



Organizing Principles

In the words of Donald Rumsfeld:

"As we know, there are known knowns, there are things we know we know.

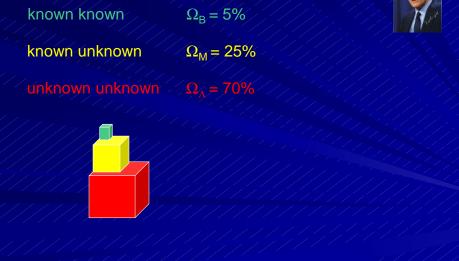
We also know there are known unknowns, that is to say we know there are some things we do not know.

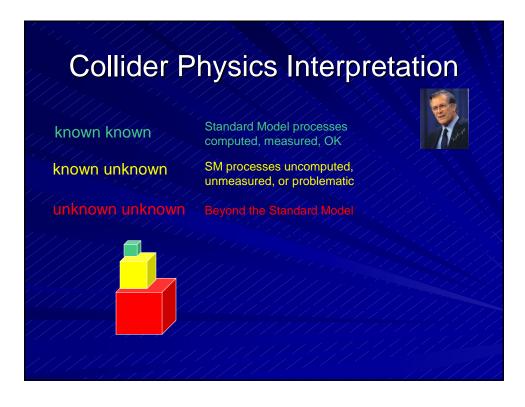
But there are also unknown unknowns -the ones we don't know we don't know."

Our science is frequently "just" about moving known unknowns into the category of known knowns -- but this work prepares us to venture into the unknown unknowns



Cosmological Interpretation



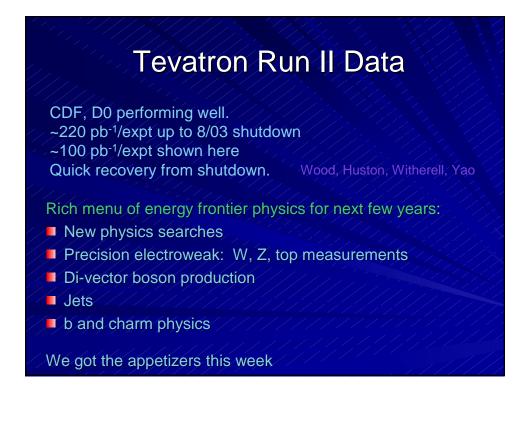


Current experimental situation

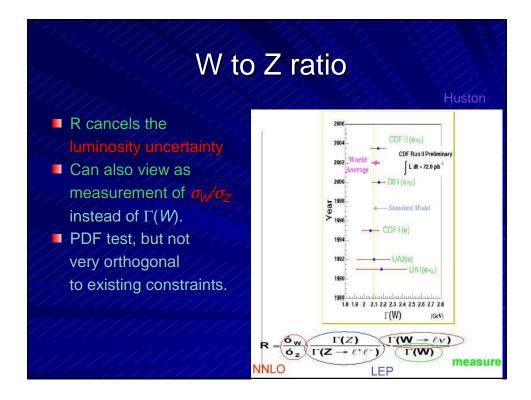
Mostly known known: Overall, Standard Model in great shape

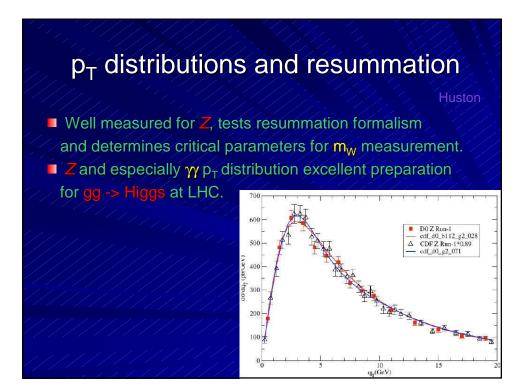
However, there are a number of $2.5-3.5\sigma$ known unknowns. Are they statistics, experimental or theoretical systematics? Or could one or more of them herald the unknown unknown?

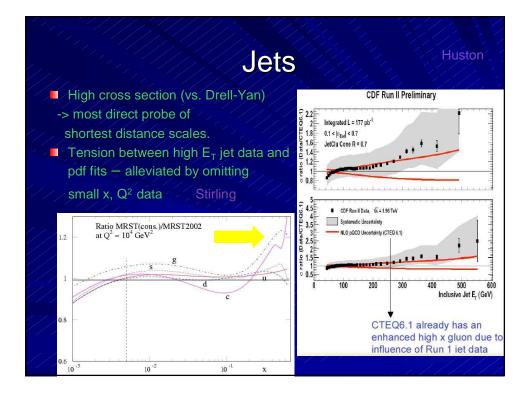
Precision Generally good agreement,		Some tension in fit			
		Winter 2			
Overall the EW precision tests support the SM and		Measurement	Pull	(Omeas_Olit)/omeas	
a light Higgs.	$\Delta \alpha_{\rm intro}^{(6)}(m_2)$	0.02761 ± 0.00038	-0.18	-3-2-10123	
	m ₇ [GeV]	91.1875±0.0021	0.02		
	F7 [GeV]	2 4 952 ± 0.0023	-0.36	Image:	
The χ^2 is not great:	of [nb]	41.540 ± 0.037	1.67		
χ^2 is not great: $\chi^2/ndof=25.5/15 (4.4\%)$	R	20.767 ± 0.025	1.01		
χ^2 /ndof=25.5/15 (4.4%)	A ^{0,1}	0.01714 ± 0.00095	0.79		
	A(P)	0.1485 ± 0.0032	-0.42		
	Rb	0.21644 ± 0.00065	0.99		
Note: includes NuTeV and	Re	0.1718 ± 0.0031	-0.15		
APV [not (g-2),,]	Atb Atb	0.0995 ± 0.0017	-2.43		
	A ^{0,c}	0.0713 ± 0.0036	-0.78		
Without NuTeV:	Ab	0.922 ± 0.020	-0.64	-	
(th. error guestionable)	Ac	0.670 ± 0.026	0.07		
(un enter questionable)	A(SLD)	0.1513 ± 0.0021	1.67		
210dot 10 7/14 (07 09/)	sin ² 0 ^{lept} (Qb)		0.82		
χ²/ndof=16.7/14 (27.3%)	m _W [GeV]	80.426 ± 0.034	1.17		ones /
Much better!	Tw [GeV]	2.139 ± 0.069	0.67		to watch
NuTe	m, [GeV]	174.3 ± 5.1	0.05		
	sin ² 0 _W (vN)	0.2277 ± 0.0016	2.94		
G. Altarelli VAPV	Q _W (Cs)	-72.83±0.49	0.12		
				-3-2-10123	
		*			



ample E 8625 6 1599 7	Back. 6 3% 2 11% 2	nple: W o·B(W>Iv_I) (nb) 2.64 ±0.01 _{stat} ±0.09 _{sys} ±0.10 2.64 ±0.02 _{stat} ±0.12 _{sys} ±0.10	∂ _{ium} →	<mark>In 2 fb⁻</mark> 1 ∼1.1M	Huston
		2.64±0.02 _{stat} ±0.12 _{sys} ±0.1			
	26%	2.62 ±0.07 _{stat} ±0.21 _{sys} ±0.1	- CORCERS - CORC	~600K ~65K	72 pb ⁻¹
ven cou le true fo er theore accepta	Inting I or <i>Z</i> etical d ance sy	lescription of lepton stematics.		l p _T would	surely help
	•••				Mangano
	ven cou le true f er theore accepta o	ven counting 1 the true for Z er theoretical c acceptance sy 0 LO, $\Gamma_{W^{=0}}$ 50(2) 0.4971(2)	ven counting luminosity uncertain the true for Z for theoretical description of lepton acceptance systematics. O $LO, \Gamma_W^{=0}$ LO, no spin corr's po(2) 0.4971(2) 0.5259(2)	ven counting luminosity uncertainty the true for Z for theoretical description of lepton rapidity and acceptance systematics.	ven counting luminosity uncertainty be true for Z er theoretical description of lepton rapidity and p_T would acceptance systematics. O $IO, \Gamma_W=0$ LO, no spin corr's $IO, PDF=CTEQ6.19$ $p_0(2)$ $0.4971(2)$ $0.5259(2)$ $0.5245(2)$







Jets

Do we need NNLO jet cross sections at hadron colliders?

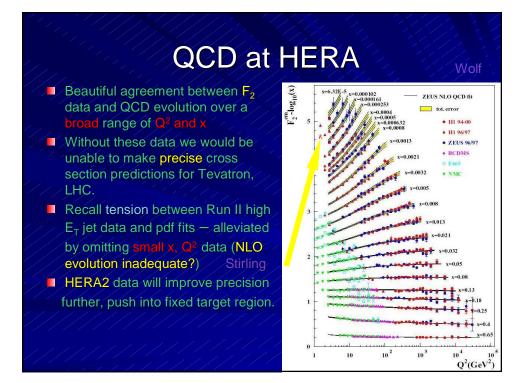
- After all, jets are rather complicated objects
- Steep É_T dépendence magnifies énérgy scale uncertainties
- Underlying event a problem (630 vs. 1800 data, energy flow studies)

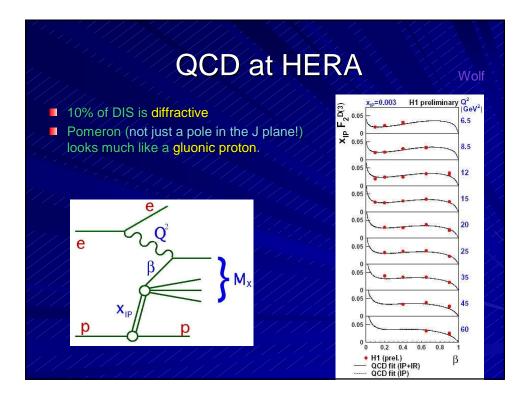
Still, I think the answer is YES ----

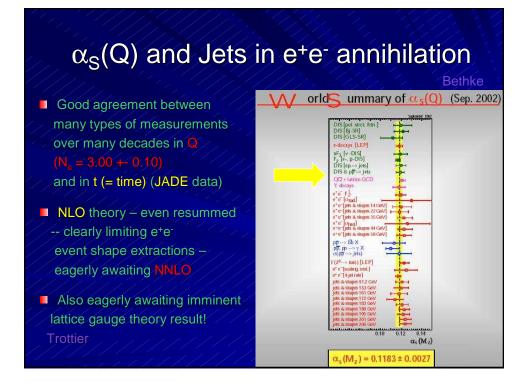
If nothing else, to focus more attention on these problems

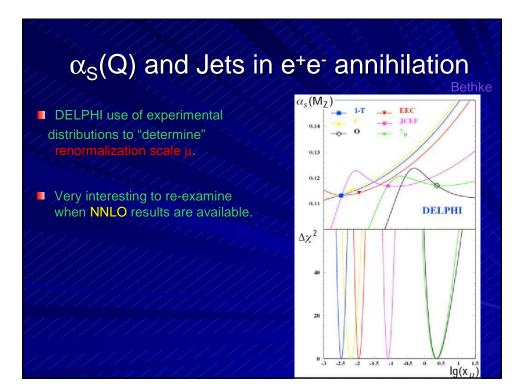
- Less μ-dependence to hide behind
- Less worries (one hopes) about matching theoretical and experimental jet algorithms, e.g. artificially introduced

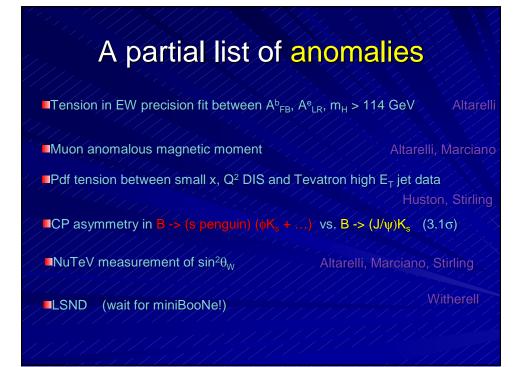
into NLO theory

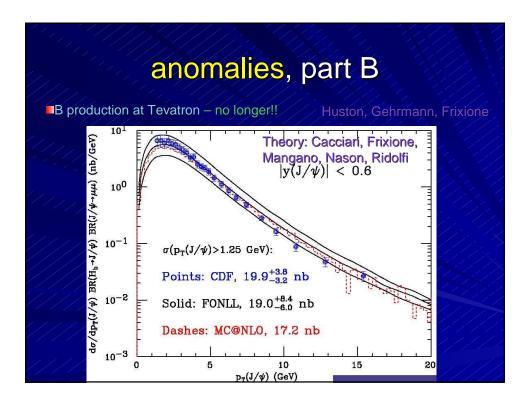


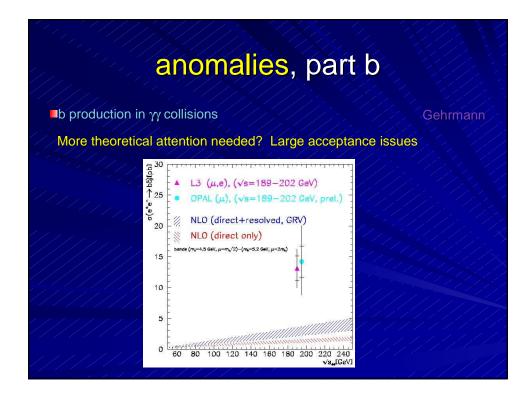




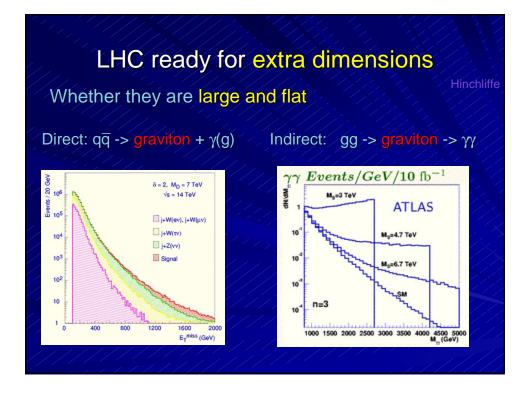


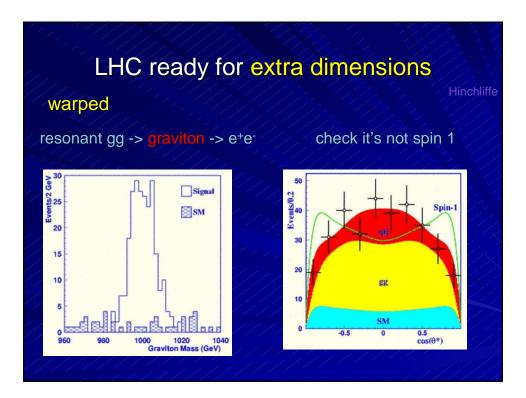


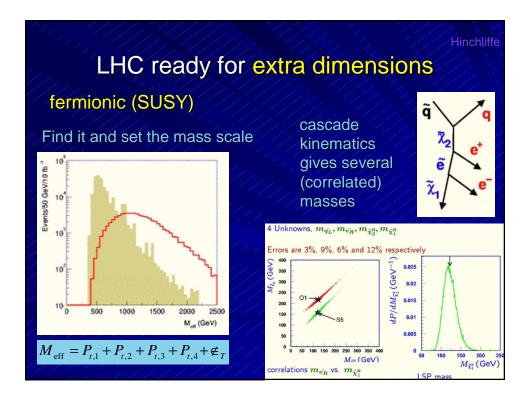


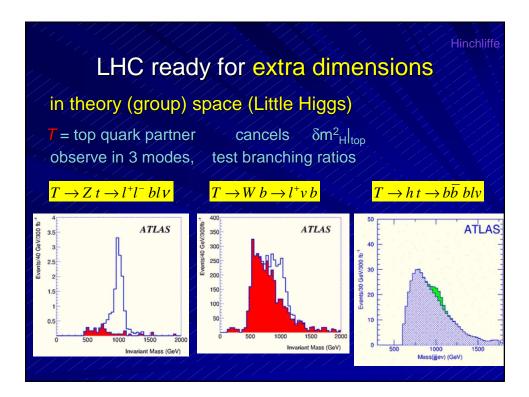


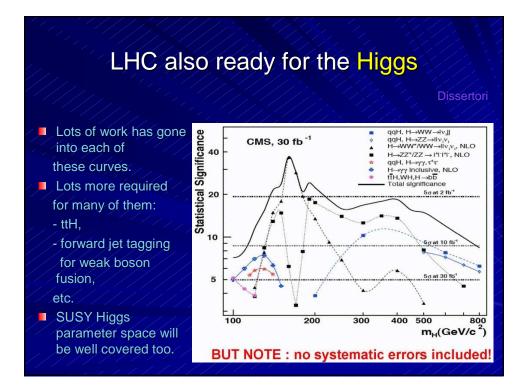


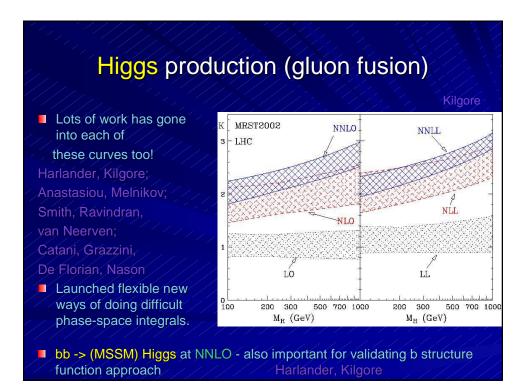




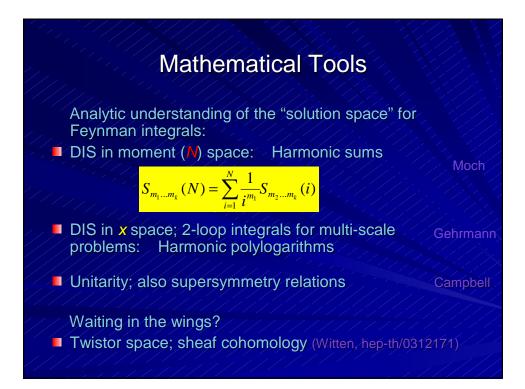


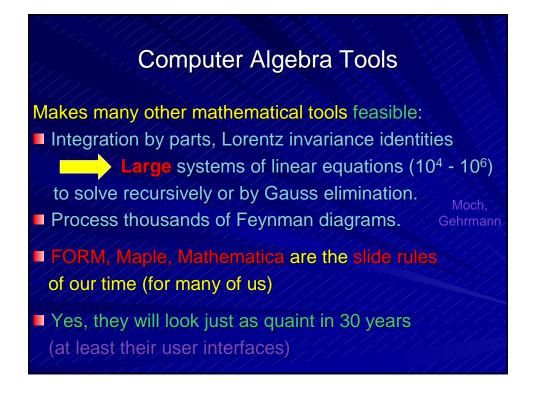








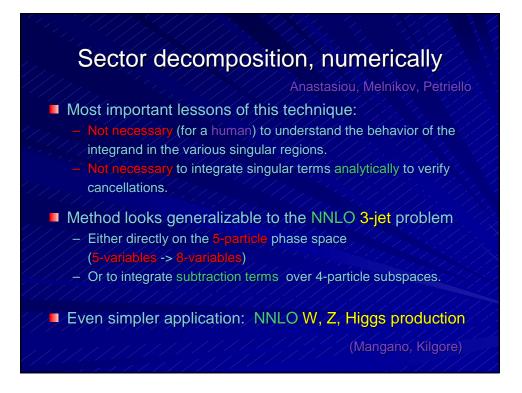


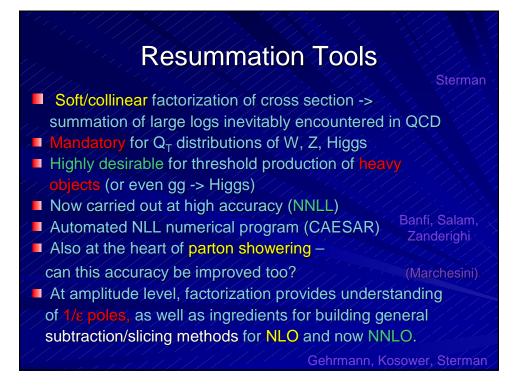




Gehrmann Kosower

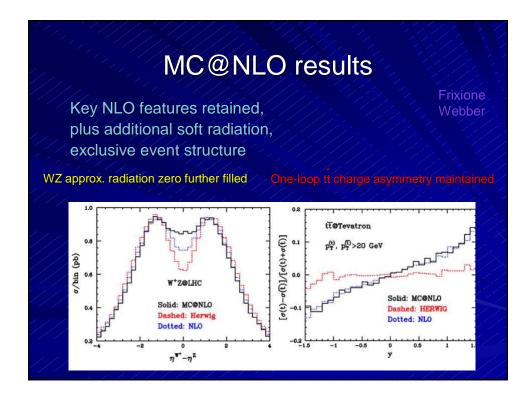
- Loop integration techniques (IBP, Lorentz invariance) can be applied to phase-space integration to give analytic results for singular integrals, order by order in dim. reg. parameter e.
- Including Gehrmann-De Ridder, Gehrmann, Heinrich those for NNLO "tripole" subtraction terms Weinzierl
- Sector decomposition, also invented for loop integrals, allows numerically stable direct evaluation of NNLO e⁺e⁻ -> 2-jet observables. Anastasiou, Melnikov, Petriello
- Efficient phase-space generation also important for accurate and fast evaluation of multi-particle cross sections with multi-channel peaking behavior: RAMBO, EXCALIBUR....





Three complementary approaches			
	ME MC's	X-sect evaluators	Shower MC's
Final state description	Hard partons jets.	Limited access to	Full information
	Describes geometry,	final state	available at the
	correlations, etc	structure	hadron level
Higher order	Hard to implement,	Straighforward	Included as vertex
effects: loop	require introduction of	to implement,	corrections
corrections	negative probabilities	when available	(Sudakov FF's)
Higher order	Included, up to high orders (multijets)	Straighforward	Approximate,
effects: hard		to implement,	incomplete phase
emissions		when available	space at large angle
Resummation of large logs	??	Possible, when available	Unitary implementation (i.e. correct shapes, but not total rates)





An experimenter's wishlist Hadron collider cross-sections one would like to know at NLO Run II Monte Carlo Workshop, April 2001			
Single boson	Diboson	Triboson	Heavy flavour
$\begin{array}{l} W+c\bar{c}+\leq 3j\\ Z+\leq 5j\\ Z+b\bar{b}+\leq 3j\\ Z+c\bar{c}+\leq 3j\\ \gamma+\leq 5j\\ \gamma+b\bar{b}+\leq 3j \end{array}$	$\begin{array}{l} WW + \leq 5j \\ WW + bb + \leq 3j \\ WW + c\bar{c} + \leq 3j \\ ZZ + \leq 5j \\ ZZ + bb + \leq 3j \\ ZZ + c\bar{c} + \leq 3j \\ \gamma\gamma + \leq 5j \\ \gamma\gamma + c\bar{c} + \leq 3j \\ WZ + c\bar{c} + \leq 3j \\ WZ + b\bar{b} + \leq 3j \\ WZ + c\bar{c} + \leq 3j \\ WZ + c\bar{c} + \leq 3j \\ Z\gamma + \leq 3j \\ Z\gamma + \leq 3j \end{array}$		

