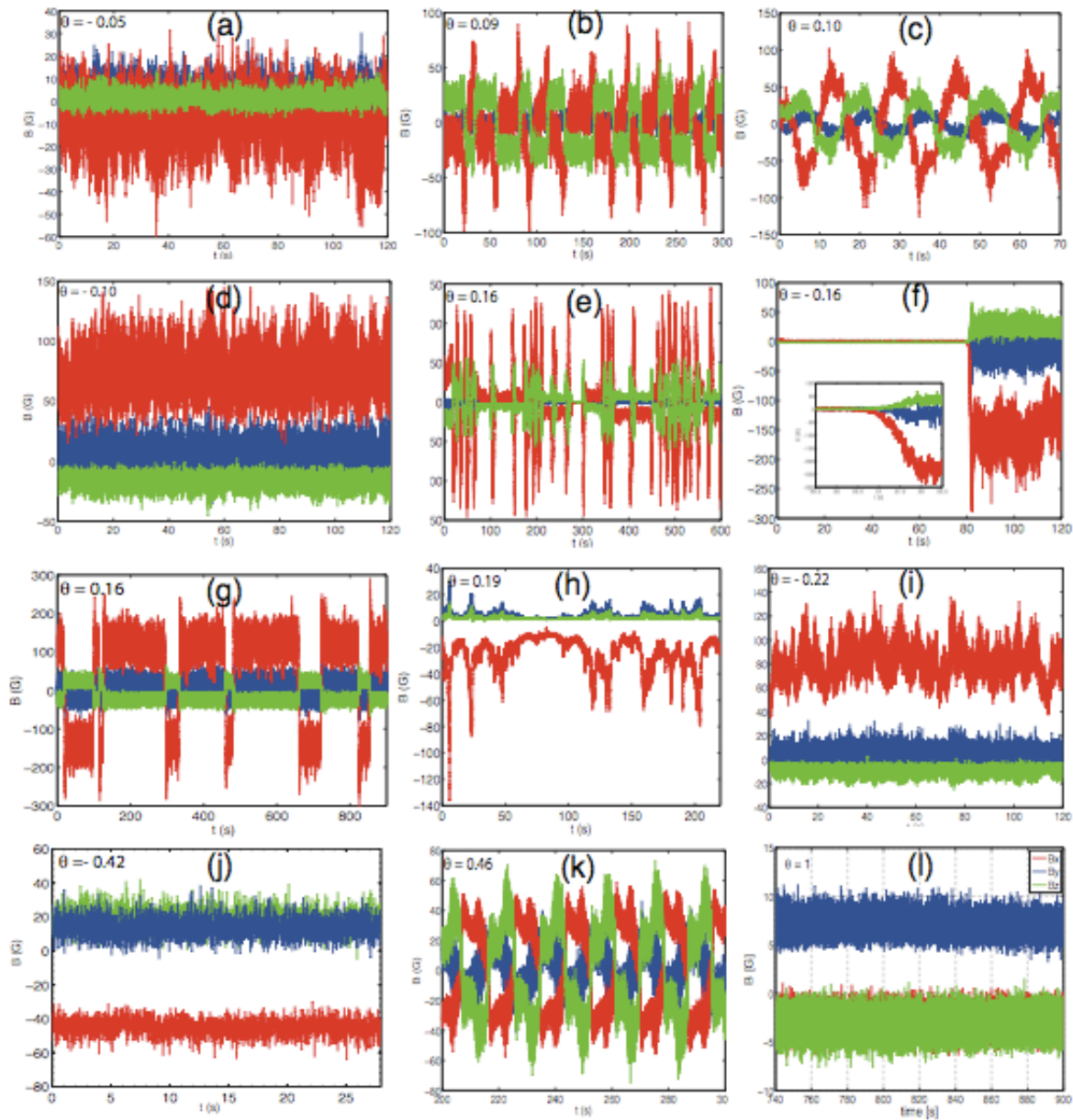
VKS / 2 disks : dynamo modes



$$\begin{aligned}\partial_t D &= f_{NL}(D, Q) \\ \partial_t Q &= g_{NL}(D, Q)\end{aligned}$$

→ Dynamics !

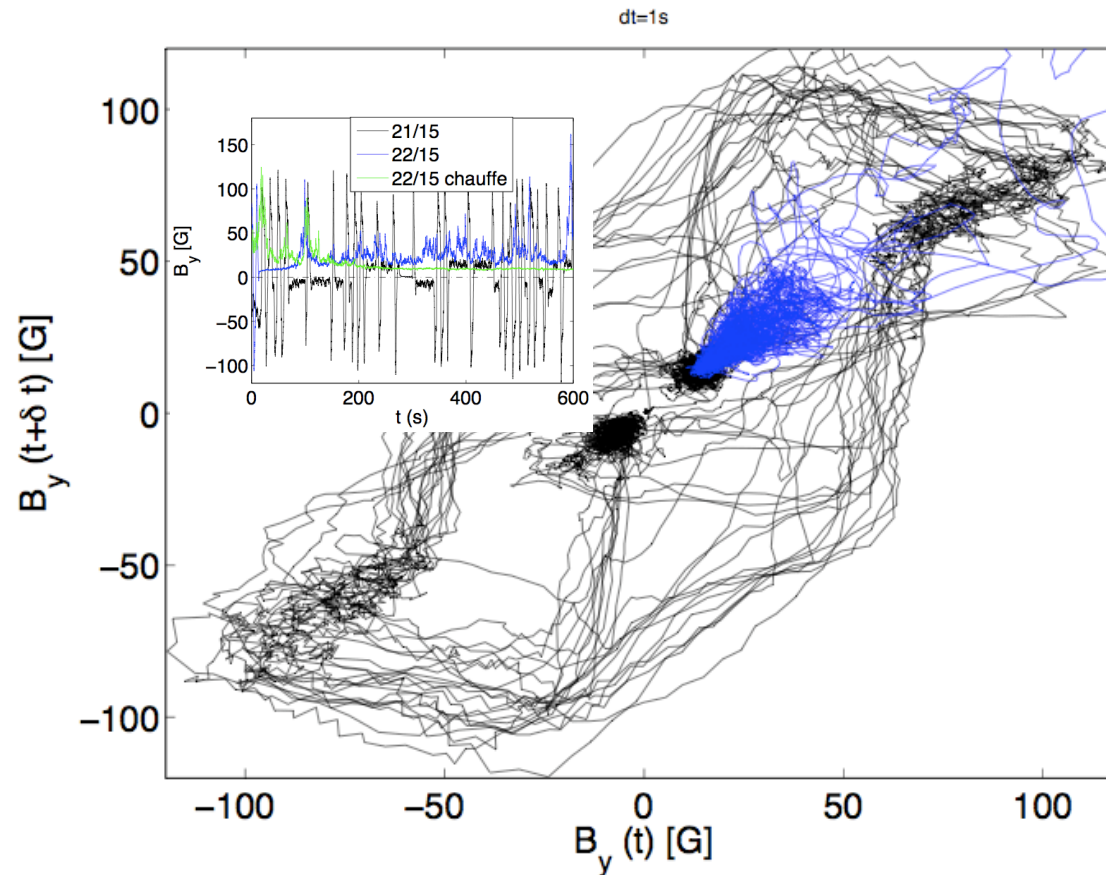
Monchaux et al., *PoF* 21 (2009)



Low dimensional behavior

R. Monchaux et al., *Phys. Rev. Lett.* **101**, 074502 (2008)

F. Petrelis & S. Fauve, *J Phys-CM* **20** (2008)

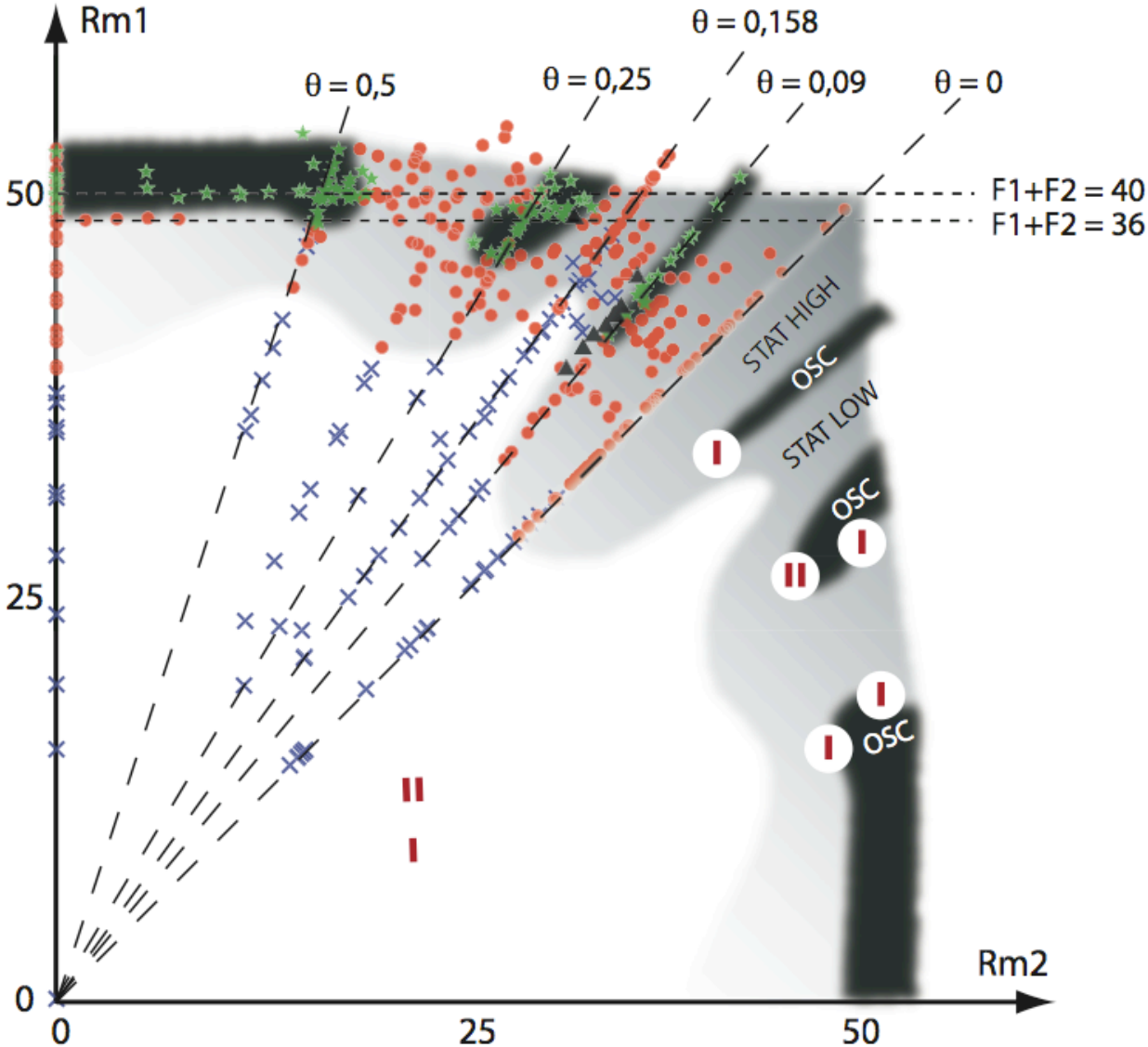


$$A = D + iQ = R \exp(i(\theta + \theta_0))$$

$$\dot{A} = \mu A + \nu \bar{A} + \beta_1 A^3 + \beta_2 A^2 \bar{A} + \beta_3 A \bar{A}^2 + \beta_4 \bar{A}^3$$

Parameter space

M. Berhanu et al., *Euro. Phys. J. B* (2010)



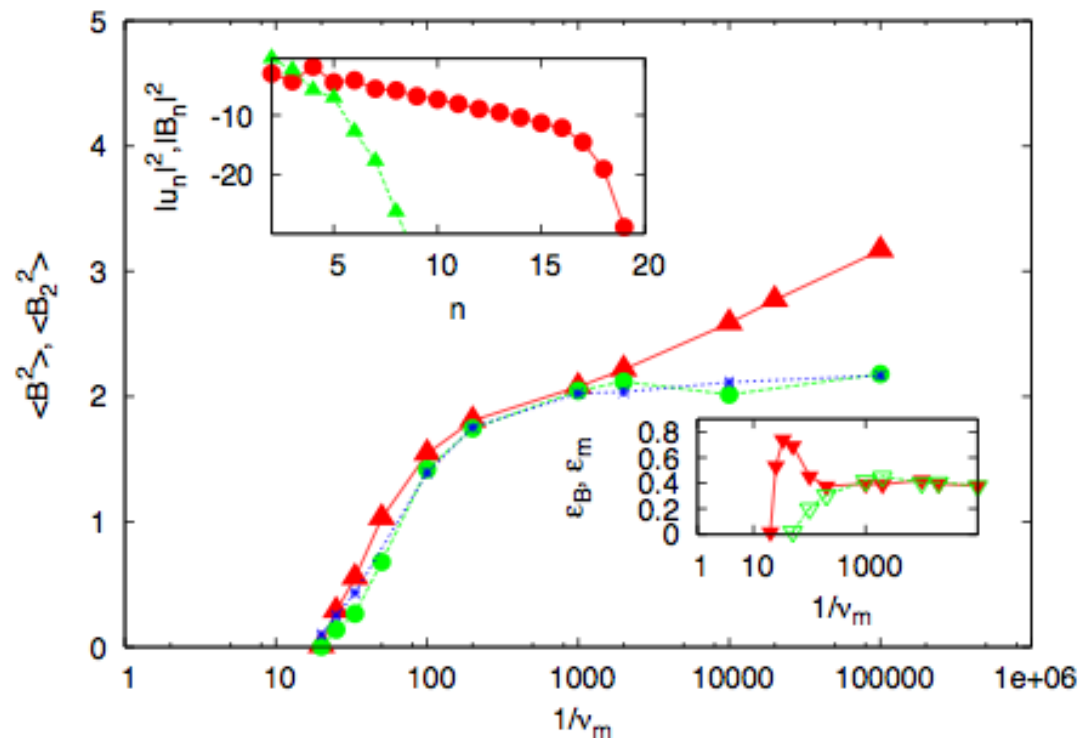
Shell model: deterministic dynamics + "turbulent" fluctuations

R. Benzi, JFP, *PRL* **105** (2010)

$$\frac{du_n}{dt} = \frac{i}{3}(\Phi_n(u, u) - \Phi_n(B, B)) - \nu k_n^2 u_n + f_n ,$$

$$\frac{dB_n}{dt} = \frac{i}{3}(\Phi_n(u, B) - \Phi_n(B, u)) - \nu_m k_n^2 B_n ,$$

$$\frac{dB_2}{dt} = F_2(u, B) - M_2(B_2) - \nu_m k_2^2 B_2$$



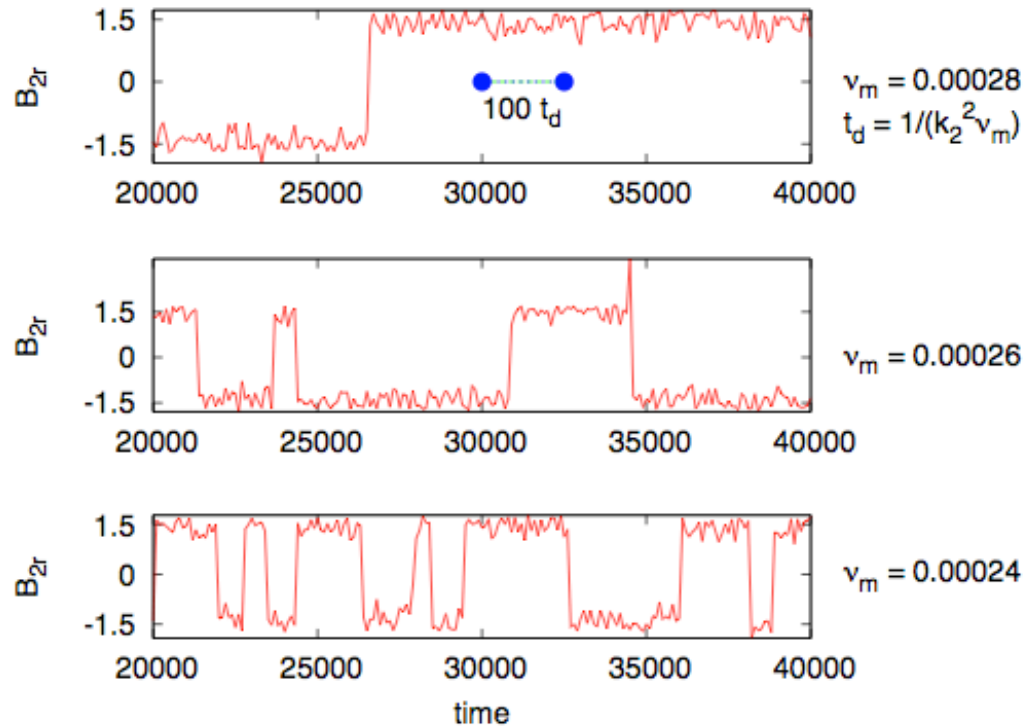
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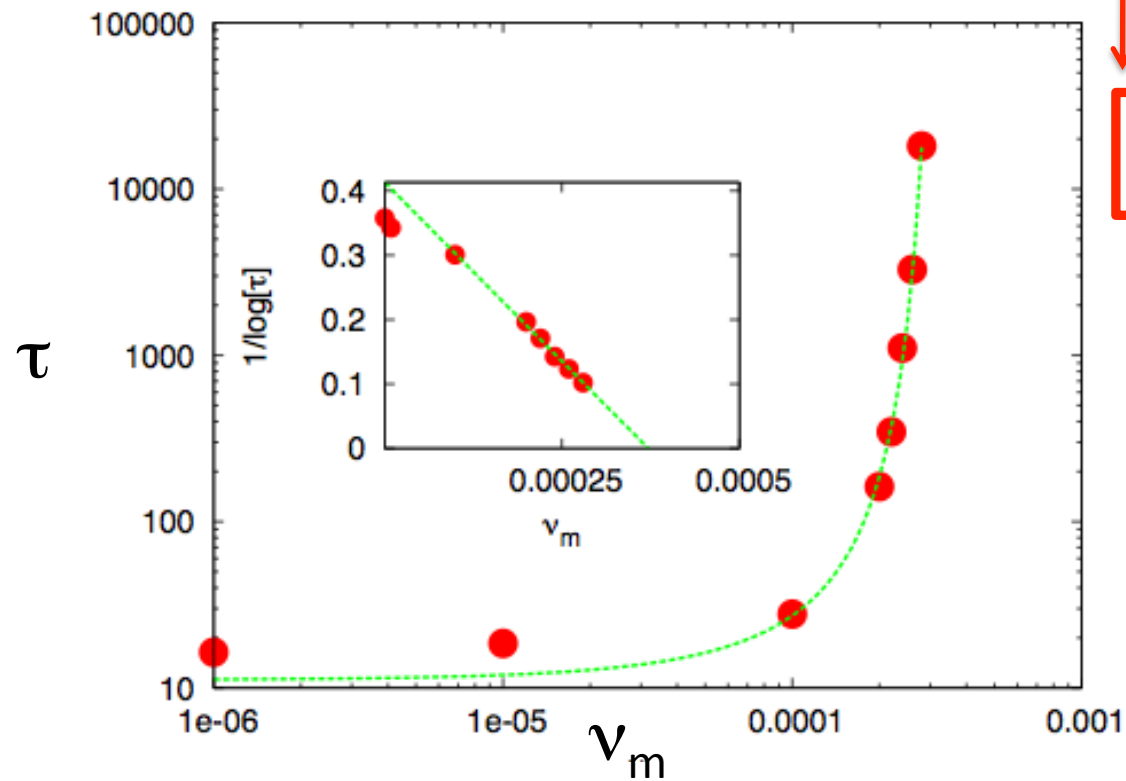


Shell model: deterministic dynamics + “turbulent” fluctuations

$$\frac{dB_2}{dt} = F_2(u, B) - M_2(B_2) - \nu_m k_2^2 B_2$$

$$F_2(u, B) = \beta B_2 + \phi'$$

$$\frac{dB_2}{dt} = \beta B_2 - a_m B_2^3 + \phi' \quad \tau \sim \exp\left(\frac{\beta^2}{a_m \sigma}\right) = \exp\left(\frac{A}{\sigma}\right)$$



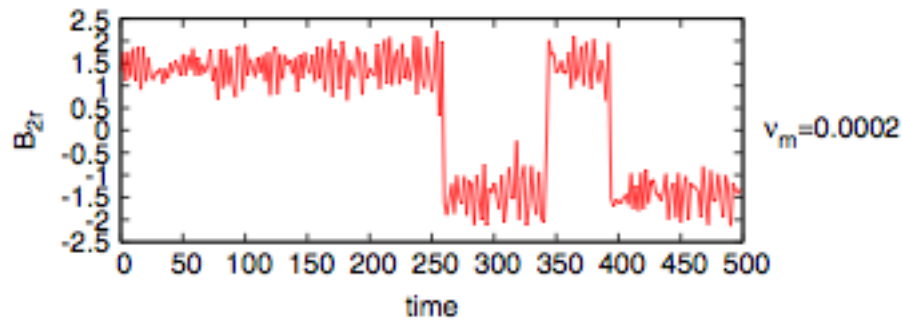
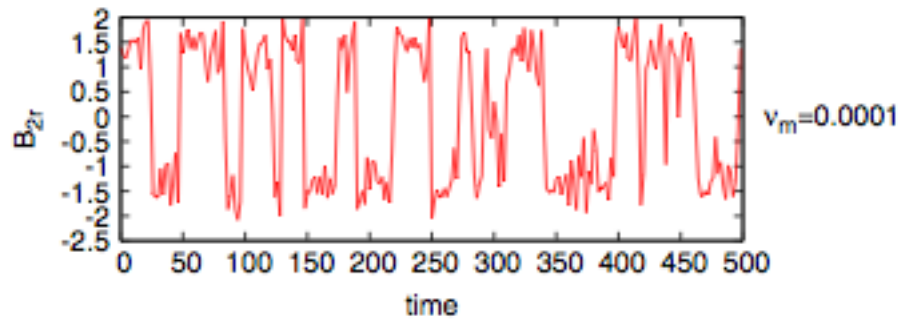
governed by variance of “noise”

Shell model: deterministic dynamics + “turbulent” fluctuations

$$\frac{dB_2}{dt} = F_2(u, B) - M_2(B_2) - \nu_m k_2^2 B_2$$

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Time-scale separation !

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the α - ω ferroDynamo.

mechanism and low dimensional behavior

turbulence and noise.

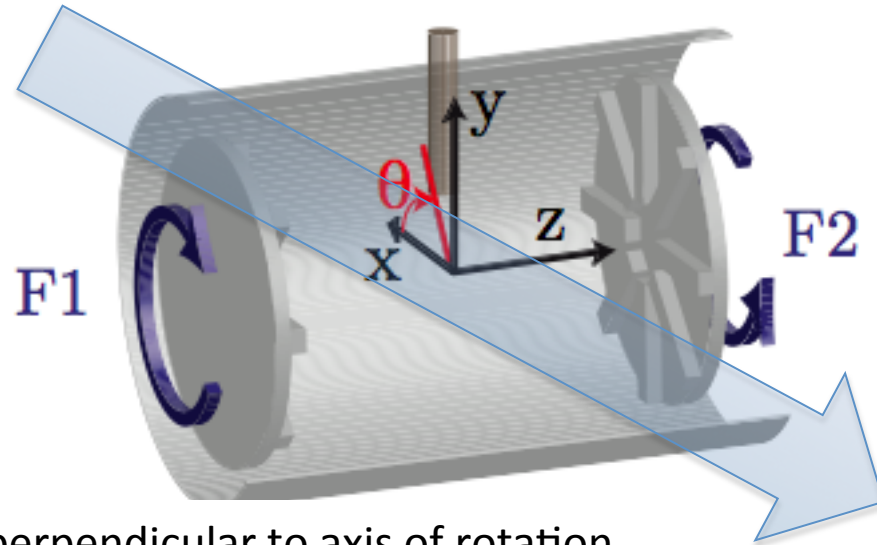
4. **Turbulent MHD**

Saturation: large B-field experiment.

5. **Next**

Magnetic field and vorticity interactions

G. Verhille, preprint (2011)



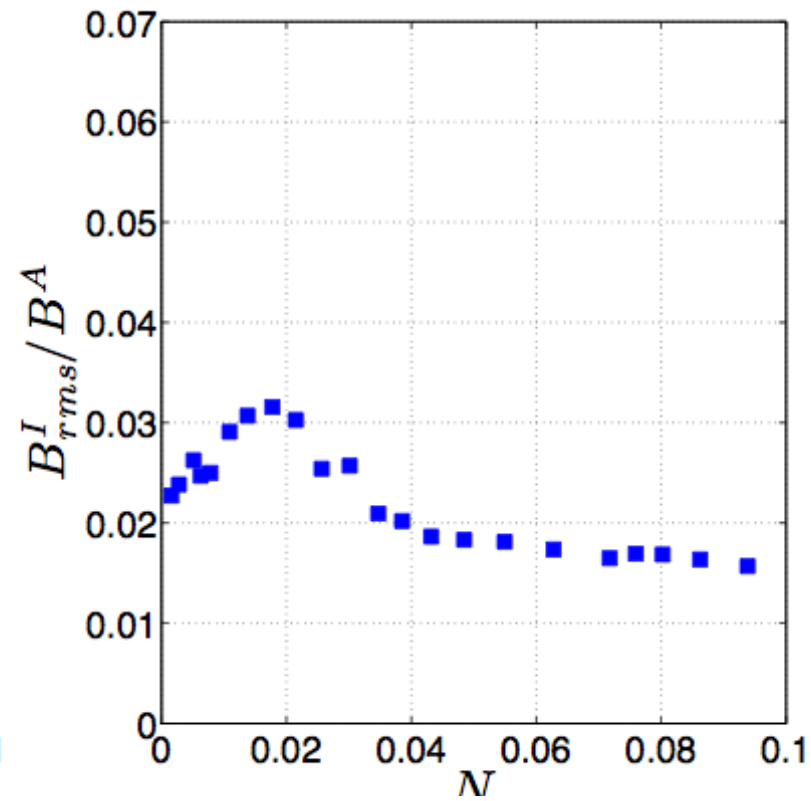
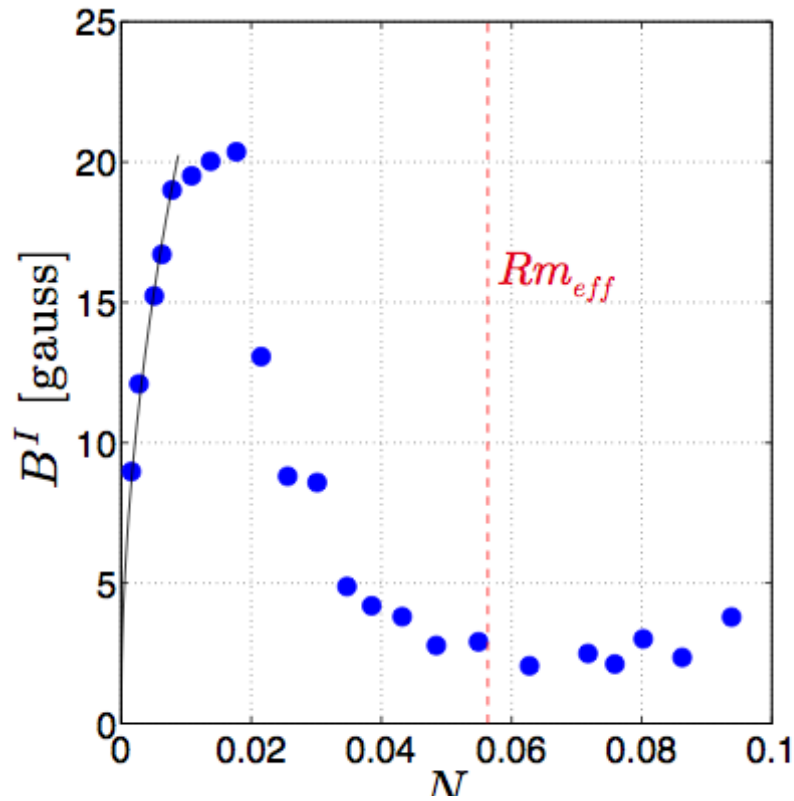
\mathbf{B}^A perpendicular to axis of rotation

\mathbf{B}^A in [0 – 1500] G

N in [0 – 0.5]

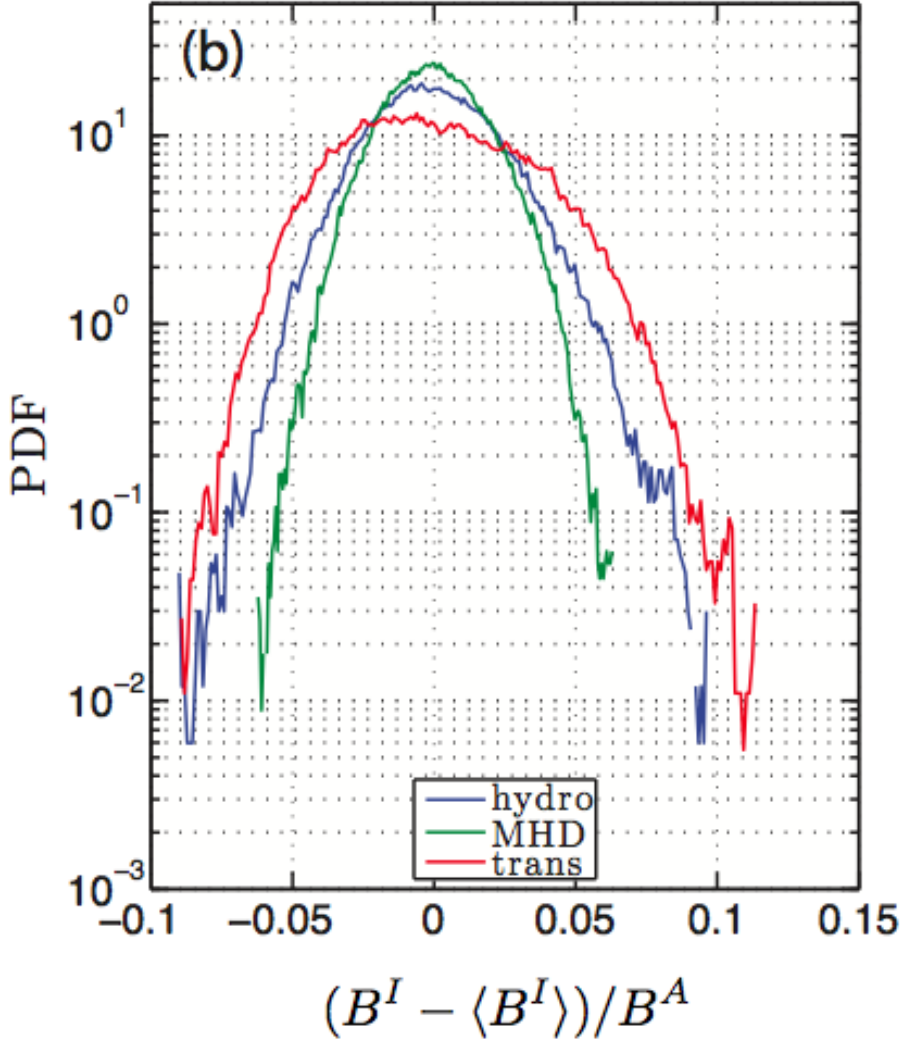
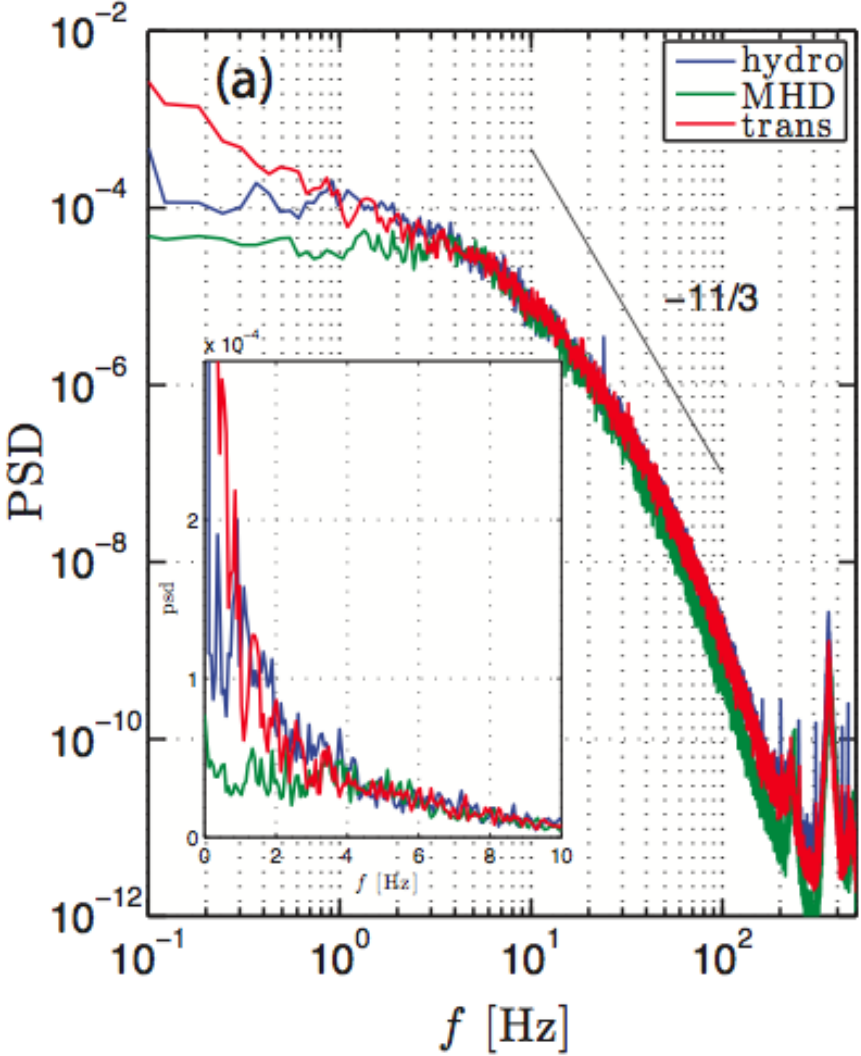
Magnetic field and vorticity interactions

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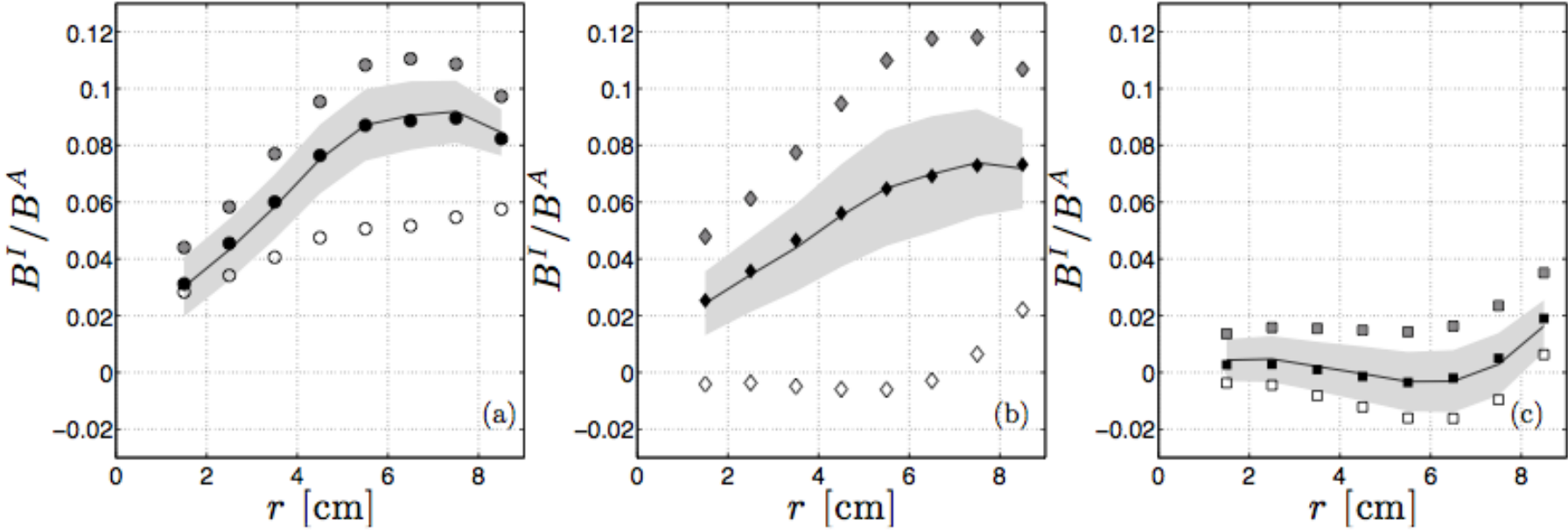


$$N = \frac{\sigma R B^A^2}{\rho u_{rms}} = \mu \sigma R u_{rms} \cdot \frac{B^A^2}{\rho \mu u_{rms}^2} = Rm \cdot \left(\frac{v_A}{u_{rms}} \right)^2$$

Magnetic field and vorticity interactions



Magnetic field and vorticity interactions



OUTLINE

1. **Dynamos**

Scalar, scalar gradient and B-equations
Is turbulence helpful or detrimental?
Experiments

2. **Mean field MHD**

easy and hard
Examples
Measurements

3. **Turbulence in VKS**

the α - ω “Duddley & James” failed dynamo
the α - ω ferroDynamo.
mechanism and low dimensional behavior
turbulence and noise.

4. **Turbulent MHD**

Saturation: large B-field experiment.

5. **Next**

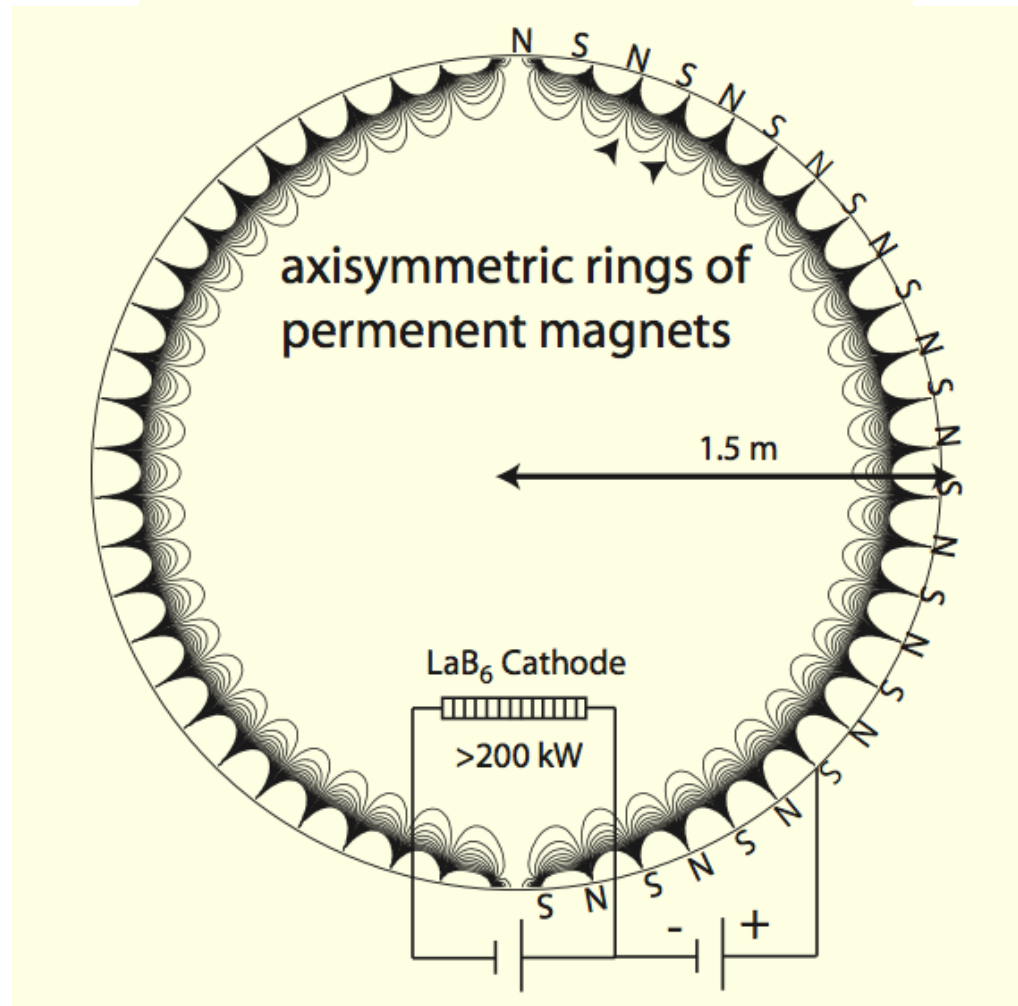
Quantity	Symbol	Value	Unit
Plasma radius	a	1.1	m
Number density	n	10^{17} – 10^{19}	m^{-3}
Ion temperature	T_i	0.5–4	eV
Electron temperature	T_e	2–10	eV
Peak speed	U_{max}	0–20	km s^{-1}
Ion species	H, He, Ar	1, 4, 40	amu
Pulse length	τ_{pulse}	5	s
Plasma beta	β	10^4	
Resistive time	τ_η	50	ms
Magnetic Reynolds number	Rm_{max}	~ 1000 – 2000	
Reynolds number	Re	2.4×10^1 to 3.8×10^6	
Magnetic Prandtl number	Pm	3.0×10^{-4} to 5.6×10^1	

$$Pm = 0.18 \frac{T_e^{3/2} T_i^{5/2}}{\mu^2 n} \quad Rm = 1.5 \frac{T_e^{3/2} U a}{Z}$$

Wisconsin plasma dynamo

Cary Forest

Rm_{max}	> 1000
Re	$24 - 3.8 \times 10^6$
Pm	$3 \times 10^{-4} - 56$



Lyon team

Gautier Verhille, Sophie Miralles

Mickael Bourgoïn, Philippe Odier, Nicolas Plihon, Romain Volk



VKS

F. Daviaud team in CEA-Saclay, S. Fauve team in ENS-Paris

