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β -testing Cosmological Holography

Based on works with Alishahika, Dong,
Horn, Karch, Matsuura, Polchinski, Tong, Tomaba
(2004 - present)

Motivations

- Accelerated expansion + exit in early \mathcal{U} (inflation: Structure seeded by quantum fluctuations) and late \mathcal{U} — now precision science, limited by cosmic variance
- Gibbons/Hawking + Covariant entropy bound suggest a holographic framework for dS + exit FRW
 - Goal: precision holographic cosmology, limited by entropy bounds

Annis '12 review

cf Bousso
Harlow
Suskind

Plan

- Concrete Uplifts of AdS/CFT

to cosmology

- Alishahia, Karch, ES, Tong '04
- Dong Horn ES Torroba '05
- "...", Matsuyama '10
- '11

- Role of basic ingredients
(β -testing):

- Flavor content + unitarity

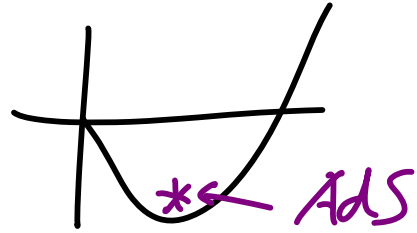
Dong Horn ES Torroba '12

- UV structure of dS/dS:

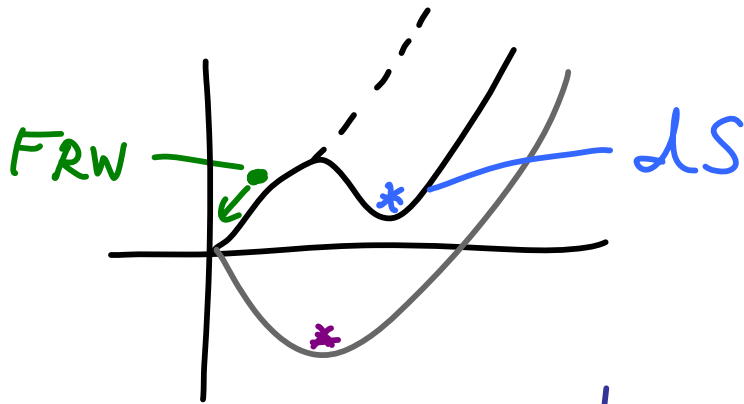
Wilsonian action & (meta-) stability
of the moduli

(In progress, a bits & branes production)

Start by asking what happens
to AdS/CFT



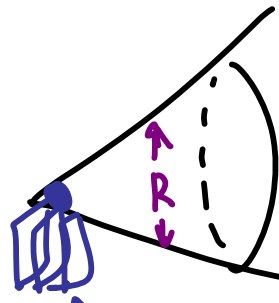
when we uplift to cosmology



in concrete examples.

- p-brane, D-brane technology \rightarrow BH entropy, AdS/CFT.
- structure of cosmo solutions in string theory provides useful guidance as well, as we'll see.

AdS/CFT brane construction



$$\left(\frac{dR}{dr}\right)^2 = \frac{1}{R^2} \rightarrow R=r \text{ (cone)}$$

$R > 0$ Einstein space

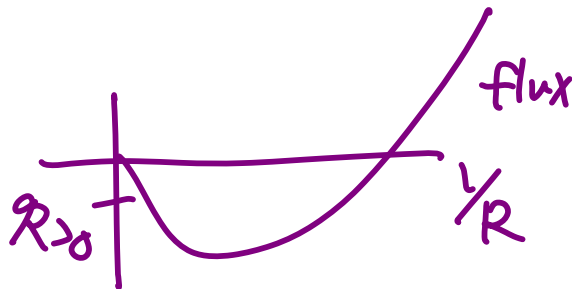
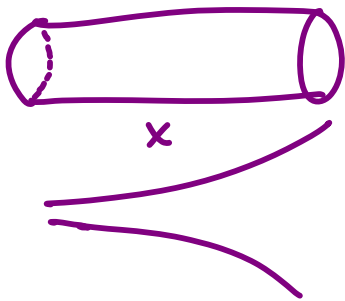
color branes

→ Gravitational Redshift $g_{00} \sim 1 - \frac{R^4}{r^4}$

⇒ Low-energy region

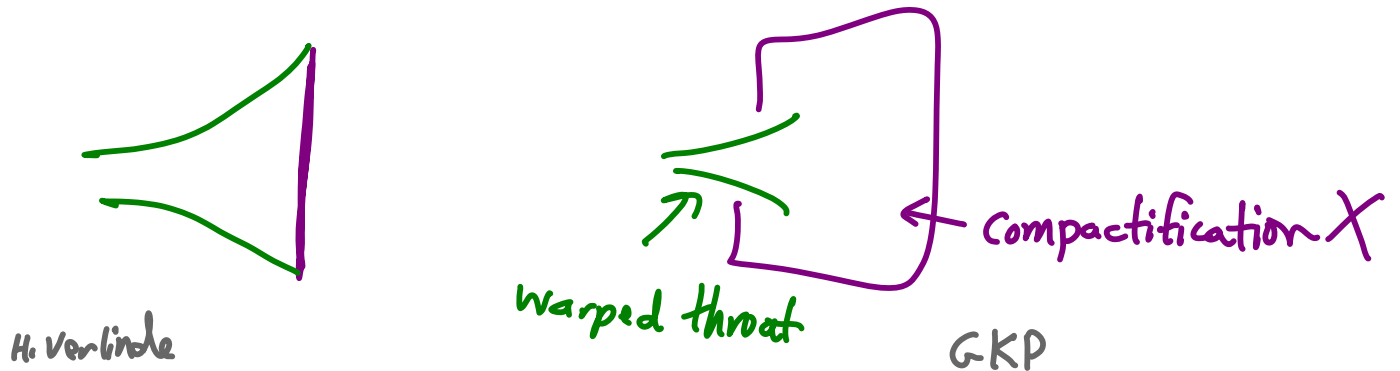
$$E = \sqrt{-g_{00}} E_{pr} \ll E_{pr}$$

→ Effective field theory (EFT) dual, complete QFT
in strict near-horizon limit



RS/warped compactifications

$$ds^2 \simeq \frac{r^2}{R^2} dx^2 + \frac{R^2}{r^2} dr^2 + \text{internal}$$



$$\simeq \underbrace{\text{CFT}}_{D-1} + \underbrace{\text{GR}_{D-1}}_{M_p^2} + \dots$$

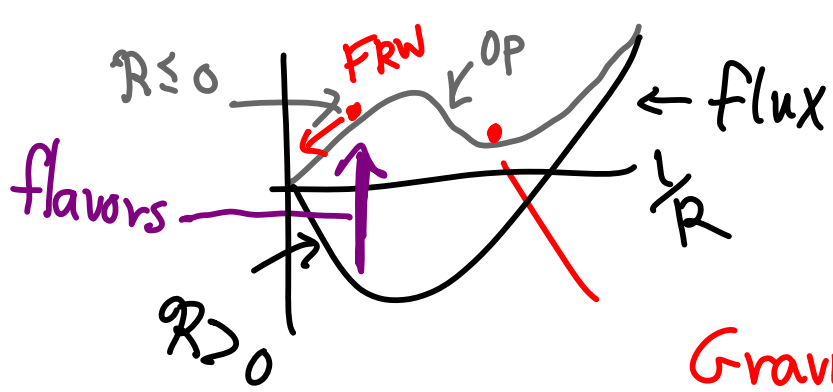
$$E < \Lambda_c = \frac{r_{uv}}{R^2} \left| M_p^2 \simeq \frac{r_{uv}^2}{R^4} N^2 + \frac{\text{Vol}(X)}{g_s^2} \right.$$

→ Complete QFT in limit

$$r_{uv} \rightarrow \infty \quad \text{with} \quad \frac{r_{uv}^2}{R^4} N^2 \gg \frac{\text{Vol}(X)}{g_s^2}$$

Turns out to be a good analogy for
 $ds \rightarrow FRW$

Uplifting AdS/CFT: Brane Constructions



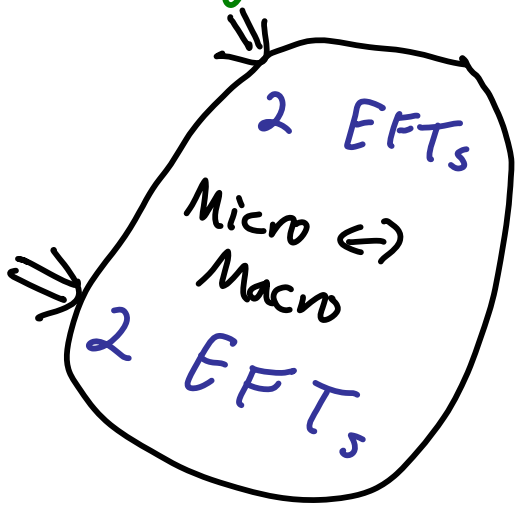
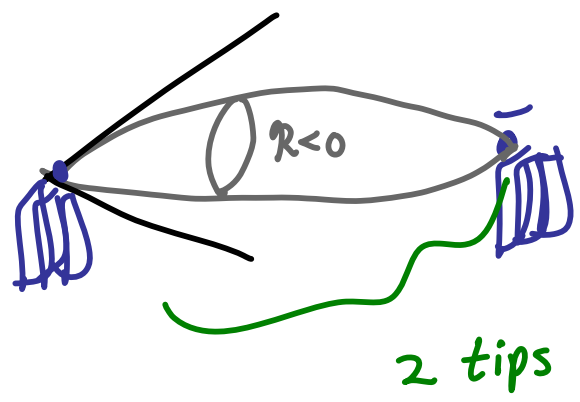
(+ other ingredients to fully meta-stabilize)

Gravity Solution:

$$ds^2 = \sin^2 \frac{w}{L} ds_{D-1}^2 + dw^2$$

2 redshifted regions

Brane Construction:



$$\left(\frac{dR}{dr}\right)^2 = \frac{-1}{R^2} + \frac{\text{const}}{R^{n>2}}$$

- $G R_{D-1}$ at finite times, but large matter sector dual to large throat (e.g. Liouville)

More specifically

dS) Uplifted AdS_3 / CFT_2 D1-D5 system with elliptic fibration, 0-planes, D-branes $\rightarrow R_{dS} \gg \sqrt{r}$

\rightarrow parametric entropy count $g_s \ll 1$

FRW) Uplift $AdS_5 \times S^5$ with $n > 36$ (p,q) 7-branes \rightarrow GR solution

$$dS_{5dE}^2 = c^2 \left(t^{\frac{2}{c}} - w^2 \right)^{c-1} dw^2 + \left(1 - \frac{w^2}{t^{\frac{1}{2c}}} \right)^{c-1} \left(-dt + c^2 t^{\frac{1}{2c}} dH_3^2 \right)$$

$c^2 = \frac{7}{3} > 1 \Rightarrow 2$ warped throats

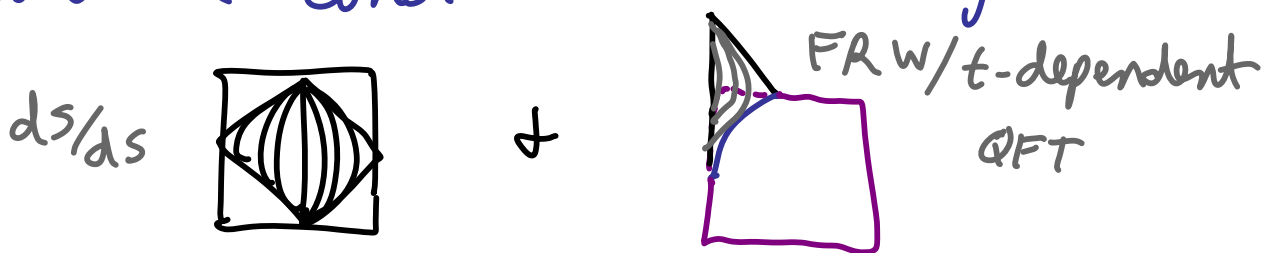
★ GR $D-1$ decouples at late times; S bound $\rightarrow \infty$

• QFT side: t -dependent theory, with $n > 36$ magnetic flavors

\rightarrow parametric count of d.o.f. on junctions

Comparative holography:

Recall in AdS/CFT that p-brane construction lands on Poincaré slicing. In dS + FRW, above brane constructions \rightarrow slicings



- inside a causal region
- a spatial direction (\leftrightarrow scale) emerges.
- \Rightarrow • # of degrees of freedom real, > 0 and unitarity more transparent
- symmetries (such as they are) less manifest

As in AdS \rightarrow Global, this may connect to other slicings (dS/CFT, FRW/CFT)

Anninos/Hartman/Strömberg
Harlow/shenker/stanford/Susskind

\nwarrow also have \nearrow
dynamical gravity

We would like to extract the essential features of the dual theories, given the concrete brane constructions

Plan :

- (1) Flavor content of FRW duals, unitarity, & time-dependent QFT couplings
- (2) dS_D duals and holographic RG

Magnetic flavors & uplifting

e.g. IIB (p,q) 7Bs wrapping $\Sigma_3 \subset S^5$

Tension $\propto \Lambda^4 \frac{1}{R^2} \frac{1}{g_s^2}$

$S^1 \rightarrow S^5 \propto 3$
 \downarrow
 $CP^2 \propto 1$

Competes with curvature

on CP^1 : 24 7Bs $\rightarrow R=0$

* CP^2 : 36 7Bs $\rightarrow R=0$

Banks
 Douglas
 Seiberg,
 Aharony
 Maldacena
 Fayazuddin
 Joe P ES

$\Delta N \equiv n - N_{R=0}$

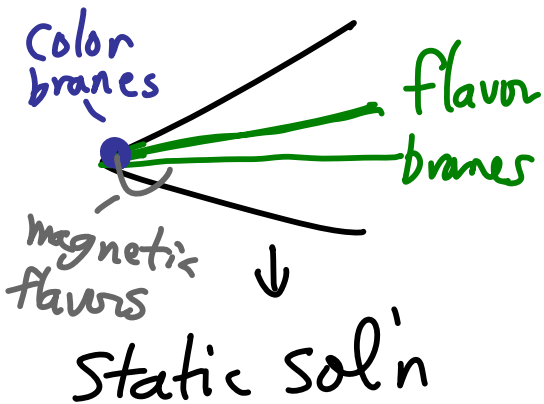
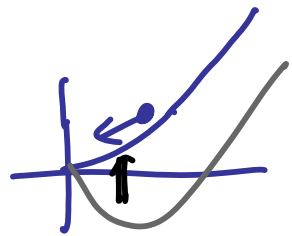
$\Delta N < 0$

AdS



$\Delta N \geq 0$

Cosmo



no static solution
 but \exists simple,
 t-dep't solutions
 cf Kleban + Redi

$\Delta N \geq 0$ distinction in dual QFT?

Unitarity bounds & t -dependent QFT.

Given well-defined QFT at
some scale, unitarity bounds
(e.g. $\Delta_{\text{scalar}} > \frac{d-2}{2}$ in CFT)
help constrain IR physics.

(e.g. SQCD conformal window ends
at $N_f = \frac{3}{2} N_c, \dots$)

In D3 - (p,q)7 system,

$\Delta n > 0 \Rightarrow$ no static solution.

* In the simplest case (with parallel 7-branes $\Leftrightarrow N=2$ susy)

this follows from unitarity.

Seiberg-Witten curve

$$y^2 = x^3 - f(u)x - g(u)$$

$(s < \frac{3}{2}r)$ $\overset{u^r}{\sim}$ $\overset{u^s}{\sim}$
 $\boxed{2\Delta_y = 3\Delta_x = 5\Delta_u}$

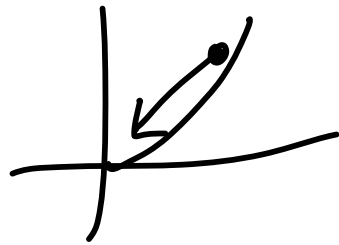
$$\frac{d\lambda_{sw}}{du} = \frac{dx}{y} \quad \uparrow \quad \int \lambda_{sw} \sim \frac{u dx}{y}$$

\uparrow
 BPS masses $\Rightarrow \left(\begin{array}{l} \Delta_u = 1 + \Delta_y \\ -\Delta_x \end{array} \right)$
 dim 1

$$\Rightarrow \Delta_u = \frac{12}{12 - N_f}, \quad N_f > 12$$

would violate unitarity

But there do exist t -dependent solutions



with the required properties
(redshift, N_{dof} , $M_p \rightarrow \infty$)

for a dual $EFT \rightarrow QFT$.

\wedge
 t -dependent

\rightarrow More general question: how do t -dependent couplings affect unitarity bounds on IR behavior \leftrightarrow field content

Consider

$$\int \Delta \mathcal{L} = \int dt d^{d-1} \vec{x} g(t, \vec{x}) \mathcal{O}$$

$$g = g_0 t^\alpha \quad \text{or} \quad g_0 (t^2 - x^2)^{\frac{\alpha}{2}}$$

as $t \rightarrow \infty$

- α can change whether $\Delta \mathcal{L}$ dominates at late times (IR)

- If $\Delta \mathcal{L}$ marginal in IR under $x^m \rightarrow \lambda x^m$, then

$$[\mathcal{O}] = d + \alpha$$

→ Expect α can shift relevance condition & unitarity bounds

We can analyze this explicitly in large- N double trace flows.

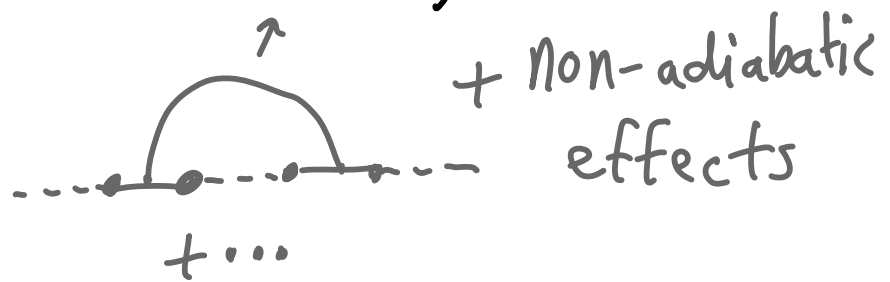
$$\Delta S = \int \frac{1}{2} \phi m^2 \phi + g(t, \vec{x}) \partial \phi$$

$$\approx \int \frac{g^2}{2m^2} \partial \partial$$

$$\langle \phi \phi \rangle = \text{---} + \text{---} \begin{matrix} \circ \circ \\ \downarrow \end{matrix} \text{---} + \text{---} \begin{matrix} \circ \circ \\ \downarrow \end{matrix} \text{---} + \dots$$

Effectively
Gaussian

$$+ \mathcal{O}\left(\frac{1}{N}\right)$$



Static Limit:

For analysis in
holographic case
see Andrade, Faulkner
Marolf ...

$$\Delta_{\pm} = \frac{d}{2} \pm \nu$$

$$\langle \mathcal{O}_{\pm}(p) \mathcal{O}_{\pm}(-p) \rangle = -i C_{\pm\nu} (p^2 - i\epsilon)^{\pm\nu}$$

$$S_{\text{CFT}}^{(+)} + \int \frac{g^2}{4m^2} \underbrace{\mathcal{O}_+ \mathcal{O}_+}_{\Delta = d+2\nu}$$

$\Delta = d+2\nu$
irrelevant

$$\langle \phi(p) \phi(-q) \rangle = \frac{-i \delta(p-q)}{m^2 - g^2 C_{\nu} (p^2)^{\nu}}$$

$$\rightarrow \langle \mathcal{O}_- \mathcal{O}_- \rangle$$

as $p^2 \rightarrow \infty$

$$\boxed{2^{-2\nu} \pi^{\frac{d}{2}} \frac{\Gamma(-\nu)}{\Gamma(\frac{d}{2} + \nu)}}$$

UV \mathcal{O}_- nonunitary for $\nu > 1$

OK as cut off QFT $\left\{ \begin{array}{l} \lambda_g^\nu \\ g^{\frac{1}{\nu-1}} \end{array} \right.$

IR \mathcal{O}_+ unitary

t-dependent case

- $\int \lambda_0 t^{2\alpha} \mathcal{O}_+^2$ is relevant
(dominates 2 pt ftns at large Δx)
when $[\lambda_0] = 2(\alpha - \nu) > 0$
- Unitarity maintained, including $\nu > 1$
- Can UV complete, e.g. SUSY models
(effects of λ lost for $\Delta t < \frac{\lambda}{\Lambda}$)

$$\Delta S = \int \frac{1}{2} \phi m^2 \phi + \underbrace{g(t, \vec{x}) \phi \partial_t \phi}_{\substack{\equiv \\ \tilde{\phi}}}$$

$$\langle \tilde{\phi}(p) \tilde{\phi}(q) \rangle \xrightarrow{\mathbb{R}} i \delta(p-q) \frac{(p^2 - i\varepsilon)^{-\nu}}{C_\nu}$$

$$\Rightarrow \langle \phi(x) \phi(x') \rangle = \frac{-1}{C_\nu C_{-\nu} g(x) g(x') [(x-x')^2]^{\Delta_-}} \quad \left\{ \begin{array}{l} \text{Not norm of} \\ \text{a state} \end{array} \right. \quad \left\{ \begin{array}{l} \Delta_- \\ * \end{array} \right.$$

Despite the Δ_- here, forward scattering* is unitary

$$\text{Im } A(\chi \rightarrow \chi) \propto -\sin \pi \nu / C_\nu > 0$$

(Integrator/Ginsparg: $\text{Im } A_\pm \propto C_\pm (\Delta - (\frac{d-2}{2}))$ in CFTs)

Scales in t -dependent case:

t -dependence of λ doesn't matter

$\lambda \rightarrow$ strong }
would get }
ghosts at } Λ_λ

\Rightarrow UV complete...

- $\lambda(t)$ dynamical field
- SUSY gauge theory e.g. in which ϕ, θ composite below Λ_λ

... or cut off

$$\frac{\dot{\lambda}}{\lambda} \sim \frac{g}{t}$$

$$\int \lambda_0 t^{2g} \theta_+^2 \text{ relevant \& unitary}$$


Note: can be more relevant than mass term, maintain long-range correlations!

② UV structure of dS_D duals

Maximal symmetry & RG

$$ds_D^2 = \sin(h)^2 \frac{w}{L} ds_{dS_{D-1}}^2 + dw^2$$

(1) In AdS_D , Moduli fixing $(\partial_{\Phi_I} V = 0)$
is dual to $\beta_{\{1\}} = 0$.

(2) In dS_D , we also (meta-)stabilize
the moduli $(\partial_{\Phi_I} V = 0)$  $M^2 \geq 0$

→ How are (1) & (2) reflected
in holographic RG? What
is the $(D-1)$ dual of (2)?

Wilsonian holographic RG (in progress)

... Heemsterk
Pohinski

$$\left\{ \begin{array}{l} ds^2 = L_D^2 \frac{dz^2}{z^2} + a(z, \phi)^2 ds_{D-1}^2 \\ \text{Scalar } \phi \text{ with potential } V(\phi) \end{array} \right.$$

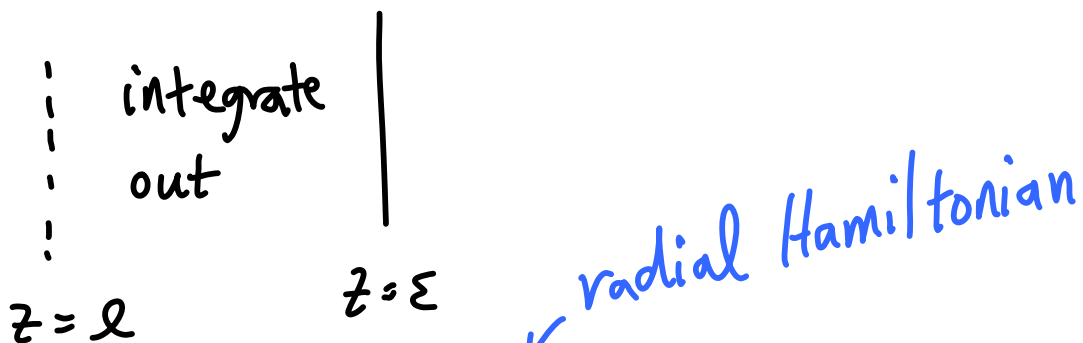
integrate
out

$$\rightarrow \text{Wilson action } S(l) = \int_{z=l}^{z=\epsilon} \sigma_{nm}(l) T^n g^m$$

• For $V'(\phi_*) = 0$, find solution
 $\partial_l \sigma_{01}^{(\vec{L}=0)} = 0$ regardless of sign
 of $V_* = V(\phi_*)$

→ single-trace coupling l -independent
 for both dS_D & AdS_D

Metastability & RG



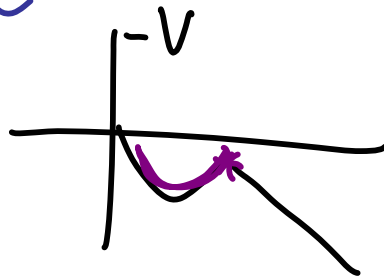
$$\psi_{uv}[\tilde{\phi}, \tilde{a}; \ell] = \langle \tilde{\phi}, \tilde{a} | e^{-H \log \ell} | \phi_0, a_0 \rangle$$

(related to Wilson action by integral transform)

Meta-stability adds bounce solution,

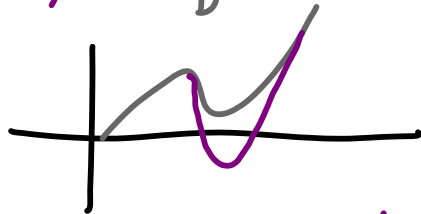
\Rightarrow non-perturbative correction to

Wilson action



- Metastability & Microscopics

Our explicit $(A)dS_D$ vacua are all meta stable



- $D3-\bar{D}3$ annihilation • decay to $R \rightarrow \infty$...

- Reminiscent of

$AdS \times S^5 / \mathbb{Z}_k$
(susy with free action)

Quiver $U(N)^k$
 QFT's

Not complete CFT, but sensible (exponentially long-lived) as warped throat in compact geometry.

Kachru ; Horowitz ; Kachru
 ES ; Orgera ; Simic
 Polchinski ; Trivedi

Can compute UV behavior of Correlators

$$ds^2 = \sin^2 w \frac{ds^2_{dS_{d-1}}}{L^2} + dw^2$$

$$Z_{\text{bulk}} = \int D\tilde{\Phi} \int_{\substack{w < \frac{\pi}{2} L \\ \Phi(\frac{\pi}{2} L) = \tilde{\Phi}}} [D\Phi] | e^{iS_<} \int_{\substack{w > \frac{\pi}{2} L \\ \Phi(\frac{\pi}{2} L) = \tilde{\Phi}}} [D\Phi] | e^{iS_>}$$

$Z_{\text{QFT}}^{(1)}[\tilde{\Phi}]$ $Z_{\text{QFT}}^{(2)}[\tilde{\Phi}]$

e.g. at Gaussian level,

$$\langle \mathcal{O} \mathcal{O} \rangle_L^{(1)} = \frac{\int \tilde{\Phi}^2 Z_1}{\int \tilde{\Phi}_1^2} = \frac{\Gamma[\frac{1}{2}(d-\hat{\Delta}+L)] \Gamma[\frac{1}{2}(1+\hat{\Delta}+L)]}{\Gamma[\frac{1}{2}(d-\hat{\Delta}+L)] \Gamma[\frac{\hat{\Delta}+L}{2}]}$$

• $\hat{\Delta} = \frac{d}{2} + \sqrt{(\frac{d}{2})^2 - m^2 L^2}$ (but $\langle \mathcal{O} \mathcal{O} \rangle_e$ real)

cf Bousso Maloney Strominger

• Δ flows to $\frac{d}{2} + \frac{1}{2}$ in UV (double trace)

• encodes (max. symmetric) shape of dS_0 warp factor

Summary

• Can uplift AdS/CFT to cosmology and derive some features of the resulting holographic duals:

— 2 EFT's + GR_{D-1}

Concrete
brane constructions \rightarrow decouples at late times
 \leftrightarrow Macro warped metric

— magnetic flavor content $N_f > N_f^{\text{CFT}}$
fits with unitarity \rightarrow more general
lesson for spacetime-dependent QFT

— Can go beyond low energy
description, e.g. capturing
the $dS_D \rightarrow dS_{D-1}$ warp factor in
RG and correlators. $d_x \sigma_{01} = 0$