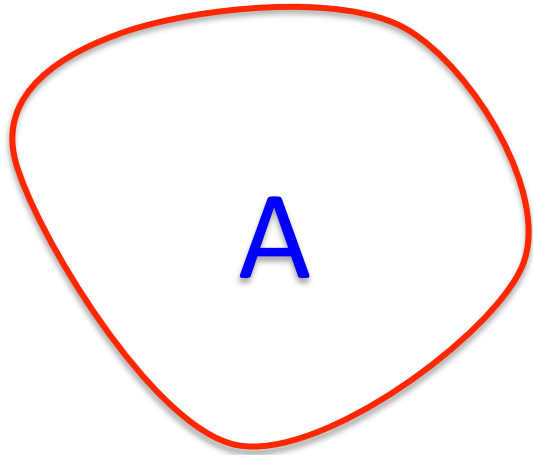


Propagation of entanglement in many-body systems

Hong Liu



Entanglement generation



Consider a many-body system in an **un-entangled** initial state:

How fast can **entanglement** be generated?

A measure: $\frac{dS_A}{dt}$

depends on size of A, total number of d.o.f.,

Not meaningful to compare it across different systems

Relativistic systems: should be constrained by causality, how?

Plan

1. Insights from holography : hints of a measure

HL and J. Suh, Phys. Rev. Lett. 112, 011601 (2014)

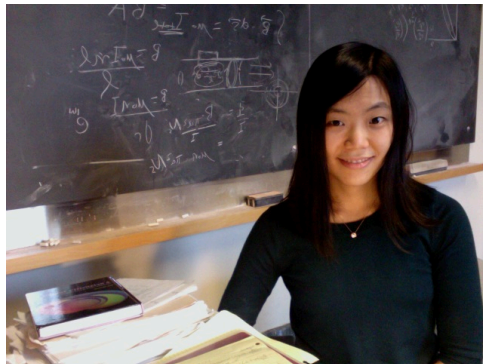
HL and J. Suh, Phys. Rev. D 89, 066012 (2014)

2. An upper bound on entanglement growth in non-interacting systems

3. A model of entanglement growth in interacting systems

H. Casini, HL,
M. Mezei, to appear

Thanks to Hubeny and Suh



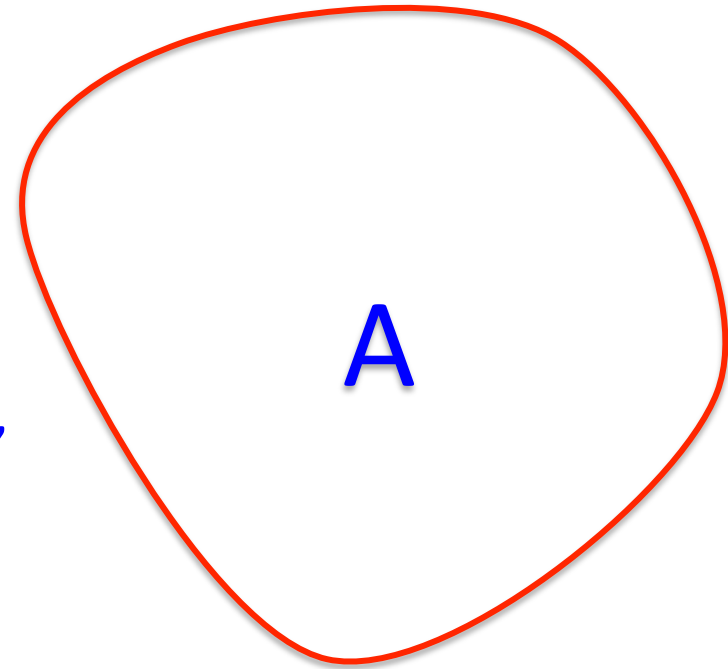
Insights from Holography

Global quenches

1. Start with a QFT in the **ground** state.
2. At $t=0$ in a **very short time** inject a **uniform** energy density
 - initial state **homogeneous, isotropic, entanglement properties as vacuum**
3. The system evolves to **(thermal) equilibrium**

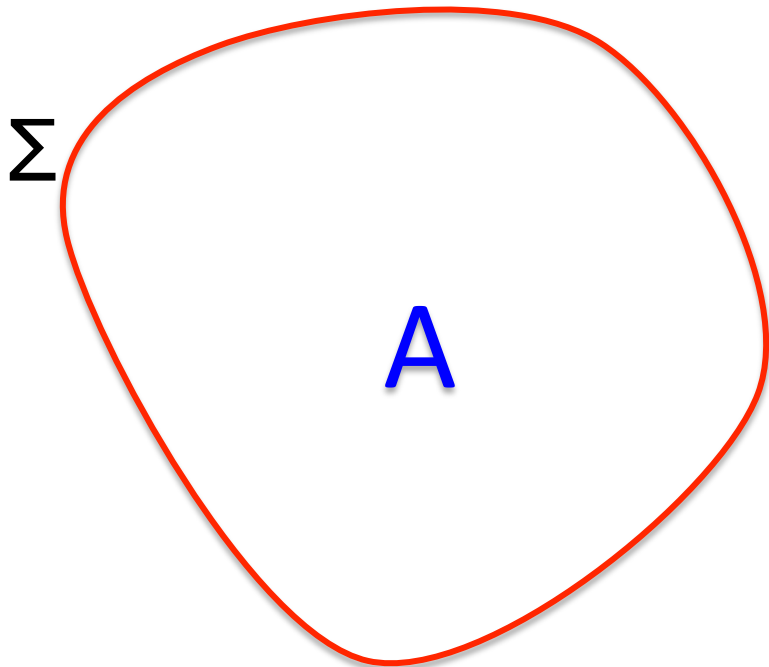
The system is in a **pure state** throughout.

$$S_A(t)?$$



$$R \gg \frac{1}{T}$$

t=0



$$S_0 = \frac{A_\Sigma}{\delta^{d-1}} + \dots$$

Typical point of A
essentially **un-entangled**
with outside

Equilibrium



$$S_{\text{eq}} = s_{\text{eq}} V_A + \dots$$

s_{eq} : equilibrium entropy density

Essentially **every point** of
A is **entangled** with outside

Full time evolution: very difficult question

$d=2$: CFTs (Calabrese-Cardy), **Linear growth**

$d > 2$: holography (for a class of strongly interaction systems)

Hubeny, Rangamani, Takayanagi: arXiv:0705.0016

Related work:

Abajo-Arrastia, Aparicio and Lopez, arXiv:1006.4090

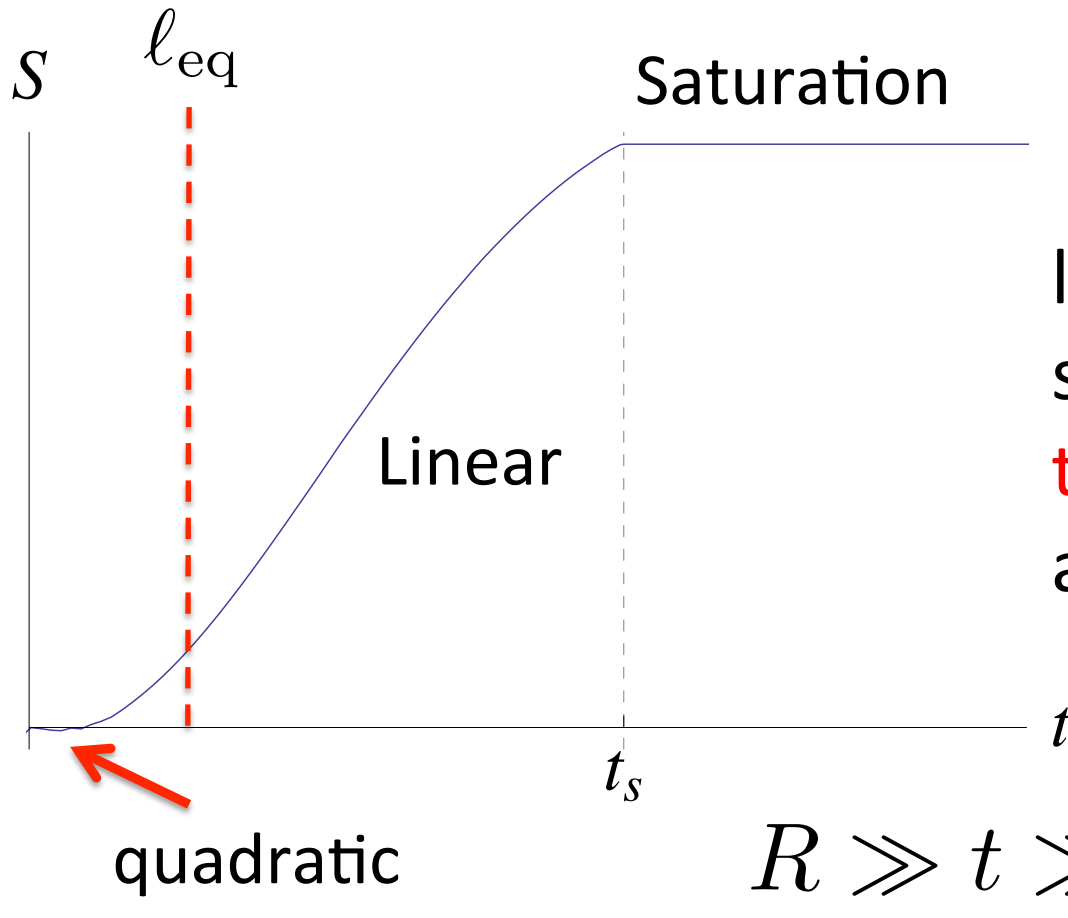
Albash and Johnson, arXiv:1008.3027

Balasubramanian, Bernamonti, de Boer, Copland, Craps, Keski-Vakkuri, Muller, Schafer, Shigemori, Staessens arXiv:1012.4753, arXiv:1103.2683

Hartman and Maldacena arXiv:1303.1080

Hubeny and Maxfield arXiv: 1312.6887

.....



$$l_{eq} \sim \frac{1}{T}$$

local **equilibration** time scale after which **thermodynamics** applies **locally**.

$$R \gg t \gg l_{eq}$$

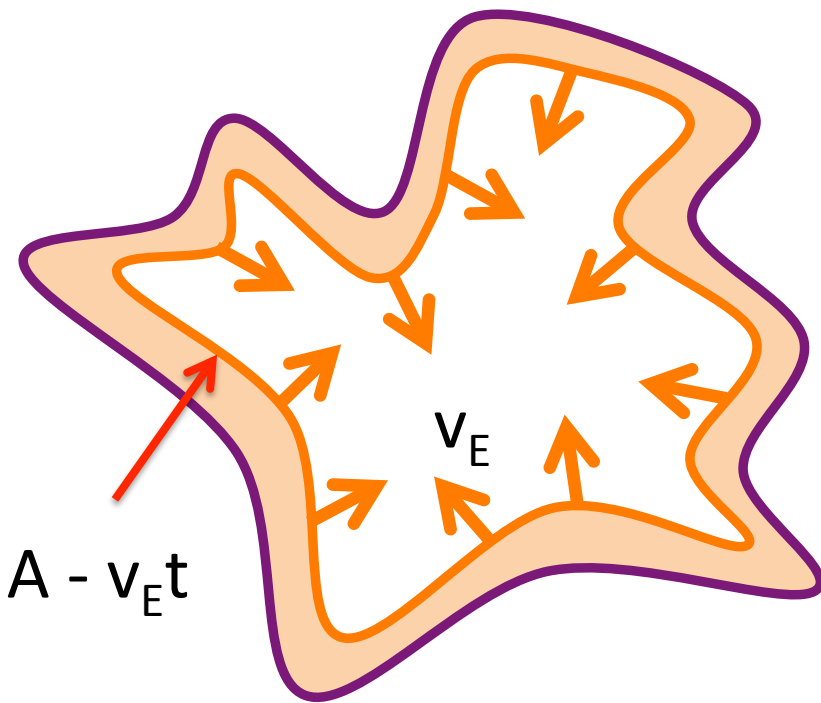
$$\Delta S_A(t) = v_E s_{eq} A_\Sigma t + \dots$$

“universal”

v_E : dimension of velocity, characterized by **final eq state**.

Entanglement Tsunami

$$\Delta S_A(t) = v_E s_{\text{eq}} A_\Sigma t = s_{\text{eq}} (V_A - V_{A-v_E t})$$



suggests a picture of
“tsunami” wave of
entanglement, moving
inward from boundary

d.o.f. in the region covered
by the wave is now entangled
with those outside A

natural with evolution from a local Hamiltonian

Tsunami velocity

$$\Delta S_A(t) = v_E s_{\text{eq}} A_\Sigma t + \dots$$

From gravity:

$$v_E \leq v_E^{(S)} = \frac{(\eta - 1)^{\frac{1}{2}}(\eta - 1)}{\eta^{\frac{1}{2}}\eta} = \begin{cases} 1 & d = 2 \\ \frac{\sqrt{3}}{2^{\frac{4}{3}}} = 0.687 & d = 3 \\ \frac{\sqrt{2}}{3^{\frac{3}{4}}} = 0.620 & d = 4 \\ \frac{1}{2} & d = \infty \end{cases}$$
$$\eta \equiv \frac{2(d - 1)}{d}$$

d=2: agree with previous Calabrese-Cardy's result

A measure of entanglement growth

$$\mathfrak{R}_A(t) \equiv \frac{1}{s_{\text{eq}} A_\Sigma} \frac{dS_A}{dt} \quad (\text{dimension: velocity})$$

can be compared among regions of different sizes,
and systems of different number of d.o.f.

From gravity: **after local equilibration**

$$\mathfrak{R}_A \leq v_E^{(S)}$$

Questions

Generality of linear growth?

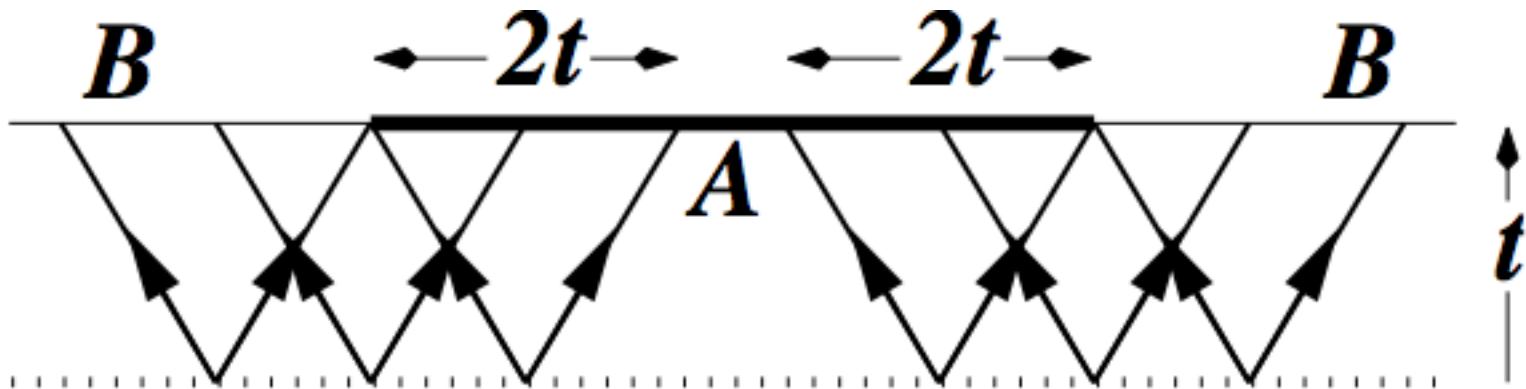
How to relate \mathfrak{R}_A, v_E directly to speed of light?

$$\text{Significance of } v_E^{(S)} = \begin{cases} 1 & d = 2 \\ \frac{\sqrt{3}}{2^{\frac{4}{3}}} = 0.687 & d = 3 \\ \frac{\sqrt{2}}{3^{\frac{3}{4}}} = 0.620 & d = 4 \\ \frac{1}{2} & d = \infty \end{cases} \quad ?$$

Free theory? Not available

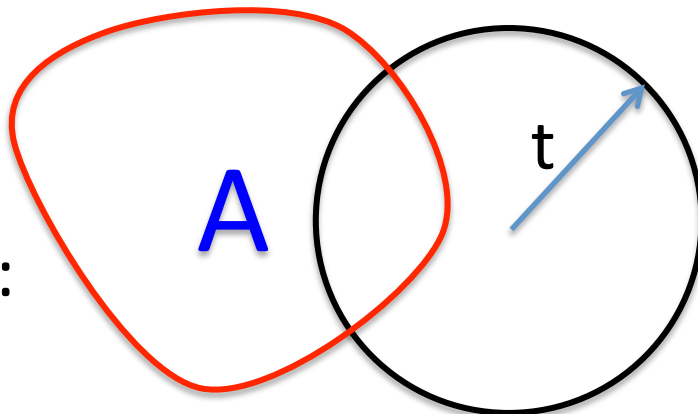
Calabrese-Cardy model

Energy injection from quench creates a finite density of **EPR pairs**, subsequently travel **freely at the speed of light isotropically**.



$d=2$: leads to **linear growth** with $v_E = 1$

Higher
Dimensions:



Entanglement spread
will now depend on
entanglement pattern
on the light cone.

An upper bound for free propagation
of entanglement

Setup

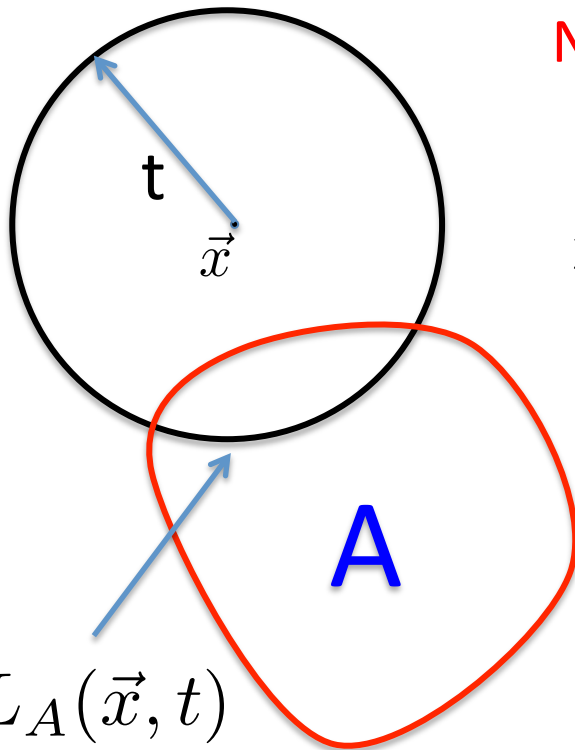
Each point is an independent source of local entanglement which subsequently spread at speed of light.

No interaction/interference among lightcones

For a region **B on the lightcone** from a point x , associate an entanglement measure $\mu[B]$: **entanglement entropy** for B in the Hilbert space of the Light cone from x

Contribution from x : $\mu[L_A(\vec{x}, t)]$

$$S_A(t) = \int d^{d-1}x \mu[L_A(\vec{x}, t)]$$



$L_A(\vec{x}, t)$

(intersection of lightcone
from x with A at time t)

Properties

$\mu[B]$ should have all the properties of entanglement entropy:

$$\mu[B] = \mu[\bar{B}], \quad \text{Strong subadditivity condition, etc.}$$

It **does not change with time** for B with fixed angular extension.

$$\lim_{B \rightarrow 0} \mu[B] = s \xi_B \quad \xi_B : \text{normalized volume for B} \quad \text{e.g. Page (1992)}$$

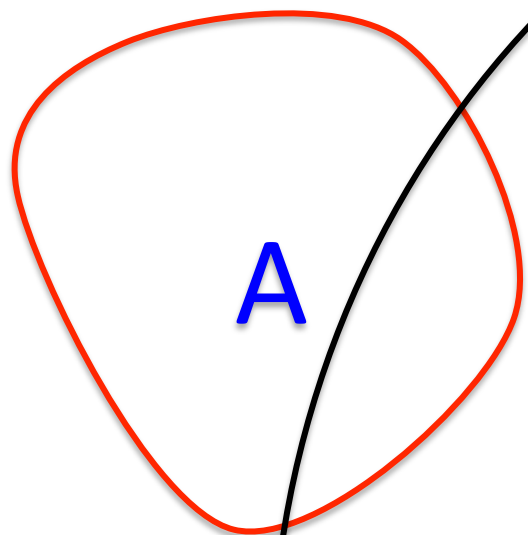
Equilibrium value:

$$S_A(t = \infty) = sV_A$$

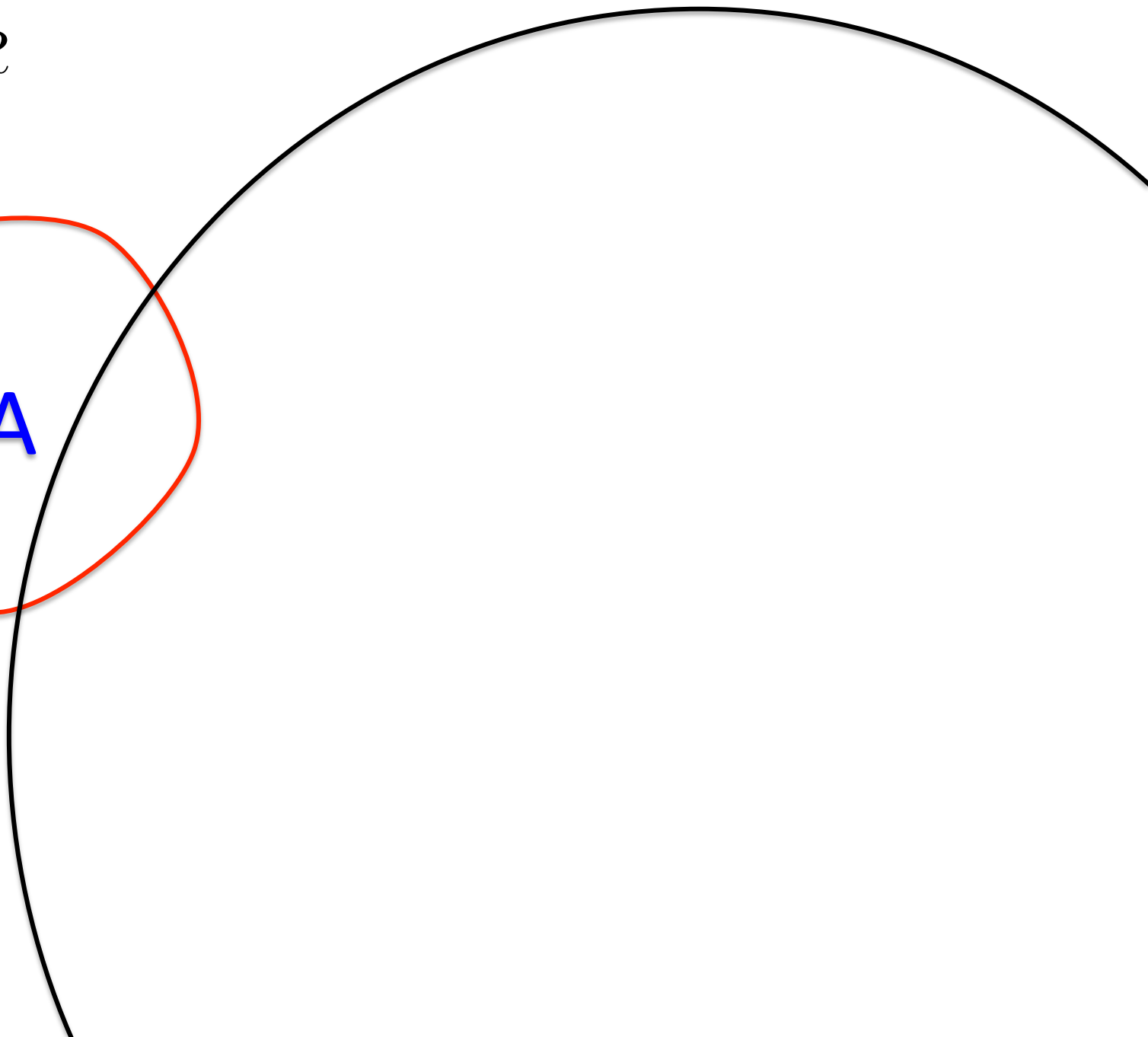


$$s_{\text{eq}} = s$$

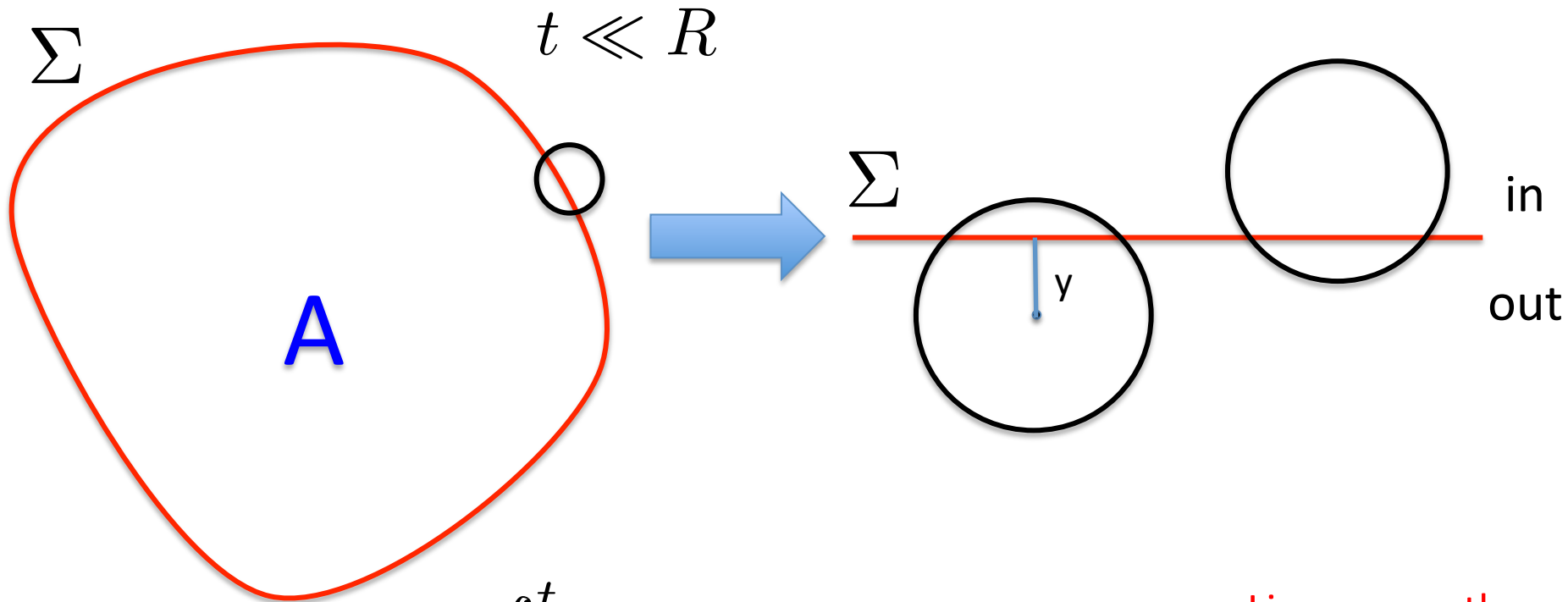
$t \gg R$



A



Linear growth



$$S_A(t) = 2A_\Sigma \int_0^t dy \mu[\text{cap}(y/t)] \propto t$$
$$v_E = \frac{2}{s} \int_0^1 dx \mu[\text{cap}(x)]$$

Linear growth
due to time
independence of
 μ

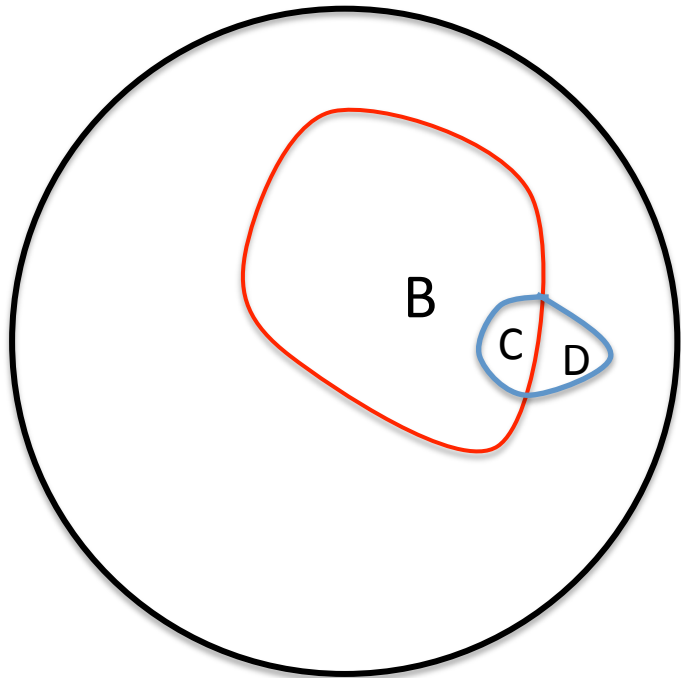
Upper bound on entanglement propagation

Random pure state measure: $\mu_R[B] \equiv s \min(\xi_B, \xi_{\bar{B}})$

Strong sub-additivity condition: $\mu[B] \leq \mu_R[B]$

$$\mu[B] + \mu[C \cup D] \geq \mu[C] + \mu[B \cup D]$$

(C,D infinitesimal)



$$\mu[B \cup D] - \mu[B] \leq s \xi_D$$

$$v_E \leq v_E^{\text{free}} \equiv 2 \int_0^1 dx \xi_{\text{cap}}(x)$$

$$\mathfrak{R}_A(t) \leq v_E^{\text{free}}$$

Free propagation

$$\frac{dS_A}{dt} \leq v_E^{\text{free}} s_{\text{eq}} A_\Sigma$$

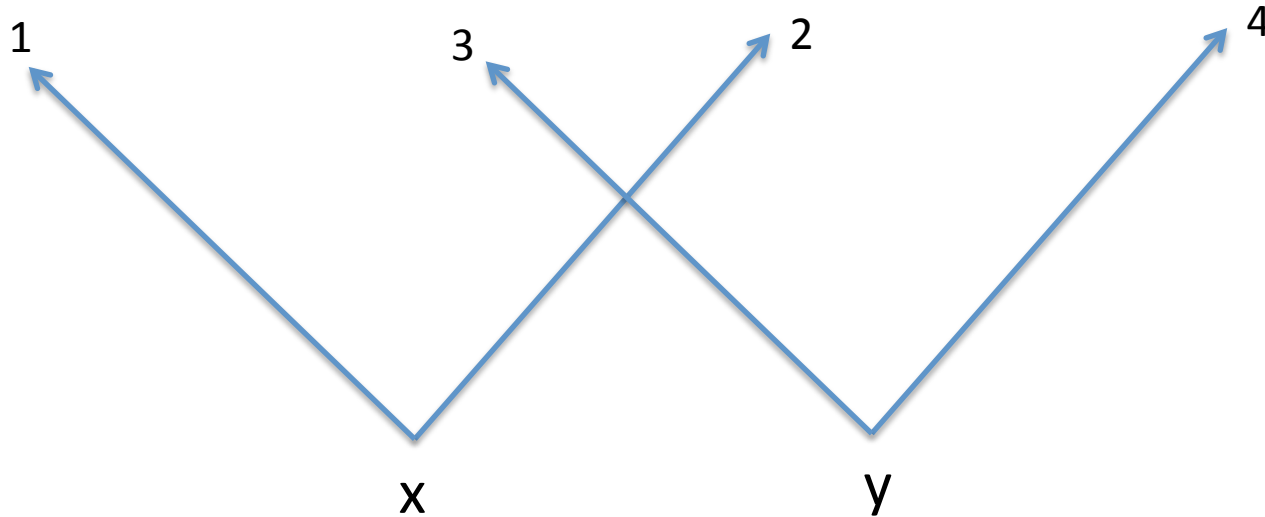
$$v_E^{\text{free}} = \frac{\Gamma(\frac{d-1}{2})}{\sqrt{\pi}\Gamma(\frac{d}{2})} = \begin{cases} 1 & d = 2 \\ \frac{2}{\pi} = 0.637 & d = 3 \\ \frac{1}{2} & d = 4 \\ \sqrt{\frac{2}{\pi d}} & d = \infty \end{cases}$$

$$v_E^{(S)} = \begin{cases} 1 & d = 2 \\ \frac{\sqrt{3}}{2^{\frac{4}{3}}} = 0.687 & d = 3 \\ \frac{\sqrt{2}}{3^{\frac{3}{4}}} = 0.620 & d = 4 \\ \frac{1}{2} & d = \infty \end{cases}$$

In strongly coupled systems, entanglement propagates **faster** than that from **free particles** at speed of light !

An interacting model

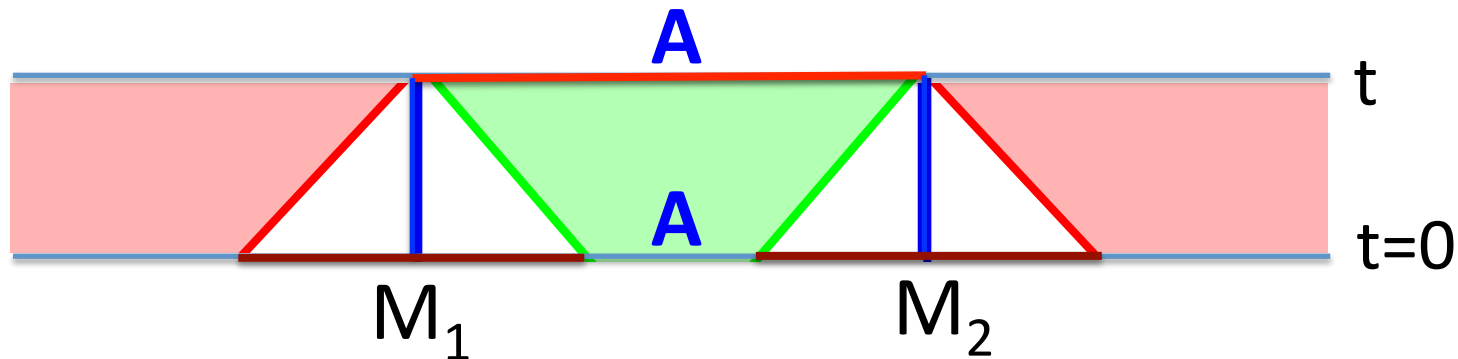
Scattering



Quantum state of the system can **no longer** be described as a **direct product** of those resulting from each point at $t=0$.

We then face the standard difficulties of how to characterize the quantum state of an interacting many-body system.

Domain of dependence



Red-shaded region: $\mathcal{D}_-(\bar{A})$ (past domain dependence of \bar{A})

Scatterings in this region
amounts to unitary
transformations in $\mathcal{H}_{\bar{A}}$

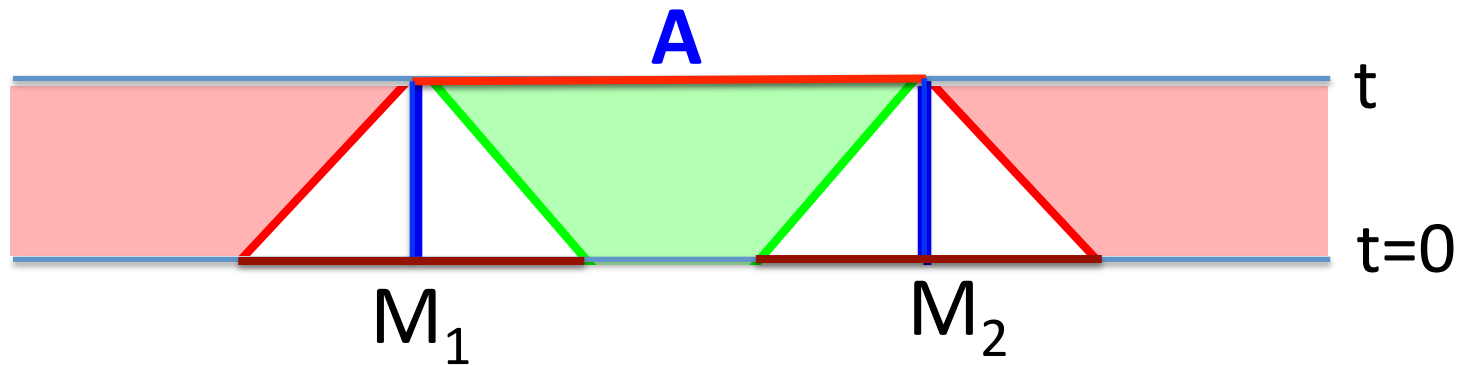
Will not affect $S_A(t)$

Green-shaded region: $\mathcal{D}_-(A)$

Scatterings in this region
amounts to unitary
transformations in \mathcal{H}_A

Will not affect $S_A(t)$ either

Only particles from M_1, M_2 and scatterings in white regions relevant!



Only particles from M_1 , M_2 and scatterings in white regions relevant!

A particle from M_1 and a particle from M_2 do not have effective scatterings. So M_1 and M_2 can be treated **independently**.

In a **strongly** coupled theory particles within M_1 scatter with one another **many times** before reaching A.

Appears natural to apply **random pure state measure** to the **full Hilbert space of all particles in M_1** (similarly with M_2):

$$S_A = s \min(N_A(t), N_{\bar{A}}(t)) \quad N_A: \text{number of particles from } M_1 \text{ falling in } A$$

General formulation

$$M(t) \equiv \mathcal{M} - (\mathcal{D}_-(A) \cap \mathcal{M}) - (\mathcal{D}_-(\bar{A}) \cap \mathcal{M})$$

\mathcal{M} : spatial manifold at $t=0$

$$M(t) = \sum_i M_i$$

$$S_A(t) = \nu_{\text{eq}} \sum_i \min \left(\int_{M_i(t)} n_A(x, t), \int_{M_i(t)} n_{\bar{A}}(x, t) \right)$$

Always **larger** than free propagation results derived earlier.
Likely an **upper** limit for interacting theories.

Results

$d=2$:

One interval: $v_E = 1$

Two intervals: precisely recover holographic results

Free propagation: not

Asplund and Bernamonti
Leichenauer and Moosa

Three intervals: generally same, but can be **larger than**
holographic results for certain time intervals

Appear to be the same as a recent proposal of
Leichenauer and Moosa in 1505.04225.

$d > 2$: $v_E = 1$

However, this might be an unachievable upper bound.

Thank You