



Holographic Entanglement Entropy

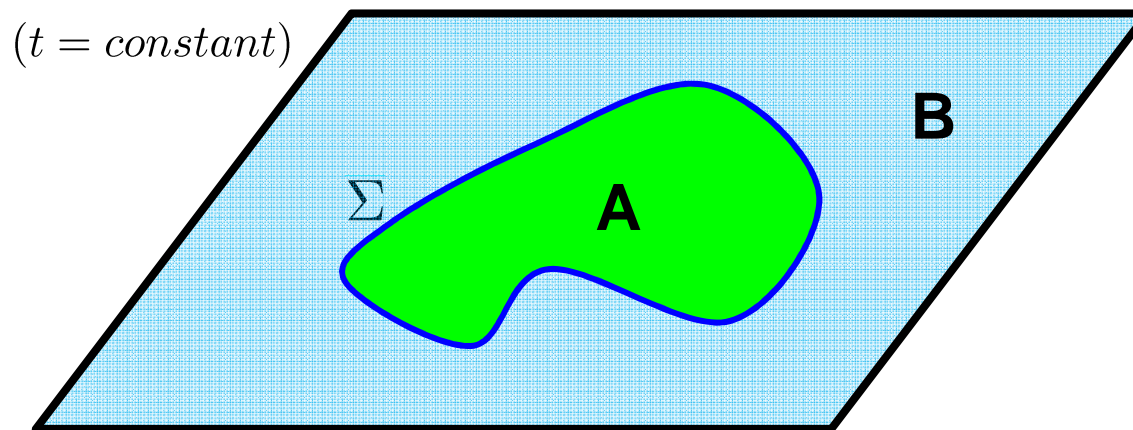
*and Renyi
Entropy*

(with H. Casini, M. Huerta, J. Hung, M. Smolkin & A. Yale)

(arXiv:1102.0440, arXiv:1110.1084)

Entanglement Entropy

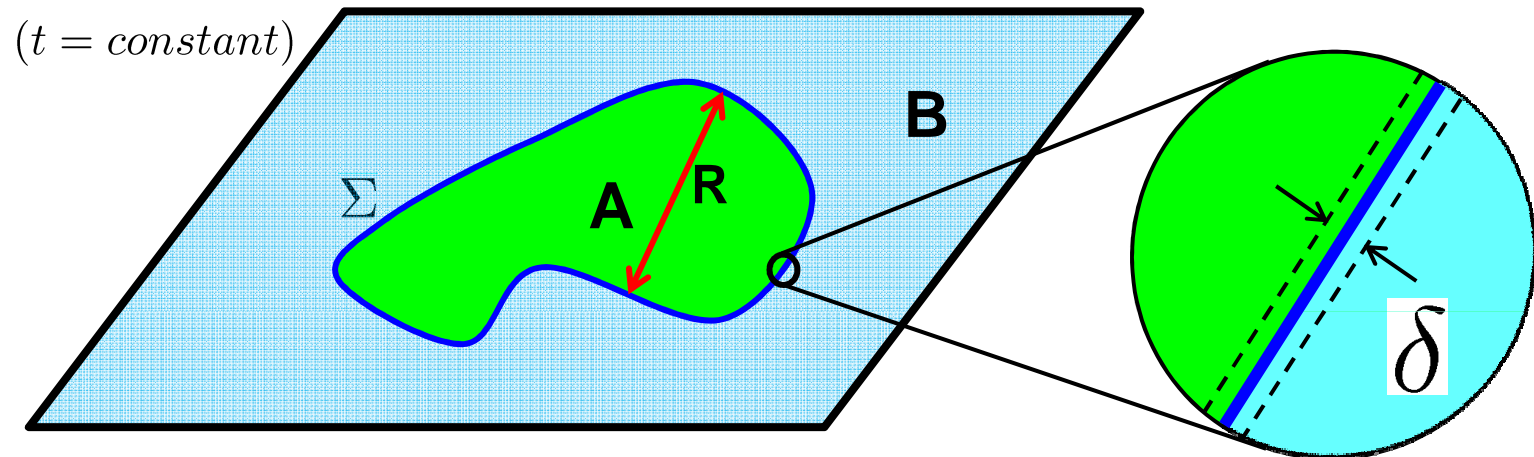
- what is entanglement entropy?
general tool; divide quantum system into two parts and use entropy as measure of correlations between subsystems
 - in QFT, typically introduce a (smooth) boundary **or entangling surface** Σ which divides the space into two separate regions
 - integrate out degrees of freedom in “outside” region
 - remaining dof are described by a density matrix ρ_A
- calculate **von Neumann entropy**: $S_{EE} = -Tr [\rho_A \log \rho_A]$



Entanglement Entropy

- remaining dof are described by a density matrix ρ_A

→ calculate von Neumann entropy: $S_{EE} = -Tr [\rho_A \log \rho_A]$



- result is **UV divergent!**
- must regulate calculation: $\delta = \text{short-distance cut-off}$

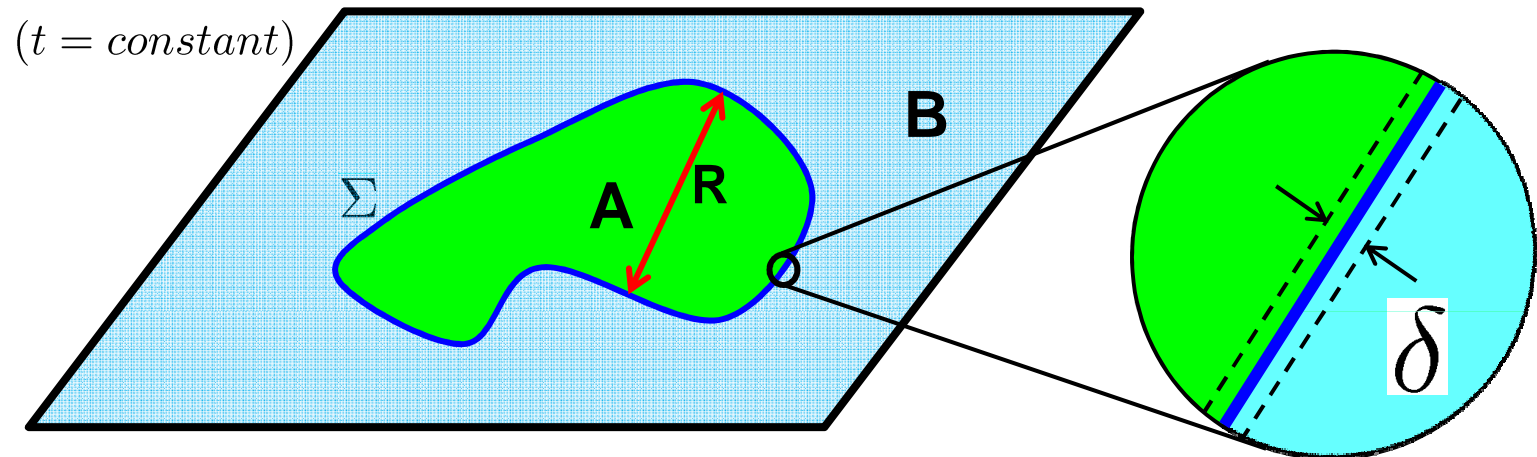
$$S = c_0 \frac{R^{d-2}}{\delta^{d-2}} + c_2 \frac{R^{d-4}}{\delta^{d-4}} + \dots \quad d = \text{spacetime dimension}$$

- careful analysis reveals geometric structure, eg, $S = \tilde{c}_0 \frac{A_\Sigma}{\delta^{d-2}} + \dots$

Entanglement Entropy

- remaining dof are described by a density matrix ρ_A

→ calculate von Neumann entropy: $S_{EE} = -\text{Tr} [\rho_A \log \rho_A]$



- must regulate calculation: $\delta = \text{short-distance cut-off}$

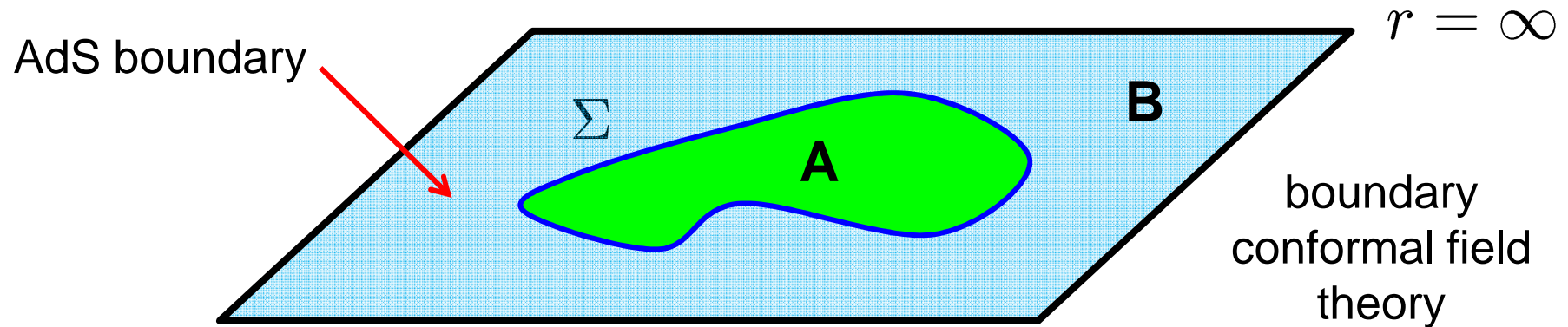
$$S = c_0 \frac{R^{d-2}}{\delta^{d-2}} + c_2 \frac{R^{d-4}}{\delta^{d-4}} + \dots \quad d = \text{spacetime dimension}$$

- leading coefficients sensitive to details of regulator, eg, $\delta \rightarrow 2\delta$
- find universal information characterizing underlying QFT in subleading terms, eg, $S = \dots + c_d \log(R/\delta) + \dots$

More general comments on **Entanglement Entropy**:

- nonlocal quantity which is (at best) very difficult to measure
→ no accepted experimental procedure
- in condensed matter theory: diagnostic to characterize quantum critical points or topological phases (eg, quantum hall fluids)
- in quantum information theory: useful measure of quantum entanglement (a computational resource)
- **black hole microphysics**: leading term obeys “area law” $S \simeq c_0 \frac{A_\Sigma}{\delta^{d-2}}$
→ suggested as origin of black hole entropy (eg, $\delta \simeq \ell_P$)
(Bombelli, Koul, Lee & Sorkin `86; Srednicki; Frolov & Novikov; Callan & Wilczek; Susskind;)
- recently considered in **AdS/CFT correspondence**
(Ryu & Takayanagi `06)

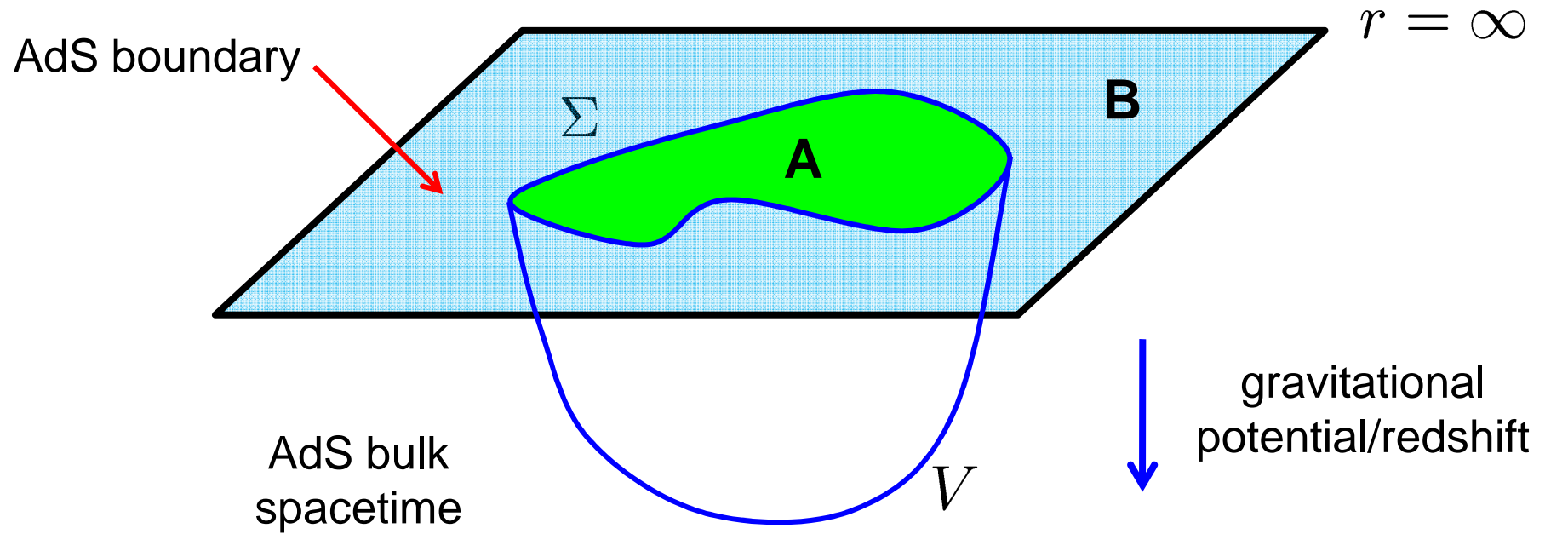
Holographic Entanglement Entropy:



AdS bulk
spacetime

$$S(A) = ??$$

Holographic Entanglement Entropy:



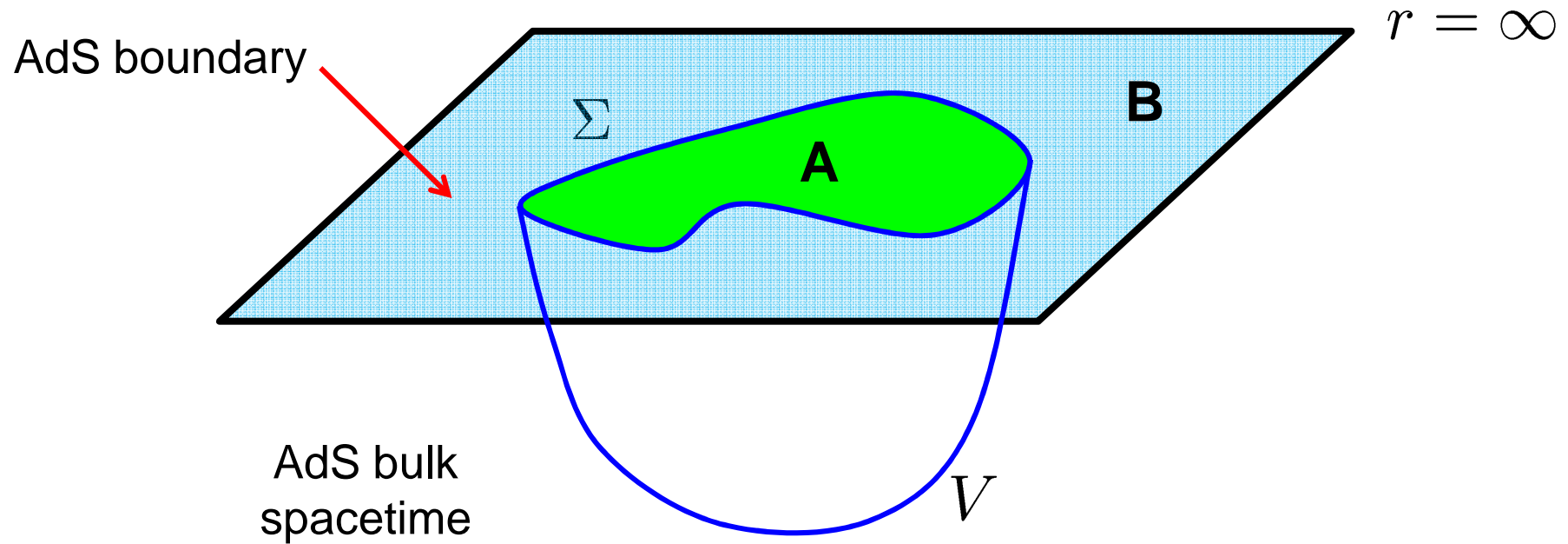
$$S(A) = \text{ext}_{\partial V = \Sigma} \frac{A_V}{4G_N}$$

$(d - 1)$ dimensional

looks like BH entropy!

$r = \infty$

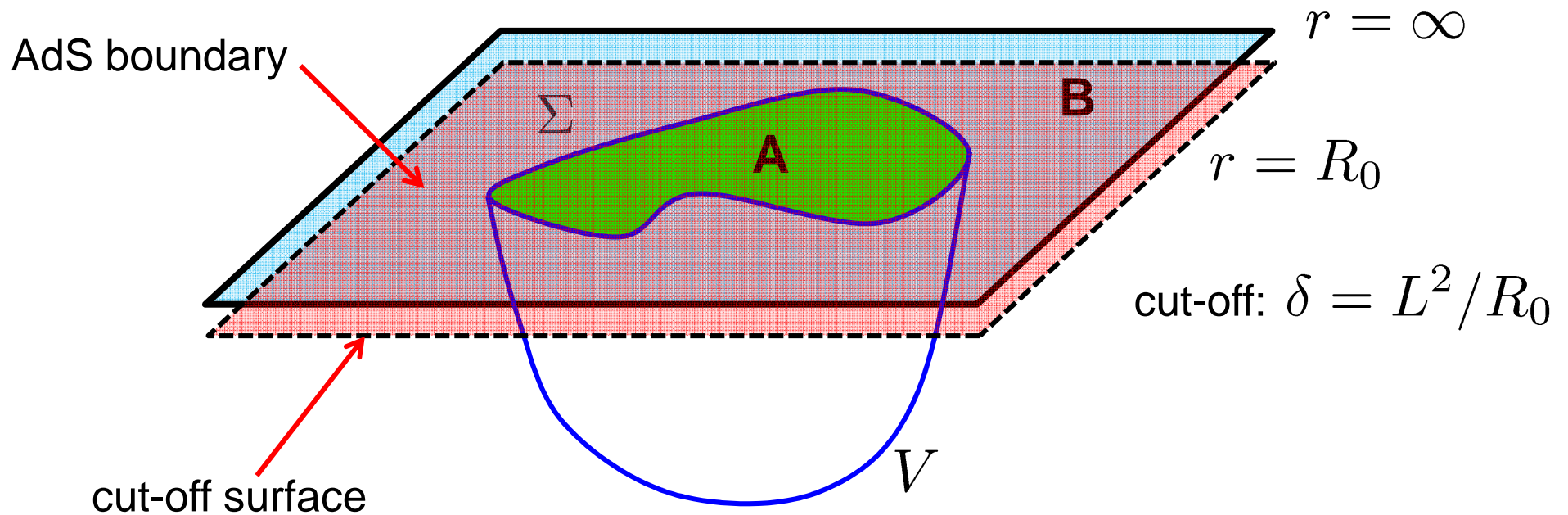
Holographic Entanglement Entropy:



$$S(A) = \text{ext}_{\partial V = \Sigma} \frac{A_V}{4G_N} = \infty!!$$

- “UV divergence” because area integral extends to $r = \infty$

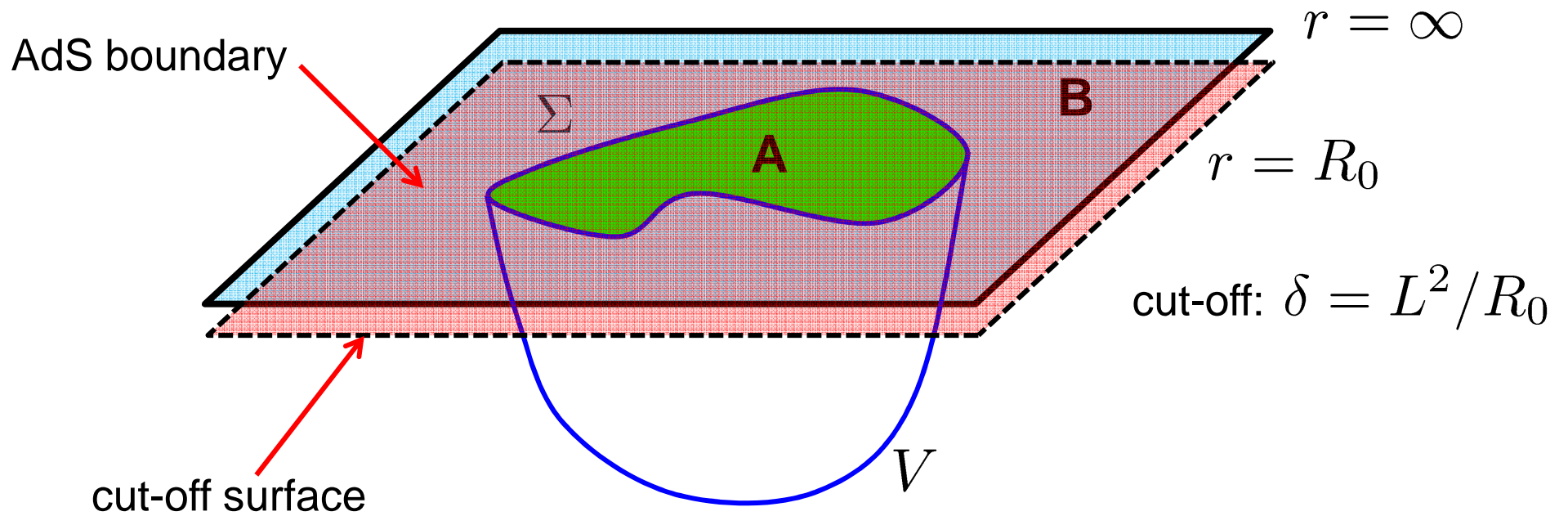
Holographic Entanglement Entropy:



$$S(A) = \text{ext}_{\partial V = \Sigma} \frac{A_V}{4G_N} \simeq \frac{L^{d-1}}{G_N} \frac{\mathcal{A}_\Sigma}{\delta^{d-2}} + \dots$$

- “UV divergence” because area integral extends to $r = \infty$
- finite result by stopping radial integral at large radius: $r = R_0$
 → short-distance cut-off in boundary theory: $\delta = L^2 / R_0$

Holographic Entanglement Entropy:

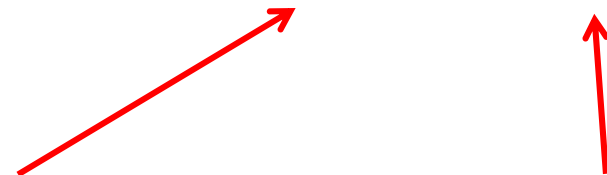


$$S(A) = \text{ext}_{\partial V = \Sigma} \frac{A_V}{4G_N} \simeq \frac{L^{d-1}}{G_N} \frac{\mathcal{A}_\Sigma}{\delta^{d-2}} + \dots$$

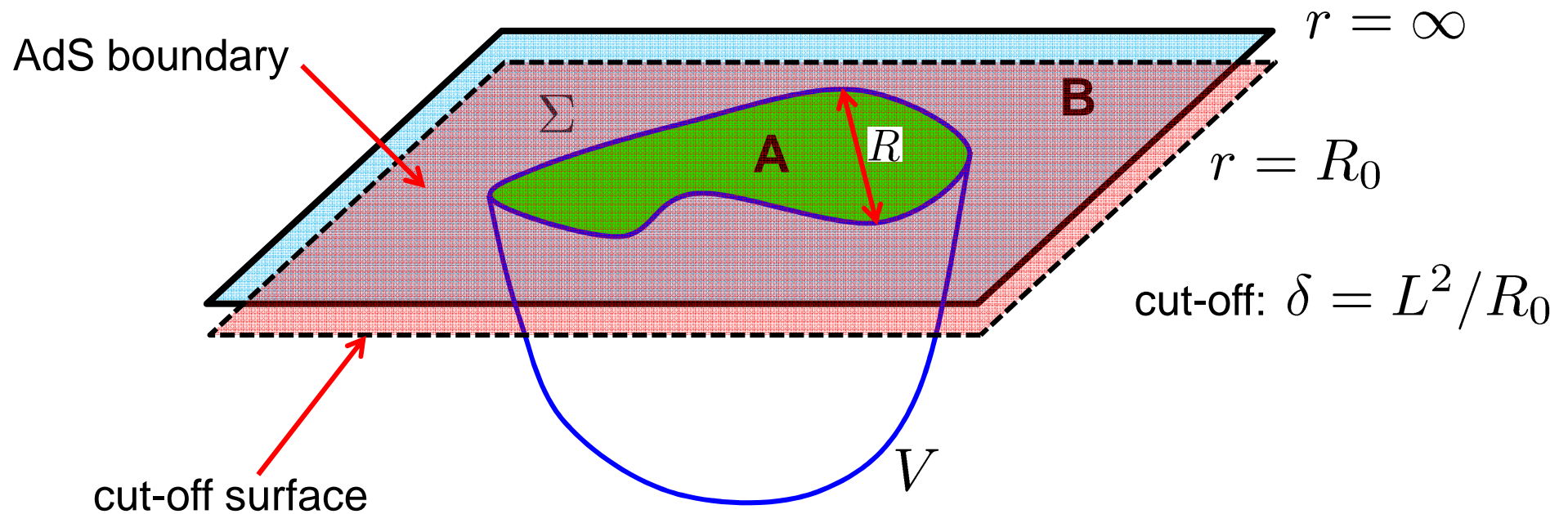
central charge
(counts dof)

$$(L/\ell_{Planck})^{d-1}$$

"Area Law"



Holographic Entanglement Entropy:



general expression (as desired):

$$S(A) \simeq c_0 (R/\delta)^{d-2} + c_1 (R/\delta)^{d-4} + \dots$$

$$\left\{ \begin{array}{l} +c_{d-2} \log(R/\delta) + \dots \quad (\text{d even}) \\ + \underbrace{c_{d-2} + \dots}_{\text{universal contributions}} \quad (\text{d odd}) \end{array} \right.$$

universal contributions

Holographic Entanglement Entropy:

$$S(A) = \text{ext}_{\partial V = \Sigma} \frac{A_V}{4G_N}$$

conjecture

Extensive consistency tests:

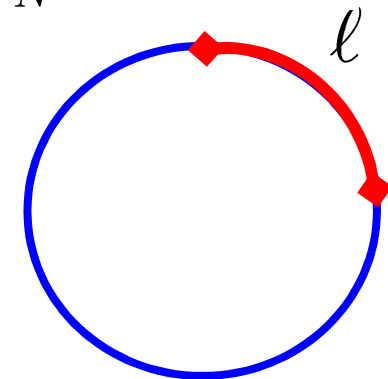
1) leading contribution yields “area law”

$$S \simeq \frac{L^{d-1}}{G_N} \frac{A_\Sigma}{\delta^{d-2}} + \dots$$

2) recover known results of Calabrese & Cardy
for d=2 CFT

$$S = \frac{c}{3} \log \left(\frac{C}{\pi \delta} \sin \frac{\pi \ell}{C} \right)$$

(also result for thermal ensemble)



$C = \text{circumference}$

Holographic Entanglement Entropy:

$$S(A) = \text{ext}_{\partial V = \Sigma} \frac{A_V}{4G_N}$$

conjecture

Extensive consistency tests:

1) leading contribution yields “area law” $S \simeq \frac{L^{d-1}}{G_N} \frac{A_\Sigma}{\delta^{d-2}} + \dots$

..... (lots of interesting tests)

7) connection to central charges of CFT for higher even d

(Hung, RCM & Smolkin, arXiv:1101.5813)

8) derivation of holographic EE for spherical entangling surfaces

(Casini, Huerta & RCM, arXiv:1102.044)

(see also: RCM & Sinha, arXiv:1011.5819)

7) connection to central charges of CFT for higher even d

(Hung, RCM & Smolkin, arXiv:1101.5813)

- trace anomaly in CFT (with even d) defines central charges

$$d=4: \quad \langle T_{\mu}^{\mu} \rangle = \frac{\mathbf{c}}{16\pi^2} I_4 - \frac{\mathbf{a}}{16\pi^2} E_4$$

$$I_4 = C_{\mu\nu\rho\sigma} C^{\mu\nu\rho\sigma} \quad \text{and} \quad E_4 = R_{\mu\nu\rho\sigma} R^{\mu\nu\rho\sigma} - 4R_{\mu\nu} R^{\mu\nu} + R^2$$

- universal/logarithmic contribution to entanglement entropy determined by central charges using trace anomaly, eg,

$$S_{uni} = \log(R/\delta) \frac{1}{2\pi} \int_{\Sigma} d^2x \sqrt{h} \left[\mathbf{c} \left(C^{ijkl} \tilde{g}_{ik}^{\perp} \tilde{g}_{jl}^{\perp} - K_a^{ib} K_b^{ia} + \frac{1}{2} K_a^{ia} K_b^{ib} \right) - \mathbf{a} \mathcal{R} \right]$$

(Takayanagi & Ryu; Schwimmer & Theisen; Solodukhin)

- R&T proposal for holographic EE exactly reproduces this result
- extends to certain higher curvature theories (eg, GB gravity)

$$S = \underset{\partial V = \Sigma}{\text{ext}} \frac{2\pi}{\ell_p^3} \int_V d^3x \sqrt{h} [1 + \lambda L^2 \mathcal{R}]$$

(see also: de Boer, Kulaxizi & Parnachev)

Holographic Entanglement Entropy:

$$S(A) = \text{ext}_{\partial V = \Sigma} \frac{A_V}{4G_N}$$

conjecture

Extensive consistency tests:

1) leading contribution yields “area law” $S \simeq \frac{L^{d-1}}{G_N} \frac{A_\Sigma}{\delta^{d-2}} + \dots$

..... (lots of interesting tests)

7) connection to central charges of CFT for higher even d

(Hung, RCM & Smolkin, arXiv:1101.5813)

8) derivation of holographic EE for spherical entangling surfaces

(Casini, Huerta & RCM, arXiv:1102.044)

(see also: RCM & Sinha, arXiv:1011.5819)

Calculating Entanglement Entropy:

$$S_{EE} = -Tr [\rho_A \log \rho_A]$$

- a “standard” approach relies on **replica trick**, first calculating **Renyi entropy** and then taking $n \rightarrow 1$ limit

$$S_n = \frac{1}{1-n} \log Tr [\rho_A^n] \qquad S_{EE} = \lim_{n \rightarrow 1} S_n$$

- **replica trick** involves path integral of QFT on **singular** n-fold cover of background spacetime
- problematic in holographic framework
 - produce singularity in dual gravity description
(resolved by quantum gravity/string theory?)

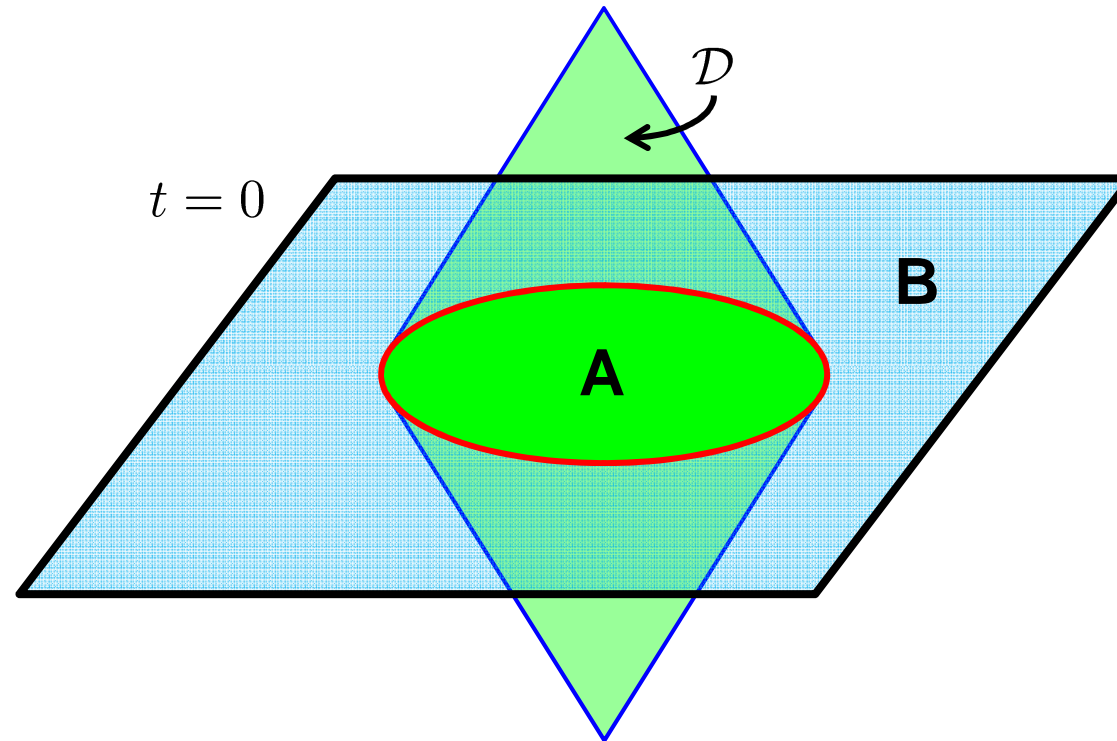
(Fursaev; Headrick)

- need another calculation with simpler holographic translation

Calculating Entanglement Entropy:

(Casini, Huerta & RCM)

- take **CFT** in d-dim. flat space and choose $\Sigma = S^{d-2}$ with radius R
→ entanglement entropy: $S_{EE} = -Tr [\rho_A \log \rho_A]$

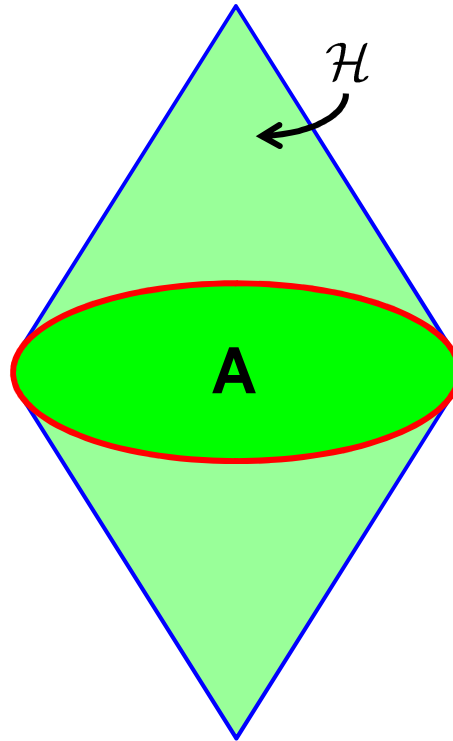


- density matrix ρ_A describes physics in entire causal domain \mathcal{D}
- conformal mapping: $\mathcal{D} \rightarrow \mathcal{H} = R_t \times H^{d-1}$

General result for any CFT

(Casini, Huerta & RCM)

- take CFT in d-dim. flat space and choose $\Sigma = S^{d-2}$ with radius R
→ entanglement entropy: $S_{EE} = -Tr [\rho_A \log \rho_A]$



- conformal mapping: $\mathcal{D} \rightarrow \mathcal{H} = R_t \times H^{d-1}$

curvature scale: $1/R$

temperature: $T=1/2\pi R$!!

- for CFT: $\rho_{thermal} = U \rho_A U^{-1} \longrightarrow \boxed{S_{EE} = S_{thermal}}$

General result for any CFT

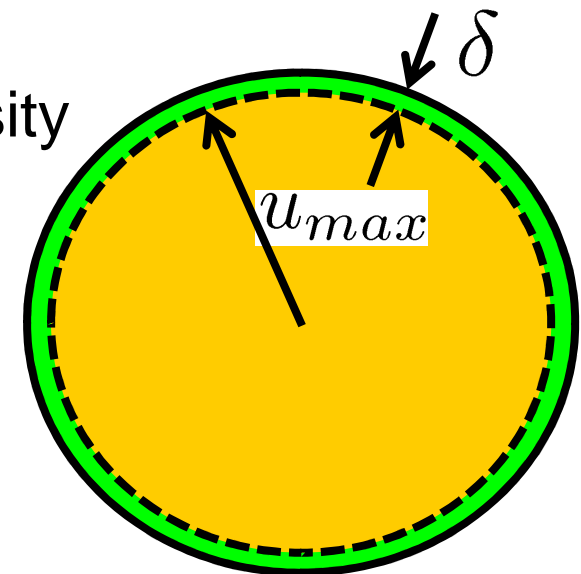
(Casini, Huerta & RCM)

- take CFT in d-dim. flat space and choose S^{d-2} with radius R
 - entanglement entropy: $S_{EE} = -Tr [\rho_A \log \rho_A]$
 - by conformal mapping relate to thermal entropy on $\mathcal{H} = R \times H^{d-1}$ with $\mathcal{R} \sim 1/R^2$ and $T=1/2\pi R$

$$S_{EE} = S_{thermal}$$

- note both sides of equality are divergent
 - $S_{thermal}$ sums constant entropy density over infinite volume
- conformal map takes original UV cut-off to IR cut-off on H^{d-1}

$$u_{max} \simeq R/\delta$$



General result for any CFT

(Casini, Huerta & RCM)

- take CFT in d-dim. flat space and choose S^{d-2} with radius R
 - entanglement entropy: $S_{EE} = -Tr [\rho_A \log \rho_A]$
 - by conformal mapping relate to thermal entropy on $\mathcal{H} = R \times H^{d-1}$ with $\mathcal{R} \sim 1/R^2$ and $T=1/2\pi R$

$$S_{EE} = S_{thermal}$$

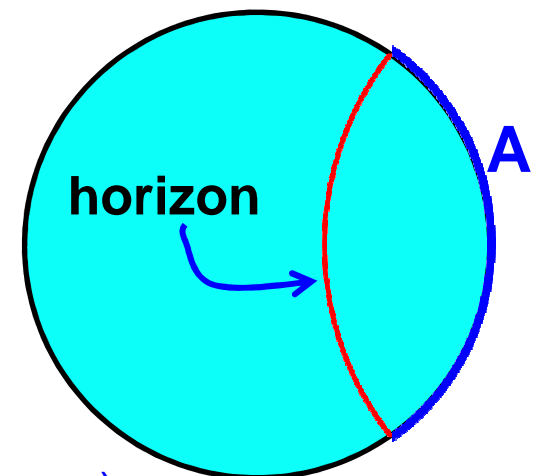
AdS/CFT correspondence:

- thermal bath in CFT = black hole in AdS

$$S_{EE} = S_{thermal} = S_{horizon}$$

- only need to find appropriate black hole
 - topological BH with hyperbolic horizon which intersects ∂A on AdS boundary

(Aminneborg et al; Emparan; Mann; ...)



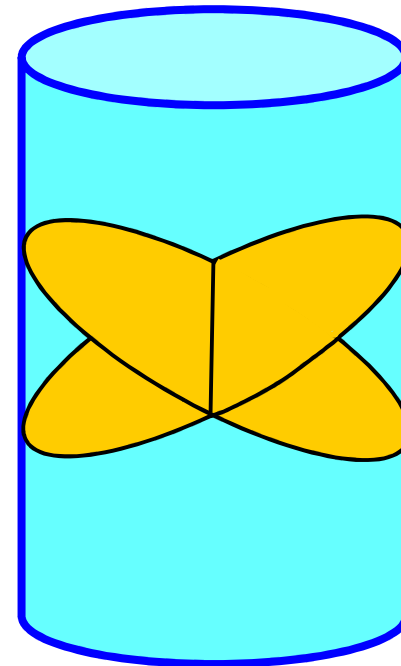
$$S_{EE} = S_{thermal} = S_{horizon}$$

- desired “black hole” is a hyperbolic foliation of empty AdS space

$$ds^2 = \frac{L^2}{z^2} (dz^2 - dt^2 + d\vec{x}^2) d\tau^2 + \rho^2 d\Sigma_2^{d-1} \quad \longrightarrow \quad T = \frac{1}{2\pi R}$$

- bulk coordinate transformation implements desired conformal transformation on boundary

- “Rindler coordinates” of AdS space:



$$S_{EE} = S_{thermal} = S_{horizon}$$

- desired “black hole” is a hyperbolic foliation of empty AdS space

$$ds^2 = \frac{L^2 d\rho^2}{(\rho^2 - L^2)} - \frac{\rho^2 - L^2}{R^2} d\tau^2 + \rho^2 d\Sigma_2^{d-1} \quad \longrightarrow \quad T = \frac{1}{2\pi R}$$

- apply Wald’s formula (for any gravity theory) for horizon entropy:

$$\begin{aligned} S &= -2\pi \int d^{d-1}x \sqrt{h} \frac{\partial \mathcal{L}}{\partial R^{\mu\nu}_{\rho\sigma}} \hat{\varepsilon}^{\mu\nu} \hat{\varepsilon}_{\rho\sigma} \\ &= \frac{2\pi}{\pi^{d/2}} \Gamma(d/2) a_d^* V(H^{d-1}) \end{aligned}$$

(RCM & Sinha)

where a_d^* contains all of the couplings from the gravity theory

$$\text{eg, } a_d^* = \frac{\pi^{d/2}}{\Gamma(d/2)} \frac{L^{d-1}}{\ell_P^{d-1}} \quad \text{for Einstein gravity}$$

$$S_{EE} = S_{thermal} = S_{horizon}$$

- desired “black hole” is a hyperbolic foliation of empty AdS space

$$ds^2 = \frac{L^2 d\rho^2}{(\rho^2 - L^2)} - \frac{\rho^2 - L^2}{R^2} d\tau^2 + \rho^2 d\Sigma_2^{d-1} \quad \longrightarrow \quad T = \frac{1}{2\pi R}$$

- apply Wald’s formula (for any gravity theory) for horizon entropy:

$$\begin{aligned} S &= -2\pi \int d^{d-1}x \sqrt{h} \frac{\partial \mathcal{L}}{\partial R^{\mu\nu}_{\rho\sigma}} \hat{\varepsilon}^{\mu\nu} \hat{\varepsilon}_{\rho\sigma} \\ &= \frac{2\pi}{\pi^{d/2}} \Gamma(d/2) a_d^* V(H^{d-1}) \end{aligned}$$

(RCM & Sinha)

where a_d^* = central charge for “A-type trace anomaly”

for even d

= entanglement entropy defines effective central charge

for odd d

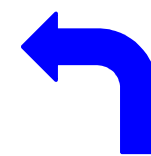
$$S_{EE} = S_{thermal} = S_{horizon}$$

- desired “black hole” is a hyperbolic foliation of empty AdS space

$$ds^2 = \frac{L^2 d\rho^2}{(\rho^2 - L^2)} - \frac{\rho^2 - L^2}{R^2} d\tau^2 + \rho^2 d\Sigma_2^{d-1} \quad \longrightarrow \quad T = \frac{1}{2\pi R}$$

- apply Wald’s formula (for any gravity theory) for horizon entropy:

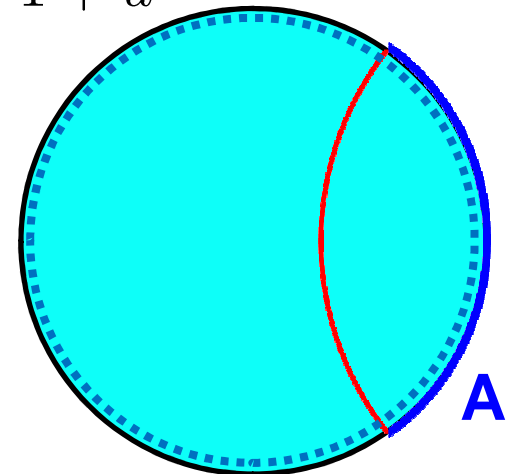
$$S = \frac{2\pi}{\pi^{d/2}} \Gamma(d/2) a_d^* V(H^{d-1})$$



intersection with standard
regulator surface: $z_{min} = \delta$

$$S = a_d^* \frac{4\pi^{\frac{d-3}{2}}}{(d-2)\Gamma(\frac{d-1}{2})} u_{max}^{d-2} + \dots$$

$$ds^2 = \frac{du^2}{1+u^2} + u^2 d\Omega_2^{d-2}$$



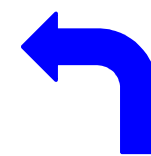
$$S_{EE} = S_{thermal} = S_{horizon}$$

- desired “black hole” is a hyperbolic foliation of empty AdS space

$$ds^2 = \frac{L^2 d\rho^2}{(\rho^2 - L^2)} - \frac{\rho^2 - L^2}{R^2} d\tau^2 + \rho^2 d\Sigma_2^{d-1} \quad \longrightarrow \quad T = \frac{1}{2\pi R}$$

- apply Wald’s formula (for any gravity theory) for horizon entropy:

$$S = \frac{2\pi}{\pi^{d/2}} \Gamma(d/2) a_d^* V(H^{d-1})$$



$$ds^2 = \frac{du^2}{1+u^2} + u^2 d\Omega_2^{d-2}$$

intersection with standard
regulator surface: $z_{min} = \delta$

$$S = a_d^* \frac{4\pi^{\frac{d-3}{2}}}{(d-2)\Gamma(\frac{d-1}{2})} \underbrace{\left(\frac{R}{\delta}\right)^{d-2}}_{\text{“area law”}} + \dots$$

“area law” for d-dimensional CFT

$$S_{EE} = S_{thermal} = S_{horizon}$$

- desired “black hole” is a hyperbolic foliation of empty AdS space

$$ds^2 = \frac{L^2 d\rho^2}{(\rho^2 - L^2)} - \frac{\rho^2 - L^2}{R^2} d\tau^2 + \rho^2 d\Sigma_2^{d-1} \quad \longrightarrow \quad T = \frac{1}{2\pi R}$$

- apply Wald’s formula (for any gravity theory) for horizon entropy:

$$S = \frac{2\pi}{\pi^{d/2}} \Gamma(d/2) a_d^* V(H^{d-1})$$

$$ds^2 = \frac{du^2}{1+u^2} + u^2 d\Omega_2^{d-2}$$

universal contributions:


$$S = \dots + (-)^{\frac{d}{2}-1} 4 a_d^* \log(2R/\delta) + \dots \quad \text{for even } d$$

$$\dots + (-)^{\frac{d-1}{2}} 2\pi a_d^* + \dots \quad \text{for odd } d$$

$$S_{EE} = S_{thermal} = S_{horizon}$$

universal contributions:

$$S = \dots + (-)^{\frac{d}{2}-1} 4 a_d^* \log(2R/\delta) + \dots \quad \text{for even } d$$
$$\dots + (-)^{\frac{d-1}{2}} 2\pi a_d^* + \dots \quad \text{for odd } d$$

- discussion extends to case with background: $R^{1,d-1} \rightarrow R \times S^{d-1}$
- for Einstein gravity, coincides with Ryu & Takayanagi result and horizon (bifurcation surface) coincides with R&T surface
  no extremization procedure?!?
- applies for classical bulk theories beyond Einstein gravity
- can imagine calculating “quantum” corrections (eg, Hawking rad)

Holographic Renyi entropy:

- recall Renyi entropies (close cousin of entanglement entropy)

$$S_n = \frac{1}{1-n} \log \text{Tr} [\rho_A^n] \qquad S_{EE} = \lim_{n \rightarrow 1} S_n$$

- universal contribution (for even d)

$$S_n = \dots + \text{constant} \times \log(R/\delta) + \dots$$

Holographic Renyi entropy:

- recall Renyi entropy (close cousin of entanglement entropy)

$$S_n = \frac{1}{1-n} \log \text{Tr} [\rho_A^n] \qquad S_{EE} = \lim_{n \rightarrow 1} S_n$$

- universal contribution (for even d)

$$d=2: \quad S_n = \frac{c}{6} \left(1 + \frac{1}{n} \right) \log (\ell/\delta) + \dots$$

(Calabrese & Cardy)

- few calculations for $d > 2$

(Metlitski, Fuertes & Sachdev; Hastings, Gonzalez, Kallin & Melko; . . .)

- standard calculation involves **singular** n-fold cover of spacetime
→ problematic for translation to dual AdS gravity



Holographic Renyi entropy:

- turn to Renyi entropy (close cousin of entanglement entropy)

$$S_n = \frac{1}{1-n} \log \text{Tr} [\rho_A^n] \qquad S_{EE} = \lim_{n \rightarrow 1} S_n$$

- recall previous derivation lead to thermal density matrix

$$\rho_A = U^{-1} \frac{e^{-H/T_0}}{\text{Tr} [e^{-H/T_0}]} U \qquad \text{with} \qquad T_0 = \frac{1}{2\pi R}$$

 $\text{Tr} [\rho_A^n] = \frac{\text{Tr} [e^{-nH/T_0}]}{\text{Tr} [e^{-H/T_0}]^n}$  partition function at new temperature, $T = T_0/n$

Holographic Renyi entropy:

- turn to Renyi entropy (close cousin of entanglement entropy)

$$S_n = \frac{1}{1-n} \log \text{Tr} [\rho_A^n] \quad S_{EE} = \lim_{n \rightarrow 1} S_n$$

- with bit more work, find some convenient formulae:

$$S_n = \frac{n}{1-n} \frac{1}{T_0} [F(T_0) - F(T_0/n)]$$

where $F(T) = -T \log Z(T)$ and $T_0 = 1/(2\pi R)$ or

$$S_n = \frac{n}{n-1} \frac{1}{T_0} \int_{T_0/n}^{T_0} S(T) dT$$

↑
Renyi entropy
for spherical Σ

↑
thermal entropy
on hyperbolic space H^{d-1}

- in holographic framework, need to know topological black hole solutions for arbitrary temperature

Holographic Renyi entropy:

- Renyi entropy of CFT for spherical entangling surface:

$$S_n = \frac{n}{n-1} \frac{1}{T_0} \int_{T_0/n}^{T_0} S(T) dT \quad \text{where} \quad T_0 = \frac{1}{2\pi R}$$


- need to know topological black holes for arbitrary temperature
- focus on gravity theories where we can calculate: Einstein, Gauss-Bonnet, Lovelock, quasi-topological,

Holographic Renyi entropy:

- for example, with Einstein gravity and (boundary) $d=4$:

$$S_n = \frac{\pi n}{4(n-1)} \left(5 + \frac{x_n}{n}\right) \left(1 - \frac{x_n}{n}\right) \frac{L^3}{\ell_P^3} V(H^3)$$

where $x_n = \left(1 + \sqrt{1 + 8n^2}\right) / (4n)$

 $(S_n)_{univ} = \frac{c}{2} \frac{n}{1-n} \left(5 + \frac{x_n}{n}\right) \left(1 - \frac{x_n}{n}\right) \log(2R/\delta)$

- compare to $d=2$ result:

$$S_n = \frac{c}{6} \left(1 + \frac{1}{n}\right) \log(\ell/\delta) + \dots$$

- might suggest simple general form for even d :

$$(S_n)_{univ} = c \times f(d, n) \times \log(2R/\delta)$$

Holographic Renyi entropy:

- for example, with GB gravity and (boundary) $d=4$:

$$(S_n)_{univ} = \log(2R/\delta) \frac{n}{2} \frac{1-x^2}{1-n} \left[(5c-a)x^2 - (13c-5a) + 16c \frac{2cx^2 - c + a}{(3c-a)x^2 - c + a} \right]$$

where $0 = x^3 - \frac{3c-a}{5c-a} \left(\frac{x^2}{n} + x \right) + \frac{1}{n} \frac{c-a}{5c-a}$

- unfortunately indicates no simple universal form:

$$(S_n)_{univ} = a \times f \left(d, n, \frac{c}{a}, t_4, \dots \right) \times \log(2R/\delta)$$

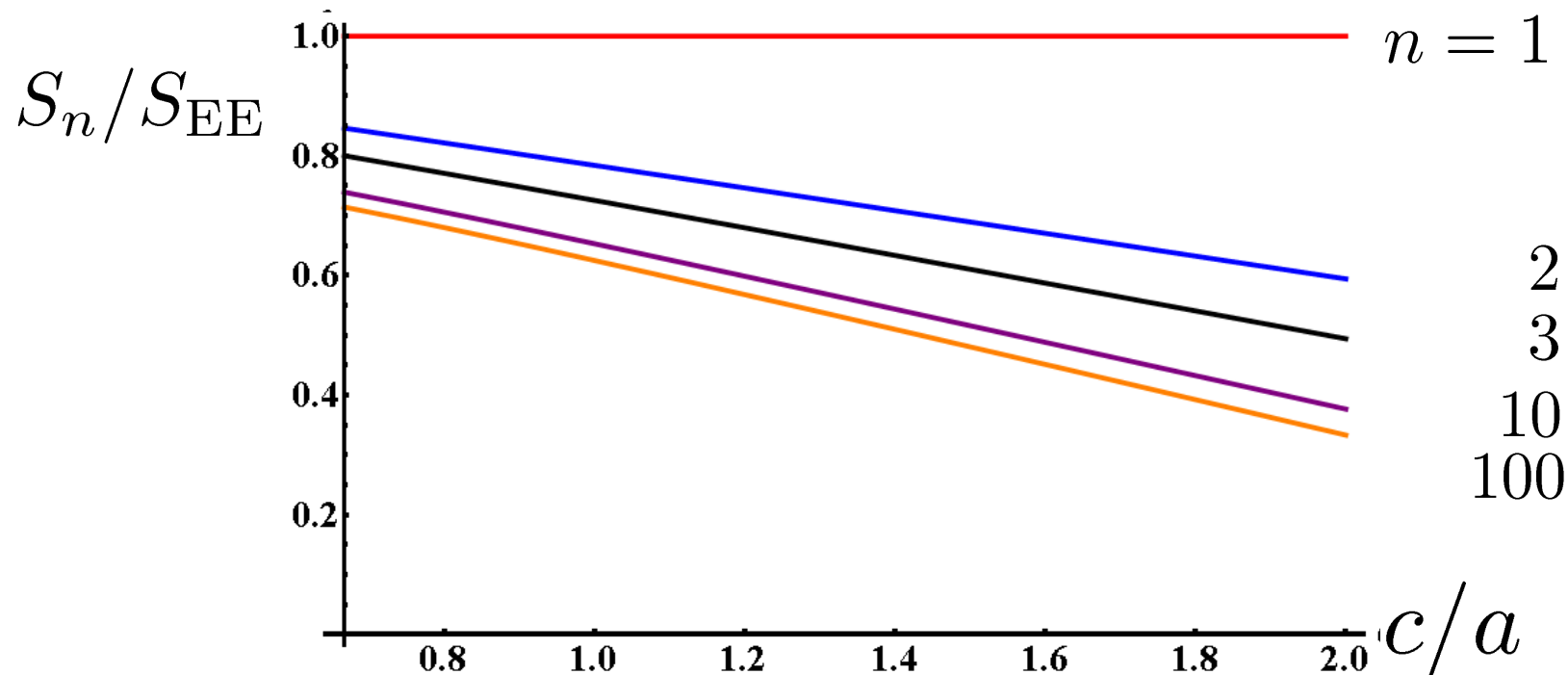
- further work (with quasi-topological gravity) shows the universal coefficient depends on more CFT data than central charges

Holographic Renyi entropy:

- for example, with GB gravity and (boundary) $d=4$:

$$(S_n)_{univ} = \log(2R/\delta) \frac{n}{2} \frac{1-x^2}{1-n} \left[(5c-a)x^2 - (13c-5a) + 16c \frac{2cx^2 - c + a}{(3c-a)x^2 - c + a} \right]$$

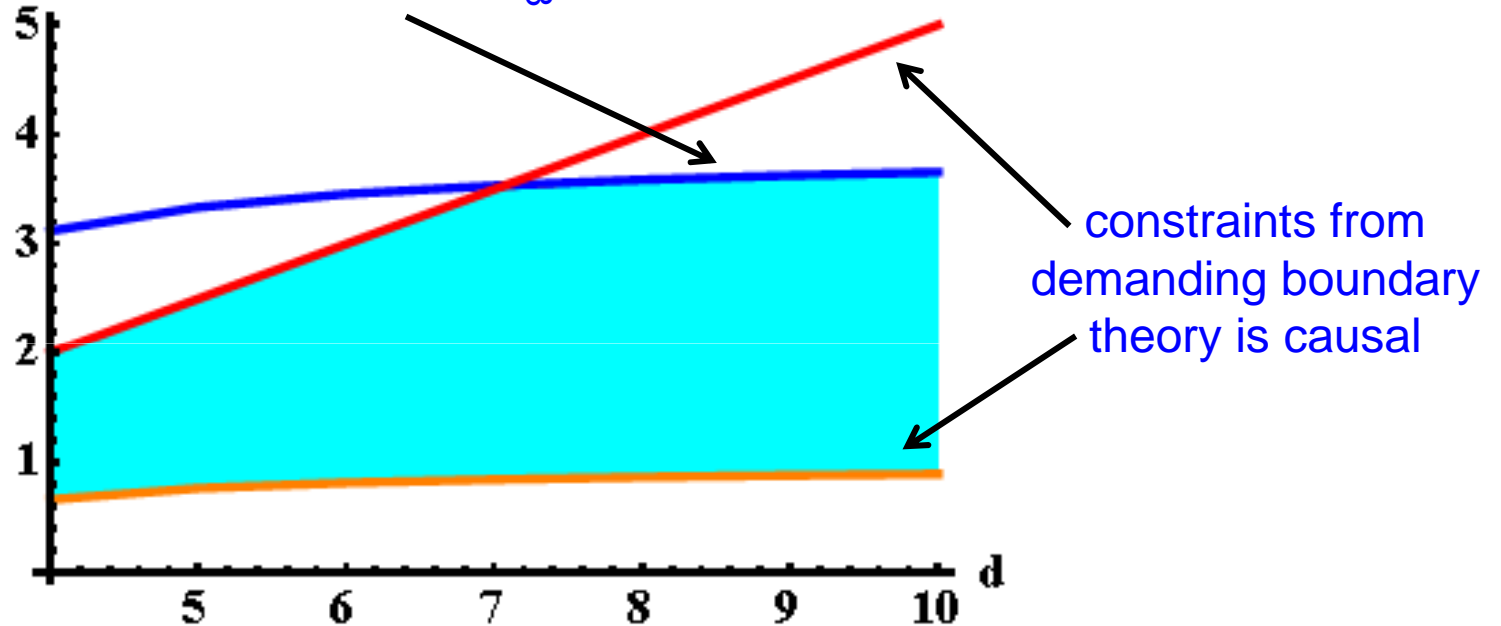
- note despite intimidating expression, results relatively simple:



⚠ Phase transitions? ⚠

- Problem: holographic S_n goes negative for large n (and large d)

GB gravity: $C_T/\alpha^* a$ surface where $S_\infty = 0$



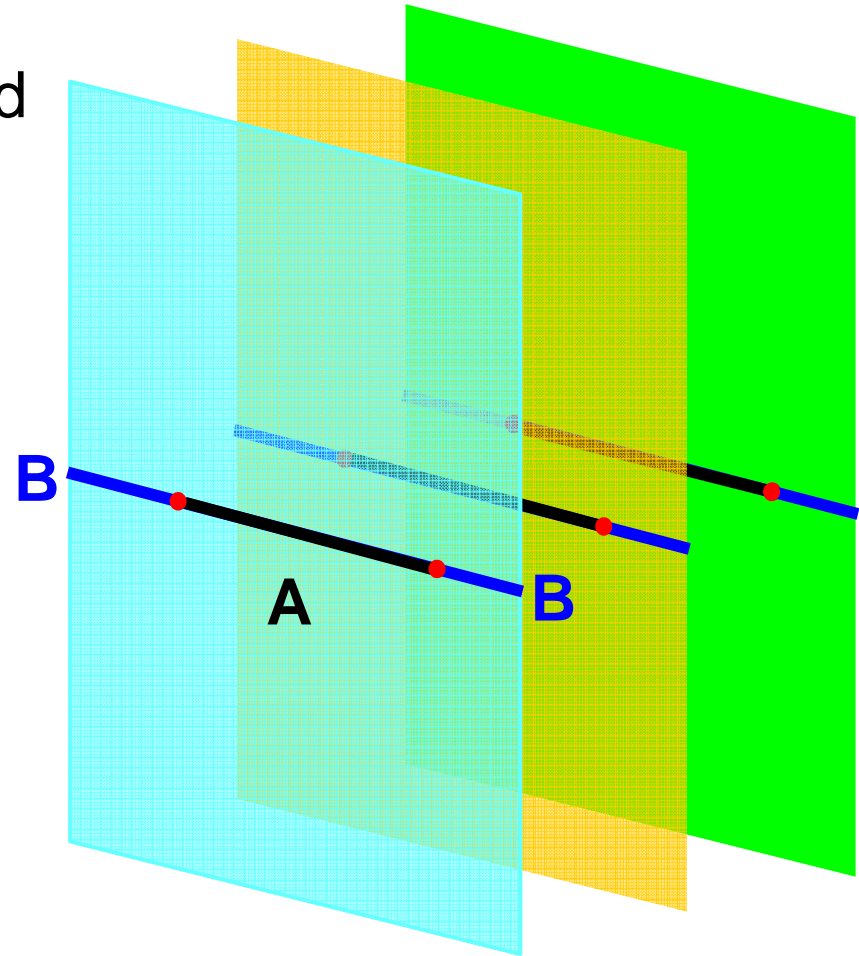
- real problem is that in certain parameter regime, entropy of hyperbolic black hole becomes negative at low temperatures

→ identified the wrong saddle-point

- right saddle-point? new constraint on theory space?

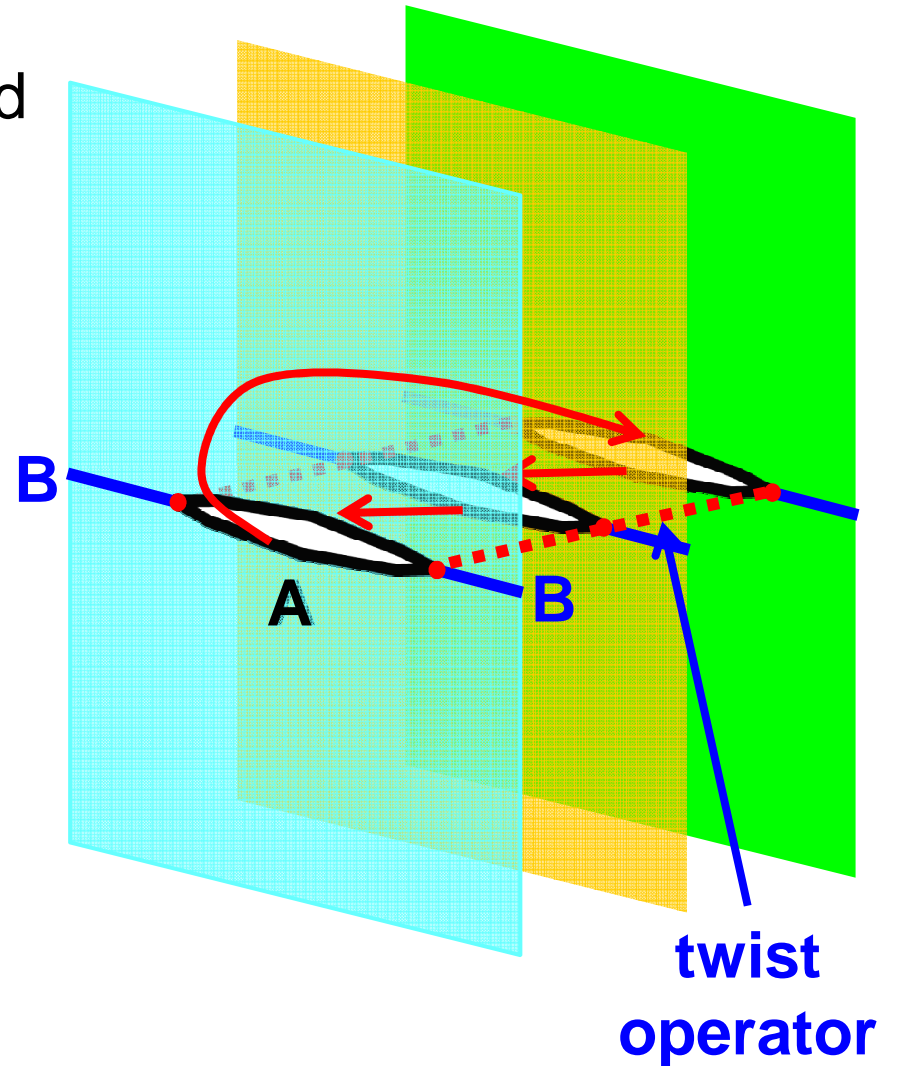
Twist Operators:

- $\text{Tr}(\rho_A^n)$ evaluated as Euclidean path integral over n copies of field theory inserting **twist operators** at boundary of region **A**



Twist Operators:

- $\text{Tr}(\rho_A^n)$ evaluated as Euclidean path integral over n copies of field theory inserting **twist operators** at boundary of region **A**
- twist operators introduce n -fold branch cuts where various copies of fields talk to each other



Twist Operators:

- $\text{Tr}(\rho_A^n)$ evaluated as Euclidean path integral over n copies of field theory inserting **twist operators** at boundary of region **A**
- twist operators introduce n -fold branch cuts where various copies talk to each other
- elegant results for $d=2$, eg, scaling dimension of twist operators

$$h_n = \frac{c}{12} \left(n - \frac{1}{n} \right)$$

(Calabrese & Cardy)

- in d dimensions, would be $(d-2)$ -dimensional surface operators but little is known about their properties

Twist Operators:

- insertion of stress tensor near planar twist operator for CFT in R^d

$$\langle T_{ab} \sigma_n \rangle = -\frac{h_n}{2\pi} \frac{\delta_{ab}}{r_{\perp}^d}, \quad \langle T_{ai} \sigma_n \rangle = 0$$

$$\langle T_{ij} \sigma_n \rangle = \frac{h_n}{2\pi} \frac{(d-1)\delta_{ij} - d n_i n_j}{r_{\perp}^d}$$

where $a, b \parallel \sigma_n$ and $i, j \perp \sigma_n$




- consider conformal mapping for spherical entangling surface
- Euclidean version gives one-to-one map: $S^1 \times H^{d-1} \rightarrow R^d$
- with $\beta = n/T_0 = 2\pi R n$ ($n \in \mathbb{Z}$) get n-fold cover of R^d
- have spherical twist operator σ_n on S^{d-2}

Twist Operators:

- evaluate $\langle T_{\alpha\beta} \sigma_n \rangle$ correlator by mapping from thermal bath

$$\langle T_{\alpha\beta} \sigma_n \rangle = \underbrace{\Omega^{d-2} \frac{\partial X^\mu}{\partial x^\alpha} \frac{\partial X^\nu}{\partial x^\beta}}_{\text{creates singularity near twist operator}} \left(\underbrace{\langle T_{\mu\nu}(T_0/n) \rangle}_{\text{uniform thermal bath}} - \mathcal{A}_{\mu\nu} \right) \langle T_0 \rangle$$

anomalous bit

(compare: Marolf, Rangamani & Van Raamsdonk)

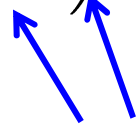
- read off h_n from short distance singularity

$$h_n = 2\pi \frac{n R^d}{d-1} \left(\mathcal{E}(T_0) - \mathcal{E}(T_0/n) \right)$$

Twist Operators:

- evaluate $\langle T_{\alpha\beta} \sigma_n \rangle$ correlator by mapping from thermal bath

$$\langle T_{\alpha\beta} \sigma_n \rangle = \underbrace{\Omega^{d-2} \frac{\partial X^\mu}{\partial x^\alpha} \frac{\partial X^\nu}{\partial x^\beta}}_{\text{creates singularity near twist operator}} \left(\underbrace{\langle T_{\mu\nu}(T_0/n) \rangle}_{\text{uniform thermal bath}} - \mathcal{A}_{\mu\nu} \right) \langle T_0 \rangle$$



 anomalous bit

(compare: Marolf, Rangamani & Van Raamsdonk)

- for example, with GB gravity and (boundary) $d=4$:

$$h_n = \frac{n}{4\pi} (x^2 - 1) [c - a - x^2(5c - a)]$$

where $0 = x^3 - \frac{3c - a}{5c - a} \left(\frac{x^2}{n} + x \right) + \frac{1}{n} \frac{c - a}{5c - a}$

- no simple universal form can be expected
- again, CFT data beyond central charges also appears

Conclusions:

- AdS/CFT correspondence (gauge/gravity duality) has proven an excellent tool to study strongly coupled gauge theories
- holographic entanglement entropy is part of an interesting dialogue has opened between string theorists and physicists in a variety of fields (eg, condensed matter, nuclear physics, . . .)
- potential to learn lessons about issues in boundary theory
eg, readily calculate Renyi entropies for wide class of theories in higher dimensions
- potential to learn lessons about issues in bulk gravity theory
eg, holographic entanglement entropy may give new insight into quantum gravity or emergent spacetime

(eg, van Raamsdonk)

Lots to explore!