Critical events and mergers

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Outline

- I. Large scale structures
- II. Critical event theory
- III. Attraction cones
- IV. Merger characteristics
- V. Conclusions and perspectives

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Large scale structures

Web like structure





~10 Mpc

The NewHorizon Simulation



~10 Mpc

The NewHorizon Simulation

Large scale structures

The cosmic web has an impact on galaxy morphology





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From above





The New Horizon Simulation

Large scale structures



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Can we predict the main characteristics of the universe directly from its initial conditions? ?



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Can we make statistical predictions based solely on the initial conditions?



12 Critical event theory









The cosmic web as critical points

At every point of the field, we take the average density in a sphere of increasing radius

This amounts to smoothing the field with a **top-hat filter**.



We may link the critical points of the field (minima, maxima, saddle points) to the particular points of the cosmic web (voids, nodes, walls and filaments)

The spherical collapse model gives a correspondance betweem the **density/smoothing** and the **cosmic time/mass** of objects.

In particular, the **smoothing radius** monotonically increases with time at the peaks, and is thus a **proxy** for time.

Mergers and critical events







Critical event theory

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Replacing the top-hat filter with a Gaussian filter, we may get analytical expressions for statistical properties of events.

22 Critical event theory



In one ...

... and two dimensions

Critical event theory 23



How can we recover the pair of peaks involved in a merging event?

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Attraction cones

We use the **distance from the critical points** to decide in which objet a critical event merges.





27 Attraction cones

What cone opening angle should we consider ?

To answer this question, we analyse the **mixed two point correlation function** between critical events and peaks.



Cone angle equal to 2

Topological mergers in 1D

In one dimension, the peak associated to the merging of a critical event is the **adjacent peak**.



Topological mergers in 1D



position

Topological mergers in 1D



31 Attraction cones

Topological mergers in 2D (and 3D), using the skeleton



Comparison with cones (in 1D)

Mean difference between number of topological mergers and number of mergers in a cone of angle β



Can we use critical event theory to make pertinent cosmological predictions?

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Mass acquired by mergers



The morphology of a galaxy is linked to its **merger history**, as well as the **characteristics** of these mergers (mass, number, time).

We use the spherical collapse model to assign a coherent mass to mergers associated to a critical event.

Mass accreted by mergers

Mass is proportional to the cube of the smoothing radius

$$M = \bar{\rho} \times V \propto \bar{\rho} \times R^3$$

The existence time of an object is a decreasing function of the overdensity of the peak that is associated to it.



We wish to compare the mass of two merging objects at a given time, taking care to not count mass twice.





position





$M = \bar{\rho} \times V \propto \bar{\rho} \times R^3 \quad \text{ In 3D}$





Density (mean)

First comparison to cosmological simulations



Simulation cosmologique

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Conclusions

Critical event theory provides a useful framework to make predictions about the geometry of mergers from first principle.

Perspectives

• Filament-filament and wall-wall mergers





 Proxy for angular momentum associated to mergers by parametrising by the direction of the vanishing eigenvalue of the Hessian

 Make merger rate evolution predictions with respect to the power spectrum



Thank you!

Étude des paramètres spéctraux des simulations 46

0.8

0.6

0.4

0.2

0.9

0.8

0.7

0.6





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Fonctions de correlation à deux points mixtes pic/événement



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