

Electroweak Radiative Corrections and Future Colliders

1. Why are Electroweak Radiative Corrections important?
2. M_W , $\sin^2 \theta_{eff}$ and M_H
3. Hadron Colliders: Electroweak Radiative Corrections to W and Z boson production
4. Linear Collider: Radiative Corrections to $e^+e^- \rightarrow 4f$ and Bhabha Scattering
5. Conclusions

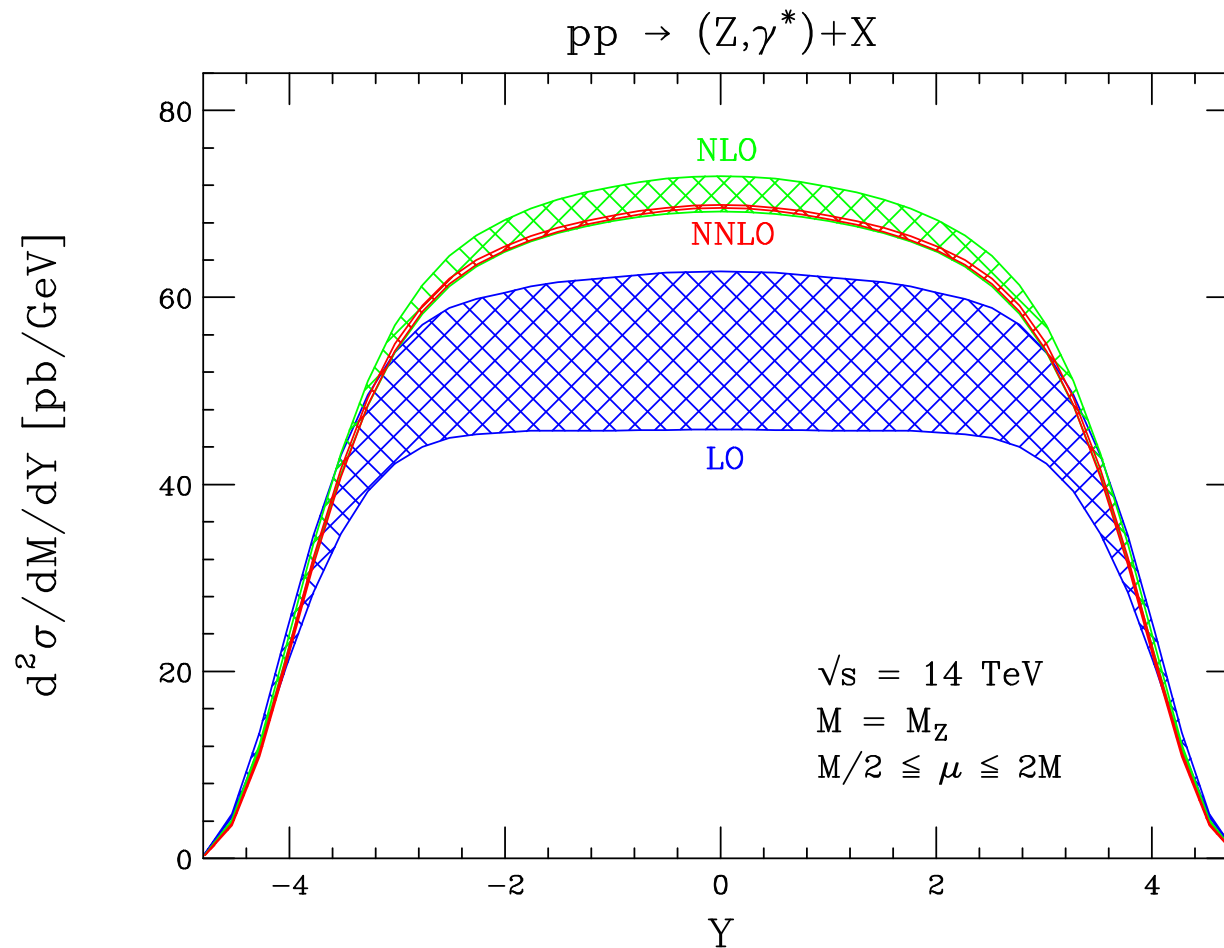
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1 – Why are Electroweak Radiative Corrections important?

- Precise measurements have to be matched by precise theoretical predictions
 - ☞ present and future collider experiments aim at measuring observables (cross section, mass, width,...) at the **% level or better**
 - ☞ need to take into account higher order corrections
- QCD corrections:
 - ☞ NLO: typically **20 – 30%**
 - ☞ NNLO: typically **a few %**
 - ☞ taking into account QCD corrections reduces (sometimes dramatically) the renormalization and factorization scale uncertainty

☞ example: Z boson rapidity distribution (Anastasiou, Dixon, Melnikov, Petriello)



- electroweak radiative corrections:

- ☞ 1-loop: naively of $\mathcal{O}(\alpha) \leq 1\%$

- ☞ why bother?

- possible exceptions:

- ☞ logarithmic enhancement factors

- collinear: $\log(\hat{s}/m_f^2)$,

- Sudakov: $\log(\hat{s}/M_{W/Z}^2)$

- ☞ QCD corrections are small (example: W/Z cross section ratio)

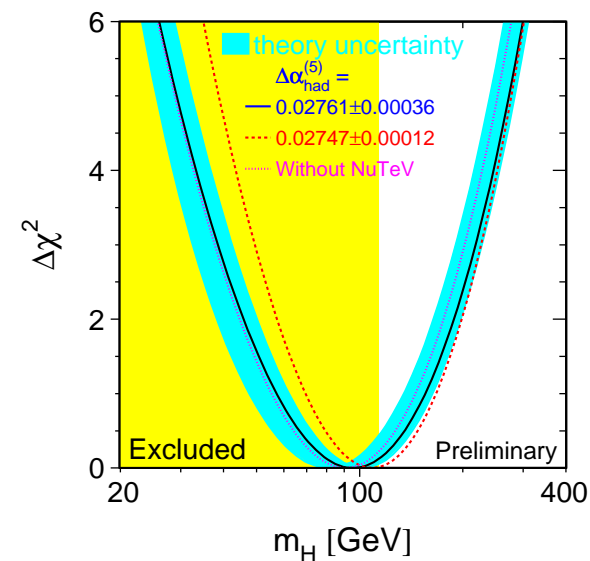
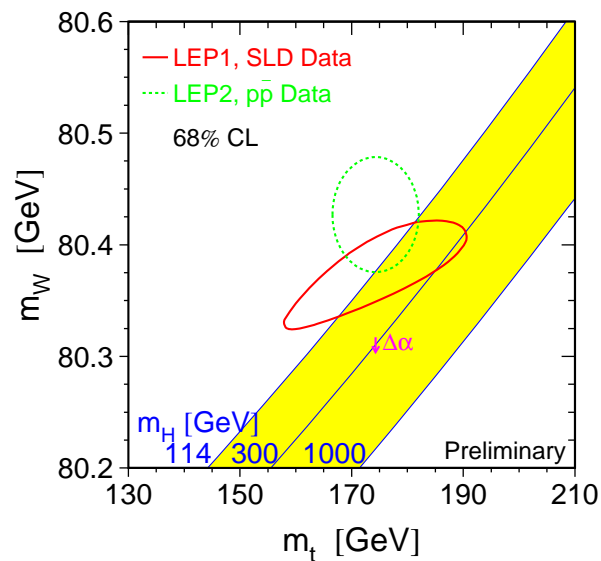
- ☞ and/or very precise measurements ($M_W, \sin^2 \theta_{eff}$)

- in some cases need \geq 2-loop EWK corrections

- ☞ should be able to make use of techniques developed for NNLO QCD corrections

2 – M_W , $\sin^2 \theta_{eff}$ and M_H

- 1-loop corrections to M_W and $\sin^2 \theta_{eff}$ depend **quadratically** on the top quark mass, m_t , and **logarithmically** on M_H
- ☞ measuring M_W ($\sin^2 \theta_{eff}$) and m_t one can extract information on M_H



- fit results depend on

- ☞ experimental uncertainties

- ☞ and theoretical uncertainties

- **primordial** theoretical uncertainties: associated with the extraction of (pseudo)observable from measured quantities

- example:** M_W from transverse mass distribution

- **intrinsic** theoretical uncertainties: from unknown higher order corrections

- example:** “blueband”

Experimental Uncertainties: Looking into the Crystal Ball

	present	Tev. run2	LHC	LC	GigaZ
$\delta \sin^2 \theta_{eff} (\times 10^{-5})$	14	63	14 – 20	6	1.3
δM_W [MeV]	34	27	10 – 15	10	7
δm_t [GeV]	5.1	2.7	1.0	0.2	0.13
$\delta M_H / M_H$ (indirect)	60%	35%	20%	15%	8%

- need intrinsic theoretical uncertainties which are considerably smaller than experimental uncertainties
- estimate size of missing higher order corrections to M_W and $\sin^2 \theta_{eff}$ (Erler)
 - ☞ collect all relevant enhancement and suppression factors
 - ☞ set remaining coefficient (from loop integrals) to unity
 - ☞ choose largest group theory factor

- estimate largest theoretical uncertainties come from

☞ $\mathcal{O}(\alpha^2\alpha_s)$ corrections for M_W (Awramik et al.)

☞ $\mathcal{O}(\alpha^2)$ corrections for $\sin^2\theta_{eff}$

- estimated intrinsic theoretical uncertainty

$$\delta M_W^{th} \approx 4 \text{ MeV} \quad \delta \sin^2\theta_{eff} \approx (6 - 7) \times 10^{-5}$$

- new results on $\mathcal{O}(\alpha^2)$ corrections for $\sin^2\theta_{eff}$ at LoopFest III?

- ultimate goal: bring intrinsic theoretical uncertainties down to

☞ $\mathcal{O}(1 \text{ MeV})$ for M_W

☞ and $\mathcal{O}(\text{few} \times 10^{-6})$ for $\sin^2\theta_{eff}$

→ if we want GigaZ option

- probably need full 3-loop corrections to $\sin^2\theta_{eff}$ and $\mathcal{O}(\alpha^2\alpha_s)$ corrections to M_W

3 – Electroweak Radiative Corrections to W and Z Boson Production

- example for primordial theoretical uncertainties
- for W mass measurement, need radiative corrections for W and Z boson production:
 - ➡ $Z \rightarrow \ell^+ \ell^-$ data constrain lepton scale and resolution
 - ➡ calibrate using using LEP data
 - ➡ need to use the same theoretical input that has been used to extract Z parameters at LEP:
 - ➔ include QED corrections (change the Z mass extracted from data)
 - ➔ include purely weak corrections
 - ➔ include $\mathcal{O}(G_F^2 m_t^2 M_W^2)$ corrections to $\sin^2 \theta_{eff}$

- Treatment of collinear singularities:

- ☞ Final state collinear singularities are regulated by finite lepton masses

- ☞ Initial state collinear singularities are **universal to all orders** and are absorbed into the parton distribution functions (PDF's), in complete analogy to QCD

- for a consistent treatment of the $\mathcal{O}(\alpha)$ initial state corrections, QED corrections should be incorporated into the global fitting of PDF's

- also need QED corrections for all data sets used to fit PDF's

- ☞ Absorbing the collinear singularities into the PDF's introduces a QED factorization scheme dependence

- ☞ **current global fits to the PDF's do not take into account QED corrections**

- ☞ fortunately initial state corrections are small:

- QED corrections to PDF's can be neglected at present stage

- 1-loop EWK corrections shift W and Z masses by $\mathcal{O}(100 \text{ MeV})$

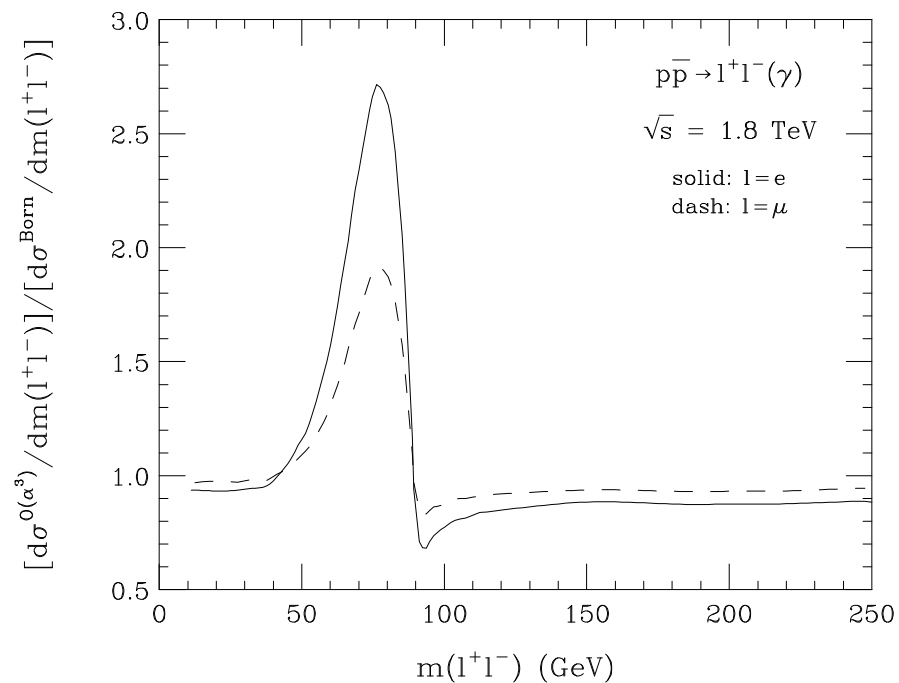
☞ most of the effect comes from final state photon radiation

☞ proportional to

$$\frac{\alpha}{\pi} \log \left(\frac{\hat{s}}{m_\ell^2} \right)$$

→ these terms significantly influence the $\ell^+ \ell^-$ inv. mass distribution

☞ taking only QED corrections into account

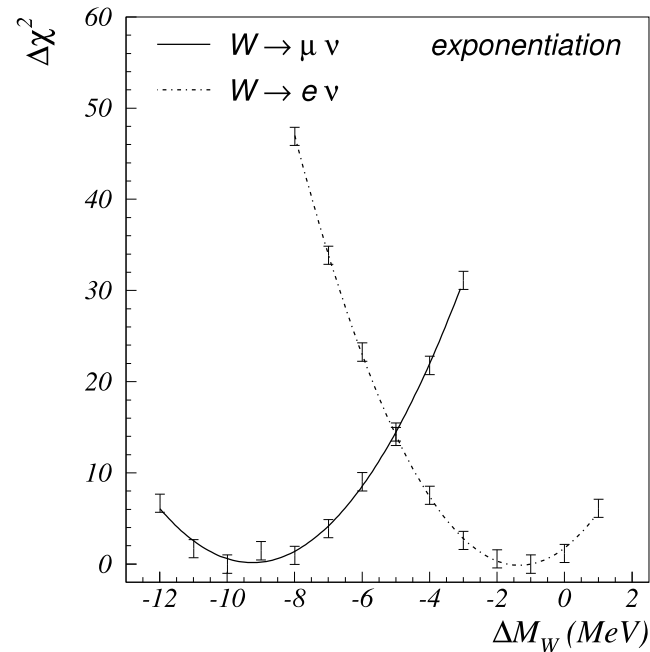
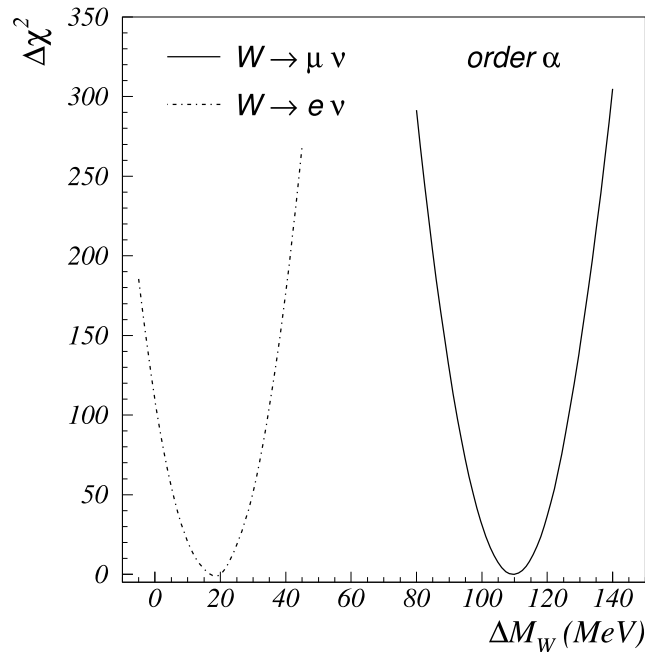


- integrating over $m(\ell\ell)$, the large positive and negative corrections cancel (**KLN theorem**)
- Detector effects may significantly influence the QED corrections:
 - ☞ It is difficult to discriminate electrons and photons which hit the same calorimeter cell
 - recombine e and γ momenta to an effective electron momentum in that case
 - an inclusive quantity is formed
 - the mass singular terms $((\alpha/\pi) \log(\hat{s}/m_\ell^2))$ disappear (**KLN again...**)
 - the effect of the QED corrections is reduced
 - ☞ Muons must be consistent with a minimum ionizing particle
 - require $E_\gamma < 2 \text{ GeV}$ in cell traversed by muon
 - this reduces the hard photon part
 - the mass singular terms survive

- calculations of the complete $\mathcal{O}(\alpha)$ EWK corrections to
 - ☞ $p\bar{p}^{(-)} \rightarrow W^\pm \rightarrow \ell^\pm \nu$ (**Dittmaier+Krämer, UB+Wackerroth**)
 - ☞ and $p\bar{p}^{(-)} \rightarrow \gamma, Z \rightarrow \ell^+ \ell^-$, including $\mathcal{O}(G_F^2 m_t^2 M_W^2)$ corrections to $\sin^2 \theta_{eff}$ (**UB et al.**) exist now
- if final state photon radiation shifts W mass by $\mathcal{O}(100)$ MeV:
 - ☞ need to worry about multiple (final) state photon radiation in W *and* Z production
 - ☞ effect should be more pronounced in Z case since both final state leptons radiate
 - ☞ two photon radiation is known to significantly change the shape of the $m(\ell\ell)$ and M_T distributions (**UB, Stelzer**)

- recent progress in taking multi-photon radiation into account: two approaches
 - ☞ YFS exclusive exponentiation (**Jadach, Placzek**)
 - currently only at parton level and for W decay
 - procedure used is gauge invariant
 - ☞ QED structure function approach (**Montagna et al.**)
 - only final state corrections are presently incorporated
 - procedure used is **not** gauge invariant
 - however, terms violating gauge invariance are numerically small ($< 0.1\%$)
- Montagna et al. calculate shift in M_W using simplified detector model:
 - combine e and γ momenta for $\Delta R(e, \gamma) < 0.2$
 - reject μ events if $E_\gamma > 2$ GeV and $\Delta R(\mu, \gamma) < 0.2$

👉 result:



👉 shift of M_W caused by multi-photon radiation is about **10%** of that caused by one photon radiation

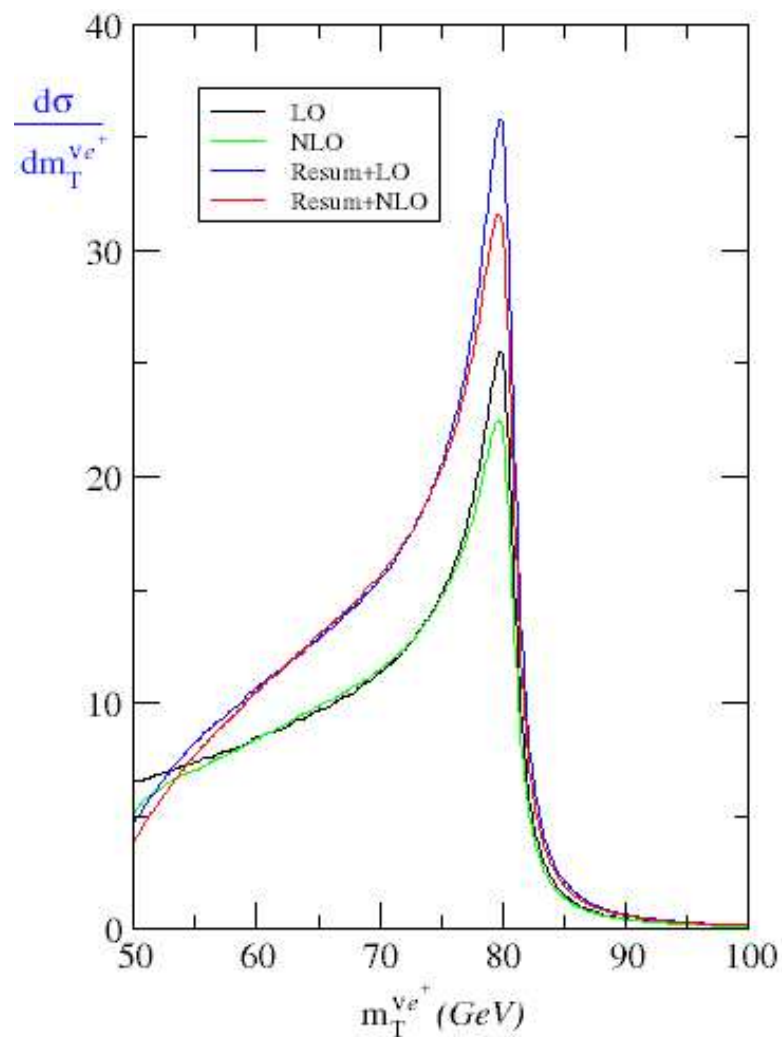
👉 **Note:** absolute value of shift caused by $\mathcal{O}(\alpha)$ corrections smaller than value observed by CDF/DØ, due to simplified detector model

👉 expect larger shifts in Z case (two final state radiators)

- for W mass analysis need calculation of W and Z production including electroweak **and** resummed QCD corrections
 - ☞ need accurate knowledge of W p_T distribution to determine \not{p}_T resolution
 - ☞ \not{p}_T resolution determines how “sharp” the edge in the M_T distribution at $M_T \approx M_W$ is
 - ☞ which in turn determines how well M_W can be measured
- first step towards this lofty goal:
 - ☞ incorporate final state photon radiation effects into RESBOS calculation (**Cao, Yuan**)

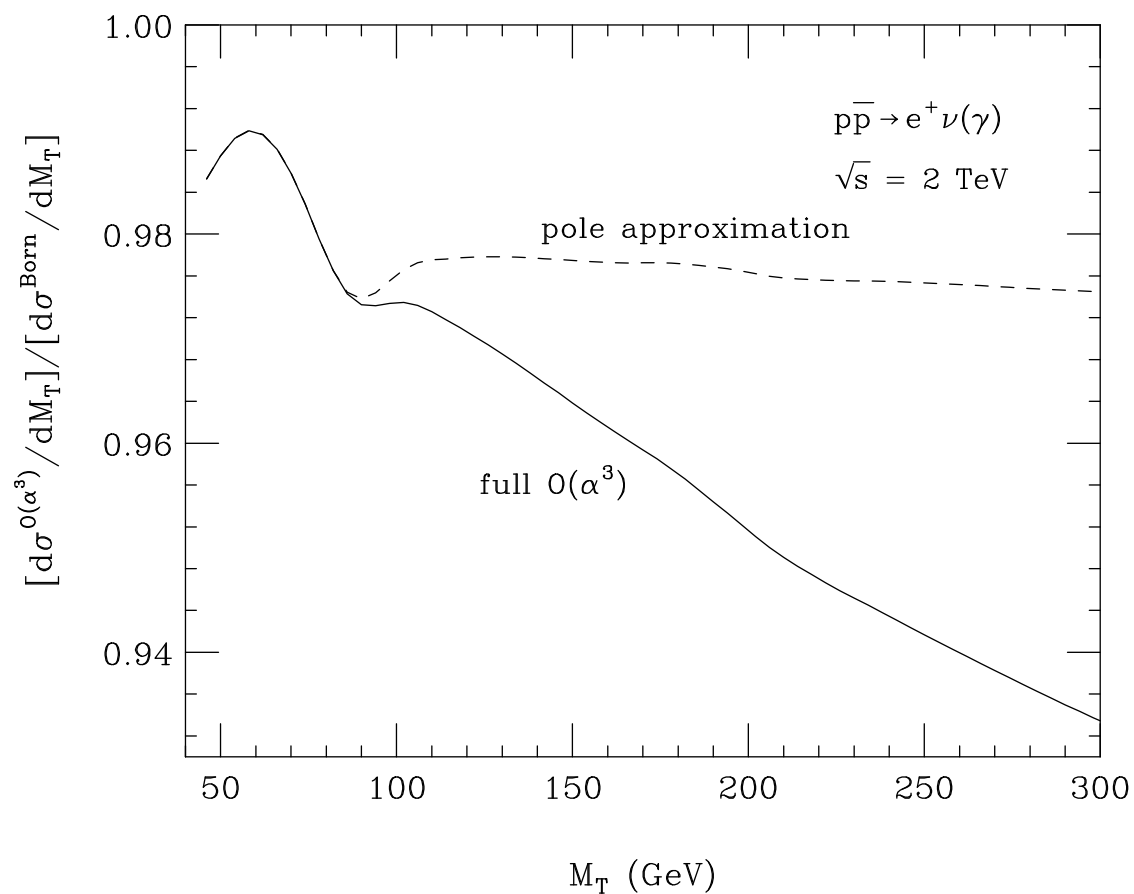
NLO: $\mathcal{O}(\alpha)$ QED final state radiation

Resum: resummed QCD corrections (RESBOS)



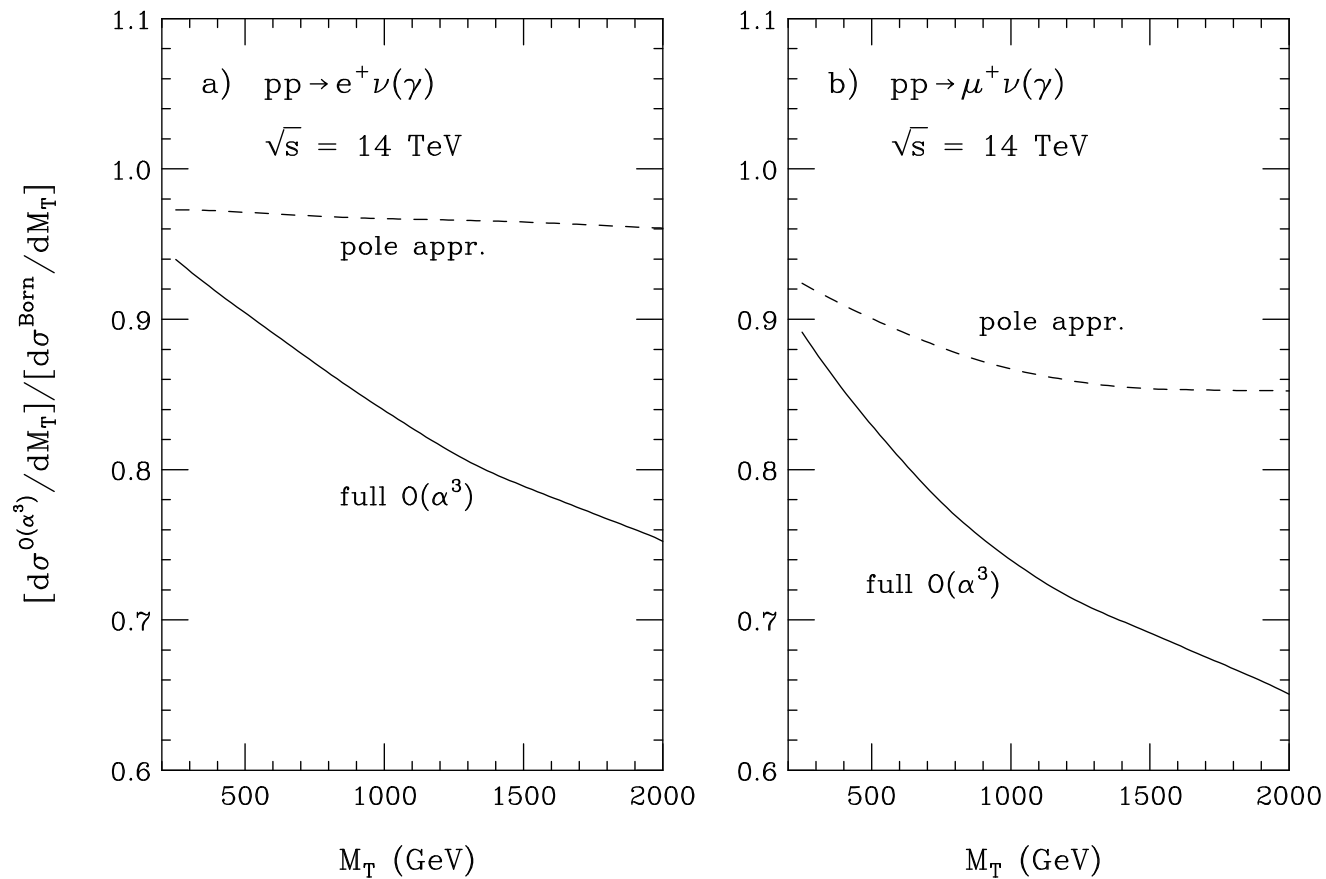
Electroweak Sudakov Logs

- for $\hat{s} \gg M_{W/Z}^2$, the weak corrections become large and negative



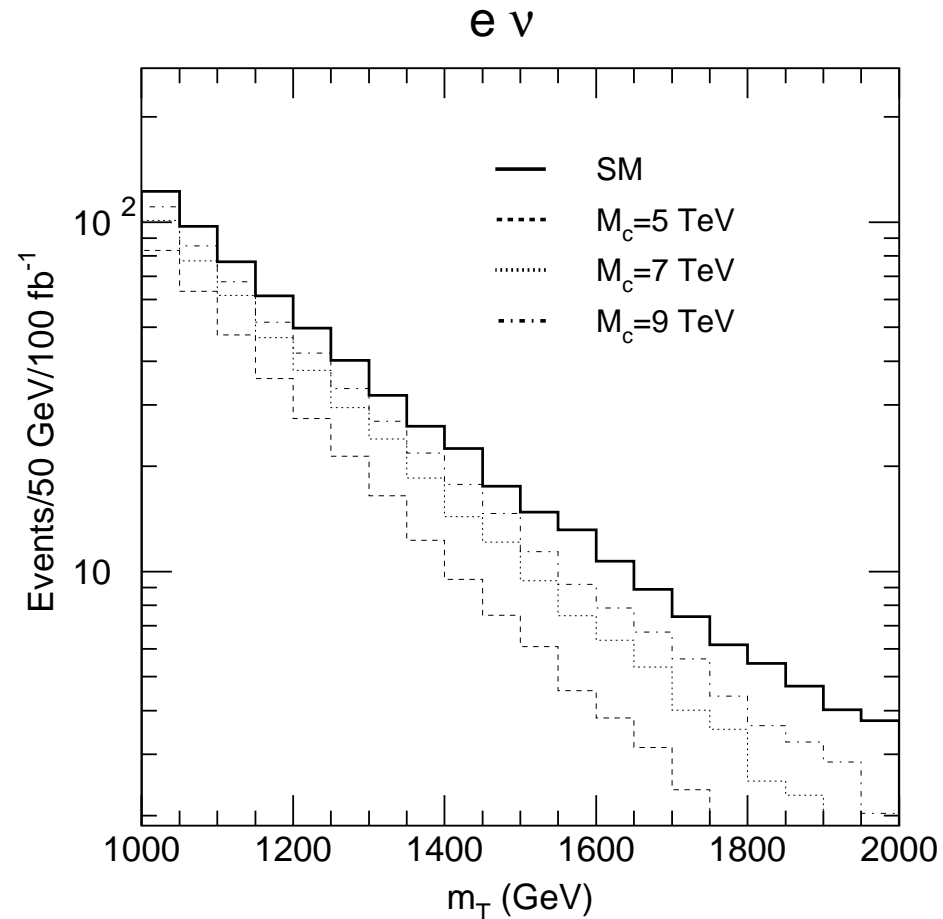
dashed: evaluate weak form factors for $\hat{s} = M_W^2$

- reason: terms $\sim \alpha \log^2(\hat{s}/M_W^2)$ from vertex and box corrections
- ☞ need to resum?
- ☞ certainly for the LHC this is necessary (**not done yet**)



- important for new physics searches:

☞ example: KK excitations of W boson: a slight reduction in cross section could signal a heavy KK excitation beyond reach for direct production (**Polesello, Prata**)



- effect on W width extracted from high M_T tail

☞ the non-resonant weak corrections which contain the Sudakov logs have not been taken into account in previous exp. analyses

☞ they change the shape of the M_T distribution

☞ performing a χ^2 analysis:

non-resonant weak corrections shift W width by

$$\delta\Gamma_W \approx -7.2 \text{ MeV}$$

☞ expected exp. precision in Tevatron run2 (2 fb^{-1} , $e + \mu$, CDF+DØ combined):

$$\Delta\Gamma_W \approx 25 - 30 \text{ MeV}$$

not negligible!

4 – Radiative Corrections to $e^+e^- \rightarrow 4f$ and Bhabha Scattering

- Measuring M_W at Linear Collider (LC):
 - ☞ continuum measurement ($\sqrt{s} > 2M_W$):
 - reconstruct W 's from decay products (similar to method employed by LEP II exps.)
 - expect to achieve $\delta M_W \approx 10$ MeV for $\int \mathcal{L} dt = 500 \text{ fb}^{-1}$ (uncertainty dominated by systematic uncertainty)
 - ☞ threshold scan: $\sqrt{s} \approx 161$ GeV (Wilson, Sitges Workshop)
 $e^+e^- \rightarrow 4$ fermion cross section is sensitive to M_W in threshold region

- the threshold scan under the magnifying glass:

👉 statistical uncertainty: **(Stirling)**

$$\delta M_W^{stat} = 90 \text{ MeV} \left[\frac{\epsilon \int \mathcal{L} dt}{100 \text{ pb}^{-1}} \right]^{-1/2}$$

for $\epsilon = 0.67$ (efficiency) and $\int \mathcal{L} dt = 100 \text{ fb}^{-1}$:

$$\delta M_W^{stat} \approx 3.5 \text{ MeV}$$

👉 add systematic errors

For a multiplicative factor C :

$$\delta M_W^{sys} = 17 \text{ MeV} \left[\frac{\Delta C}{C} \times 100\% \right]$$

assume $\Delta\epsilon \approx 0.25\%$, $\Delta\mathcal{L} \approx 0.1\%$:

$$\delta M_W \approx 6 \text{ MeV}$$

👉 detailed simulations yield $\delta M_W \approx 7 \text{ MeV}$ (**Mönig**)

- theoretical uncertainties:

- ☞ if one wishes to achieve $\delta M_W \approx 7$ MeV, one needs $\delta M_W^{theor} \sim 1$ MeV

- ☞ need to know cross section in threshold region with

$$\frac{\Delta\sigma}{\sigma} \approx 0.05\%$$

- ☞ present situation: only calculation valid in threshold region: GENTLE, includes full (improved) Born $e^+e^- \rightarrow 4$ fermion cross section, including non-resonant graphs, finite W width, Coulomb corrections and ISR effects

- ☞ uncertainty of GENTLE cross section in threshold region (**CERN LEP2 Yellow Report**):

$$\frac{\Delta\sigma}{\sigma} \approx 1.4\%$$

- ☞ need full $\mathcal{O}(\alpha)$ corrections in threshold region

- ☞ finite W width effects are important in threshold region:

- ➔ must go beyond double pole approximation

- ☞ current status: $\mathcal{O}(\alpha)$ corrections to $e^+e^- \rightarrow 4$ fermions known in double pole approximation (RACOONWW and YFSWW3)
- ☞ full $\mathcal{O}(\alpha)$ calculation extremely problematic, results out of sight
 - many channels contribute
 - $\mathcal{O}(1000-10000)$ Feynman diagrams per channel (Vicini, Jegerlehner)
 - technical problems with 4, 5 and 6-point functions and phase space (lots of peaks)
 - no practicable solution of gauge invariance problem associated with finite width and radiative corrections
- ☞ If theoretical uncertainties of cross section do not improve:

$$\delta M_W \approx \delta M_W^{theor} \approx 24 \text{ MeV}$$

sic transit gloria mundi...

- recall: need to measure luminosity to 0.1% to be able to achieve $\delta M_W = 7 \text{ MeV}$ at GigaZ
 - ☞ an accurate determination of the luminosity is important for many measurements at a LC
- LEP and SLC: used small angle Bhabha scattering, $e^+e^- \rightarrow e^+e^-$, to measure luminosity
 - ☞ has large cross section
 - ☞ accurately calculable (QED + contributions from hadronic vacuum polarization)
- due to beamstrahlung, resulting in variable e^+e^- energy, small angle Bhabha scattering cannot be used to determine luminosity at a LC
 - ☞ use large angle Bhabha scattering
 - ☞ measure acollinearity angle, caused by the asymmetric collision of electrons and positrons (Frary, Miller)

- to achieve desired precision need
 - ☞ 2-loop QED corrections
 - ☞ need to include Z exchange diagrams (log enhanced: $\log(s/m_e^2)$ and $\log(s/m_Z^2)$)
- 2-loop QED matrix elements for Bhabha scattering exist
(Bern, Ghinkulov, Dixon)
- still need to construct numerically stable code for performing singular phase space integrations
 - ☞ similar to NNLO QCD predictions

5 – Conclusions

- controlling electroweak radiative corrections is essential for future high precision tests of the SM
- significant progress has been made over the last few years
- a long shopping list of things to do remains:
 - ➡ higher order corrections to M_W and $\sin^2 \theta_{eff}$
 - ➡ higher order EWK corrections to W and Z production in the pole region
 - ➡ resummation of EWK Sudakov logs
 - ➡ full 1-loop corrections to $e^+e^- \rightarrow 4$ fermions
 - ➡