Recent Updates of Holographic Entanglement Entropy

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- (1) Introduction
- 2 Holographic Entanglement Entropy (HEE)
- (3) Can HEE Probe Fermi Surfaces?

[Ogawa-Ugajin-TT, in preparation]

4 HEE and Holographic Dual of BCFT

[TT Phys.Rev.Lett 107:101602,2011 [arXiv:1105.5165]

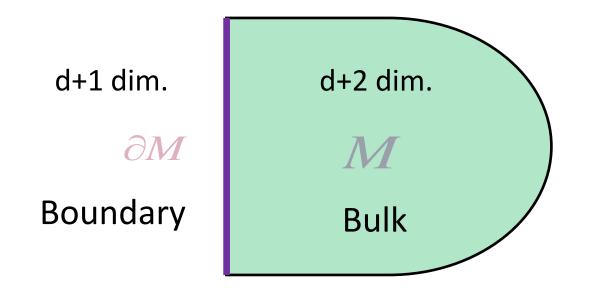
Fujita-Tonni-TT, arXiv:1108.5152, to appear in JHEP]

5 Conclusions

1 Introduction

Holography (e.g. AdS/CFT [Maldacena 97])

⇒ Non-perturbative Definition of Quantum Gravity



$$Z_{QM} (\partial M) = Z_{Gravity} (M)$$

To understand the holography better and to explore the holography in general setups, we need suitable universal physical quantities.

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Stationary BH ⇒ Mass M, Charge Q, Spin J. (Thermodynamics)
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Generic spacetime ⇒ We need much more quantities!

(Non-equilibrium, Topological, etc.)

For this purpose, the entanglement entropy (EE) is a very useful quantity. [⇔ In cond-mat., EE is useful to characterize quantum structures of ground states.]

<u>Definition of Entanglement Entropy</u>

Divide a quantum system into two subsystems A and B.

$$H_{tot} = H_A \otimes H_B$$
.

Example: Spin Chain

A

B

We define the reduced density matrix ρ_A for A by

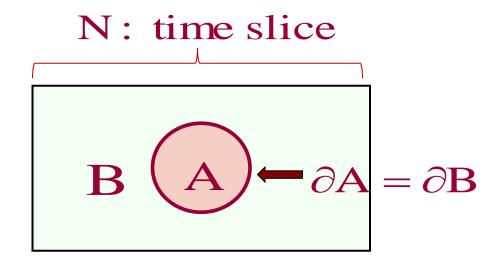
$$\rho_{A} = \operatorname{Tr}_{B} \rho_{tot}$$
,

taking trace over the Hilbert space of B.

Now the entanglement entropy S_A is defined by the von-Neumann entropy

$$S_A = -\operatorname{Tr}_A \rho_A \log \rho_A$$
.

In QFTs, it is defined geometrically:



Various Applications in other subjects

Quantum Information and Quantum Computing

EE = the amount of quantum information

[see e.g. Nielsen-Chuang's text book 00]

Condensed Matter Physics

EE = Efficiency of a computer simulation (DMRG) [Gaite 03,...]

Divergent at quantum critical points!

[G. Vidal, J. I. Latorre, E. Rico, and A. Kitaev, 02,...]

A new quantum order parameter!

[Topological entanglement entropy: Kitaev-Preskill 06, Levin-Wen 06]

Basic property: Area law

EE in d+1 dim. QFTs (in ground states) includes UV divergence

[Bombelli-Koul-Lee-Sorkin 86, Srednicki 93]

$$S_A \sim \frac{\text{Area}(\partial A)}{\varepsilon^{d-1}} + \text{(subleading terms)},$$

where ε is a UV cutoff (i.e. lattice spacing).

A comment: The holographic EE predicts that the area law is always true for any QFT if there is a UV fixed point.

Similar to the Bekenstein-Hawking formula of black hole entropy

$$S_{BH} = \frac{\text{Area(horiz} \quad \text{on)}}{4G_{N}}$$
. [EE = loop corrections to BH entropy, Susskind-Uglum 94,...]

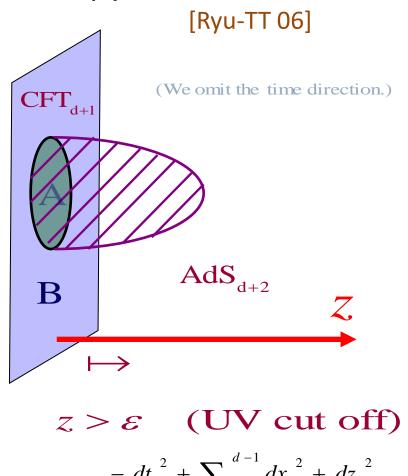
2 Holographic Entanglement Entropy

(2-1) Holographic Entanglement Entropy Formula

$$S_{A} = \frac{Area(\gamma_{A})}{4G_{N}}$$

\(\cert_A\) is the minimal area surface(codim.=2) such that

$$\partial A = \partial \gamma_A$$
 and $A \sim \gamma_A$. homologous



$$ds_{AdS}^{2} = R_{AdS}^{2} = \frac{-dt^{2} + \sum_{i=1}^{d-1} dx_{i}^{2} + dz^{2}}{z^{2}}$$

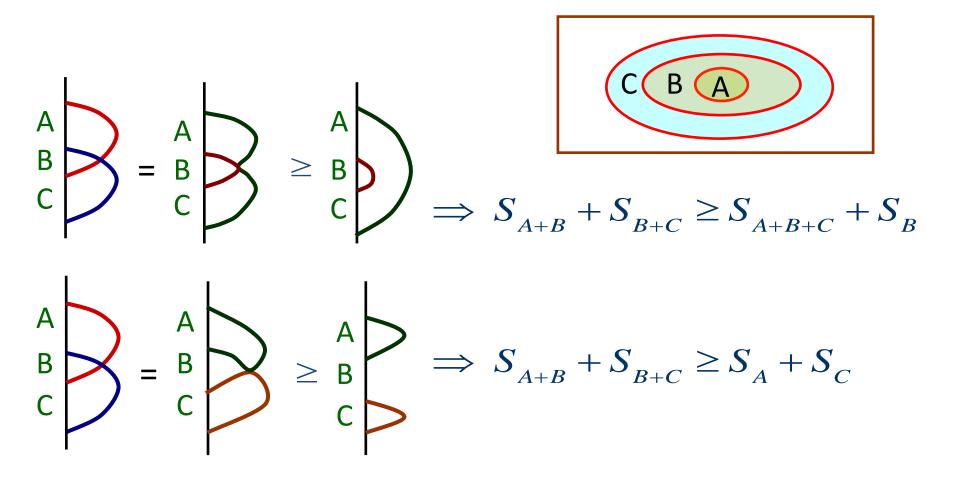
 In spite of a heuristic argument [Fursaev, 06], there has been no complete proof. But, so many evidences and no counter examples.

[A Partial List of Evidences]

- Area law follows straightforwardly [Ryu-TT 06]
- Agreements with analytical 2d CFT results for AdS3 [Ryu-TT 06]
- Holographic proof of strong subadditivity [Headrick-TT 07]
- Consistency of 2d CFT results for disconnected subsystems
 [Calabrese-Cardy-Tonni 09] with our holographic formula [Headrick 10]
- Agreement on the coefficient of log term in 4d CFT (~a+c) [Ryu-TT 06, Solodukhin 08,10, Lohmayer-Neuberger-Schwimmer-Theisen 09, Dowker 10, Casini-Huerta, 10, Myers-Sinha 10]
- > A direct proof when A = round ball [Casini-Hueta-Myers 11]
- ➤ Holographic proof of Cadney-Linden-Winter inequality
 [Hayden-Headrick-Maloney 11]

Holographic Proof of Strong Subadditivity [Headrick-TT 07]

We can easily derive the *strong subadditivity*, which is known as the most important inequality satisfied by EE. [Lieb-Ruskai 73]



Tripartite Information [Hayden-Headrick-Maloney 11]

Recently, the holographic entanglement entropy is shown to have a special property called *monogamy*.

$$S_{AB} + S_{BC} + S_{AC} \ge S_A + S_B + S_C + S_{ABC}$$
$$\Leftrightarrow I(A:B) + I(A:C) \le I(A:BC)$$

Comments:

- (i) HEE argues that this is true for large N gauge theories.
- (ii) This property is not always true for QFTs.
- (iii) In 2+1 dim. mass gapped theories, this argues that the topological entanglement entropy is non-negative.
- (iv) This property is also confirmed in time-dependent examples.

 [Balasubramanian-Bernamonti-Copland-Craps-Galli 11, Allais-Tonni 11]

- The area formula of HEE assumes the supergravity approximation (i.e. strongly coupled limit and large N limit).
- ⇒ The holographic formula is modified by higher derivatives.
 (deviations from strongly coupled limit, but still large N)
- ⇒ A precise formula was found for Lovelock gravities.

[Hung-Myers-Smolkin 11, de Boer-Kulaxizi-Parnachev 11]

Ex. Gauss-Bonnet Gravity

$$S_{GBG} = -\frac{1}{16 G_{N}} \int dx^{d+2} \sqrt{g} \left[R - 2 \Lambda + \lambda R_{AdS}^{2} L_{GB} \right]$$

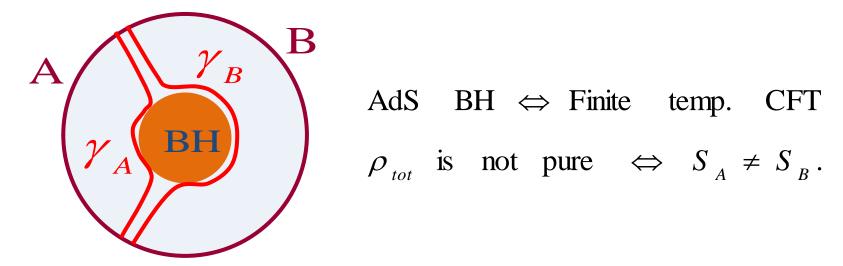
$$L_{GB} \equiv R_{\mu\nu\rho\sigma} R^{\mu\nu\rho\sigma} - 4 R_{\mu\nu} R^{\mu\nu} + R^{2}.$$

$$S_{A} = \text{Min}_{\gamma_{A}} \left[\frac{1}{4 G_{N}} \int_{\gamma_{A}} dx^{d} \sqrt{h} \left(1 + 2 \lambda R_{AdS}^{2} R \right) \right].$$

[But for general higher derivative theories, this is hard!]

Comments

- In the presence of a black hole horizon, the minimal surfaces typically wrap the horizon.
 - ⇒ Reduced to the Bekenstein-Hawking entropy, consistently.



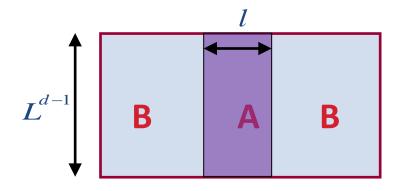
 We need to replace minimal surfaces with extremal surfaces in the time-dependent spacetime. [Hubeny-Rangamani-TT 07]

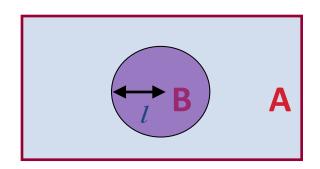
(2-2) Explicit Calculations of HEE

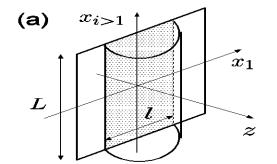
Two analytical examples of the subsystem A:

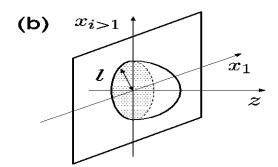
(a) Infinite strip

(b) Circular disk









Entanglement Entropy for (a) Infinite Strip from AdS

$$S_{A} = \frac{R^{d}}{2(d-1)G_{N}^{(d+2)}} \left[\left(\frac{L}{\varepsilon} \right)^{d-1} - C \cdot \left(\frac{L}{l} \right)^{d-1} \right],$$
where $C = 2^{d-1} \pi^{d/2} \left(\Gamma \left(\frac{d+1}{2d} \right) / \Gamma \left(\frac{1}{2d} \right) \right)^{d}.$

Area law divergence

This term is finite and does not depend on the UV cutoff.

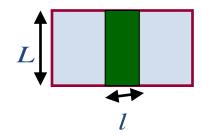
d=1 (i.e. AdS3) case:

$$S_A = \frac{R}{2G_N^{(3)}} \log \frac{l}{\varepsilon} = \frac{c}{3} \log \frac{l}{\varepsilon}.$$

Agrees with 2d CFT results [Holzhey-Larsen-Wilczek 94; Calabrese-Cardy 04]

Basic Example of AdS5/CFT4

$$AdS_5 \times S^5 \Leftrightarrow N = 4 SU(N) SYM$$



CFT:
$$S_A^{freeCFT} = K \cdot \frac{N^2 L^2}{\varepsilon^2} - 0.087 \cdot \frac{N^2 L^2}{l^2}$$
.

Gravity:
$$S_A^{AB} = K' \cdot \frac{N^2 L^2}{\varepsilon^2} - 0.051 \cdot \frac{N^2 L^2}{l^2}$$
.

The order one deviation is expected since the AdS result corresponds to the strongly coupled Yang-Mills.

[cf. 4/3 in thermal entropy, Gubser-Klebanov-Peet 96]

Entanglement Entropy for (b) Circular Disk from AdS

[Ryu-TT 06]

$$S_{A} = \frac{\pi^{d/2}R^{d}}{2G_{N}^{(d+2)}\Gamma(d/2)} \left[p_{l} \left(\frac{l}{\varepsilon} \right)^{d-1} + p_{3} \left(\frac{l}{\varepsilon} \right)^{d-3} + \cdots \right]$$

$$P_{d-1} \left(\frac{l}{\varepsilon} \right) + p_{d} \quad \text{(if } d = \text{even)}$$

$$p_{d-2} \left(\frac{l}{\varepsilon} \right)^{2} + q \log \left(\frac{l}{\varepsilon} \right) \quad \text{(if } d = \text{odd)}$$
where $p_{1} = (d-1)^{-1}, p_{3} = -(d-2)/[2(d-3)], \dots$

$$\dots = q = (-1)^{(d-1)/2} (d-2)!!/(d-1)!!$$

A universal quantity in odd dimensional CFT?

(Note: in mass gapped theories this is the topological EE.)

Conformal Anomaly (~central charge)

2d CFT $c/3 \cdot \log(1/\epsilon)$

4d CFT -4a • log(l/ε)

Comments: EE in Odd dim. CFT (e.g. 2+1 dim.)

Recently, [Casini-Hueta-Myers 11] proved for any odd dim. CFTs:

$$S_A = \log Z(S^{d+1})$$

if the subsystem A = d dim. round ball.

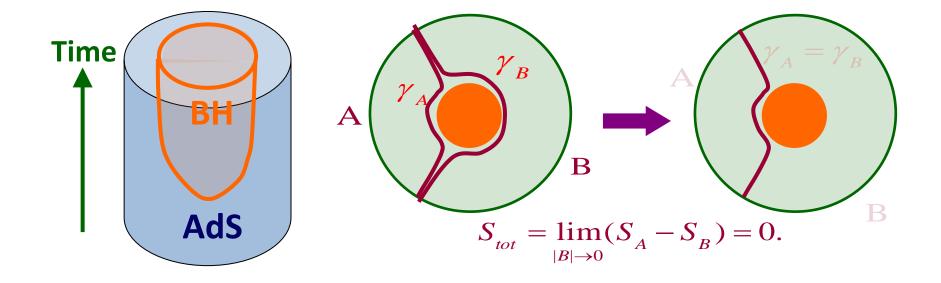
Also, in this setup, [Myers-Sinha 10] proved that the holography tells us that the finite part of S_A monotonically decreases under the RG flow.

⇒ A `c'-theorem for odd dim. QFTs
EE = `c-function'

(2-3) HEE and Thermalization

BH formation \Leftrightarrow Thermaliza tion [e.g. Chesler-Yaffe 08', ρ_{tot} is pure (unitary evolution) Bhattacharyya-Minwalla 09', ...] $\Rightarrow S_{tot} = 0, \text{ even if } \exists \text{ dynamical horizon.}$ [Arrastia-Aparicio-Lopez 10, Ugajin-TT 10]

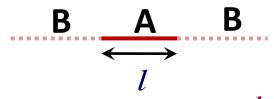
On the other hand, $S_A^{finite} \propto \text{Size of BH} \neq 0$, $\rightarrow \text{EE} = \text{Coarse} - \text{grained} = \text{entropy}$



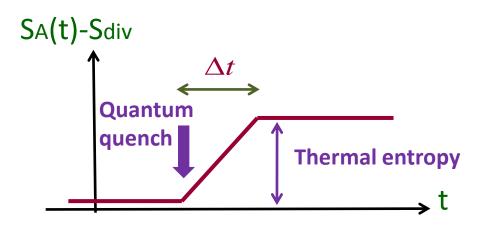
Time Evolutions of HEE under Quantum Quenches

In 1+1 dim. CFTs, we expect a linear growth of EE after a quantum quench.

[Calabrese-Cardy 05]



Causality
$$\rightarrow \Delta t \approx \frac{l}{2}$$



HEE reproduced the same result.

[Arrastia-Aparicio-Lopez 10]

In higher dim., the thermalization time depends on the shape of A.

HEE predicts: A = strip
$$\rightarrow \Delta t > \frac{l}{2}$$
, A = round disk $\rightarrow \Delta t \approx \frac{l}{2}$

[Albash-Johnson 10, Balasubramanian-Bernamonti-de Boer-Copland-Craps- Keski-Vakkuri-Müller-Schäfer-Shigemori-Staessens 10, 11]

3 Can HEE probe Fermi Surfaces?

[Ogawa-Ugajin-TT, in preparation]

(3-1) Logarithmic Violation of Area Law

In d dim. lattice models that the area law of EE is violated logarithmically in free fermion theories. [Wolf 05, Gioev-Klich 05]

$$S_A \sim L^{d-1} \log L$$
, $(L = size \ of \ A)$.

Comments:

- (i) This property can be understood from the logarithmic EE in 2D CFT, which approximates the radial excitations of fermi surface.
- (ii) It is natural to expect that this property is true for non-Fermi liquids. [Swingle 09,10, Zhang-Grover-Vishwanath 11 etc.]

Note in this lattice calculation assumes

$$\varepsilon^{-1}$$
 (UV cut off) ~ k_F .

Instead, in our holographic context which corresponds to a continuous limit, we are interested in the case $\ensuremath{arepsilon}^{-1}>>k_{_F}$.

In this case, we expect

$$S_A = (div.) + \eta \cdot (L \cdot k_F)^d \log Lk_F + \cdots$$

Below we would like to see if we can realize this behavior in HEE. We assume that all physical quantities can be calculable in the classical gravity limit ($\Leftrightarrow_{\exists O(N^2)}$ Fermi surfaces).

(3-2) Holographic Construction

The metric ansatz:
$$ds^2 = \frac{R_{AdS}^2}{z^2} \left(-f(z)dt^2 + g(z)dz^2 + dx^2 + dy^2 \right)$$
.

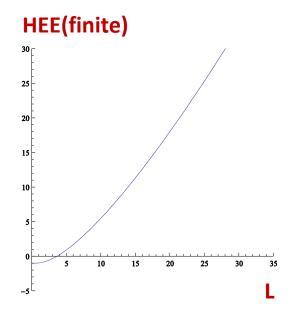
Asymp. AdS
$$\Rightarrow f(0) = g(0) = 1$$
.

(Below we work d=2 i.e. AdS4/CFT3 setup.)

The logarithmical behavior of EE occurs iff

$$g(z) \rightarrow \left(\frac{z}{z_F}\right)^2 \qquad (z \rightarrow \infty).$$

Note: f(z) does not affect the HEE. z_F^{-1} is dual to the fermi energy.



(3-3) Null Energy Condition

To have a sensible holographic dual, a necessary condition is known as the null energy condition:

$$T_{\mu\nu}N^{\mu}N^{\nu} \ge 0$$
 for any null vector N^{μ} .

In the IR region, the null energy condition argues

$$g(z) \propto z^2$$
, $f(z) \propto z^{-2m} \implies m \ge 2$.

The specific heat behaves like

$$C \propto T^{\alpha}$$
 with $\alpha \leq \frac{2}{3}$.

Notice that this excludes standard Landau fermi liquids.

In summary, we find that *classical gravity duals only allow non-fermi liquids*.

Comments:

- (i) Our definition of classical gravity duals is so restrictive that it does not include either the emergent AdS2 geometry [Faulkner-Liu-McGreevy-Vegh 09, Cubrovic-Zaanen-Schalm 09] nor the electron stars (or Lifshitz) [Hartnoll-Polchinski-Silverstein-Tong 09, Hartnoll-Tavanfar 10].
- (ii) More generally, the background with $g(z) \propto z^{2n}$ leads to

$$S_A^{finite} \propto L^{rac{2n}{n+1}}.$$

(iii) We can embed this background in an effective gravity theory:

$$S_{EMS} = \frac{1}{16 G_N} \int dx^{d+2} \sqrt{-g} [R - 2\Lambda - W(\phi) F_{\mu\nu} F^{\mu\nu} - \partial_{\mu} \phi \partial^{\mu} \phi - V(\phi)].$$

if W and V behave in the large φ limit as follows

$$V(\phi) + 2\Lambda \approx -\frac{(p^{2} + 12 p + 32)}{4R_{AdS}^{2}} \cdot e^{-\sqrt{\frac{2}{p-2}}\phi},$$

$$W(\phi) \approx \frac{8A^{2}}{z_{F}^{2} p(8+p)R^{2}} e^{3\sqrt{\frac{2}{(p-2)}}\phi},$$

$$\Rightarrow f(z) \propto z^{-p}, \quad g(z) \propto z^{2}, \quad (p > 2).$$

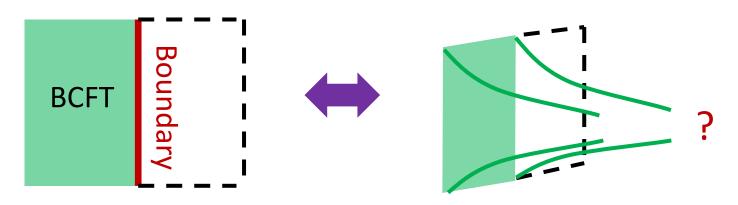
4 HEE and Holographic Dual of BCFT

(4-1) AdS/BCFT

What is a holographic dual of CFT on a manifold with Boundary (BCFT)?

CFTd: $SO(d,2) \Leftrightarrow AdS_{d+1}$

BCFTd: $SO(d-1,2) \Leftrightarrow AdSd$



[Earlier studies: Karch-Randall 00 (BCFT,DCFT),...

Bak-Gutperle-Hirano 03, Clark-Freedman-Karch-Schnabl 04 (Janus CFT)]

AdS/BCFT Proposal [Fujita-Tonni-TT 11]

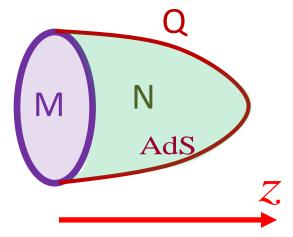
In addition to the standard AdS boundary M, we include an extra boundary Q, such that $\partial Q = \partial M$.

$$I_{E} = -\frac{1}{16\pi G_{N}}\int_{N}\sqrt{g}\left(R - 2\Lambda - L_{matter}\right) - \frac{1}{8\pi G_{N}}\int_{Q}\sqrt{h}\left(K - L_{matter}^{Q}\right).$$

EOM at boundary leads to the Neumann b.c. on Q:

$$K_{ab} - Kh_{ab} = 8\pi G_N T_{ab}^Q .$$

Conformal inv.
$$\Rightarrow T_{ab}^{Q} = -Th_{ab}$$
.



(4-2) Simplest Example

Consider the AdS slice metric:

$$ds_{AdS(d+1)}^{2} = d\rho^{2} + \cosh^{2}(\rho/R)ds_{AdS(d)}^{2}$$
.

Restricting the values of ρ to $-\infty < \rho < \rho_*$ solves the boundary condition with $T = \frac{d-1}{\tan h} \frac{\rho_*}{-} \ .$

(4-3) Holographic Boundary Entropy

The boundary entropy [Affleck-Ludwig 91]

Sbdy measures the degrees of freedom at the boundary.

The g-theorem:

Sbdy monotonically decreases under the RG flow in CFT.

[proved by Friedan -Konechny 04]

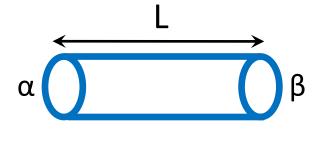
<u>Definition 1</u> (Disk Amplitude)

It is simply defined from the disk amplitude

$$S_{bdy(\alpha)} = \log g_{\alpha}, \qquad g_{\alpha} \equiv \langle 0 | B_{\alpha} \rangle.$$

<u>Definition 2</u> (Cylinder Amplitude)

$$Z_{(\alpha,\beta)}^{cylinder} = \left\langle B_{\alpha} \left| e^{-HL} \right| B_{\beta} \right\rangle \underset{L \to \infty}{\underset{\mathcal{L} \to \infty}{\underbrace{g_{\alpha}g_{\beta}}}} \underbrace{e^{-E_{0}L}}_{\text{Boundary Bulk Part}}.$$

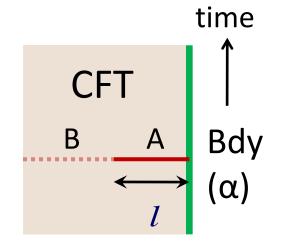


<u>Definition 3</u> (Entanglement Entropy)

$$S_{A} = -\text{Tr}[\rho_{A} \log \rho_{A}],$$

$$\rho_{A} = \text{Tr}_{B} \rho_{tot}.$$

In 2D BCFT, the EE generally behaves like



$$S_A = \frac{c}{-\log \frac{l}{-} + \log g_{\alpha}} .$$

$$\frac{6}{6} \underbrace{\varepsilon}_{\text{Boundary}} .$$

$$\frac{1}{6} \underbrace{\log \frac{l}{-} + \log g_{\alpha}}_{\text{Boundary}} .$$

[Calabrese-Cardy 2004]

In our setup, HEE can be found as follows

$$S_A = \frac{\text{Length}}{4G_N} = \frac{1}{4G_N} \int_{-\infty}^{\rho_*} d\rho = \frac{c}{6} \log \frac{l}{\varepsilon} + \frac{\rho_*}{4G_N} .$$
Boundary Entropy

[Earlier calculations: Azeyanagi-Karch-Thompson-TT 07 (Non-SUSY Janus), Chiodaroli-Gutperle-Hung, 10 (SUSY Janus)]

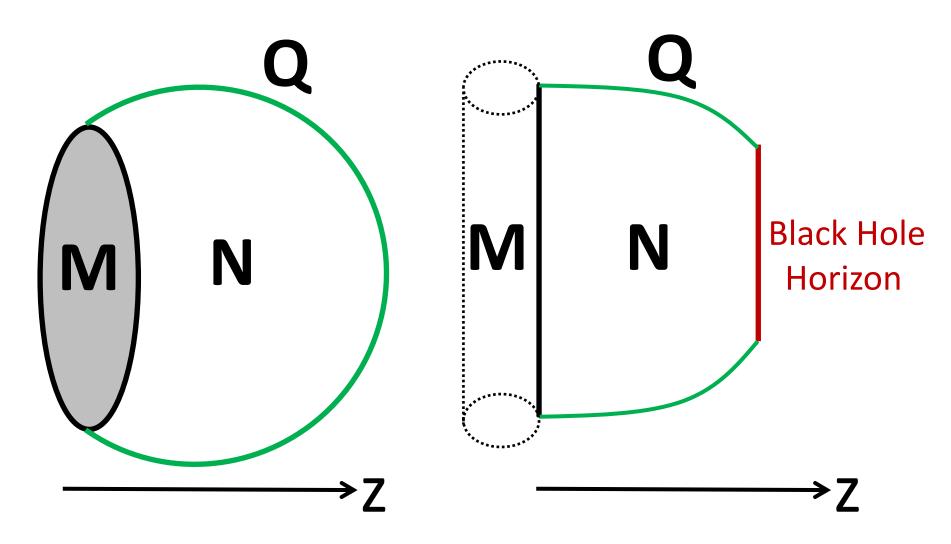
Also $S_{bdy} = \rho_* / 4G_N$ can be confirmed in other two definitions.

$$I_{Disk} = \frac{R}{4G_N} \left(\frac{r^2}{2\varepsilon^2} + \frac{r \sinh(\rho_*/R)}{\varepsilon} + \log \frac{\varepsilon}{r} - \frac{\rho_*}{R} - \frac{1}{2} \right).$$

$$I_{Cylinder} = \frac{\pi}{3} c \cdot l \cdot T_{BH} + \frac{\rho_*}{2G_N}.$$

Holographic Dual of Disk

Holographic Dual of Cylinder



(4-4) Holographic g-Theorem

Consider the surface Q defined by x = x(z) in the Poincare metric

$$ds^{2} = R^{2} \left(\frac{dz^{2} - dt^{2} + dx^{2} + (d\vec{w})^{2}}{z^{2}} \right).$$

We impose the null energy condition for the boundary matter i.e. $T_{ab}^{Q} N^{a} N^{b} \ge 0$ for any null vector N^{a} .

[cf. Hol. C-theorem: Freedman-Gubser-Pilch-Warner 1999, Myers-Sinha 2010]

For the null vector, $N^t=1$, $N^z=1/\sqrt{1+(x')^2}$, $N^x=x'/\sqrt{1+(x')^2}$, we find the constraint

$$(K_{ab} - Kh_{ab})N^{a}N^{b} = -\frac{R \cdot x''}{z(1+(x')^{2})^{3/2}} \ge 0.$$

Thus we simply get $x''(z) \leq 0$ from the null energy condition. Define the holographic g-function:

$$\log g(z) = \frac{R^{d-1}}{4G_N} \cdot \operatorname{Arcsinh}\left(\frac{x(z)}{z}\right) = \frac{R^{d-2}}{4G_N} \cdot \rho_*(z).$$

Then we find
$$\frac{\partial \log g(z)}{\partial z} = \frac{x'(z)z - x(z)}{\sqrt{z^2 + x(z)^2}} \le 0 ,$$

because $(x'z-x)'=x''z\leq 0$.

For d=2, at fixed points $\log g(z)$ agrees with the boundary entropy. For any d, $\rho_*(z)$ is a monotonically decreasing function w.r.t. z.



This is our holographic g-theorem!

Example: AdS4/BCFT3

In this case, we obtain

$$I_{E} = \frac{R^{2}}{2G_{N}} \left[\frac{\pi}{2} + \arctan\left(\sinh\frac{\rho_{*}}{R}\right) - \frac{1}{24}\sinh\frac{3\rho_{*}}{R} - \left(\sinh\frac{\rho_{*}}{R}\right) \log r_{B} + \left(\log\cosh\frac{\rho_{*}}{R} - \frac{33}{24} - \log 2\right) \sinh\frac{\rho_{*}}{R} \right].$$

Conformal anomaly in odd dim. CFT?



This should come from the 2 dim. boundary!

Boundary central charge

As the usual central charge in 2 dim. CFT, we can define a boundary central charge in BCFT3 as follows:

$$r_{B} \frac{\partial \log Z_{Ball}}{\partial r_{B}} = -\frac{1}{2\pi} \left\langle \int_{\Sigma} dx^{2} \sqrt{g_{b}} T_{\mu}^{\mu} \right\rangle = \frac{c_{bdy}}{6} \chi(\Sigma).$$

In our holographic calculation, we obtain

$$c_{bdy} = \frac{3R^2}{2G_N} \sinh \frac{\rho_*}{R}.$$

Our holographic g-theorem leads to a c-theorem for c_{pdy} .



Our conjecture: this is true for all BCFT3.

5 Conclusions

 The entanglement entropy (EE) is a useful bridge between gravity (string theory) and cond-mat physics.



- Classical gravity duals ⇒ Non-fermi liquids
- AdS/BCFT can be useful to probe the boundary physics.
 The boundary central charge etc.

Future Problems

- Proof of HEE ?
- Complete Higher derivative corrections to HEE ?
- 1/N corrections to HEE ?
- More on HEE and Fermi Liquids ?
- HEE for non-AdS spacetimes ?
- What is an analogue of the Einstein eq. for HEE?
- •
- .
- Superconductors in AdS/BCFT?
- Topological Insulators and AdS/BCFT?
- •