

Strongly Interacting Fermi Gases: Rotation and Optical Lattices

Hui Zhai

in collaboration with Tin-Lun Ho

Department of Physics, Ohio-State University

**Poster on “ Correlated States in Degenerate Atomic Gases ”
KITP, UC Santa Barbara
April 2007**

Ways to achieve strong interactions in quantum gases

Feshbach resonance:

scattering length diverges,
Fermi energy is the only energy scale
====> non-perturbative

Rotation:

each Landau level is highly degenerate,
====> within each Landau level interaction
is the only energy scale

Optical Lattices:

wave function localized at each site,
====> tunneling suppressed and
interaction enhanced

More Intriguing Physics:

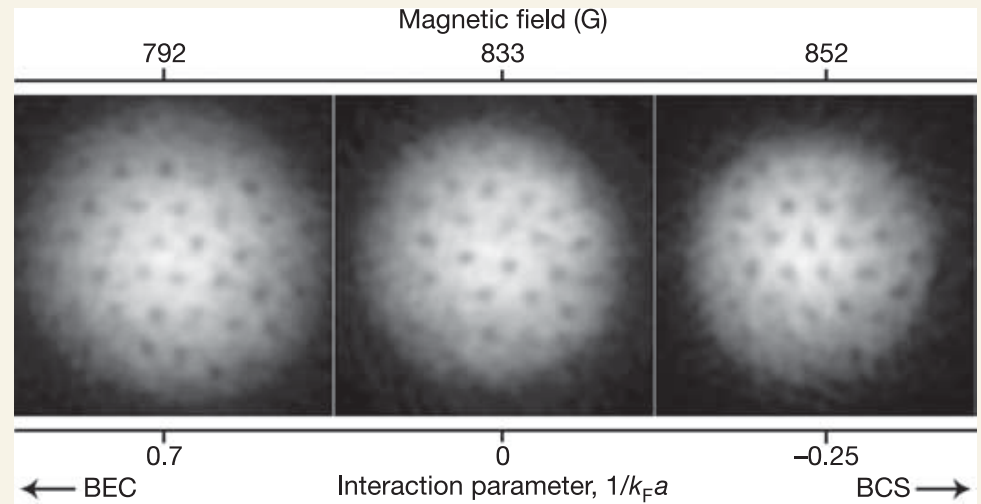
resonance + rotation
resonance + optical lattices

Experimental Motivation

Rotating Fermion Superfluid:

Stable vortex lattices created

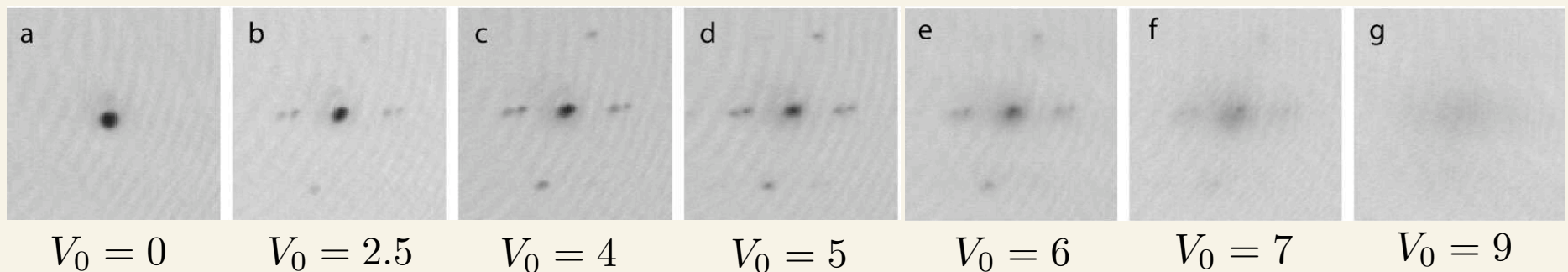
M. W. Zwierlein, J. R. Abo-Shaeer, A. Schirotzek, C. H. Schunck, W. Ketterle:
Nature, 435, 1047 (2005)



Fermion Superfluid in Optical Lattices:

Loss of superfluidity as lattice increases

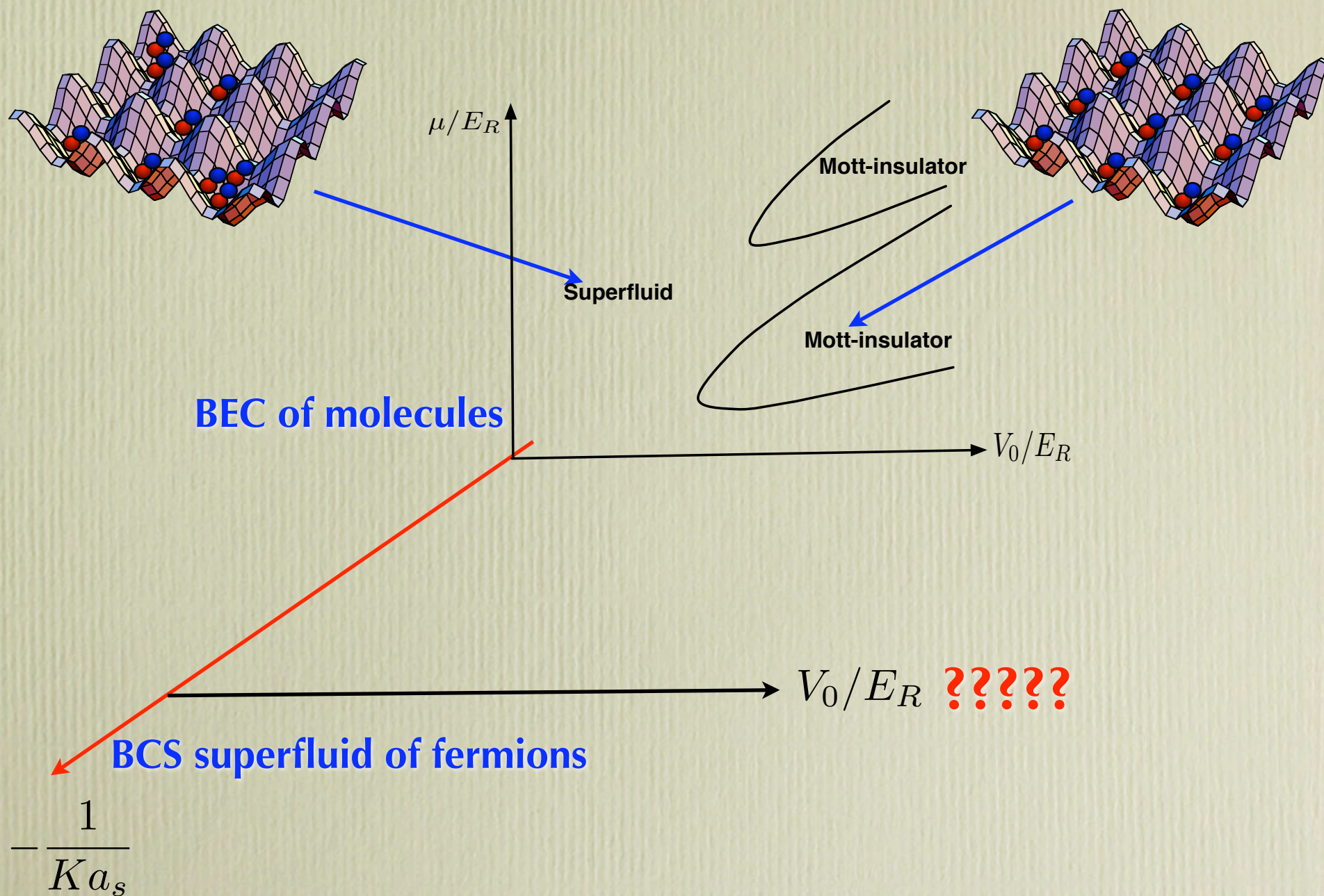
J. K. Chin, D. E. Miller, Y. Liu, C. Stan, W. Setiawan, C. Sanner, K. Xu and W. Ketterle,
Nature, 443, 961 (2006)



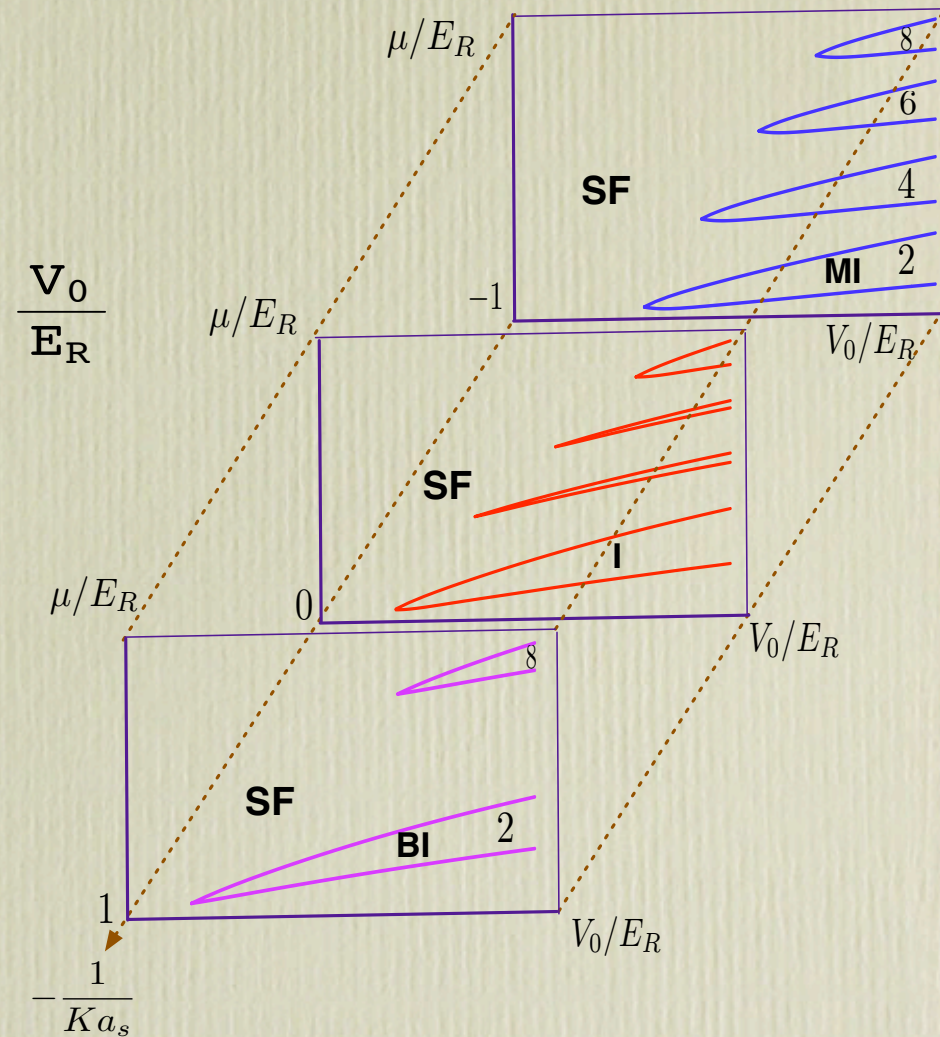
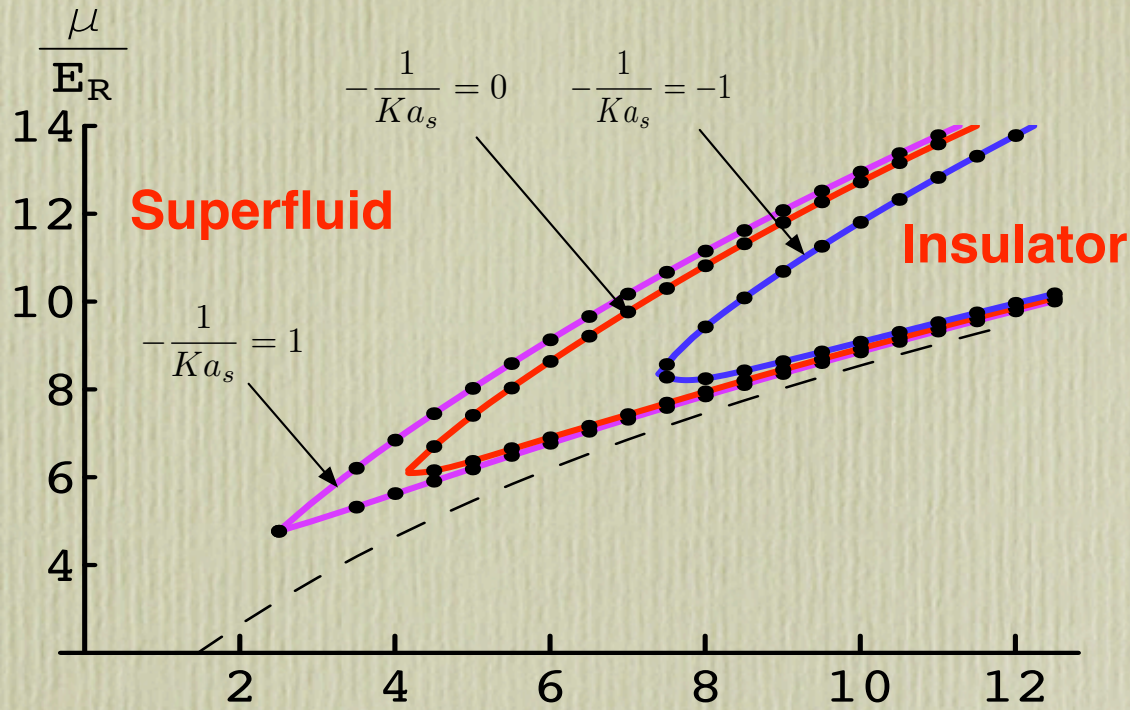
Part I:
**Superfluid-Insulator Transition of Strongly
Interacting Fermi Gases**

Hui Zhai and Tin-Lun Ho, arXiv: 0704.2957

Question: How SF-Insulator transition depends on a_s



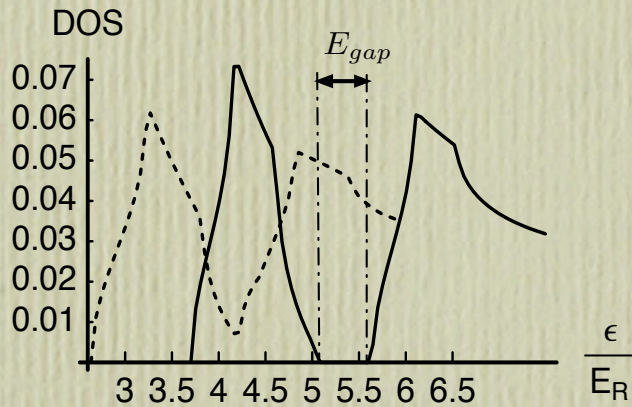
Answer from mean-field theory



Model

Single particle physics:
band structure

$$H_0 = -\frac{\hbar^2}{2m}\nabla^2 + V_0(\sin^2(Kx) + \sin^2(Ky) + \sin^2(Kz))$$

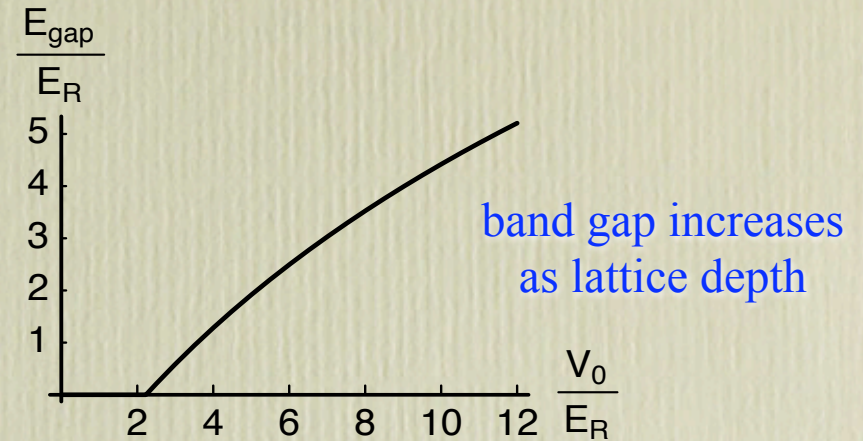


$$V_0 = 2E_R$$

dash line, no band gap

$$V_0 = 3E_R$$

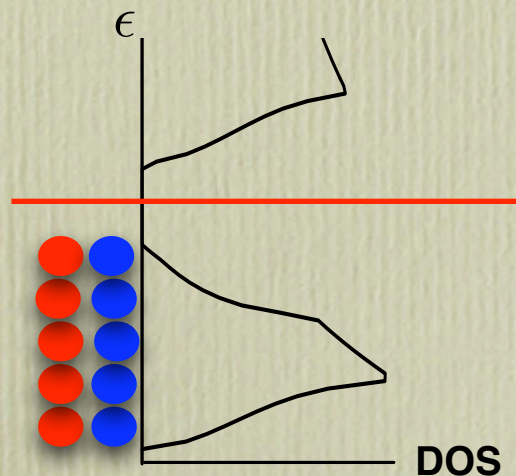
solid line, band gap opens



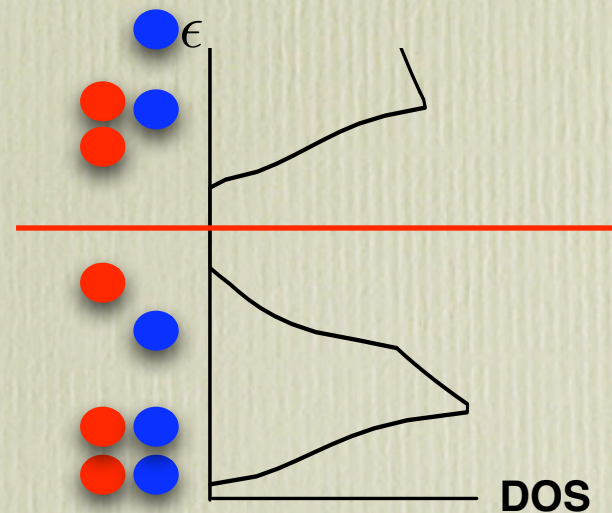
band gap increases
as lattice depth

Many-body physics:
Cooper instability

competition between **Energy Gain from Forming Pairs**
and **Energy Cost to Overcome Band Gaps**



Band Insulator



Fermion Pairs Superfluid

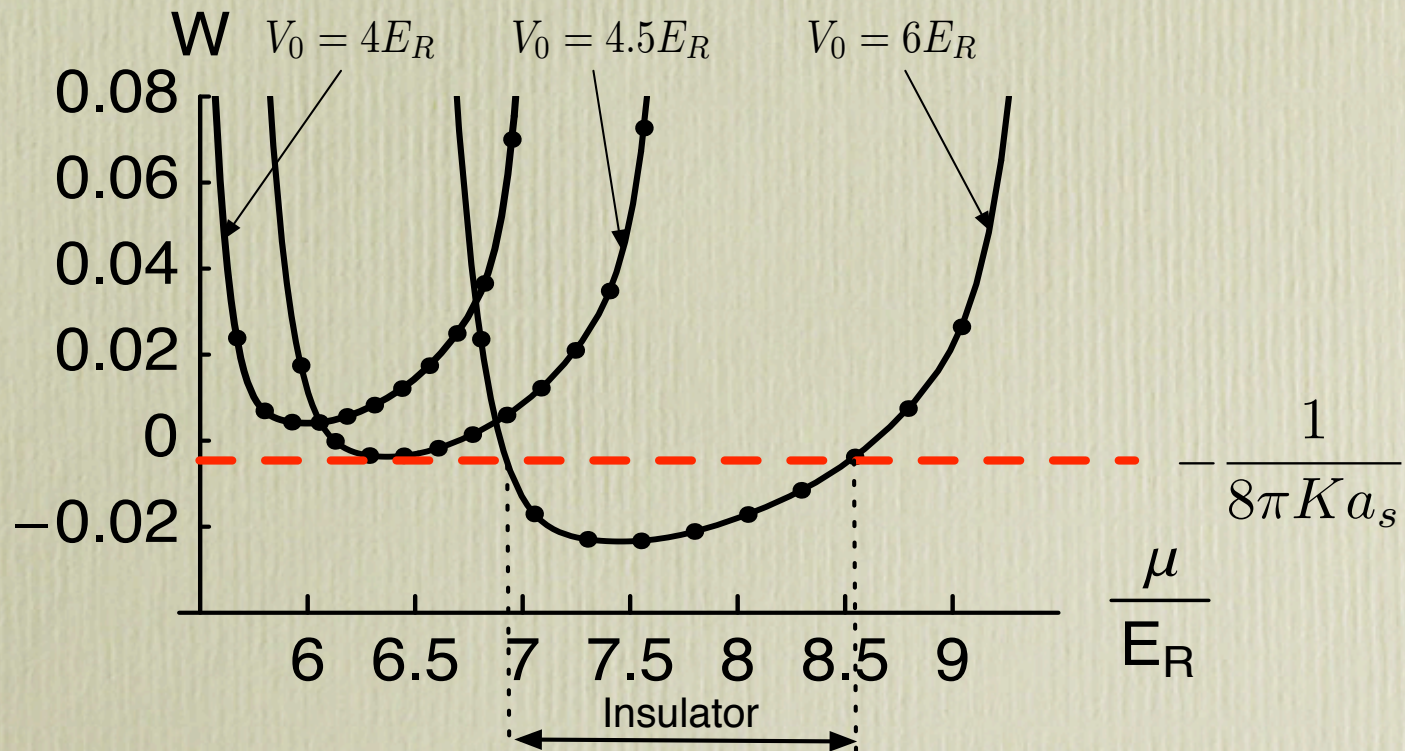
Outline of the calculation: Pairing susceptibility

$$\alpha_{\mathbf{G}} = W_{\mathbf{G}} + \frac{m}{4\pi\hbar^2 a_s}$$

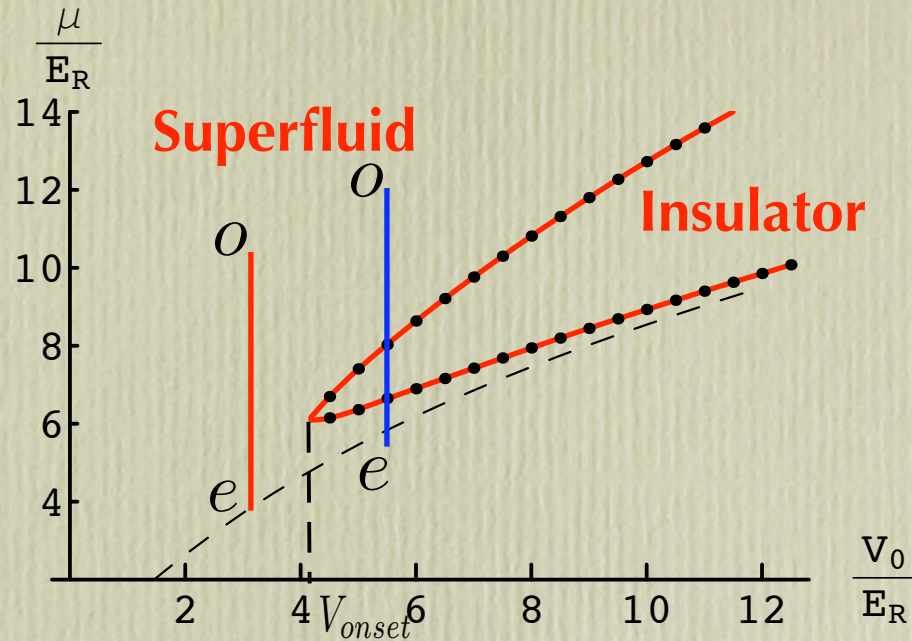
Condition for the onset of superfluidity

At least one of $\alpha_{\mathbf{G}} > 0 \iff W_{\mathbf{G}} > -\frac{m}{4\pi\hbar^2 a_s}$

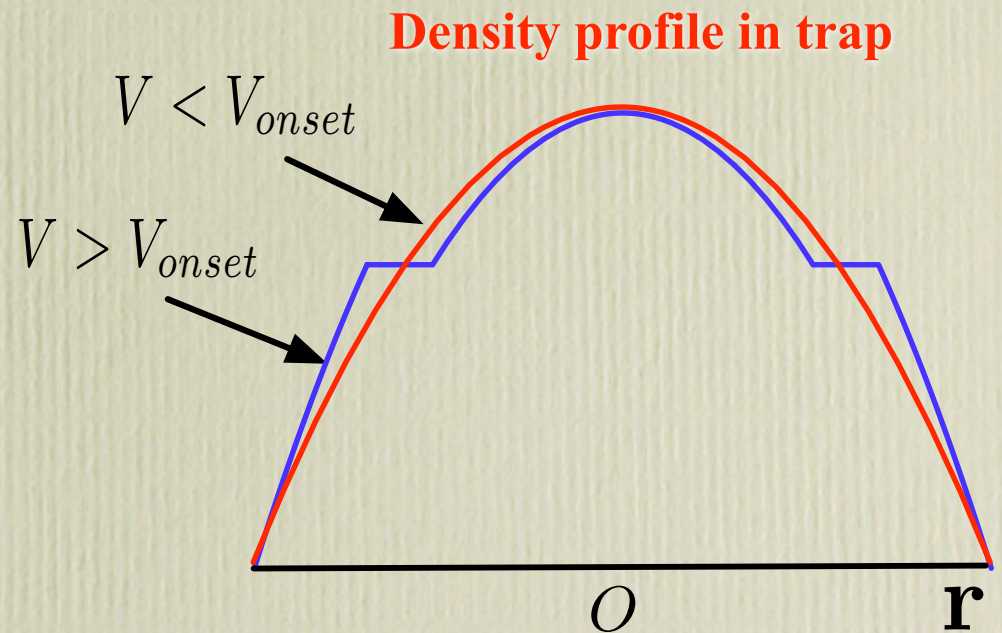
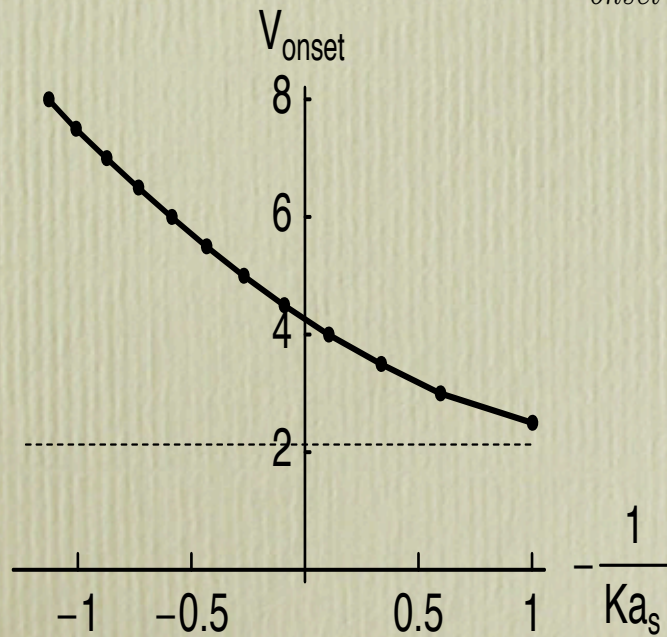
$$W_{\mathbf{G}} = \sum_{\mathbf{k}} \left(\sum_{n,n'} \frac{|Q_{n,n',\mathbf{k}}^{\mathbf{G}}|^2}{\xi_{n\mathbf{k}} + \xi_{n'-\mathbf{k}}} (\Theta(\xi_{n\mathbf{k}}) + \Theta(\xi_{n'-\mathbf{k}})) - \sum_{\mathbf{G}} \frac{1}{\hbar^2(\mathbf{k} + \mathbf{G})^2/m} \right) \quad Q_{nn'k}^{\mathbf{G}=0} = \delta_{nn'}$$



Critical Lattice Depth for the Onset of Insulator



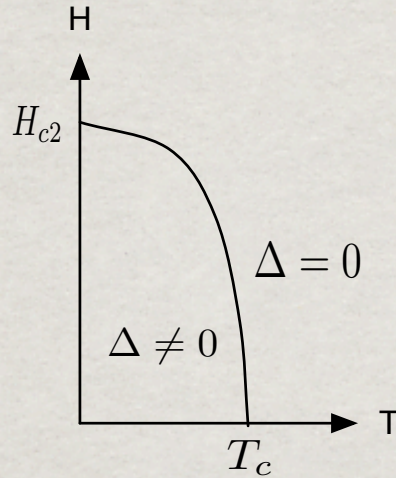
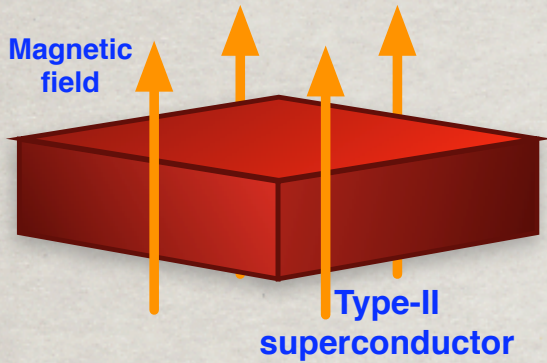
$$\mu(\mathbf{r}) = \mu_0 - V(\mathbf{r})$$



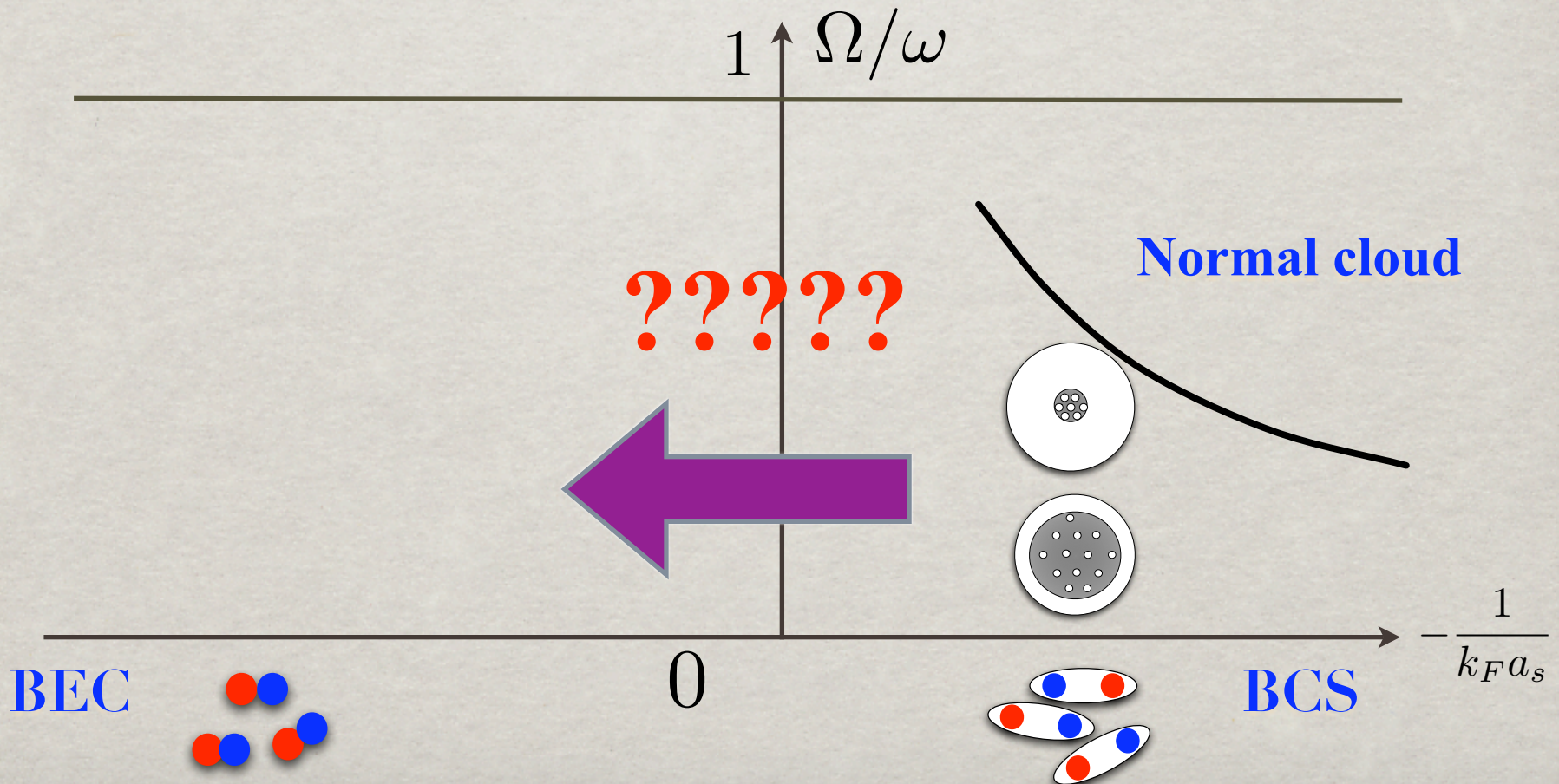
Part II:
**Critical Rotational Frequency of Superfluid
Fermi Gases across Feshbach Resonance**

Hui Zhai and Tin-Lun Ho, Physical Review Letters, 97, 180414 (2006)

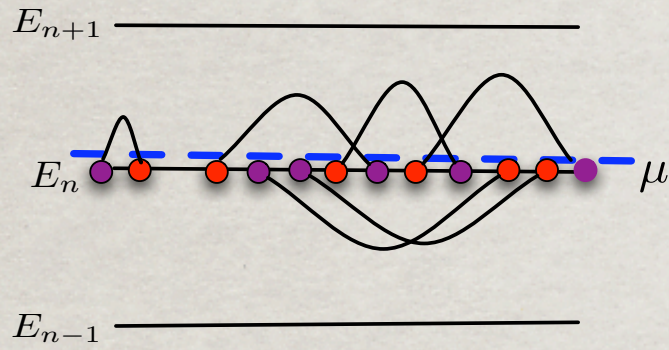
Question: How critical frequency depends on a_s



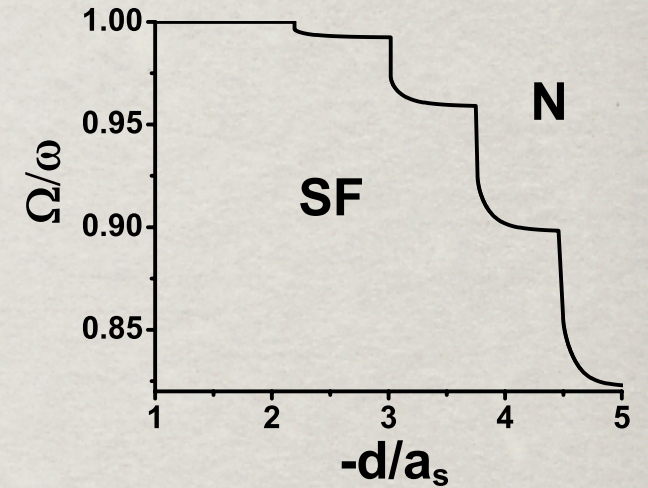
Critical rotational frequency \iff Analogy of H_c in superconductor



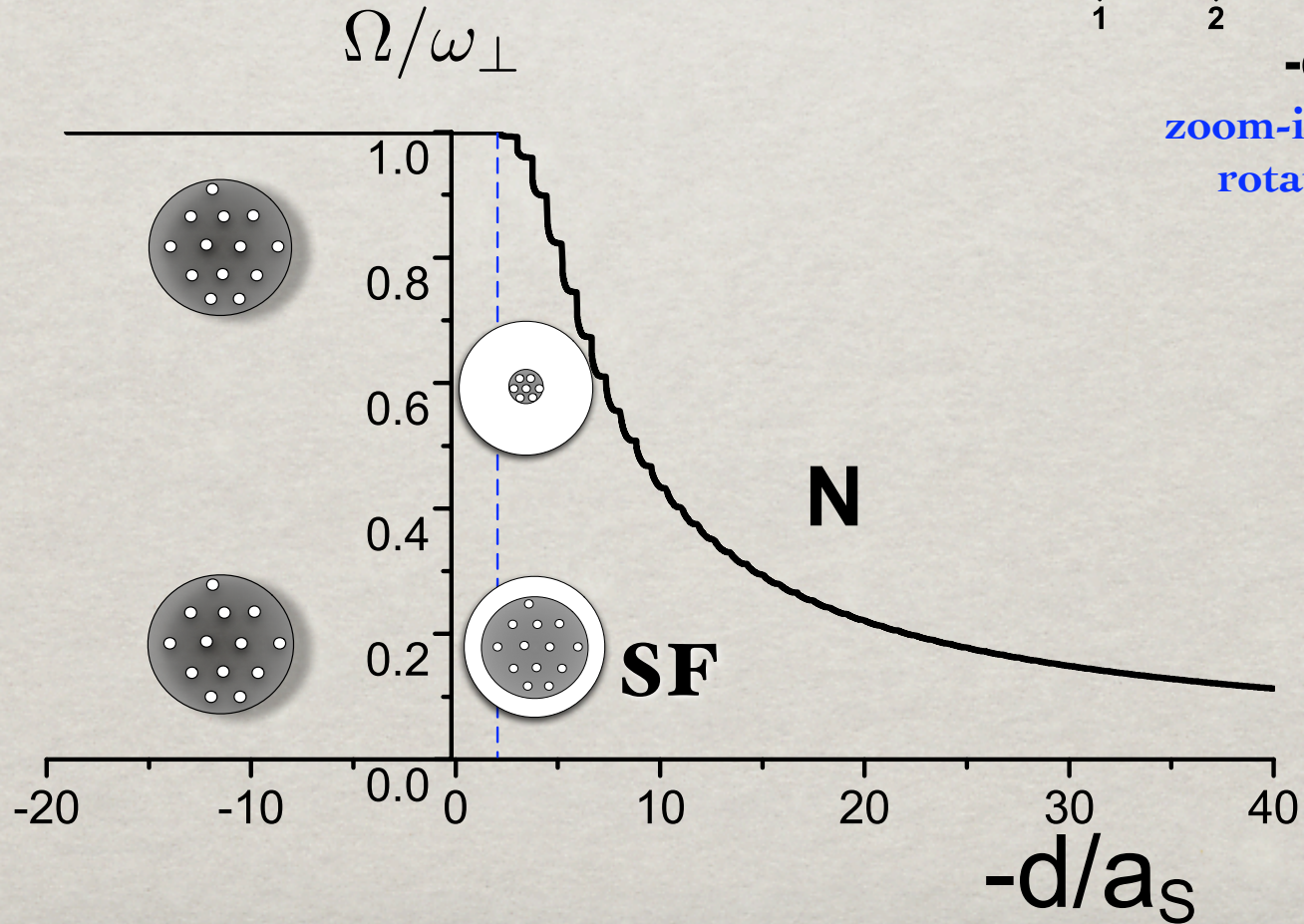
Answer: Seeing integer Landau levels



A sequence of jumps
reveal the features of
Landau levels



zoom-in plot of high
rotational limit



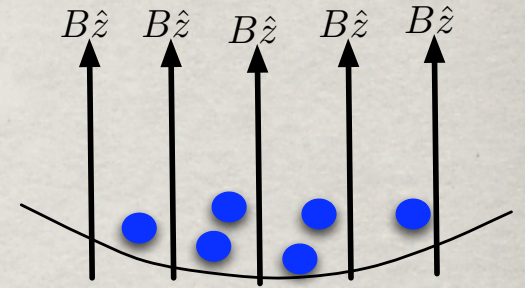
Model

Single particle Hamiltonian

$$H_0 = \frac{\hbar^2}{2m} \nabla^2 - \Omega L_z + \frac{1}{2} m \omega^2 r^2 + \frac{1}{2} m \omega_z^2 z^2$$



$$H_0 = \frac{1}{2m} (\mathbf{p} - m\Omega\hat{z} \times \mathbf{r})^2 + \frac{1}{2} m (\omega^2 - \Omega^2) r^2 + \frac{1}{2} \omega_z^2 z^2$$



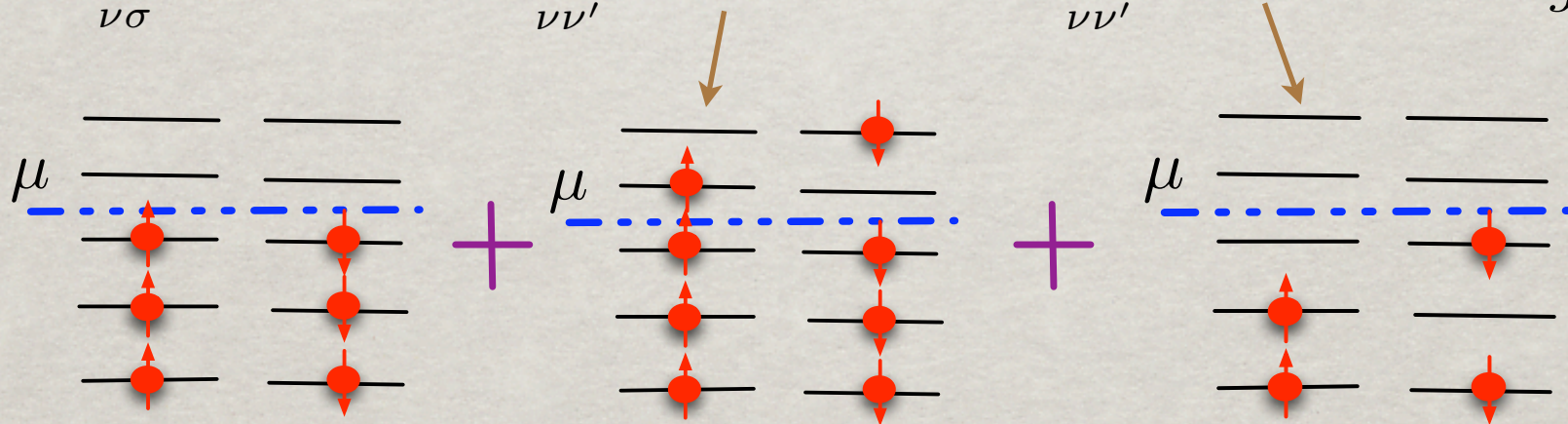
“electrons” in magnetic field:
Landau levels

local density approximation

$$\mu \rightarrow \mu - \frac{1}{2} m (\omega^2 - \Omega^2) r^2 - \frac{1}{2} \omega_z^2 z^2$$

Many-body Hamiltonian

$$\mathcal{H} = \sum_{\nu\sigma} \xi_\nu c_{\nu\sigma}^\dagger c_{\nu\sigma} + \Delta \sum_{\nu\nu'} \alpha_{\nu\nu'} c_{\nu\uparrow}^\dagger c_{\nu'\downarrow}^\dagger + \Delta^* \sum_{\nu\nu'} \alpha_{\nu\nu'}^* c_{\nu\uparrow} c_{\nu'\downarrow} + \frac{|\Delta|^2}{g}$$



$$F = F_0 + \alpha |\Delta|^2 + \beta |\Delta|^4 + \dots$$

$$\alpha < 0$$

Superfluid

$$\alpha > 0$$

Normal

Outline of the calculation: Pairing susceptibility

$$\alpha = -W \left(\frac{\mu}{\hbar\Omega} \right) - \frac{a}{a_s} \quad a = \sqrt{\frac{\hbar}{2m\Omega}}$$

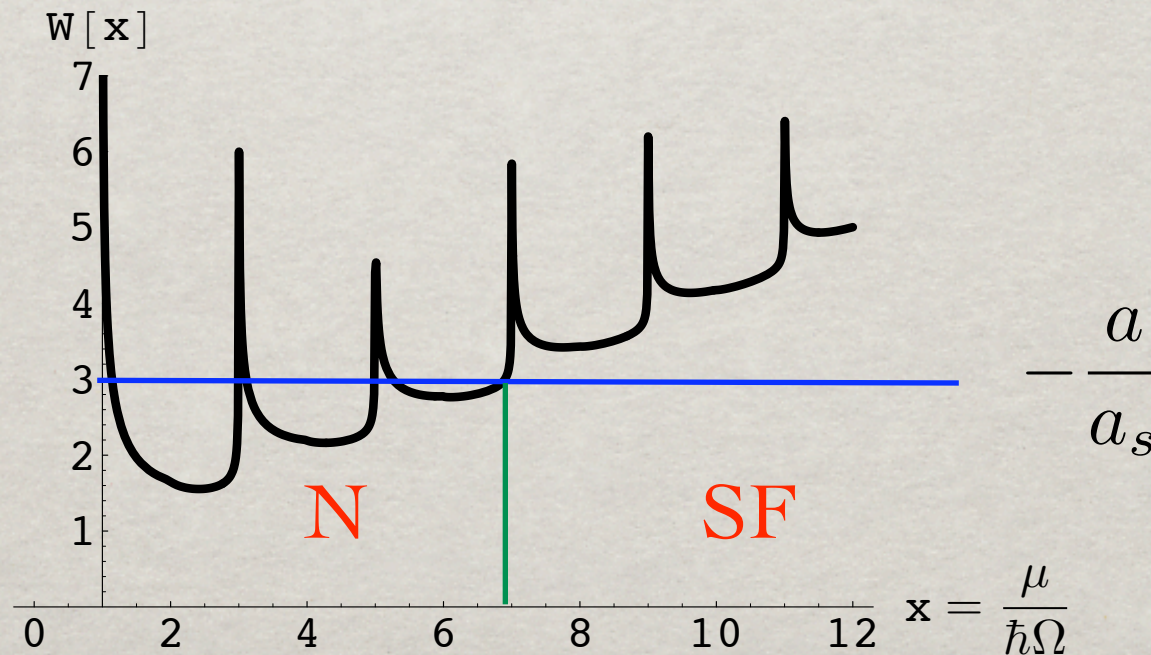
$$W = \frac{\pi\sqrt{2}a\hbar^2}{2m} \sum_{\nu\nu'} \frac{\tanh(\frac{\beta\xi_\nu}{2}) + \tanh(\frac{\beta\xi_{\nu'}}{2})}{\xi_\nu + \xi_{\nu'}} |\alpha_{\nu\nu'}|^2 - \sum_N \frac{1}{4\sqrt{N+1}}$$

$$W > -\frac{a}{a_s}$$

Superfluid

$$W < -\frac{a}{a_s}$$

Normal



Thank You Very Much for Your Attention

