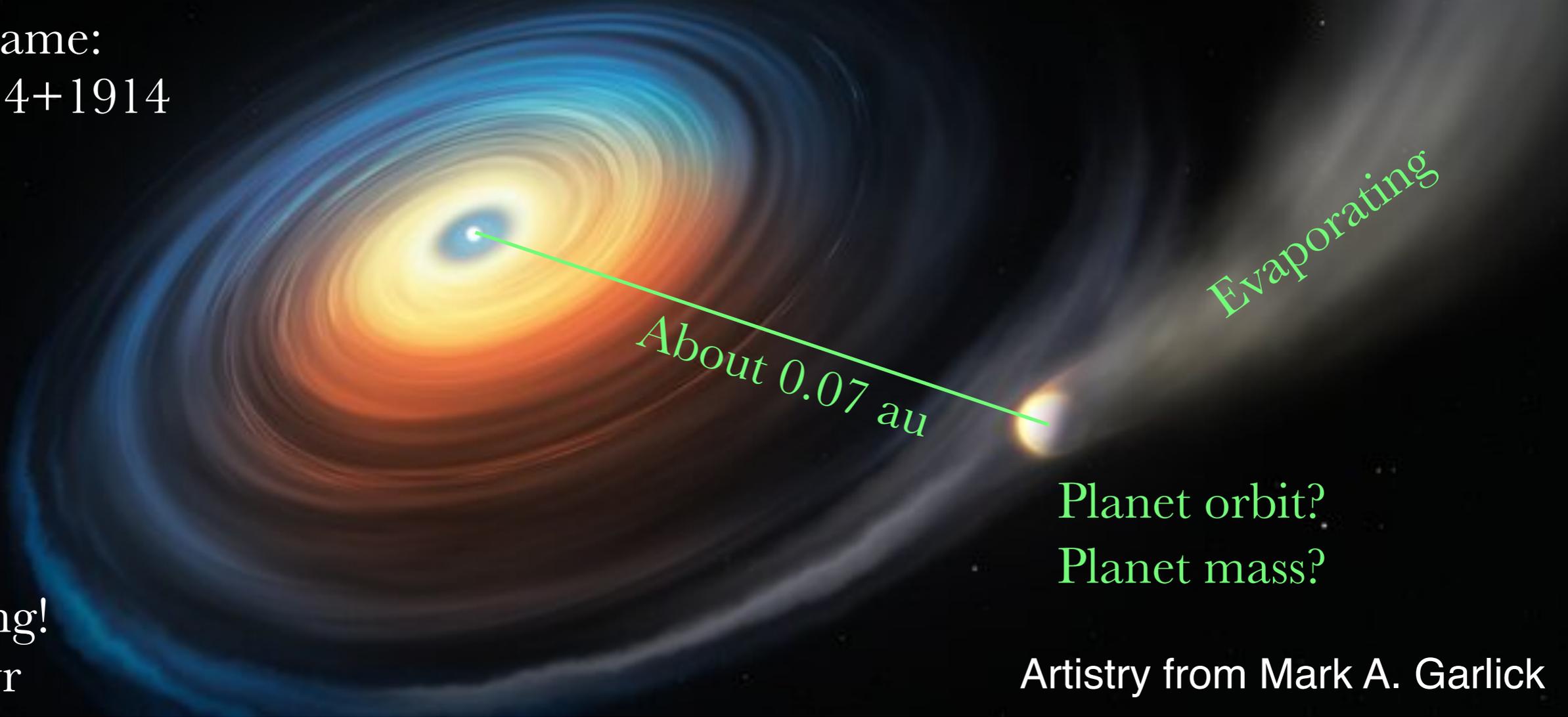


Star name:
WD J0914+1914



Very young!
13.3 Myr

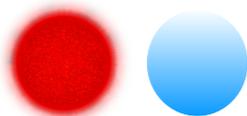
Artistry from Mark A. Garlick

Planets around white dwarfs: physical and dynamical constraints



Dimitri Veras
(University of Warwick)

Results from *Exostar*



THE ASTROPHYSICAL JOURNAL, 885:31 (12pp), 2019 November 1
TESS Asteroseismology of the Known Red-giant Host Stars HD 212771 and HD 203949

Tiago L. Campante^{1,2,3} , Enrico Corsaro⁴ , Mikkel N. Lund^{3,5} , Benoît Mosser⁶ , Aldo Serenelli^{3,7,8} ,
Dimitri Veras^{3,9,10,35} , Vardan Adibekyan¹ , H. M. Antia¹¹ , Warrick Ball^{5,12} , Sarbani Basu¹³ ,
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Maria Tsantaki¹ , and Margaret C. Turnbull³⁴

MNRAS 489, 2941–2953 (2019)

Tidal circularization of gaseous planets orbiting white dwarfs

Dimitri Veras ^{1,2}★† and Jim Fuller³

MNRAS 488, 153–163 (2019)

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Dimitri Veras ^{1,2}★† and Alexander Wolszczan^{3,4}

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Constraining planet formation around 6–8 M_⊙ stars

Dimitri Veras ^{1,2}★†, Pier-Emmanuel Tremblay,² J. J. Hermes,³
Catriona H. McDonald ,^{1,2} Grant M. Kennedy ,^{1,2}† Farzana Meru ^{1,2}§
and Boris T. Gänsicke ^{1,2}

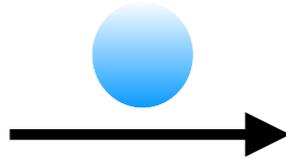
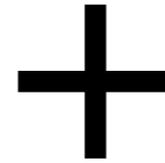
Synergy with Gänsicke et al. planet



Nature 576, 61–64(2019)

Accretion of a giant planet onto a white dwarf star

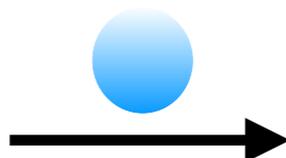
Boris T. Gänsicke^{1,2*}, Matthias R. Schreiber³, Odette Toloza¹, Nicola P. Gentile Fusillo¹, Detlev Koester⁴ & Christopher J. Manser¹



MNRAS 489, 2941–2953 (2019)

Tidal circularization of gaseous planets orbiting white dwarfs

Dimitri Veras ^{ID}1,2★† and Jim Fuller³



MNRAS 492, 6059–6066 (2020)

The dynamical history of the evaporating or disrupted ice giant planet around white dwarf WD J0914+1914

Dimitri Veras ^{ID}1,2★† and Jim Fuller³

Problem:

Shrinking post-scattered orbits in short times

$$a \gtrsim 2 \text{ au} \longrightarrow a \lesssim 0.1 \text{ au}$$

$$e \gtrsim 0.95 \longrightarrow e \approx 0.0$$

(One) solution: Invoke chaotic tides

Stochastic orbital evolution which occurs due to excitation of f modes within the planet and the resulting exchange of energy between those modes and the angular orbital momentum

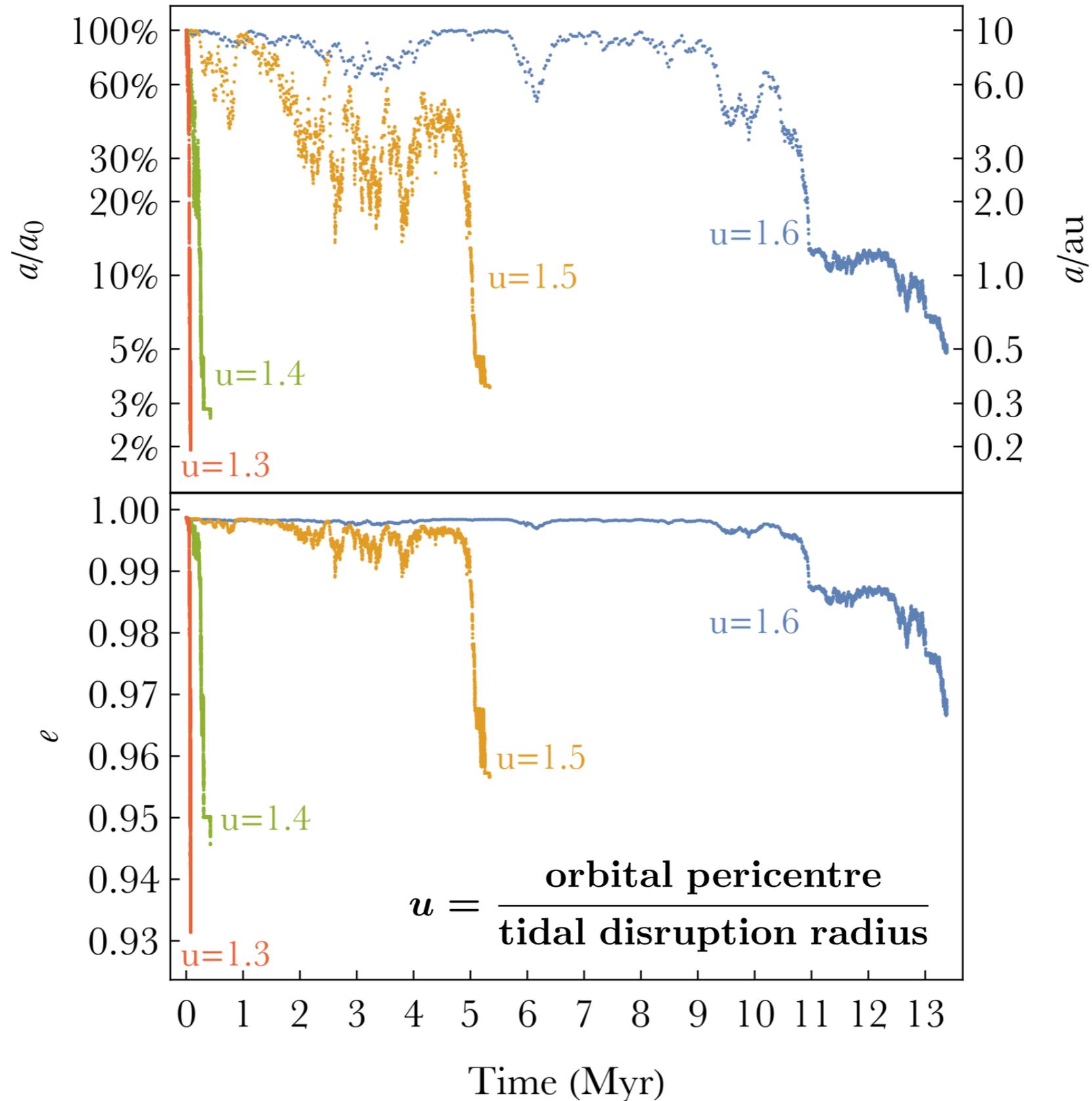
Activates only for highly eccentric orbits

- **Mardling (1995a,b)**
- **Ivanov & Papaloizou (2004, 2007)**
- **Wu (2018)**
- **Vick & Lai (2018)**
- **Vick, Lai & Anderson (2019)**
- **Teyssandier, Lai & Vick (2019)**

Chaotic tides examples

Veras & Fuller (2019)

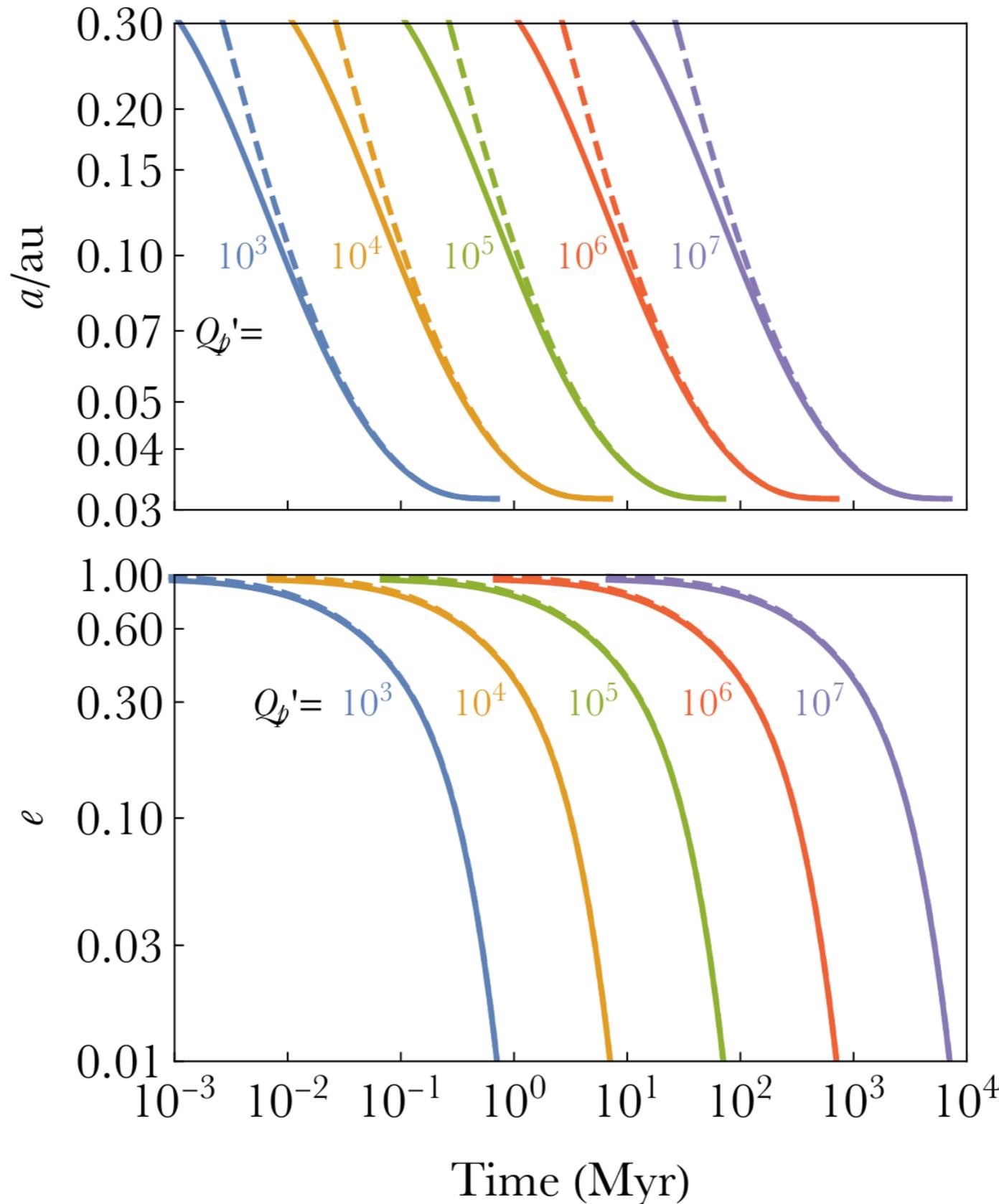
Chaotic orbital evolution for exo-Jupiters



Post-chaotic evolution

Veras & Fuller (2019)

Non-chaotic evolution



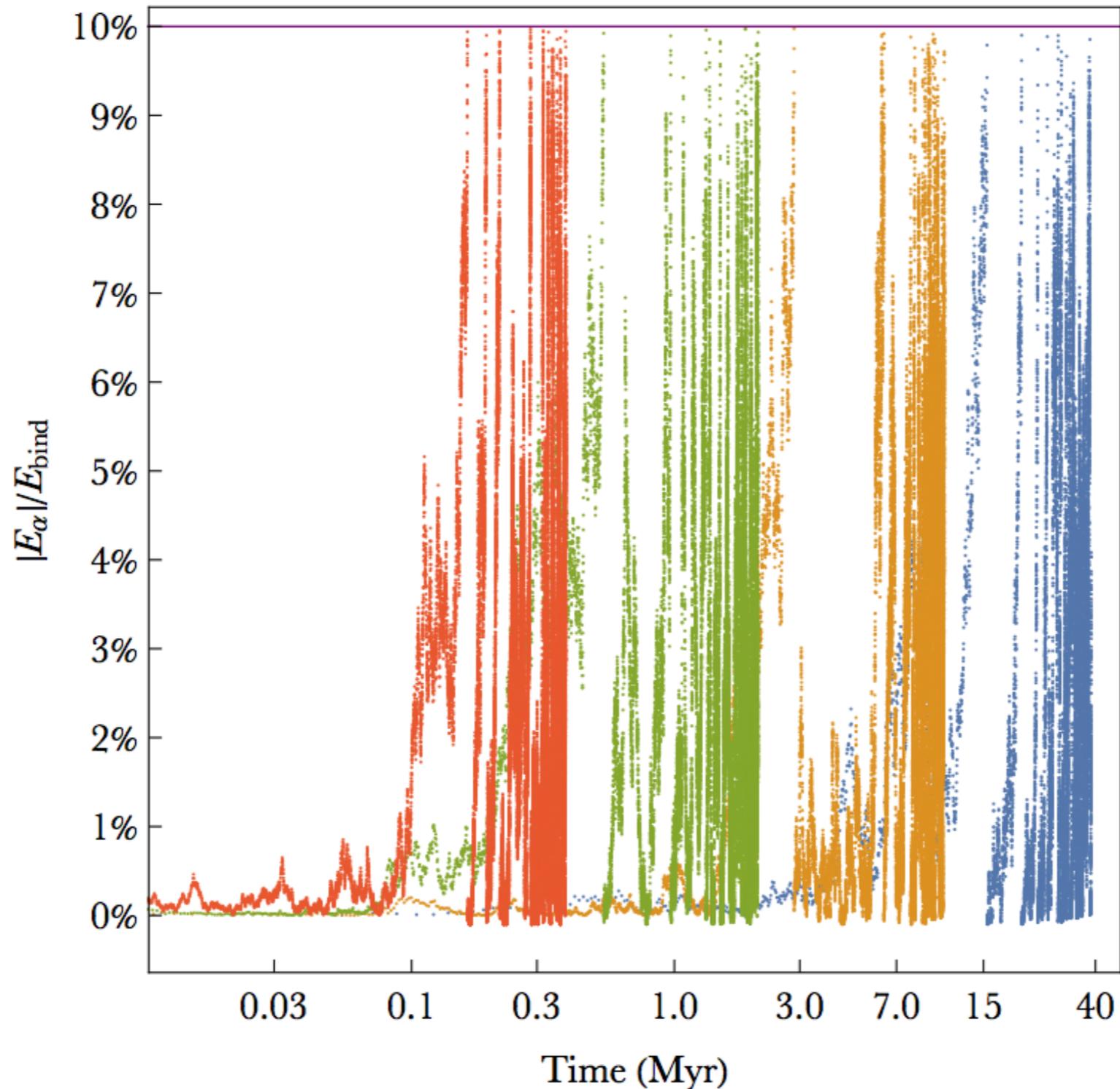
Standard
equilibrium
weak friction
tidal
approximation
from Hut
(1981)

Q'_p
equals the
modified
planetary
quality factor

Interesting twist: Planet potentially disruption

Veras & Fuller (2019)

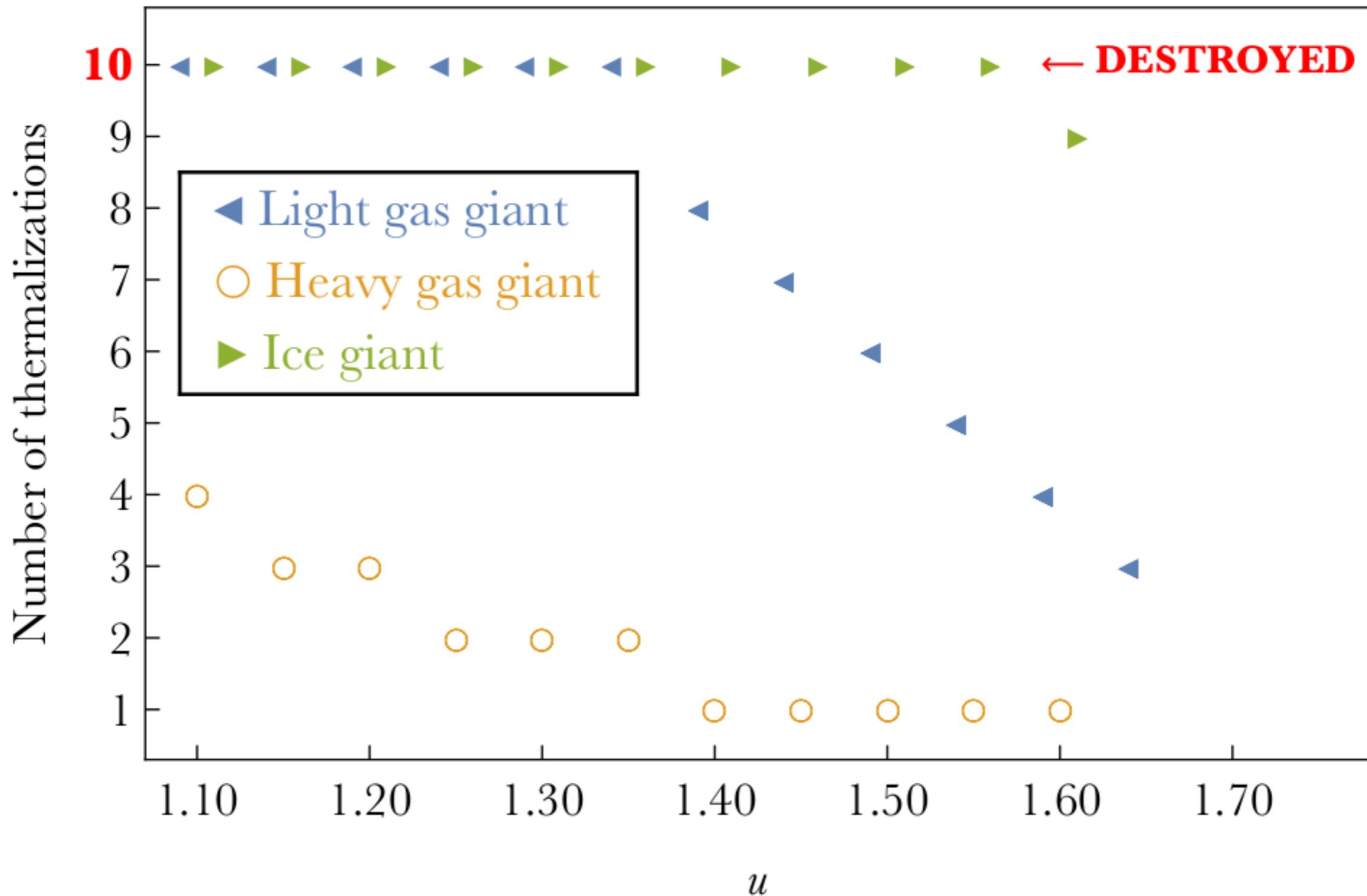
Mode energy evolution for exo-Neptunes



Interesting twist: Planet potentially disruption

Veras & Fuller (2019)

$a_0=10$ au



Gravitational scattering inwards



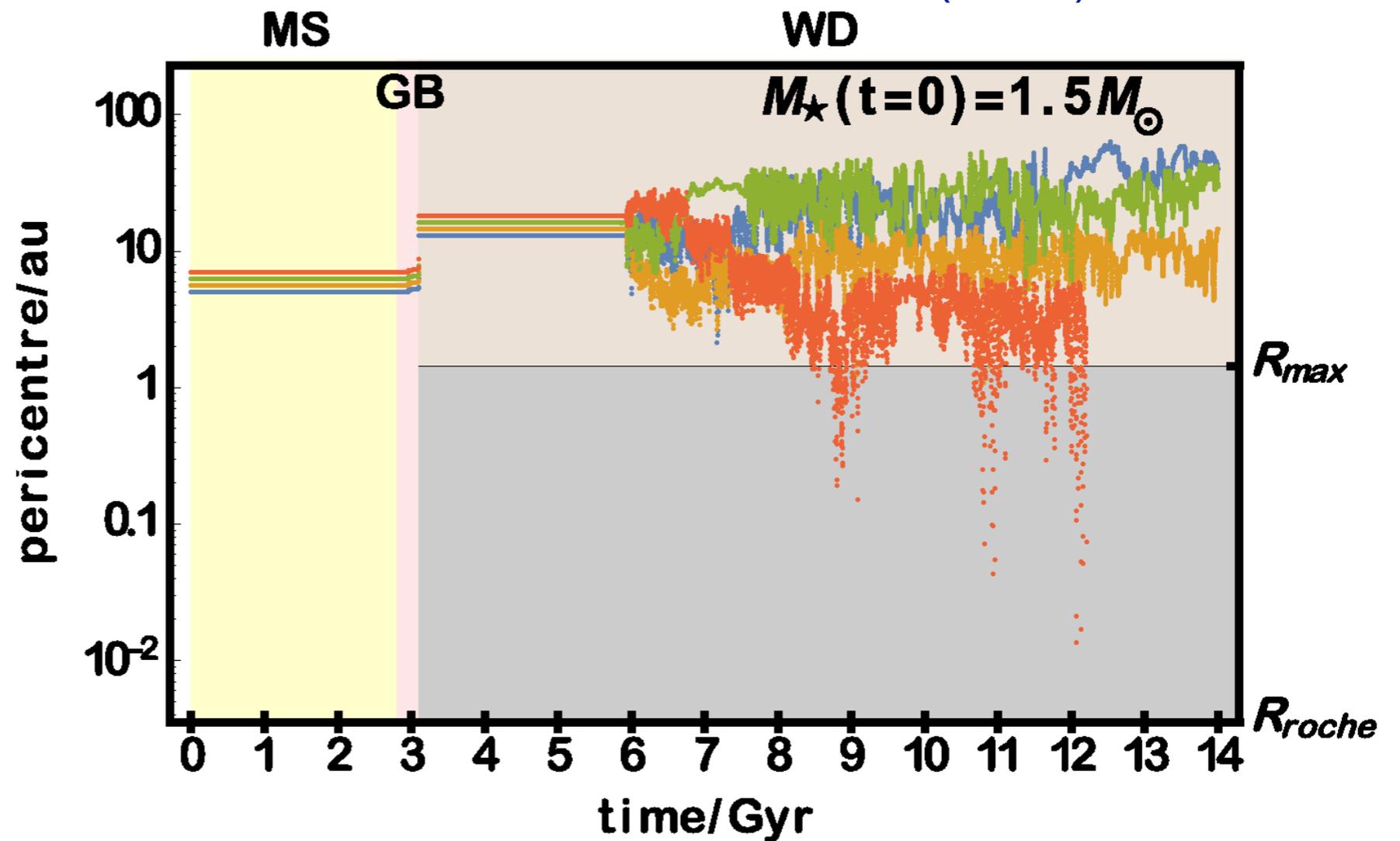
Nature 576, 61–64(2019)

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Veras & Gänsicke (2015)

Scattering
with other
planets



Scattering
with other stars

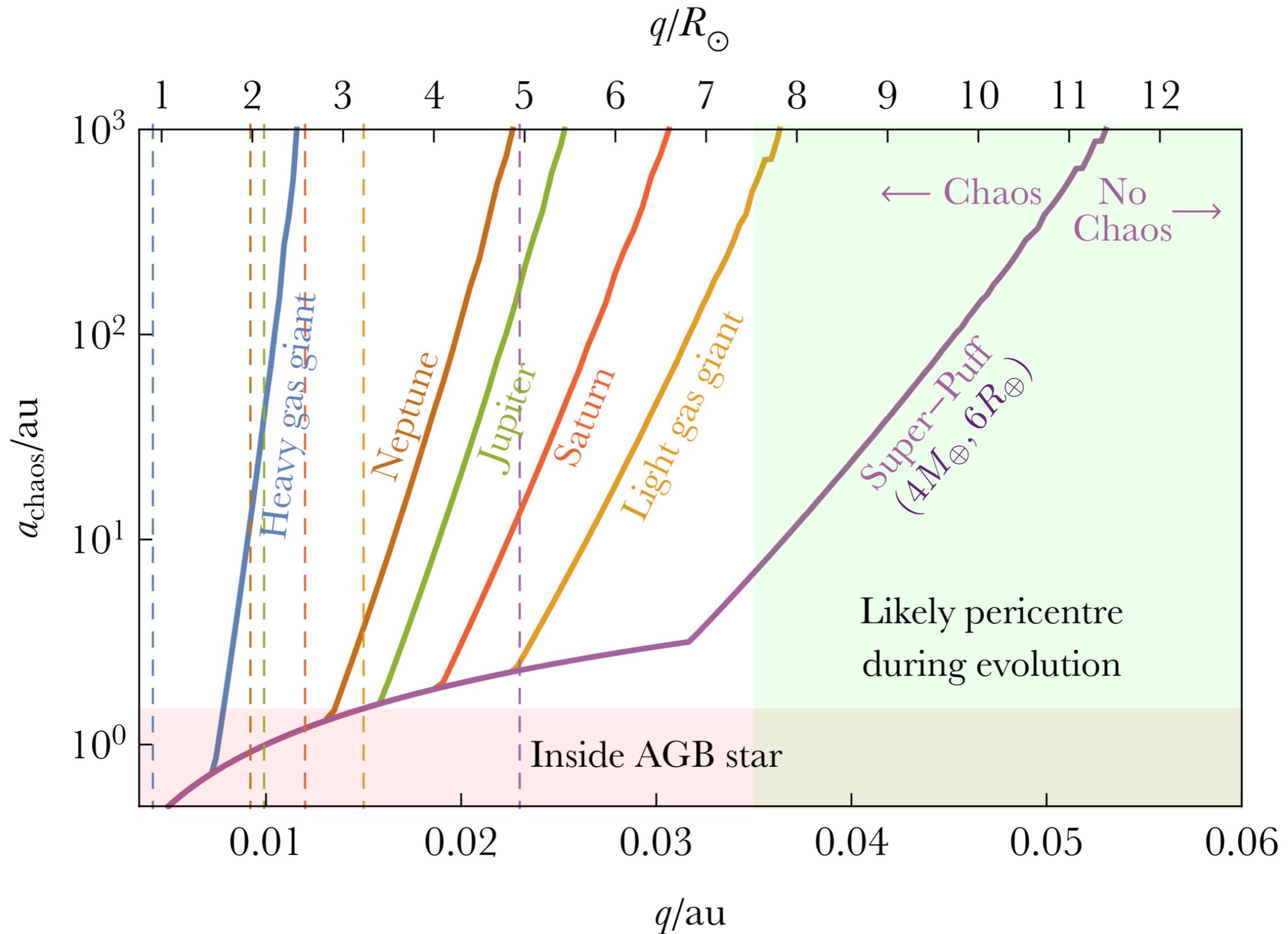


Stephan, Naoz & Gaudi (In Prep)

Physical properties of planet

Veras & Fuller (2020)

Boundary to initiate chaotic evolution around WD J0914+1914



Follow-up studies of this system

—Exploring why no rocky pollutants

MNRAS **493**, 4692–4699 (2020)

The white dwarf planet WD J0914+1914 b: barricading potential rocky pollutants?

Dimitri Veras ^{ID}1,2★†

—Exploring stability regions of gas disc

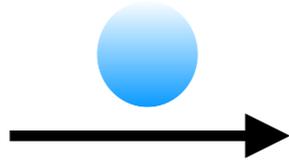
MNRAS in press, arXiv: 2007.16174

Short-term stability of particles in the WD J0914+1914 white dwarf planetary system

Euaggelos E. Zotos¹, Dimitri Veras^{2,3★†}, Tareq Saeed⁴, Luciano A. Darriba^{5,6}

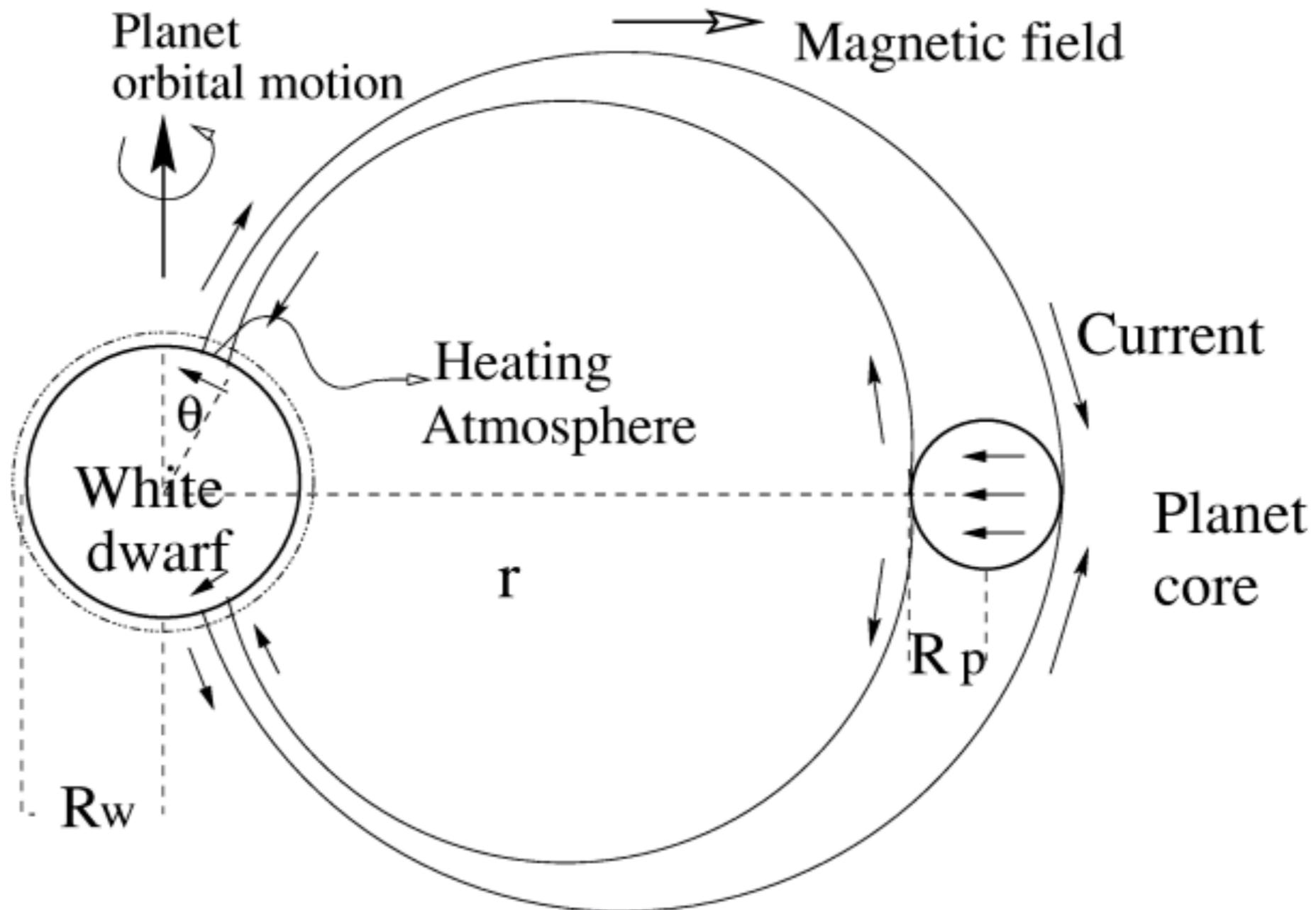
—Exploring binary scattering

Stephan, Naoz & Gaudi (In Prep)



MNRAS 488, 153–163 (2019)
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Dimitri Veras ^{1,2}★† and Alexander Wolszczan ^{3,4}

Li, Ferrario & Wickramasinghe (1998)

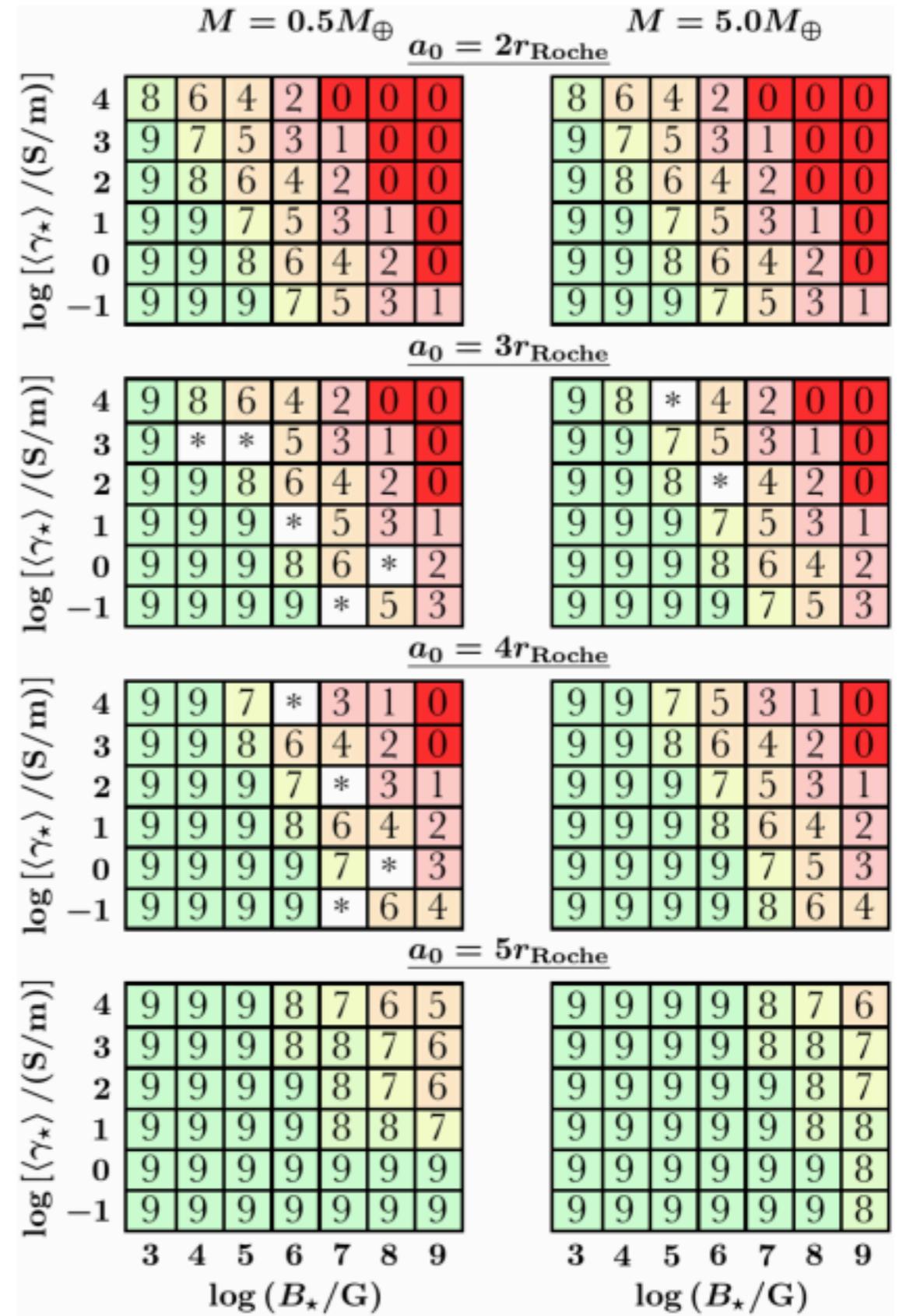
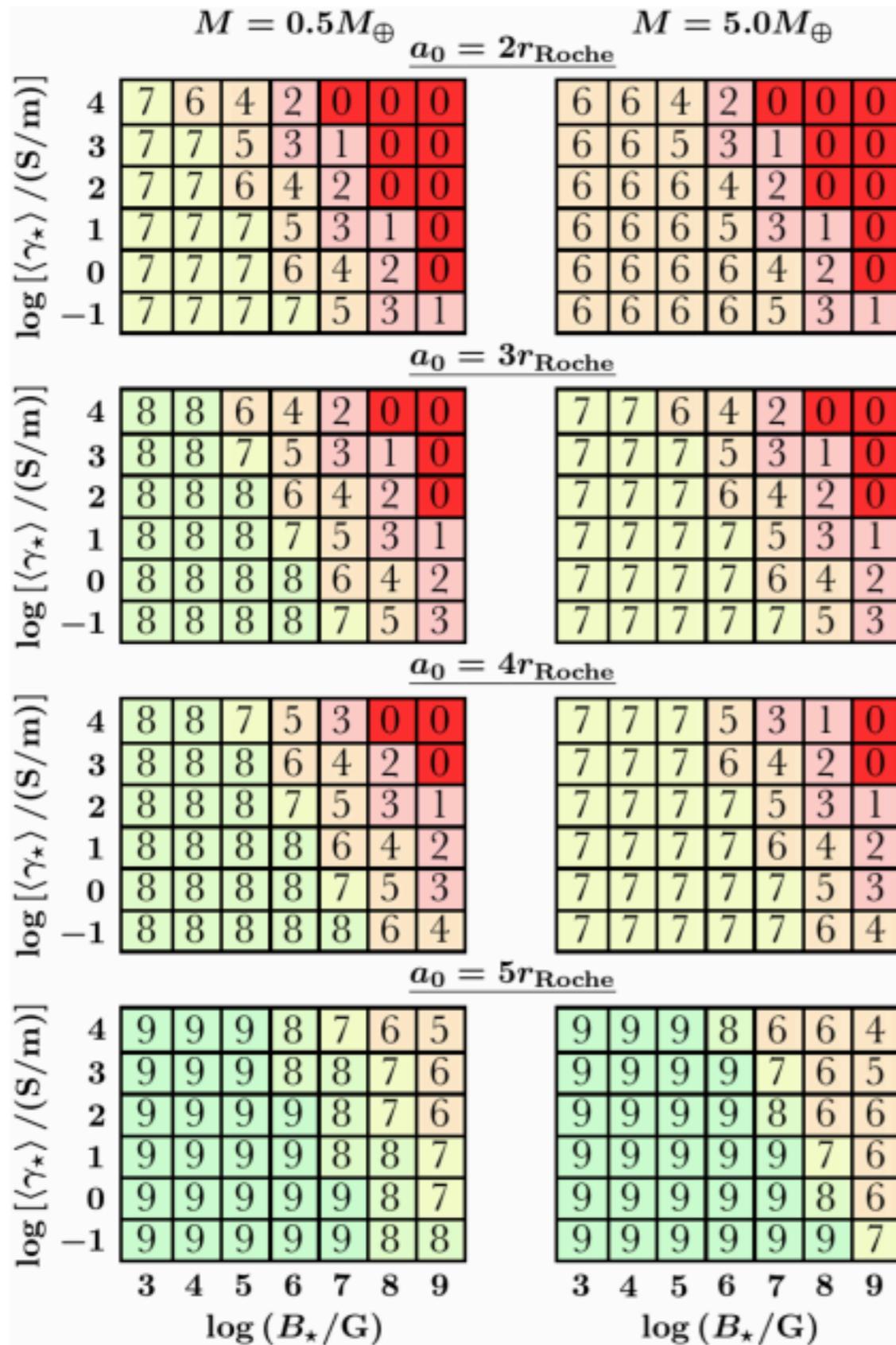


Radio-loud white dwarf planets

Veras & Wolszczan (2019)

Dynamic viscosity = 10^{21} Pa s

Dynamic viscosity = 10^{24} Pa s



Radio-loud white dwarf planets

Survival timescales even longer from circuit breakage due to magnetic flux tube twisting?

Lai (2012)

$$\frac{a}{|\dot{a}|} = 5.7 \times 10^{15} \zeta_{\phi}^{-1} \left(\frac{M_{\star}}{1 M_{\odot}} \right)^{1/2} \left(\frac{R_{\star}}{1 R_{\odot}} \right)^{-6} \left(\frac{B_{\star}}{1 \text{ G}} \right)^{-2} \\ \times \left(\frac{R_p}{1 R_J} \right) \left(\frac{\bar{\rho}_p}{1 \text{ g cm}^{-3}} \right) \left(\frac{a}{0.04 \text{ AU}} \right)^{11/2} \text{ yr}$$

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