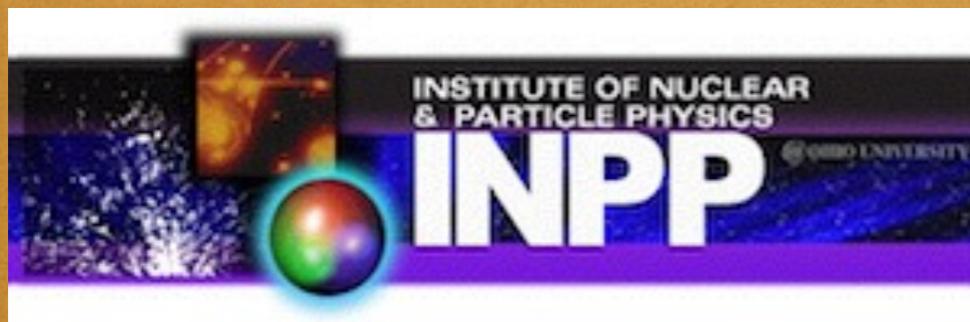


Do halo nuclei exhibit universality?

Daniel Phillips,
Institute of Nuclear and Particle Physics and
Department of Physics and Astronomy
Ohio University, Athens, Ohio



Research supported by the US Department of Energy

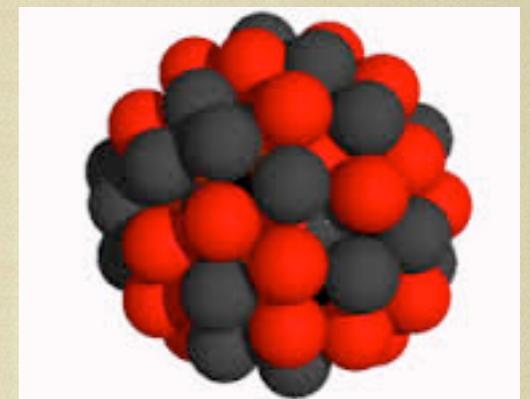
Outline

- The system: halo nuclei
- The tool: Effective field theory for short-range interactions
- Application 1: dipole response of a two-body system in the universal regime (^{19}C)
- Application 2: Radii of three-body systems in the universal regime (Radii of two-neutron halos)
- Application 3: E1 response of three-body systems in the universal regime (E1 response of ^{11}Li)
- Conclusion

Ordinary vs. halo nuclei

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- In nuclei, each nucleon moves in the potential generated by the others
- The nuclear size grows as $A^{1/3}$; cross sections like $A^{2/3}$

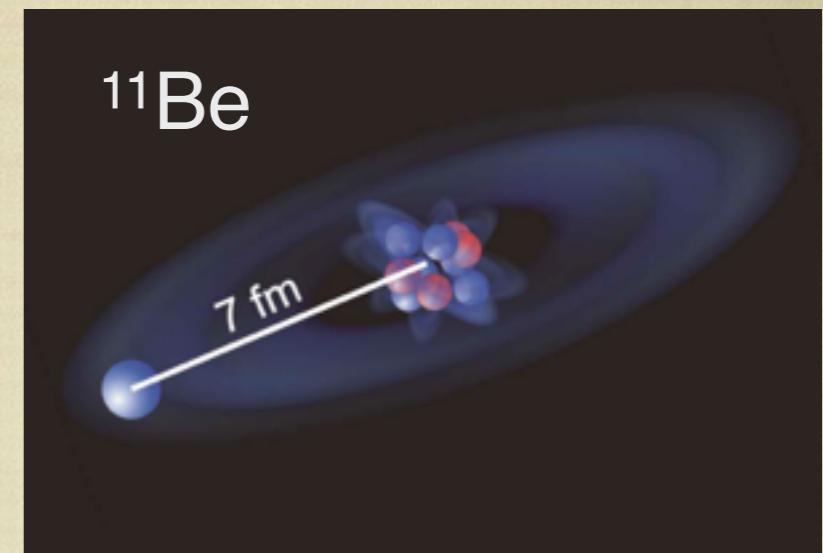


<http://alternativephysics.org>

- Nuclear binding energies are on the order of 8 MeV/nucleon

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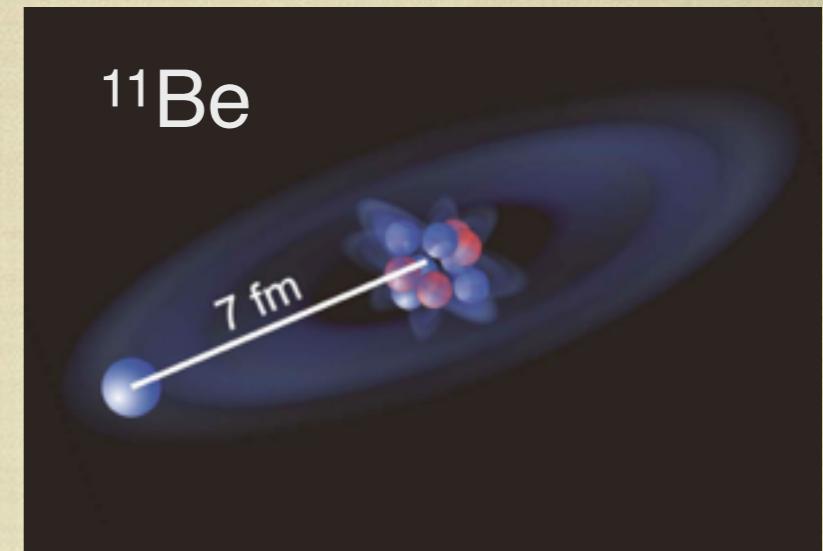
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<http://www.uni-mainz.de>

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- Halo nuclei: the last few nucleons “orbit” far from the nuclear “core”
- Characterized by small nucleon binding energies, large radii, large interaction cross sections, large E1 transition strengths.



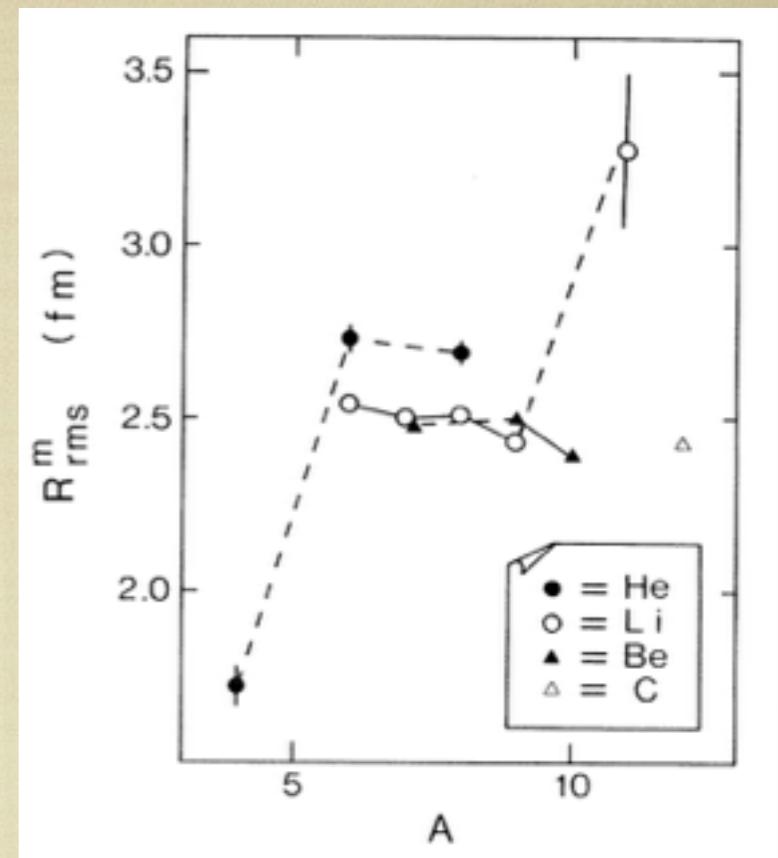
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Halo nuclei: history & examples

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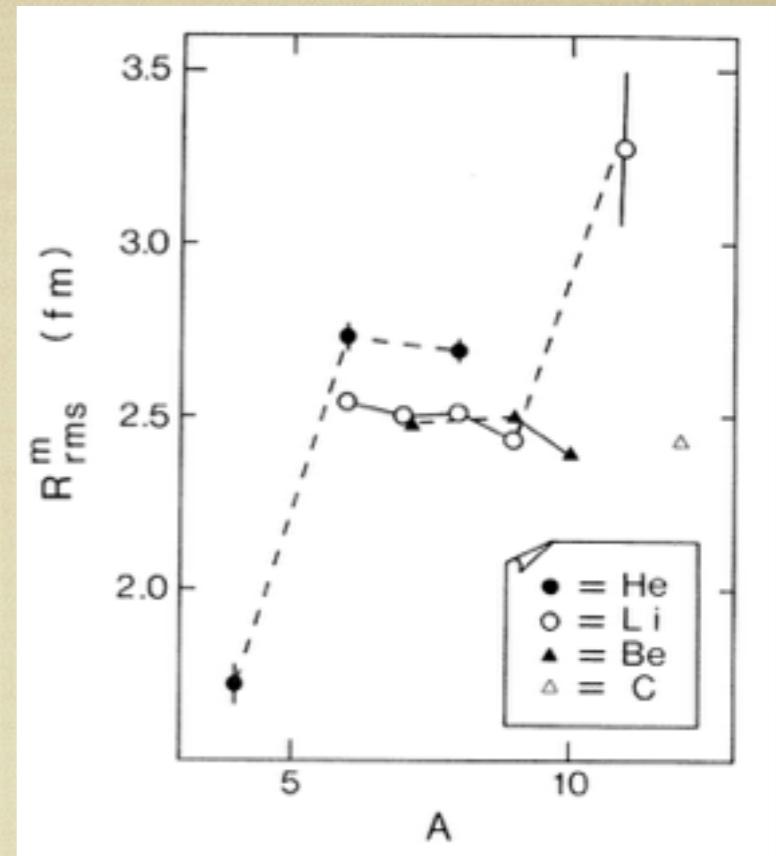
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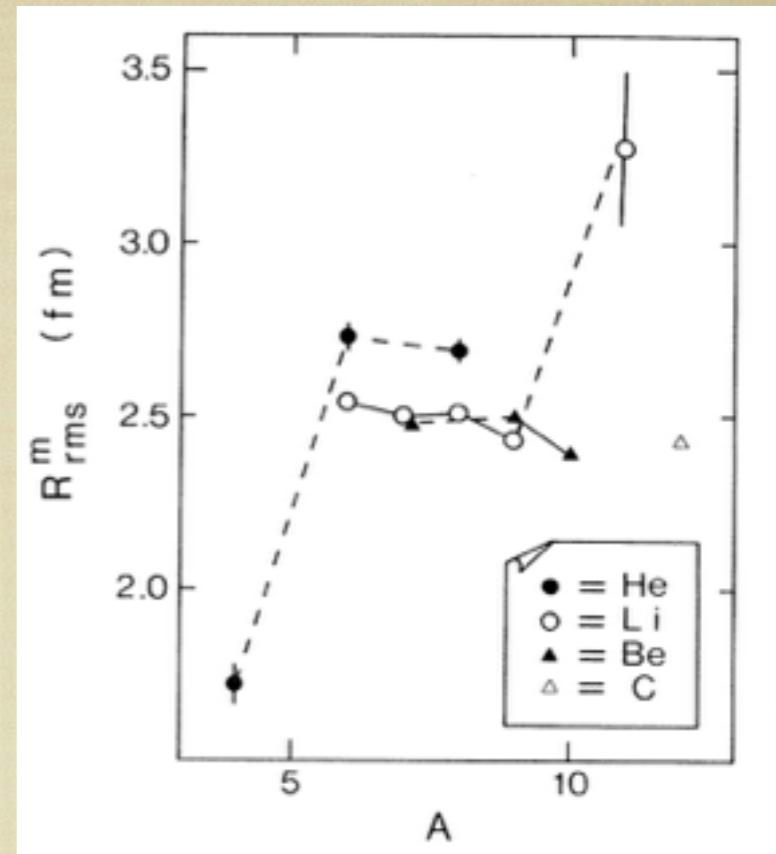
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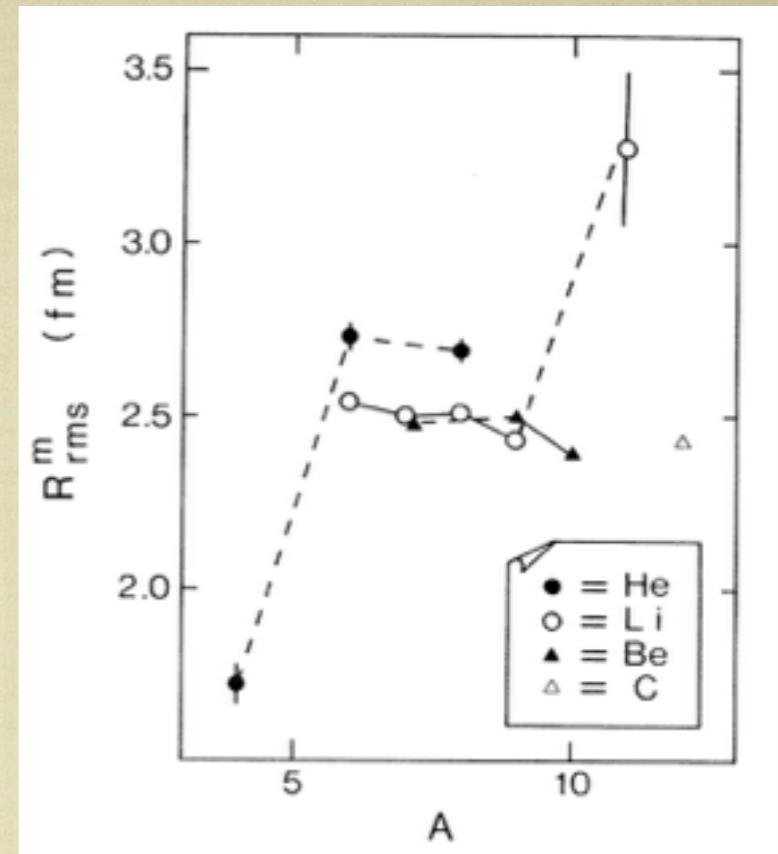
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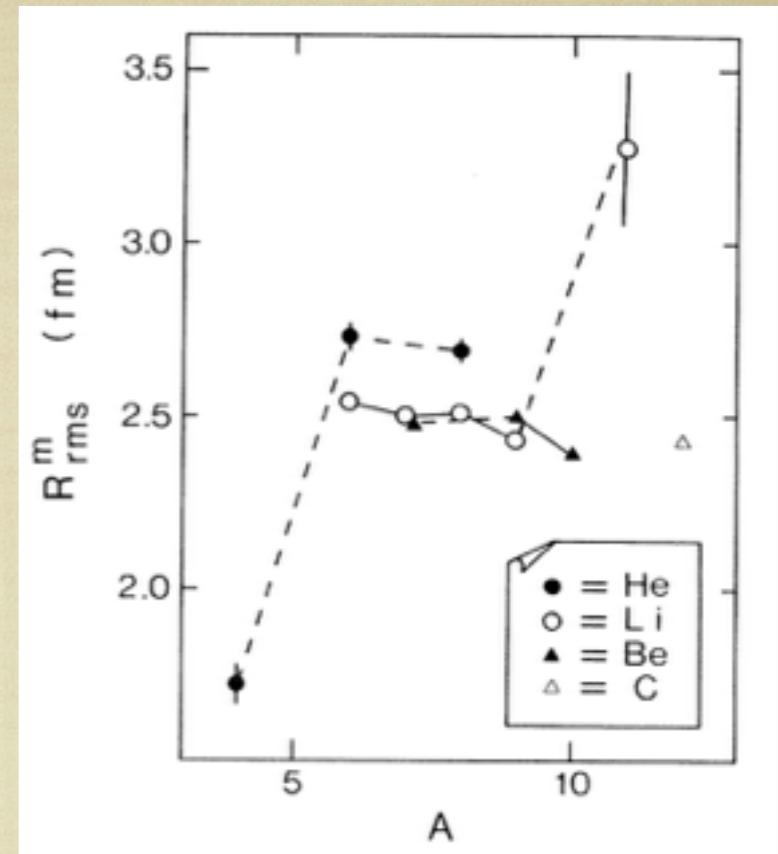
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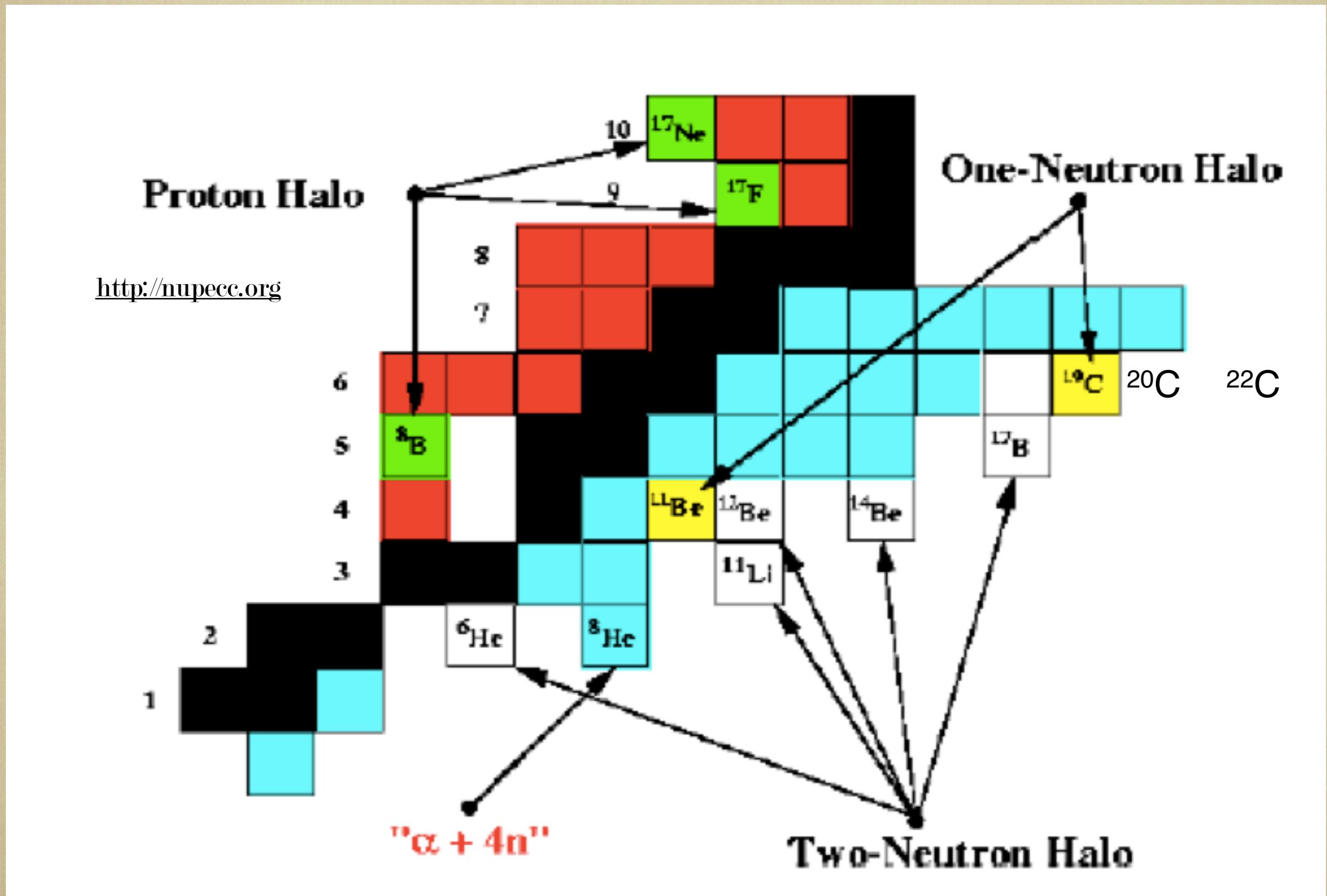
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- Universality: common features of weakly-bound quantum systems



Halo nuclei: history & examples



Universality

Systems with $|a| \gg R$ exhibit the same correlations between low-energy observables

System	R	$ a $	Observables
He atom clusters	7 Å	104 Å	Binding energies, distributions
Cold atoms	$100 a_B$	Varies	Bound states; recombination
X(3872)	1.5 fm	~10 fm	Spectrum, decays
Halo nuclei	3 fm	~10 fm	Spectrum, scattering, EM excitation
NN, NNN, ...	1.7 fm	5.4 fm	Phase shifts; EM props...

EFT formulation for two-body

$$-\frac{\hbar^2}{2m_R}\nabla^2\psi+V(\mathbf{r})\psi(\mathbf{r})=E\psi(\mathbf{r})$$

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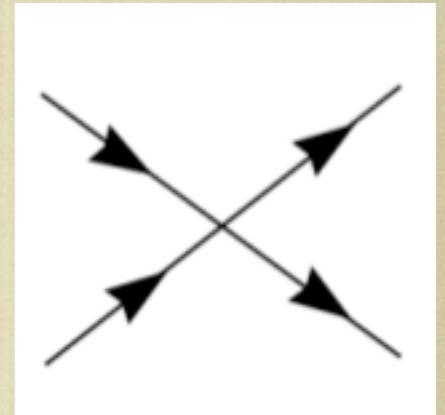
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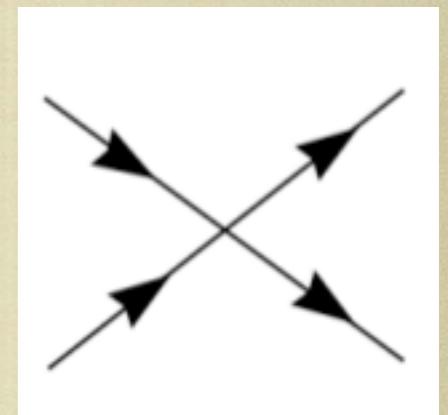


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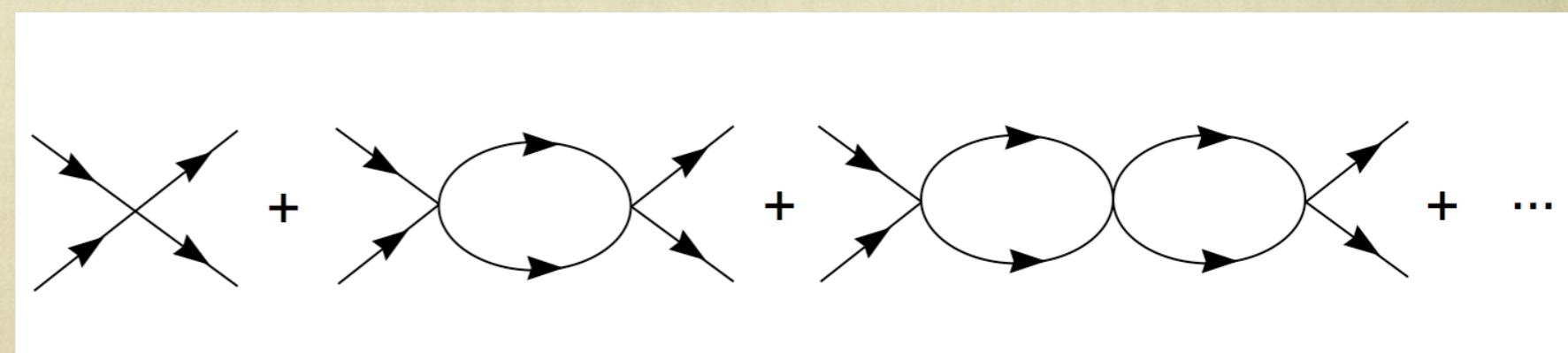
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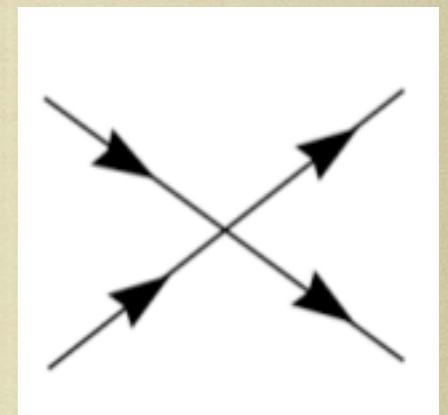
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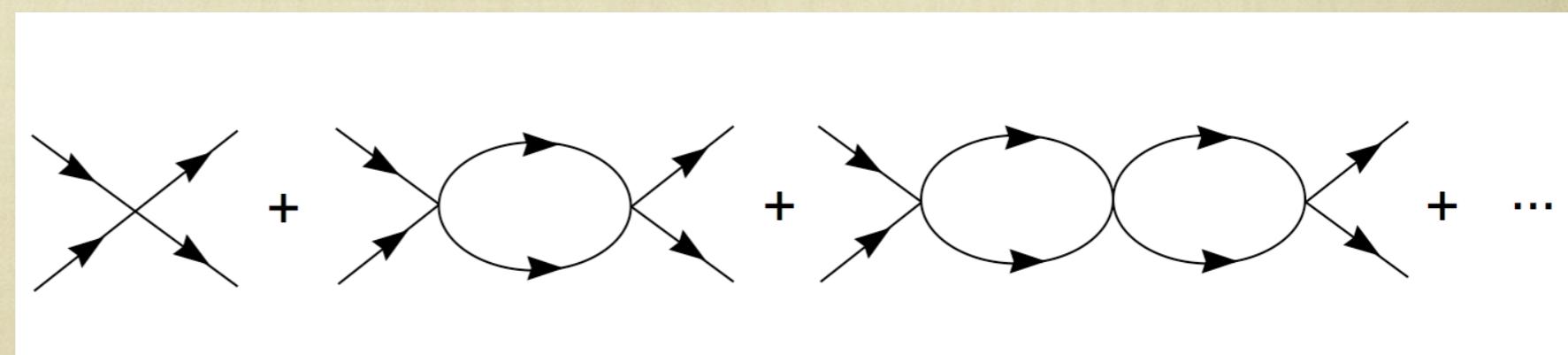
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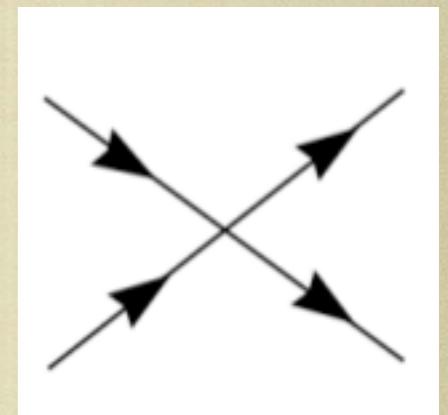


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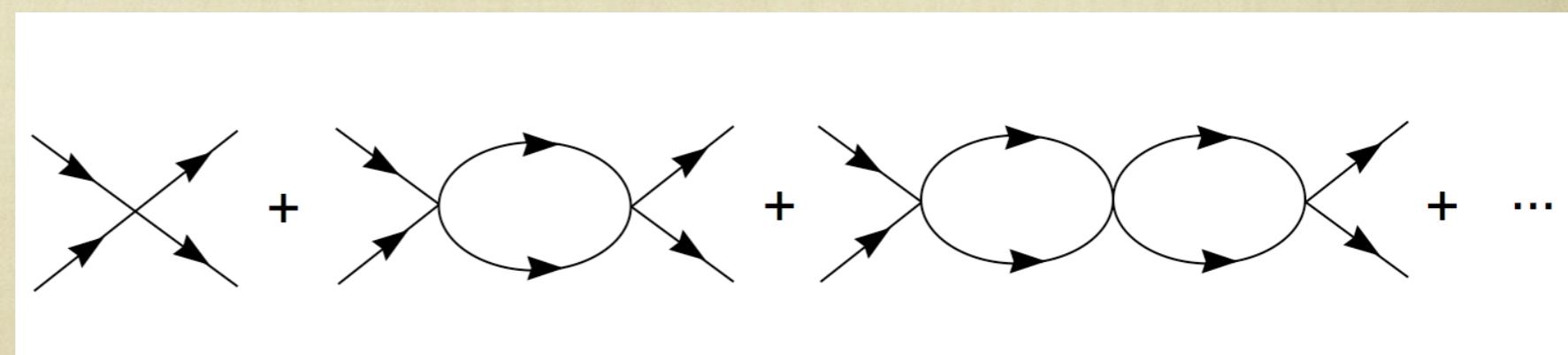
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Leading in systematic EFT expansion \Rightarrow Estimate theory uncertainty

Two-body t beyond LO

$$t(E) = -\frac{4\pi}{m} \frac{1}{k \cot \delta(k) - ik}; \quad k = \sqrt{mE}$$

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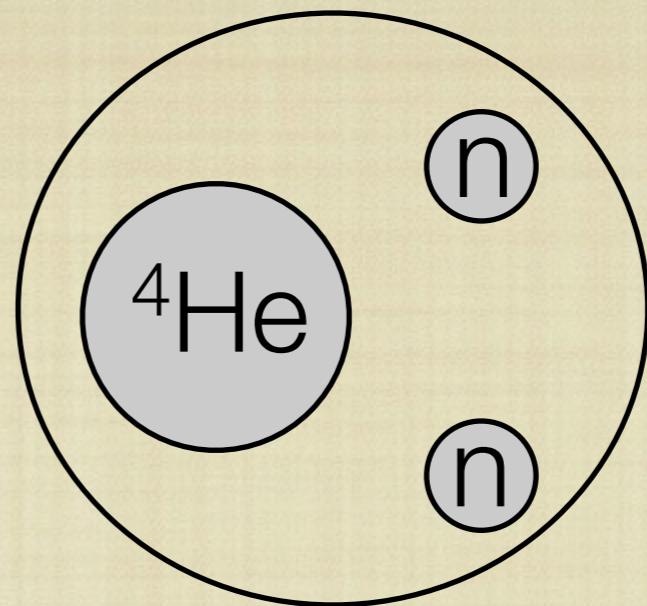
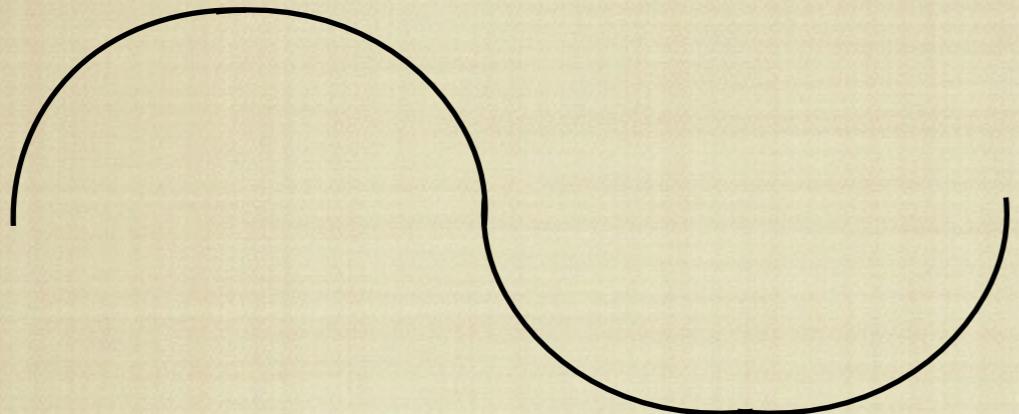
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LO NLO

- ...provided $k \sim 1/a$. As good as ERE?

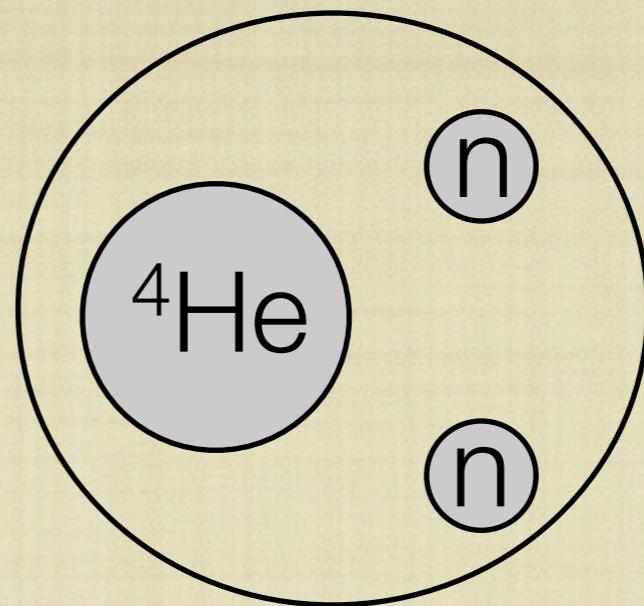
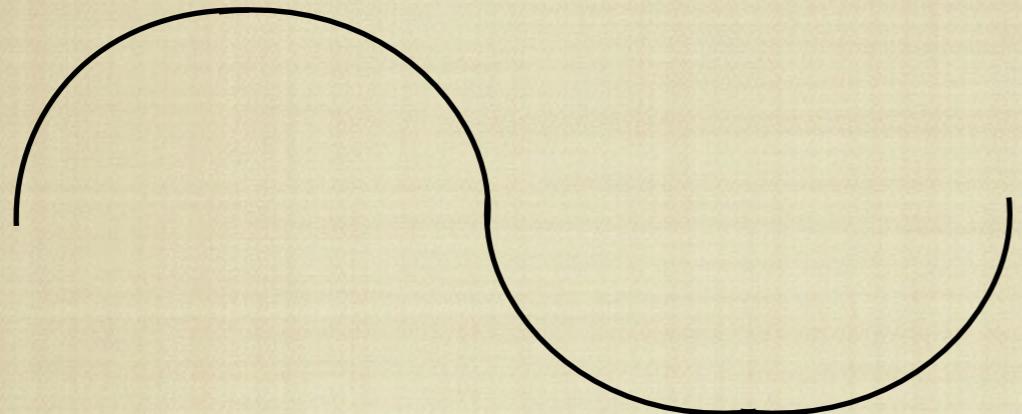
Scales in halo nuclei

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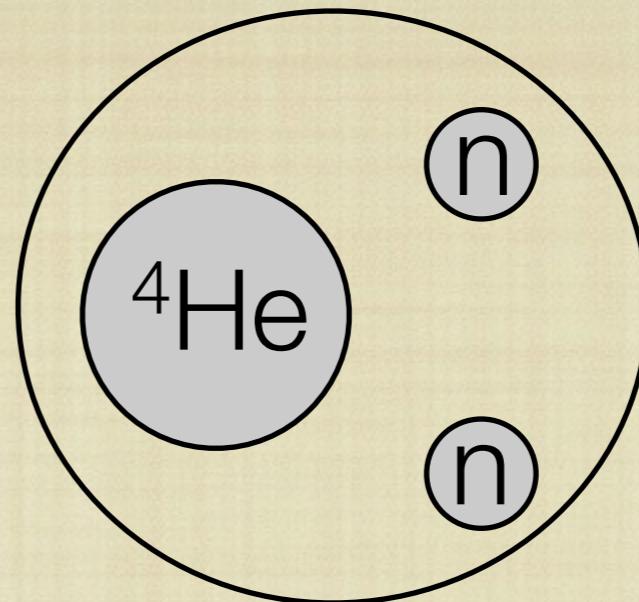
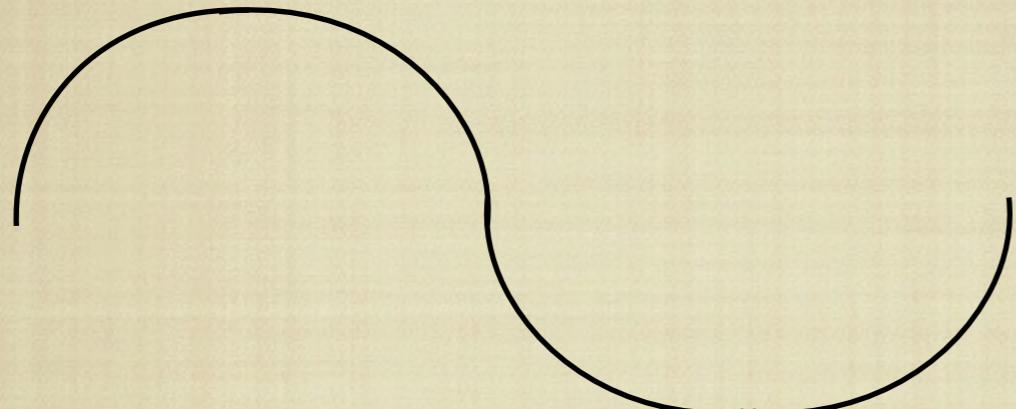
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- To be in the universal regime need $R_{\text{core}} \ll R_{\text{halo}}$
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- By this definition the deuteron is the lightest halo nucleus, and few-nucleon systems are a specific case of halos

Lagrangian for shallow bound states

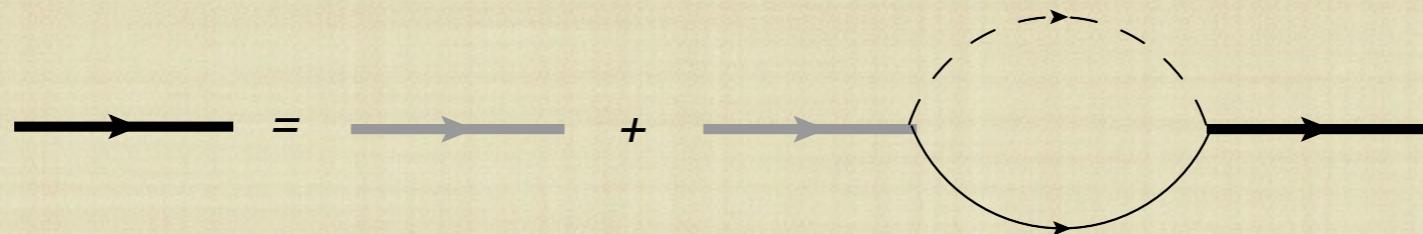
$$\begin{aligned}\mathcal{L} = & c^\dagger \left(i\partial_t + \frac{\nabla^2}{2M} \right) c + n^\dagger \left(i\partial_t + \frac{\nabla^2}{2m} \right) n \\ & + \sigma^\dagger \left[\eta_0 \left(i\partial_t + \frac{\nabla^2}{2M_{nc}} \right) + \Delta_0 \right] \sigma + \pi_j^\dagger \left[\eta_1 \left(i\partial_t + \frac{\nabla^2}{2M_{nc}} \right) + \Delta_1 \right] \pi_j \\ & - g_0 [\sigma n^\dagger c^\dagger + \sigma^\dagger n c] - \frac{g_1}{2} \left[\pi_j^\dagger (n \stackrel{\leftrightarrow}{i\nabla}_j c) + (c^\dagger \stackrel{\leftrightarrow}{i\nabla}_j n^\dagger) \pi_j \right] \\ & - \frac{g_1}{2} \frac{M-m}{M_{nc}} \left[\pi_j^\dagger \stackrel{\rightarrow}{i\nabla}_j (n c) - \stackrel{\leftrightarrow}{i\nabla}_j (n^\dagger c^\dagger) \pi_j \right] + \dots,\end{aligned}$$

- c, n: “core”, “neutron” fields. c: boson, n: fermion
- σ, π_j : S-wave and P-wave fields
- Minimal substitution generates leading EM couplings

Dressing the S-wave state

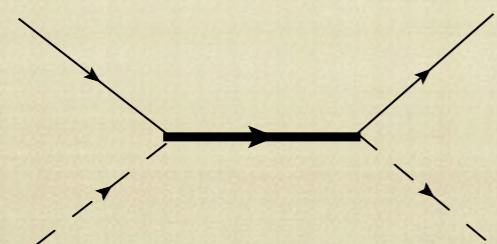
Kaplan, Savage, Wise; van Kolck; Gegelia;
Birse, Richardson, McGovern

- $\sigma n c$ coupling g_0 of order R_{halo} , nc loop of order $1/R_{\text{halo}}$. Therefore need to sum all bubbles:



$$D_\sigma(p) = \frac{1}{\Delta_0 + \eta_0[p_0 - \mathbf{p}^2/(2M_{nc})] - \Sigma_\sigma(p)}$$

$$\Sigma_\sigma(p) = -\frac{g_0^2 m_R}{2\pi} \left[\mu + i\sqrt{2m_R \left(p_0 - \frac{\mathbf{p}^2}{2M_{nc}} + i\eta \right)} \right] \quad (\text{PDS})$$



$$t = \frac{2\pi}{m_R} \frac{1}{\frac{1}{a_0} - \frac{1}{2}r_0 k^2 + ik}$$

$$D_\sigma(p) = \frac{2\pi\gamma_0}{m_R^2 g_0^2} \frac{1}{1 - r_0\gamma_0} \frac{1}{p_0 - \frac{\mathbf{p}^2}{2M_{nc}} + B_0} + \text{regular}$$

Counting in S waves:
 $a_0 \sim R_{\text{halo}} \sim 1/\gamma_0$; $r_0 \sim R_{\text{core}}$.
 $r_0 = 0$ at LO.

Radii of s-wave 1n halos

Wave function: $u(r) = C \exp(-\gamma_0 r) \Rightarrow \langle r_{nc}^2 \rangle^{1/2} = \frac{C}{2} \left(\frac{A+1}{2AM_N S_{1n}} \right)^{3/4}$

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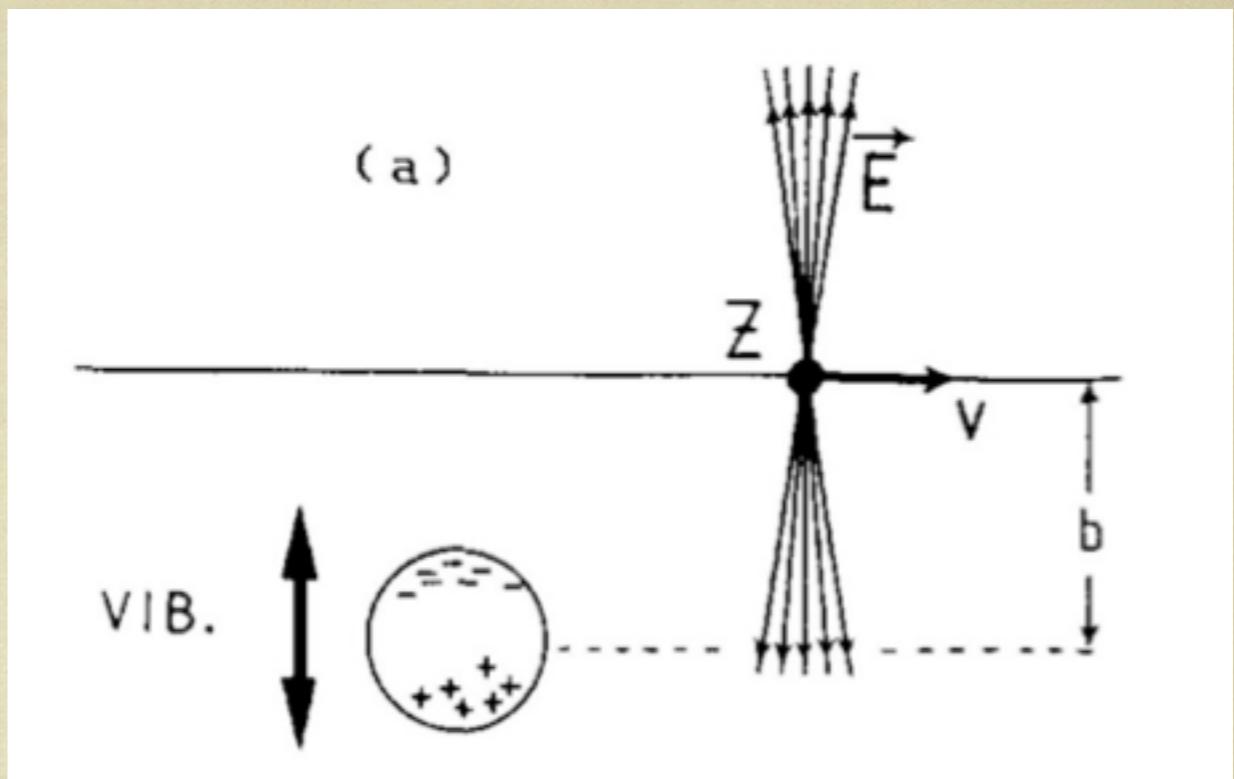
	S_{1n} (MeV)	R_{core}/R_{halo}	$\langle r_{nc}^2 \rangle^{1/2}$ (fm) NNLO	$\langle r_{nc}^2 \rangle^{1/2}$ (fm) Expt
2H	2.224573(2)	0.33	3.954	3.9270(90)
^{11}Be	0.50164(25)	0.4	6.16	5.7(4)
^{15}C	1.2181(8)	0.45	4.93	4.5(5)
^{19}C	0.58(9)	0.33	5.72	6.8(7)

All radii are substantially smaller at LO: range corrections are crucial to obtaining agreement with experiment

Photodissociation: experiments

Bertulani, arXiv:0908.4307

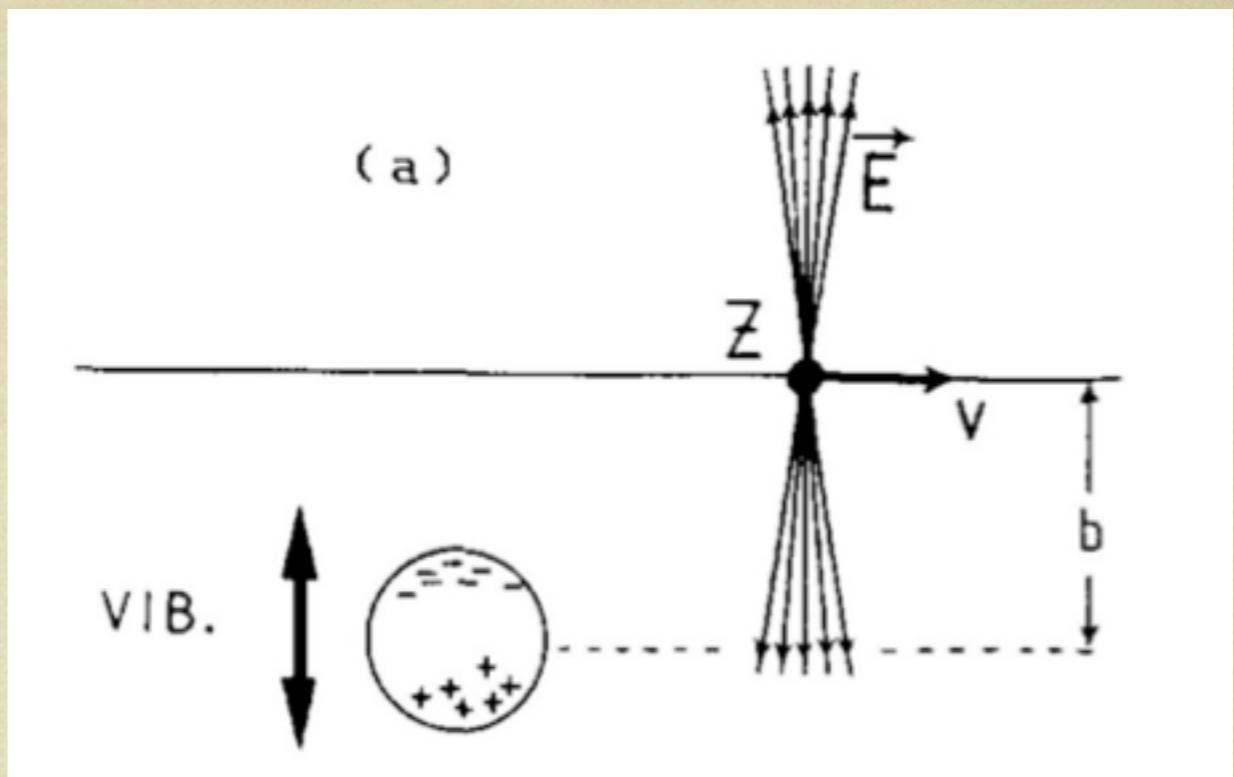
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- Do with different Z , different nuclear sizes, different energies to test systematics
- C.f. trimer photoassociation



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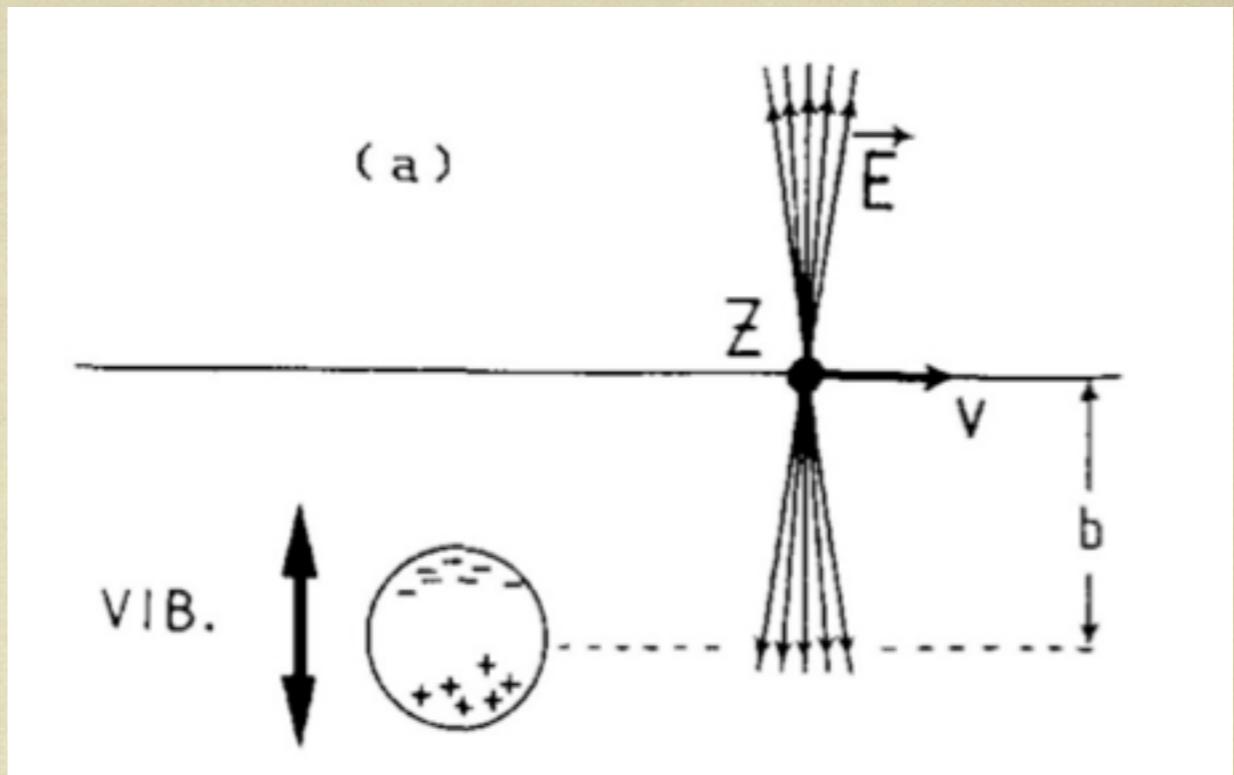


Bazak, Liverts & Barnea, PRL (2012), PRA (2013), Bazak & Barnea, arXiv:1502.07119

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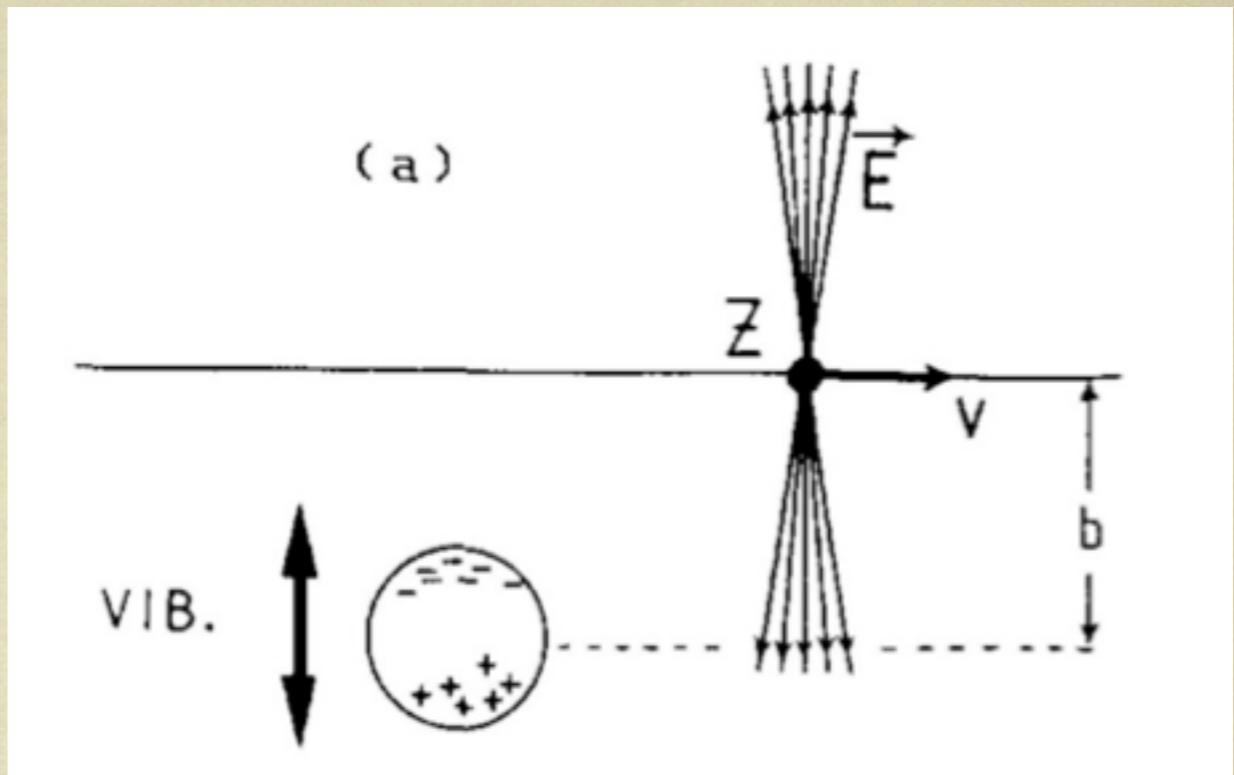
- Coulomb excitation dissociation cross section (p.v. $b \gg R_{\text{target}}$)

$$\frac{d\sigma_C}{2\pi b db} = \sum_{\pi L} \int \frac{dE_\gamma}{E_\gamma} n_{\pi L}(E_\gamma, b) \sigma_\gamma^{\pi L}(E_\gamma)$$

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- $n_{\pi L}(E_\gamma, b)$ virtual photon numbers, dependent only on kinematic factors.
Number of equivalent (virtual) photons that strike the halo nucleus.
- $\sigma_\gamma^{\pi L}(E_\gamma)$ can then be extracted: it's the (total) cross section for dissociation of the nucleus due to the impact of photons of multipolarity πL .

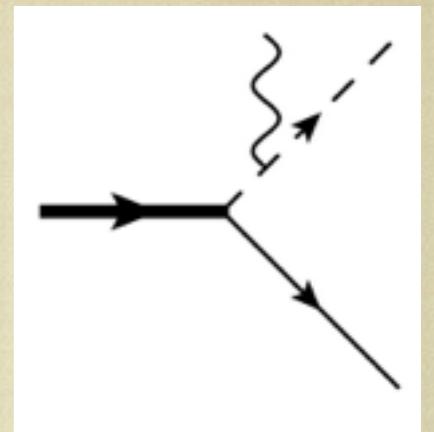
Universal dissociation

Chen, Savage (1999)

- Leading order: no FSI, $r_0=0 \Rightarrow \gamma_0$ is only free parameter

$$\gamma_0^2 = 2 m_R S_{1n}$$

$$\mathcal{M} = \frac{eQ_c g_0 2m_R}{\gamma_0^2 + \left(\mathbf{p}' - \frac{m}{M_{nc}} \mathbf{k} \right)^2}$$



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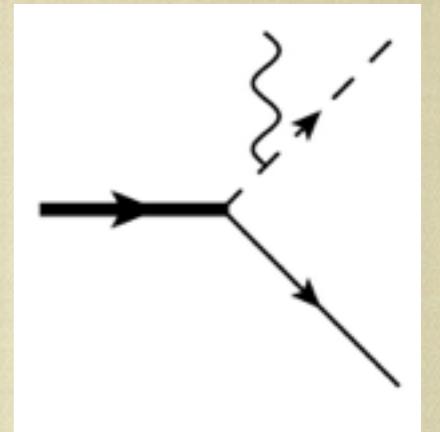
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$$Z_{eff} = Z/(A+1)$$

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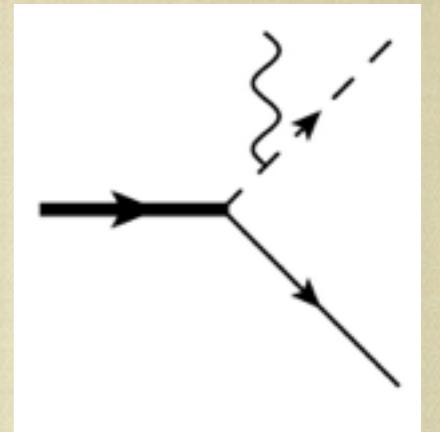
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Universal E1 strength formula for 2B systems

Universal dissociation

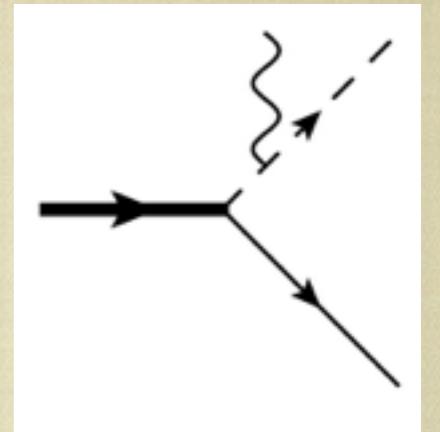
Chen, Savage (1999)

- Leading order: no FSI, $r_0=0 \Rightarrow \gamma_0$ is only free parameter

$$\gamma_0^2 = 2 m_R S_{1n}$$

$$\mathcal{M} = \frac{eQ_c g_0 2m_R}{\gamma_0^2 + \left(\mathbf{p}' - \frac{m}{M_{nc}} \mathbf{k} \right)^2}$$

Up to NNLO: $\frac{dB(E1)}{e^2 dE} = \frac{12m_R}{\pi^2} Z_{eff}^2 \frac{\gamma_0}{1 - r_0 \gamma_0} \frac{p^3}{(\gamma_0^2 + p^2)^4}$



$$Z_{eff} = Z/(A+1)$$

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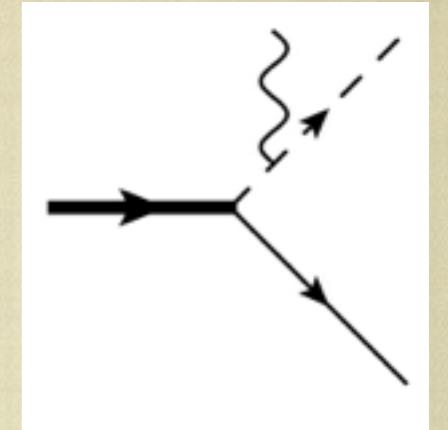
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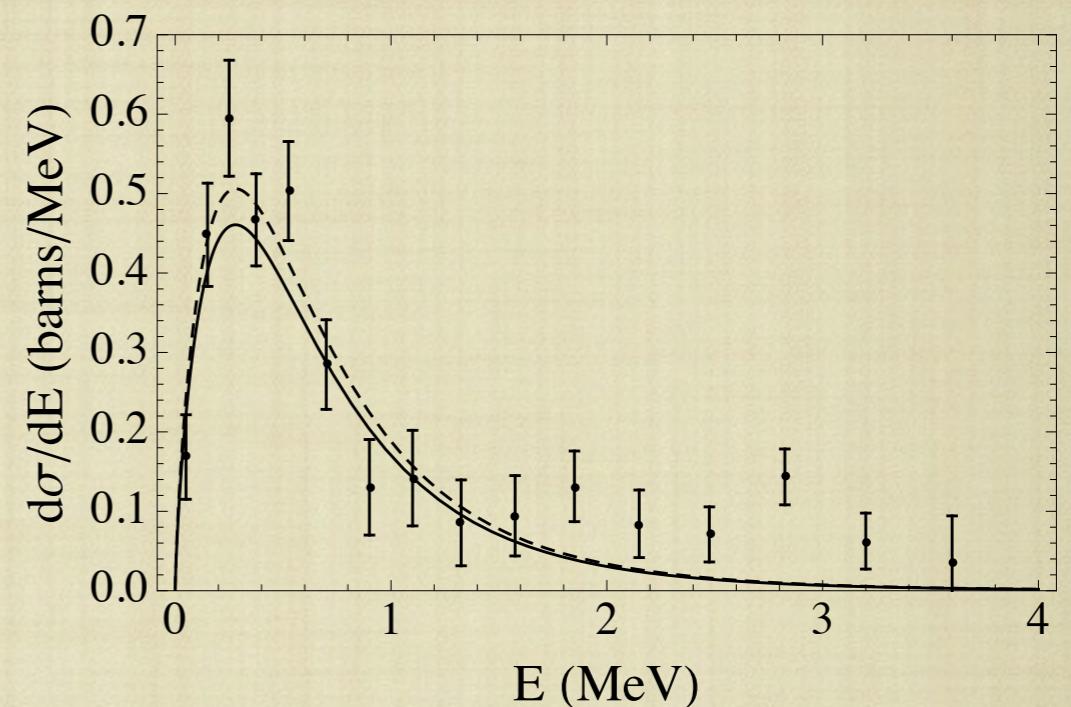
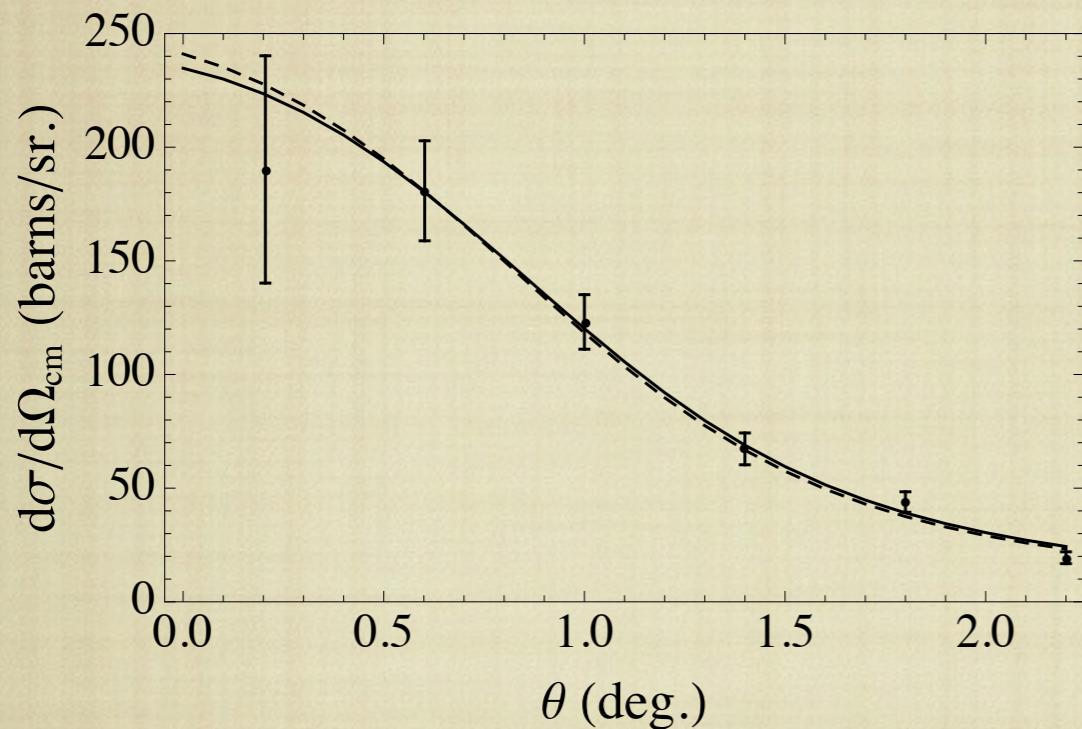
Universal E1 strength formula for 2B systems

- Corresponds to $u_0(r) = C \exp(-\gamma_0 r)$: $C^2 = \frac{2\gamma_0}{1 - r_0 \gamma_0}$
- Final-state interactions suppressed by $(R_{core}/R_{halo})^3$
- First gauge-invariant contact operator: $L_{E1} \sigma^\dagger \mathbf{E} \cdot (n \stackrel{\leftrightarrow}{\nabla} c) + h.c.$

Results

Data: Nakamura et al., 1999, 2003
Analysis: Acharya, Phillips. 2013
cf. Singh et al., 2008

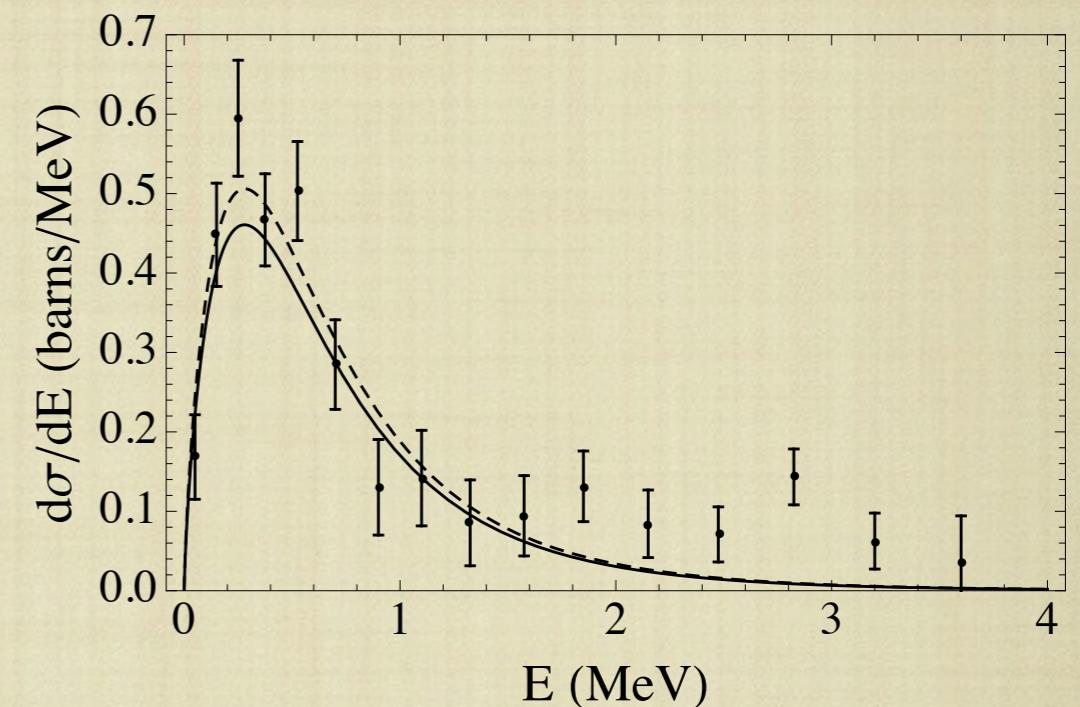
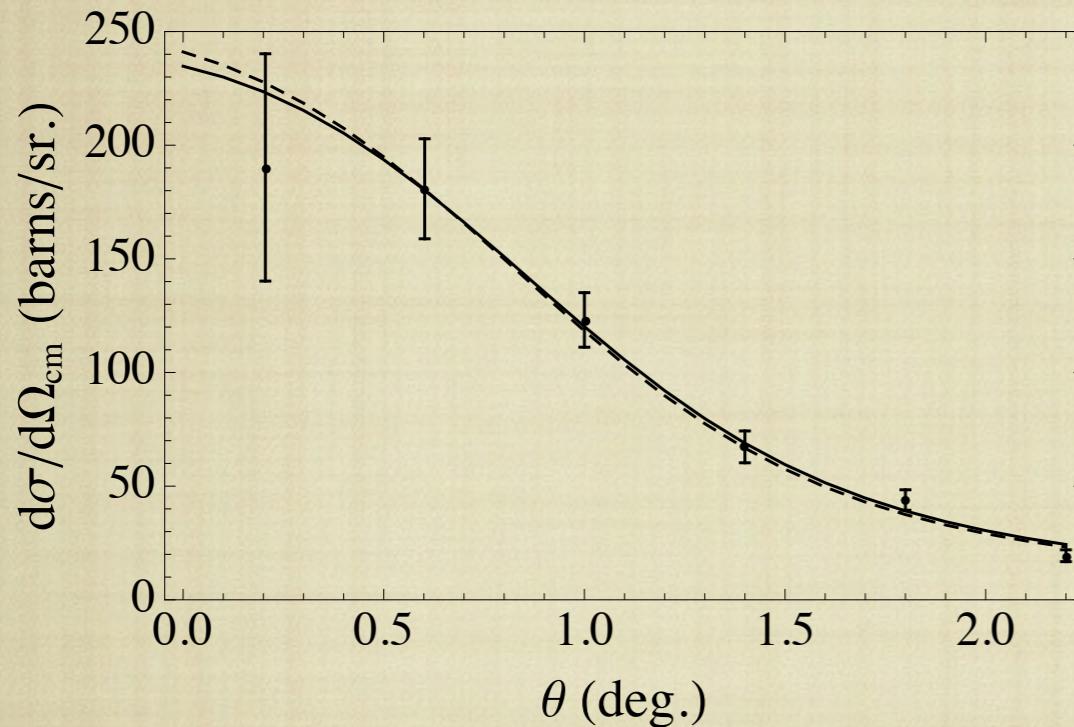
- Integrate this E1 strength for transition to a core + neutron state, per unit energy per unit solid angle, as function of energy of the outgoing nc pair over differential photon numbers and over angle.



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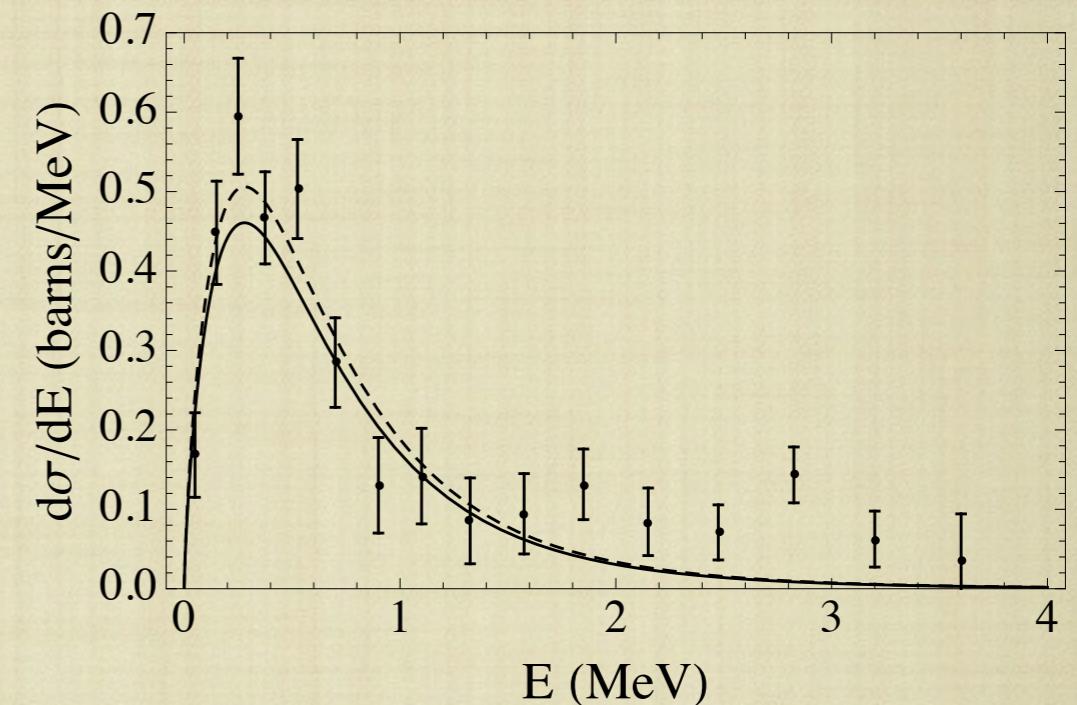
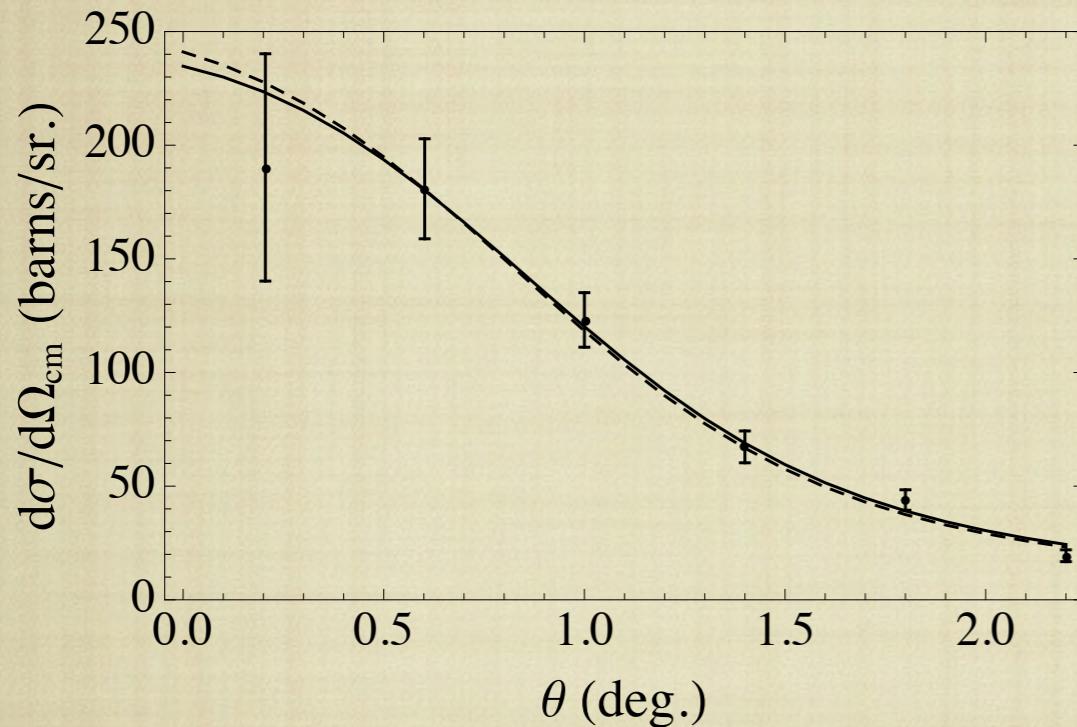


- $\gamma_0 \equiv a$ determines peak position and fall off of angular distribution
- r_0 fixed from fitting height of peak

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- Integrate this E1 strength for transition to a core + neutron state, per unit energy per unit solid angle, as function of energy of the outgoing nc pair over differential photon numbers and over angle.



$$\begin{aligned} a &= (7.75 \pm 0.35(\text{stat.}) \pm 0.3(\text{EFT})) \text{ fm;} \\ r_0 &= (2.6^{+0.6}_{-0.9}(\text{stat.}) \pm 0.1(\text{EFT})) \text{ fm.} \end{aligned}$$

Determine S-wave $^{18}\text{C}-\text{n}$ scattering
parameters \leftrightarrow ANC s from dissociation data.

P-waves: $\gamma_{E1} + {}^{11}\text{Be} \rightarrow {}^{10}\text{Be} + n$

TypeI & Baur, Phys. Rev. Lett. 93, 142502 (2004); Nucl. Phys. A759, 247 (2005); Eur. Phys. J. A 38, 355 (2008)

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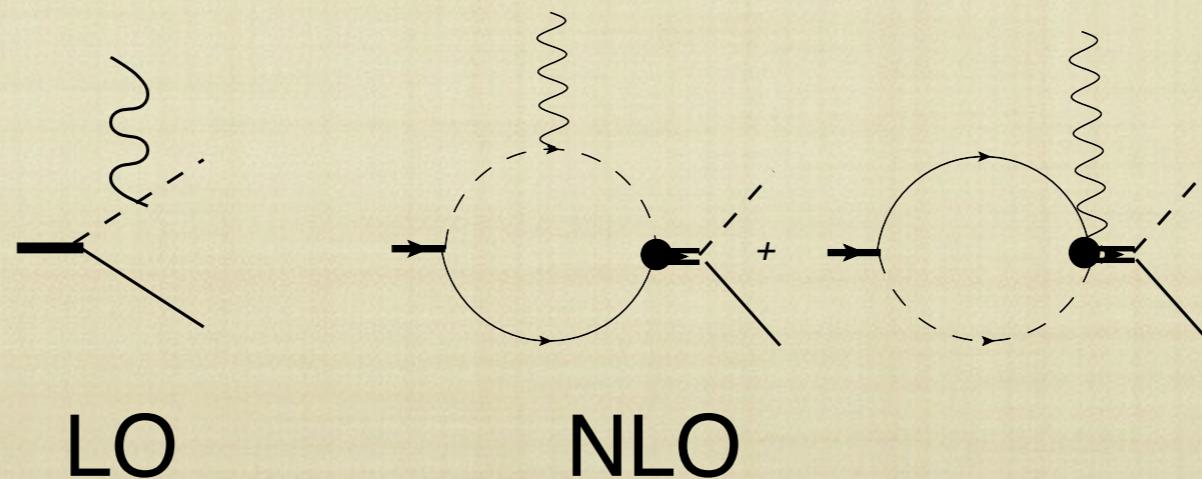
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- FSI in spin-1/2 channel: stronger, but “kinematic” nature of P-wave state means it’s perturbative away from resonance.

Bertulani, Hammer, van Kolck (2002); Bedaque, Hammer, van Kolck (2003)



- Need γ_1 and $r_1 \equiv A_1$ at NLO in this observable. Coulomb dissociation of ${}^{11}\text{Be}$

How-to: three-body system

$$\begin{aligned} -\frac{\hbar^2}{2m_1}\nabla_1^2\Psi-\frac{\hbar^2}{2m_2}\nabla_2^2\Psi-\frac{\hbar^2}{2m_3}\nabla_3^2\Psi+\sum_{i<j}V_{ij}(\mathbf{r}_{ij})\Psi(\mathbf{r}_{ij},\mathbf{r}_{ij,k}) \\ +V_{123}(\mathbf{r}_{ij},\mathbf{r}_{ij,k})\Psi(\mathbf{r}_{ij},\mathbf{r}_{ij,k})=E\Psi(\mathbf{r}_{ij},\mathbf{r}_{ij,k}) \end{aligned}$$

- Remember: most of Ψ occurs outside range of V 's
- Construct two-body and three-body potentials as limiting sequence of functions: you can take whatever's easiest to solve!
- Strength of V_{ij} set to a , strength of V_{ijk} set to lowest 3B binding energy
- Favorite solution method
- Perturbative evaluation of R/a (or kR) corrections: EFT expansion

Universal three-body relations

Universal three-body relations

- Energies of two states: $B_{n+1} = e^{-2\pi/s_0} B_n$
- Features on the Efimov plot: $a_{0,n} = -0.210 a_{-,n}$

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Ann. Phys. (2012)
- Unification via universality: in what ways are all halo nuclei similar?
- Diagnosing via universality: determine unmeasured properties of halo nuclei through universal relationships

Equations for s-wave $2n$ halo

Canham, Hammer (2008)

Equations for s-wave $2n$ halo

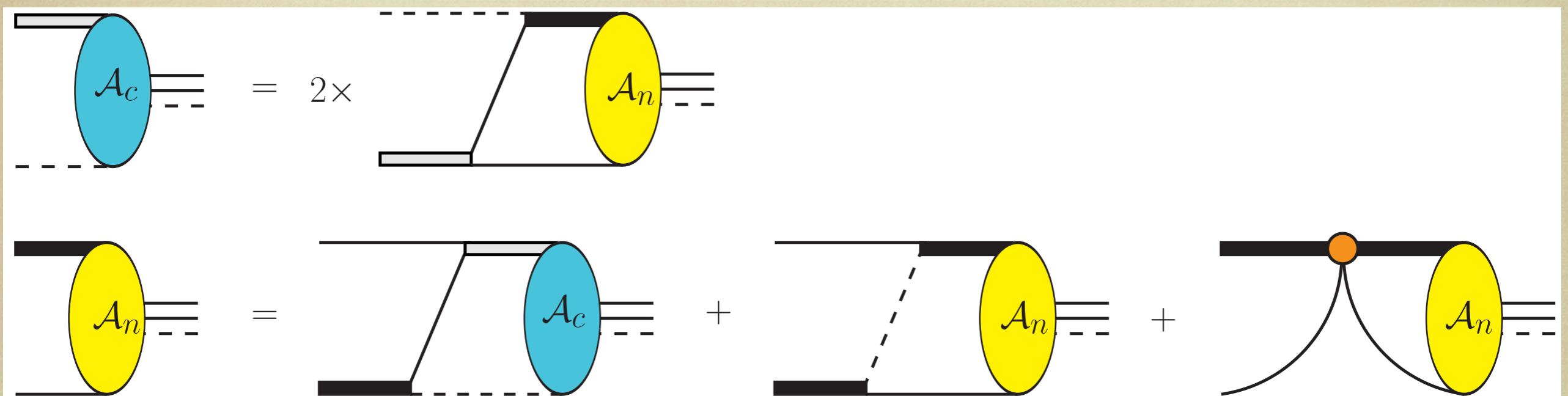
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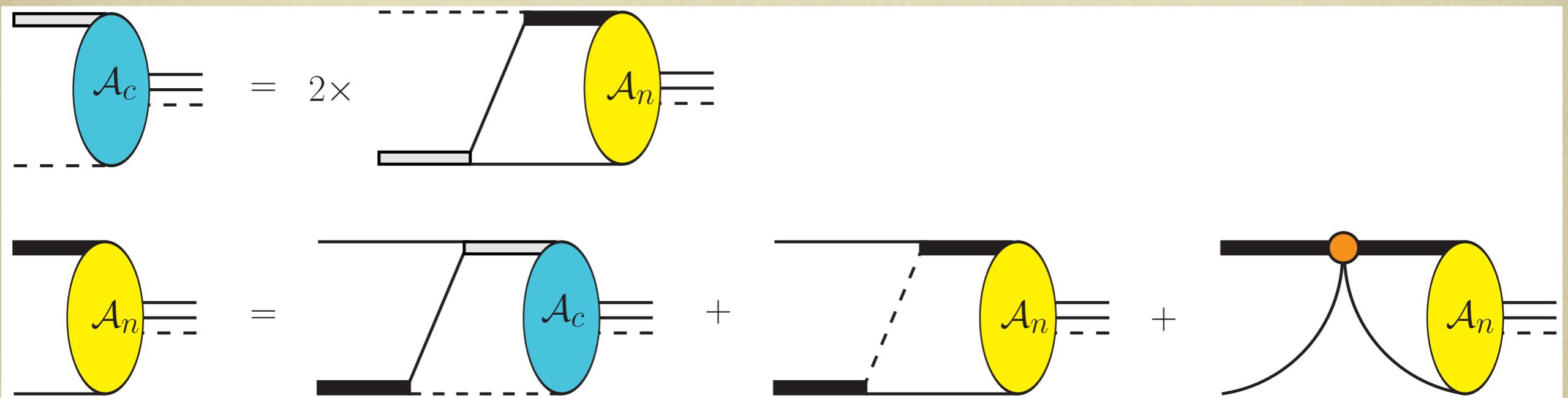
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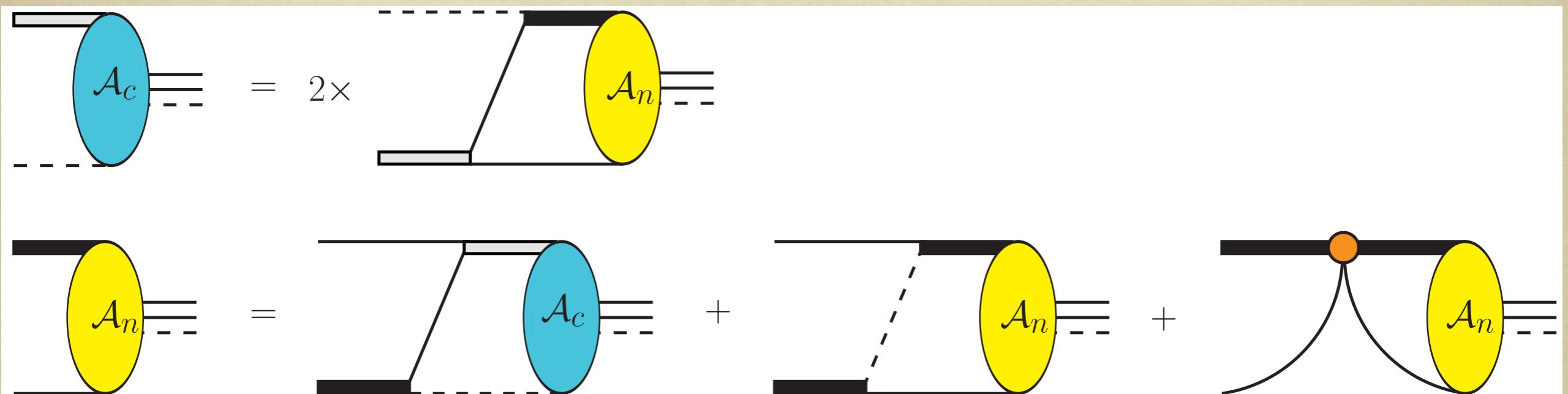


- (cn)-n contact interaction to stabilize three-body system

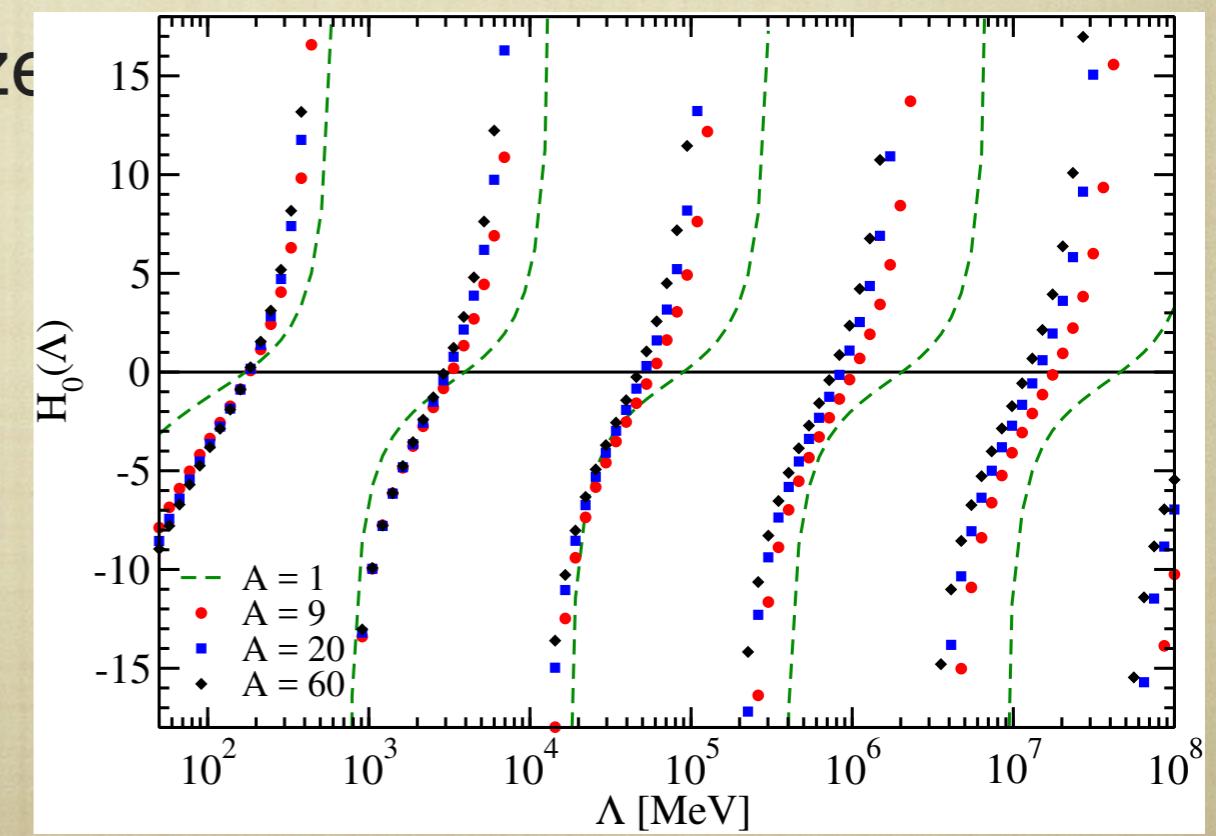
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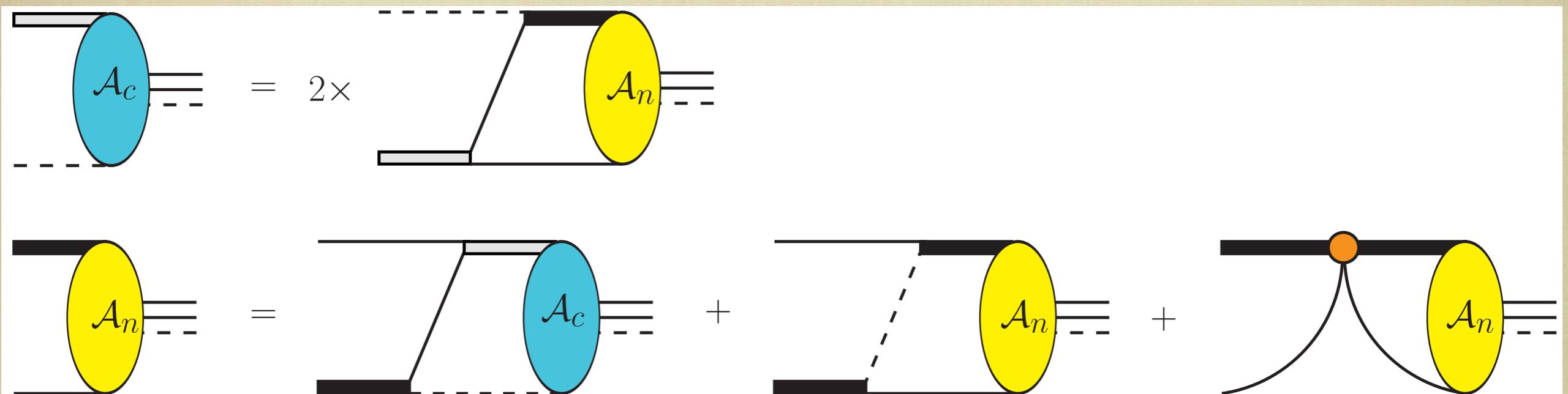
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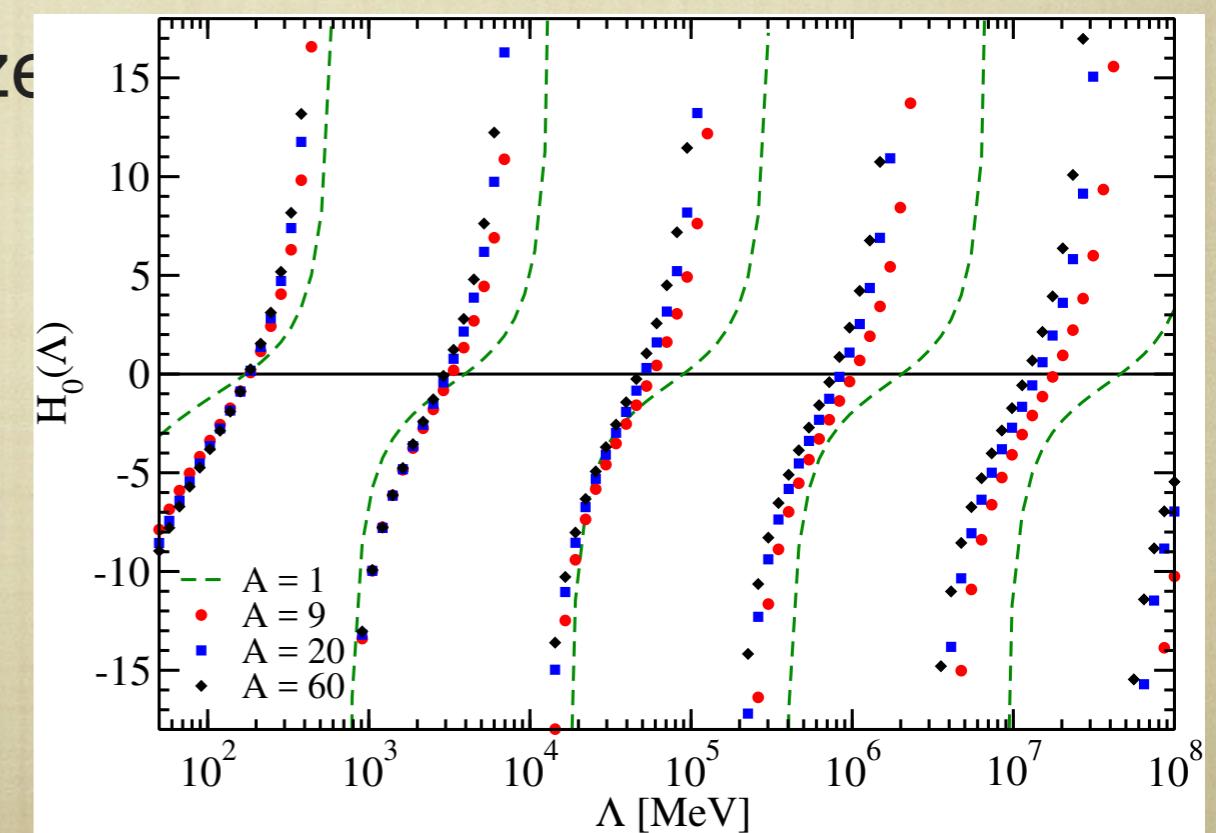
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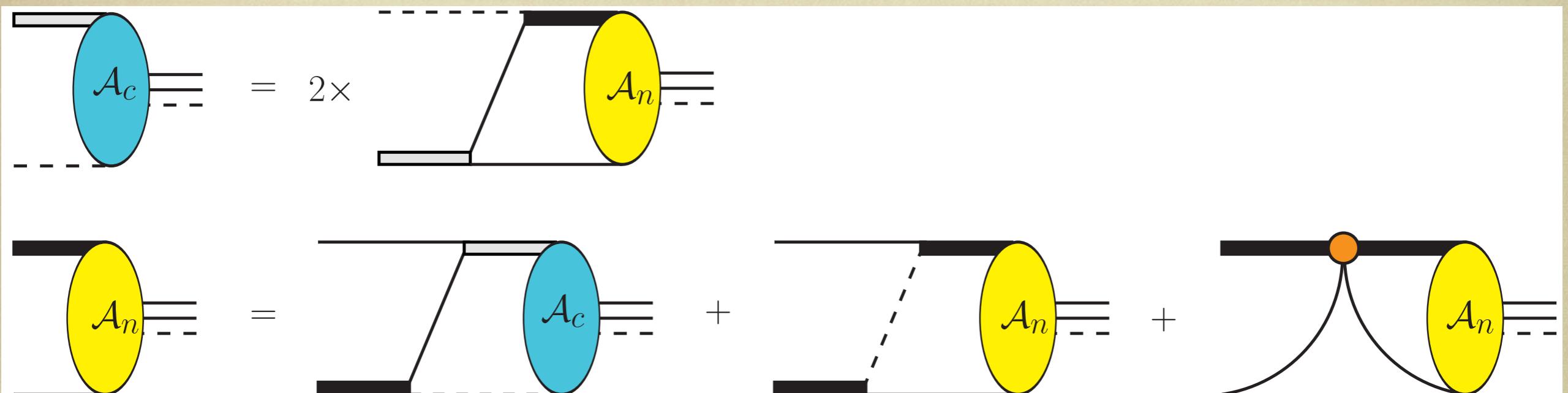
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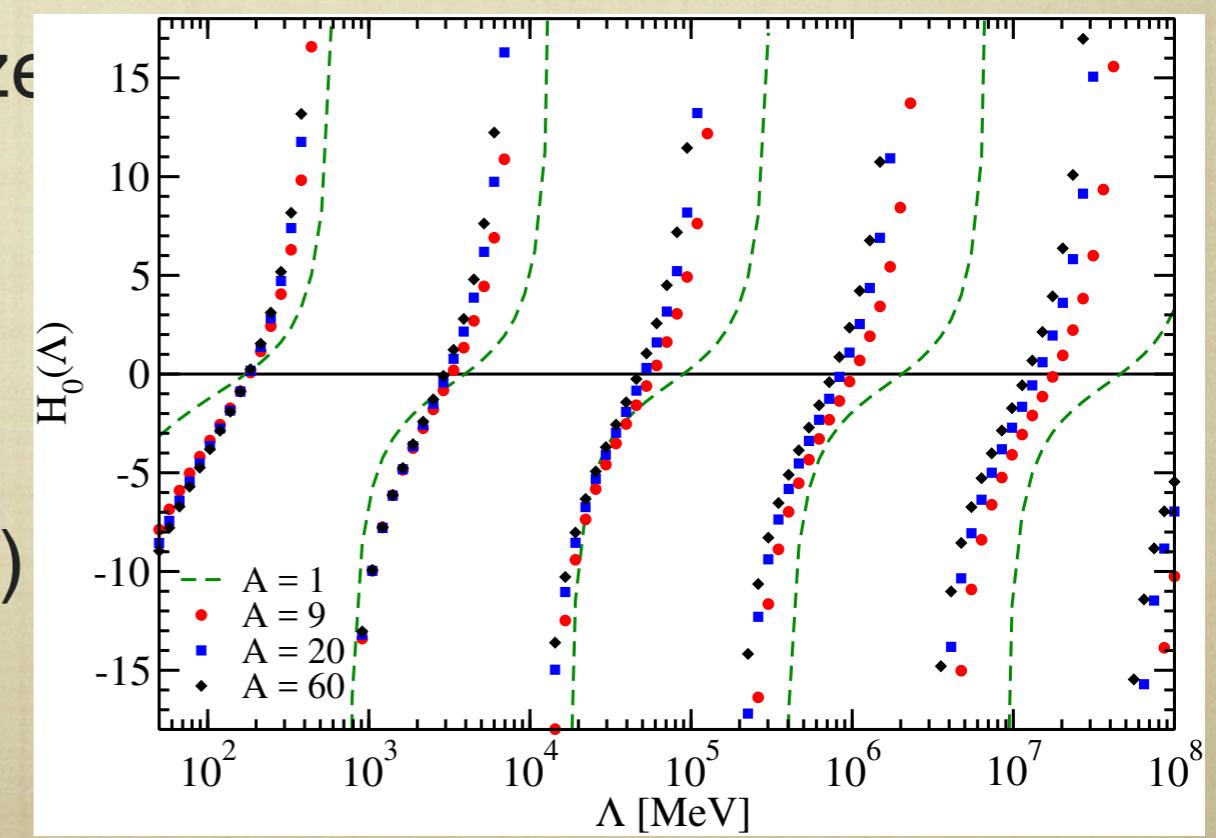
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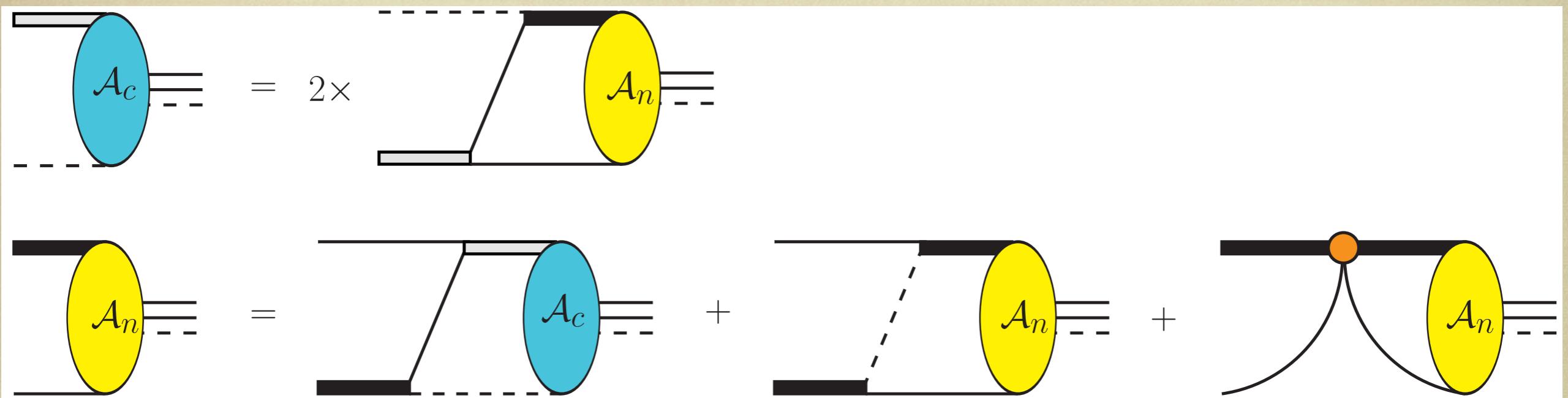
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Equations for s-wave 2n halo

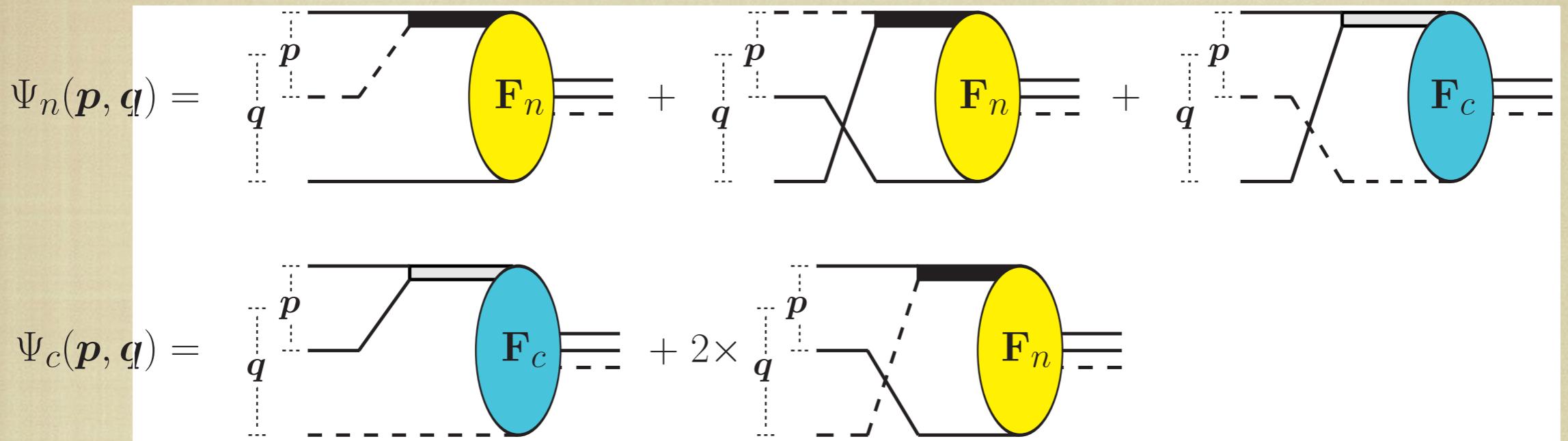
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- (cn)-n contact interaction to stabilize three-body system
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- Inputs: $E_{nn} = 1/(m a_{nn}^2)$, E_{nc} , S_{2n} (=B)
- Output: everything. Up to R_{core}/R_{halo} corrections.

Matter radii of 2n s-wave halos



Canham, Hammer (2008)

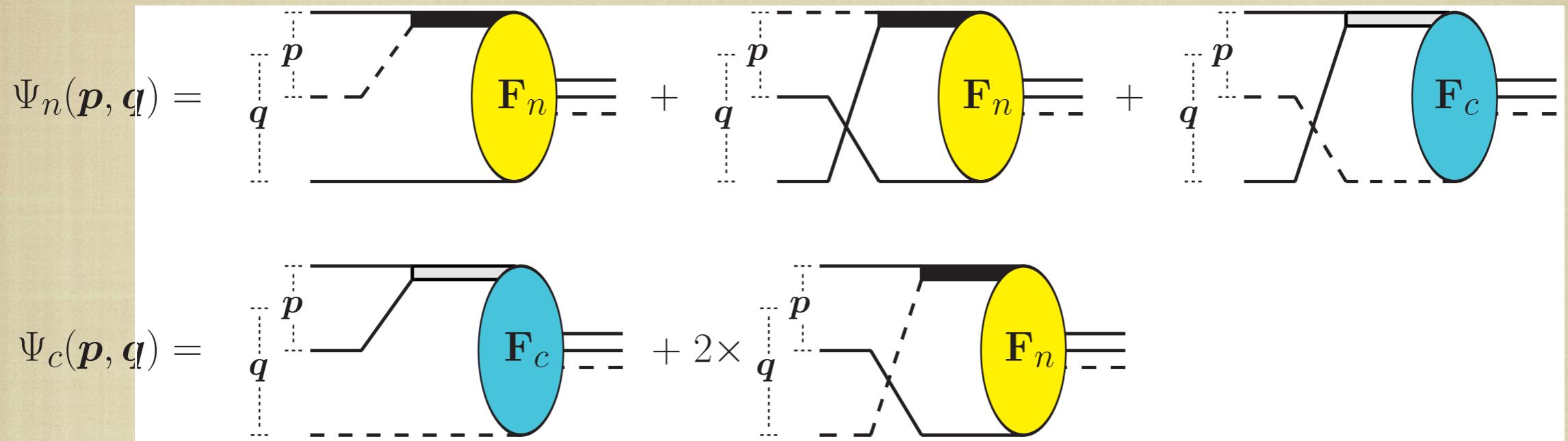
- One-body form factors:

$$\mathcal{F}_x(k^2) = \int_0^\infty dp \, p^2 \int_0^\infty dq \, q^2 \int_{-1}^1 d(\hat{q} \cdot \hat{k}) \, \Psi_x(p, q) \, \Psi_x(p, |\vec{q} - \vec{k}|).$$

- Radii: $\mathcal{F}_x(k^2) = 1 - \frac{1}{6} \langle r_x^2 \rangle k^2 + O(k^4)$

- Matter radius: $\langle r_0^2 \rangle = \frac{2(A+1)^2}{(A+2)^3} \langle r_n^2 \rangle + \frac{4A}{(A+2)^3} \langle r_c^2 \rangle$

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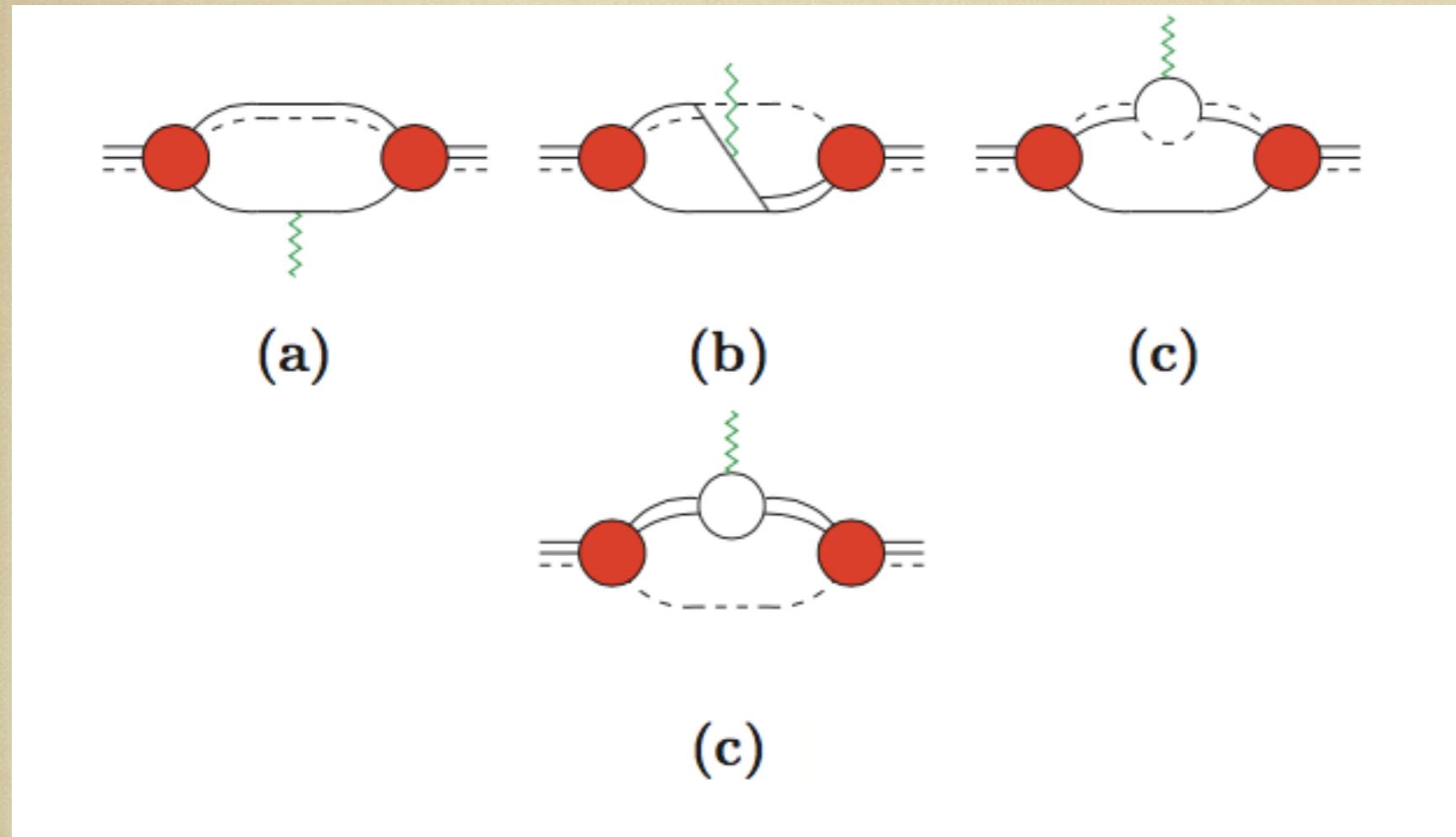
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Input: E_{nn} , E_{cn} and S_{2n}
Output: all radii

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...or use field theory

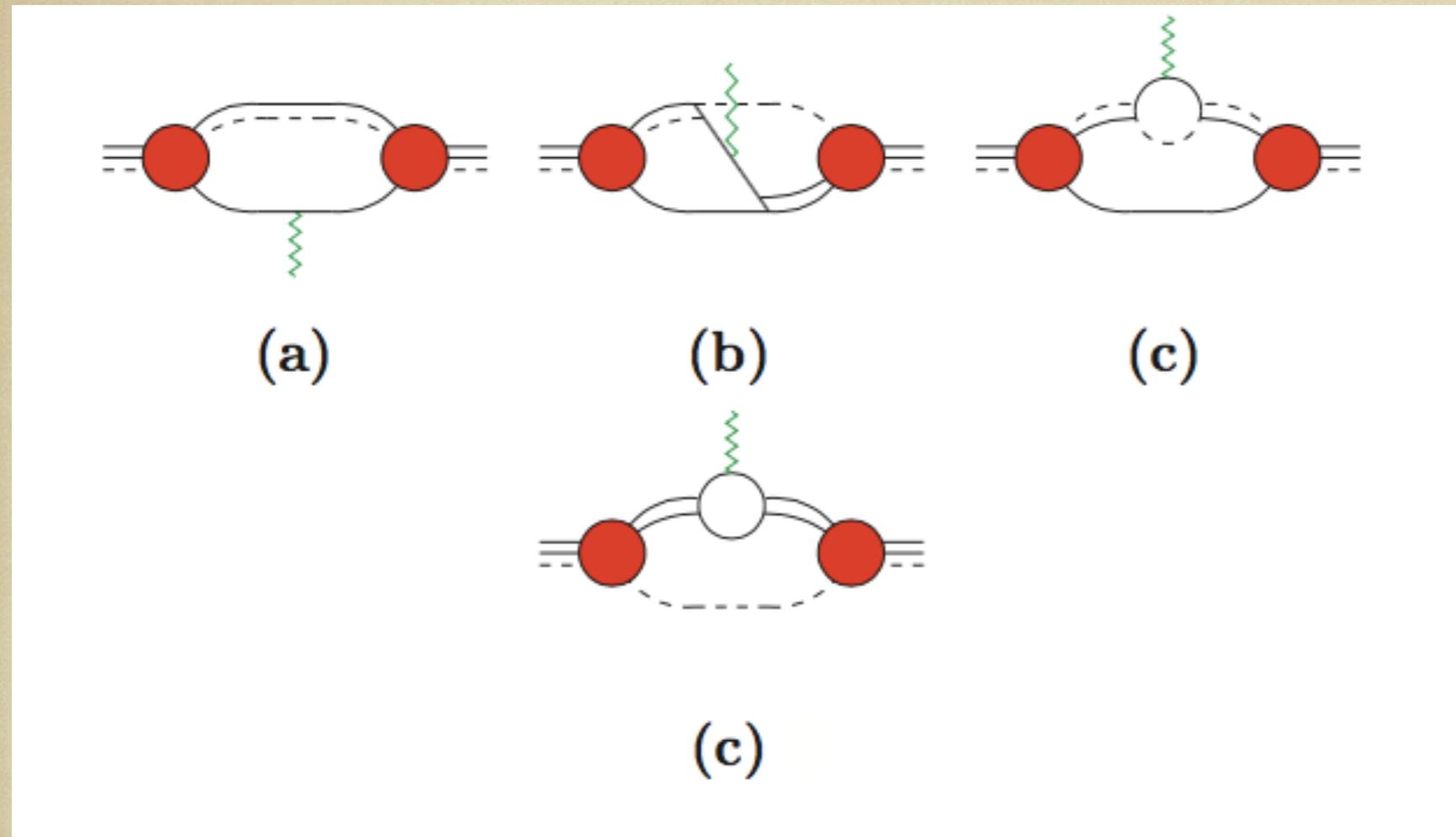
Hagen et al., (2013), Vanasse arXiv:1609.08552



- Same Lagrangian as shown before for S-waves, now with three-body force added
- Introduce “trimer” field to compute 3B state properties

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Matter radii of $2n$ halos at LO

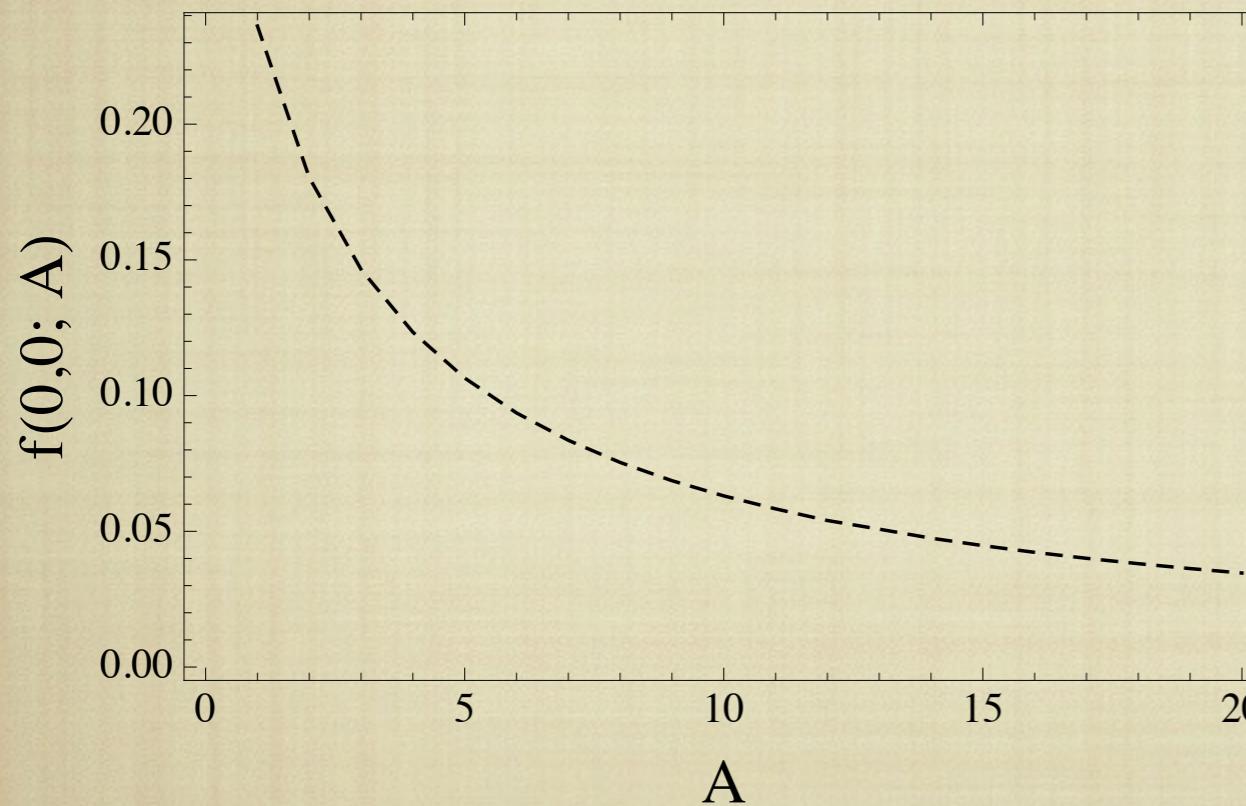
Acharya, Ji, Phillips (2013)

- Define: $f\left(\frac{E_{nn}}{B}, \frac{E_{nc}}{B}; A\right) \equiv mB\langle r_0^2 \rangle$
- Unitary limit, $E_{nn}=E_{nc}=0$: f becomes a number depending solely on A

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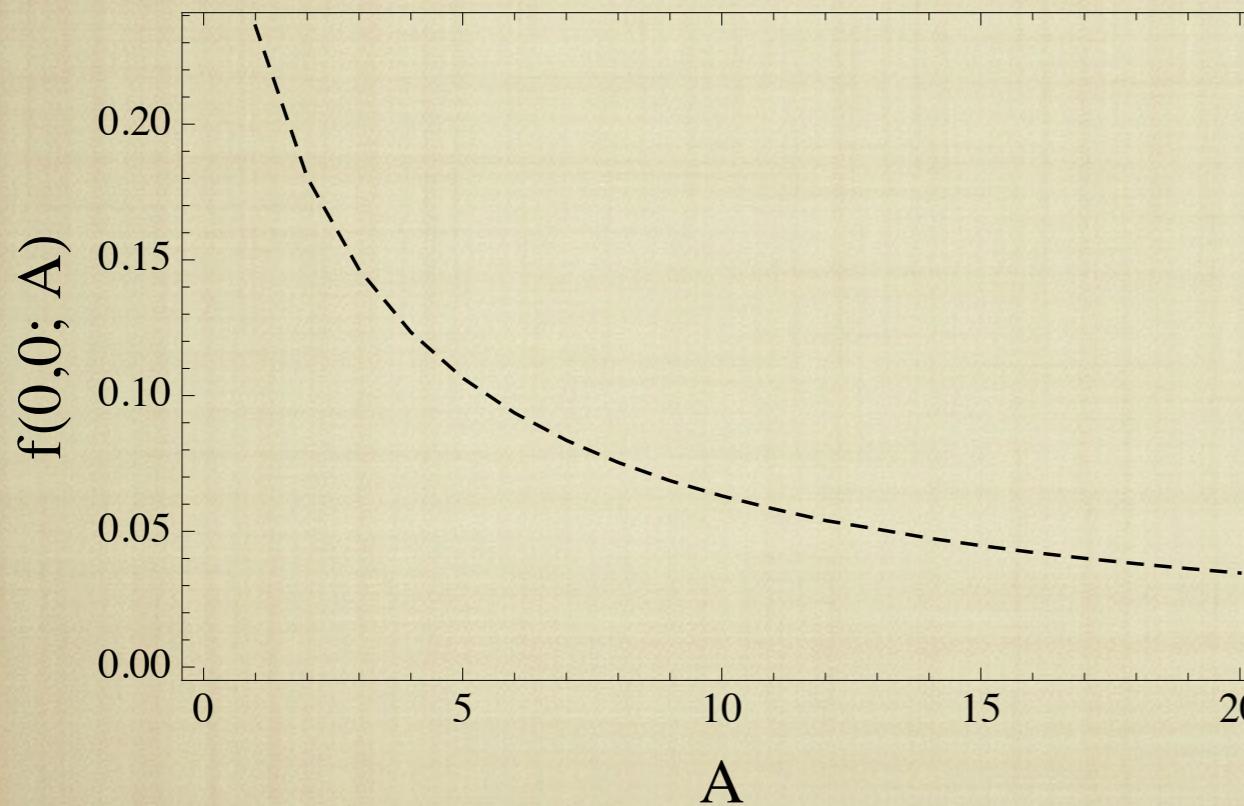
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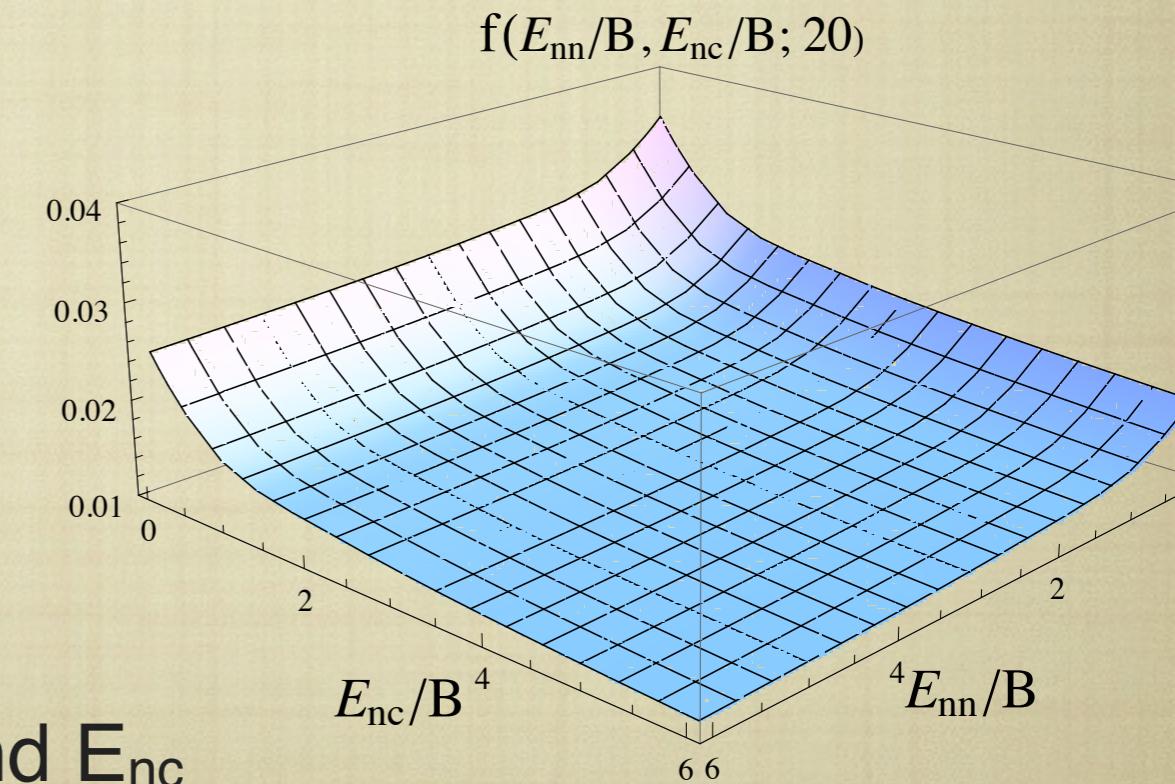
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- Fix $A=20$, plot f as a function of E_{nn} and E_{nc}

LO results for radii of 2n halos

Canham, Hammer (2011); Hagen, Platter, Hammer (2014); Acharya, Ji, Phillips (2013)
Vanasse (2016)

Are these systems “universal enough”?

	E_{nc} (MeV)	S_{2n} (MeV)	R_{core}/R_{halo}	$\langle r_0^2 \rangle$ (fm ²) LO	$\langle r_0^2 \rangle$ (fm ²) Expt
¹¹ Li	-0.026(13)	0.3693(6)	0.37	5.76 ± 2.13	5.34 ± 0.15
¹⁴ Be	-0.510	1.27(13)	0.78	1.23 ± 0.96	4.24 ± 2.42 2.90 ± 2.25
²² C	-0.01(47)	0.11(6)	0.26	$3.99 - \infty$	21.1 ± 9.7 3.77 ± 0.61

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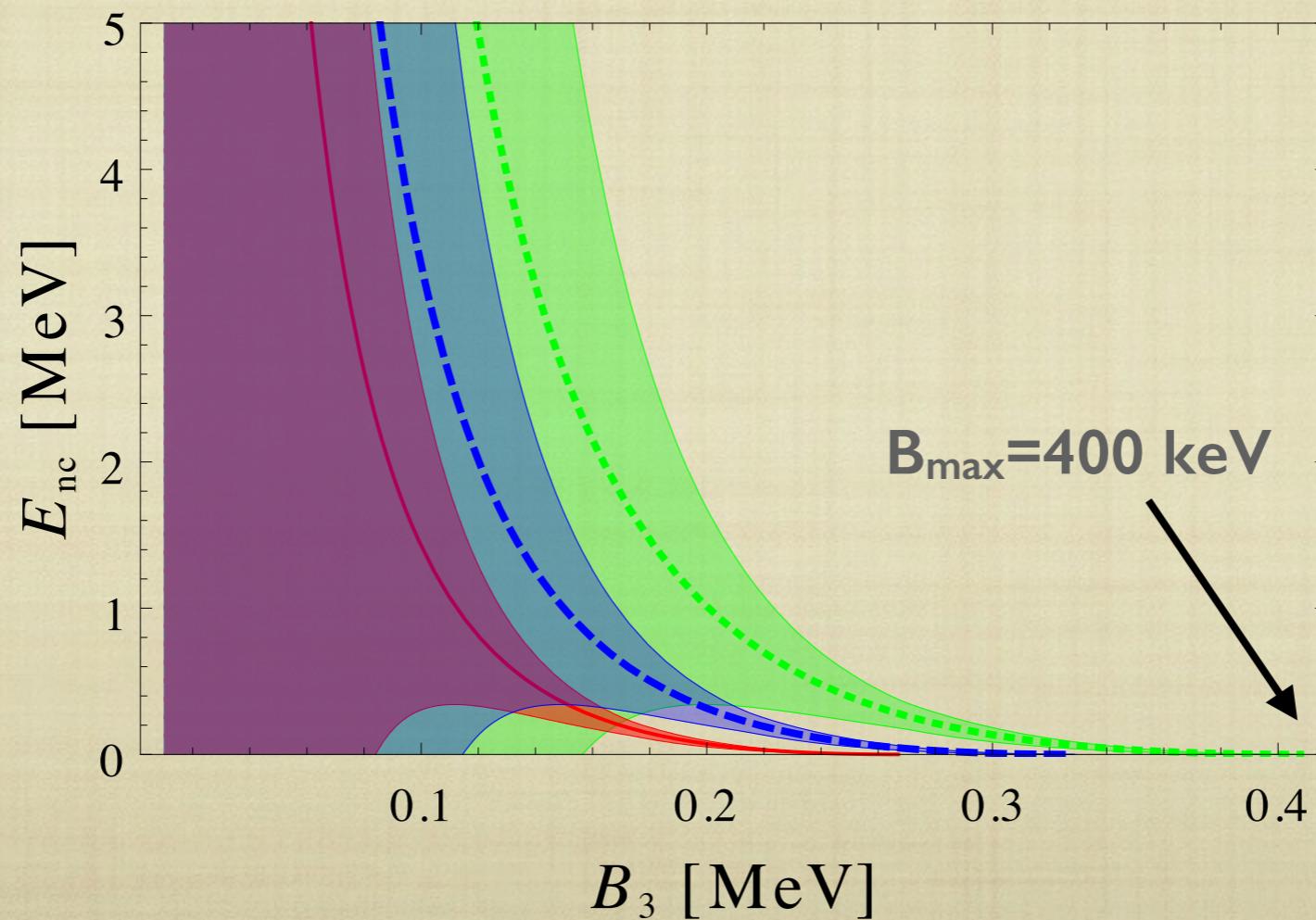
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Errors tend to be dominated by EFT uncertainty \Rightarrow need ranges
to become more accurate

Application to ^{22}C

- Include finite size of ^{20}C
- Consider uncertainty due to NLO effects:

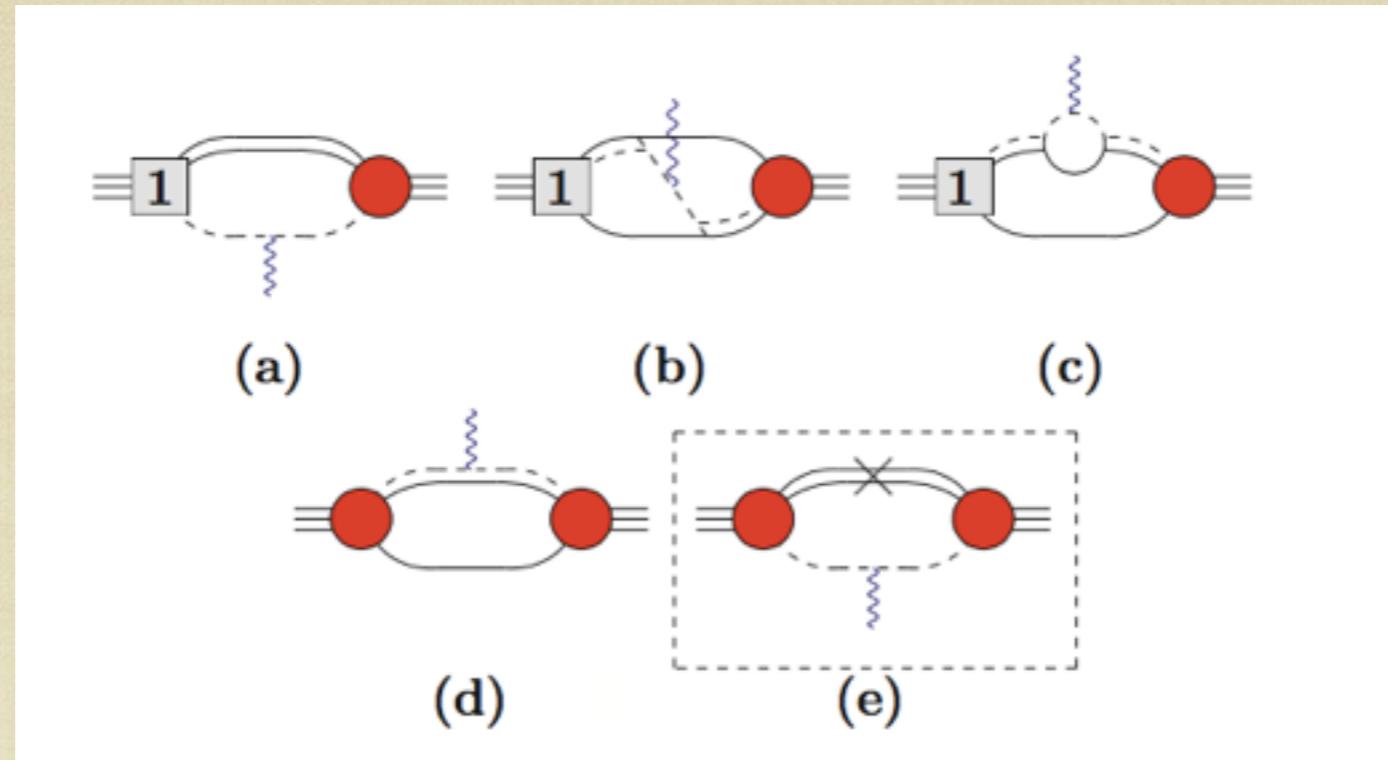
Relative size \sim largest of $(mE_{nn})^{1/2} R_{\text{core}}$; $(2mE_{nc})^{1/2} R_{\text{core}}$; $(2mB)^{1/2} R_{\text{core}}$



cf. Yamashita et al. (2011);
Fortune & Sherr (2012)

Next-to-leading order

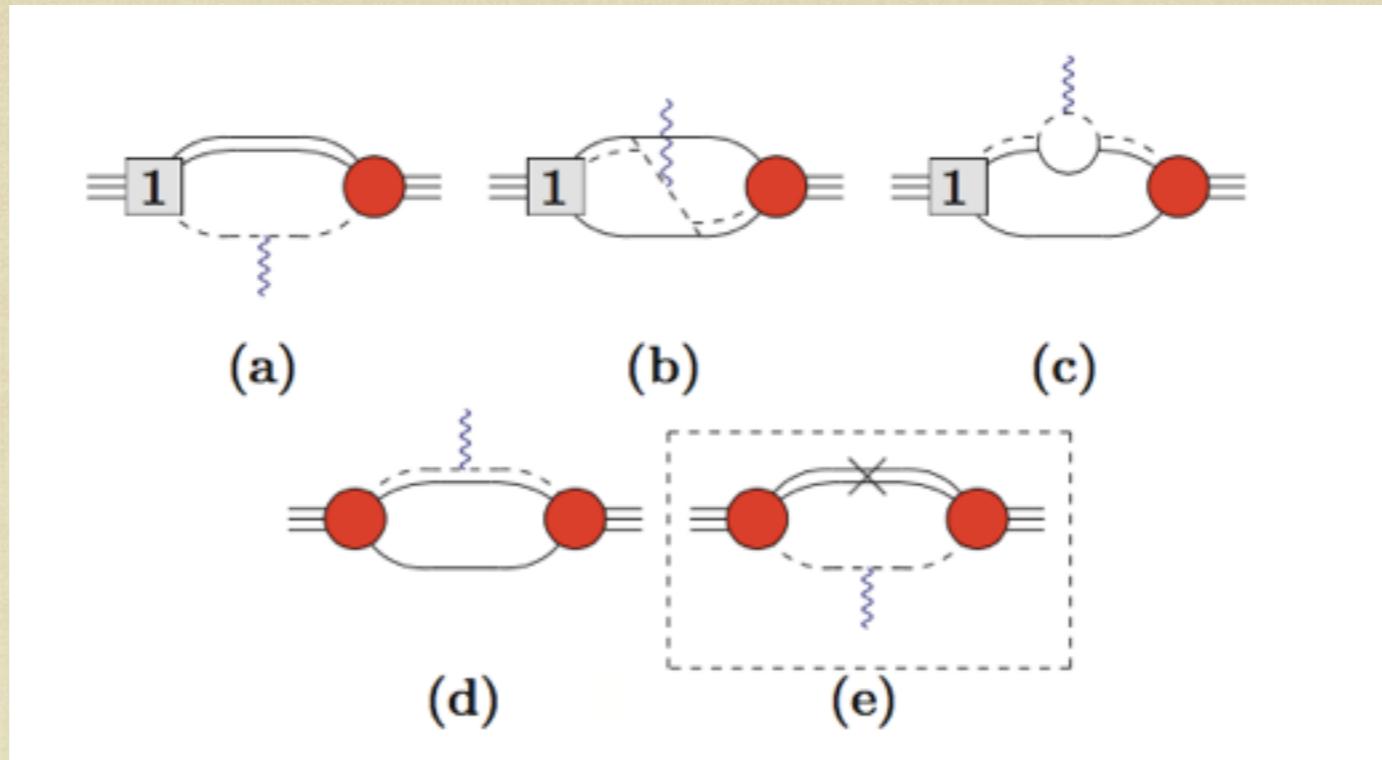
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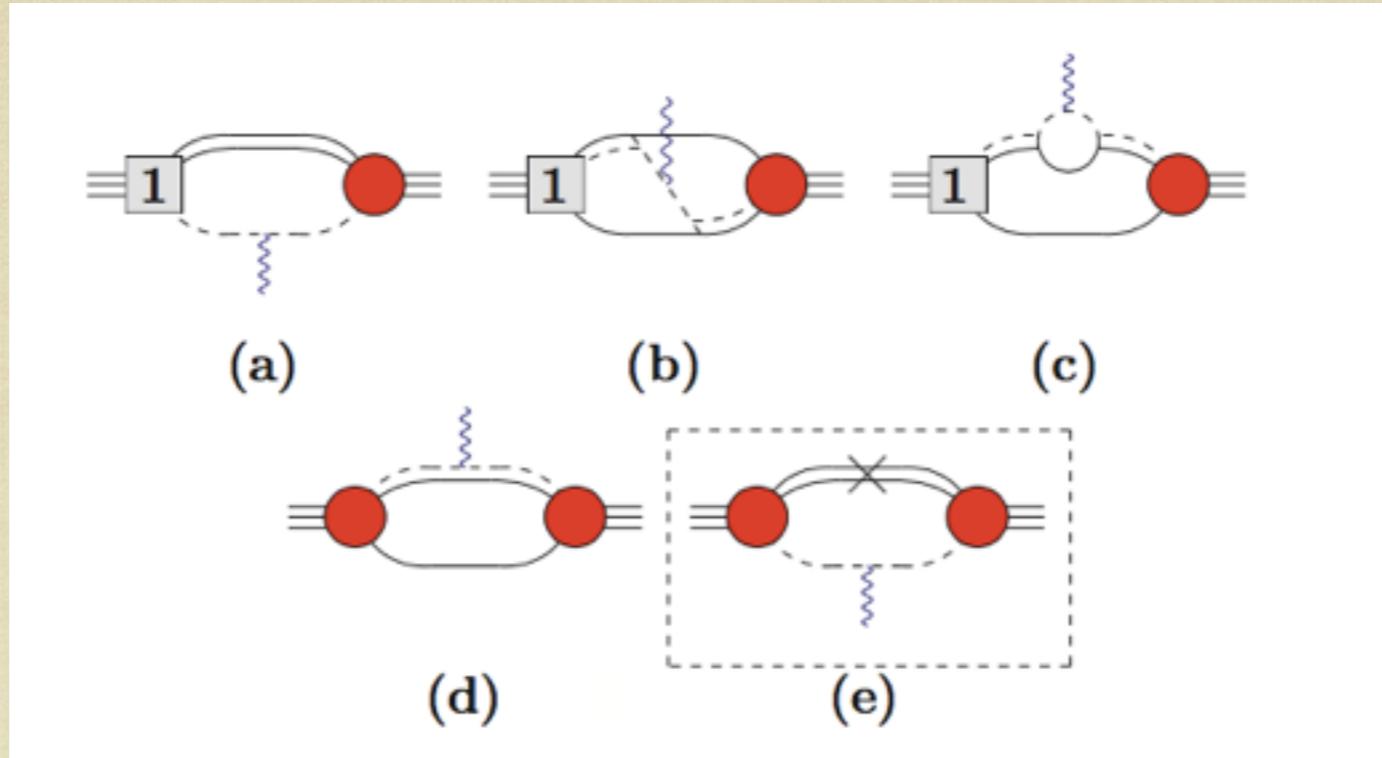
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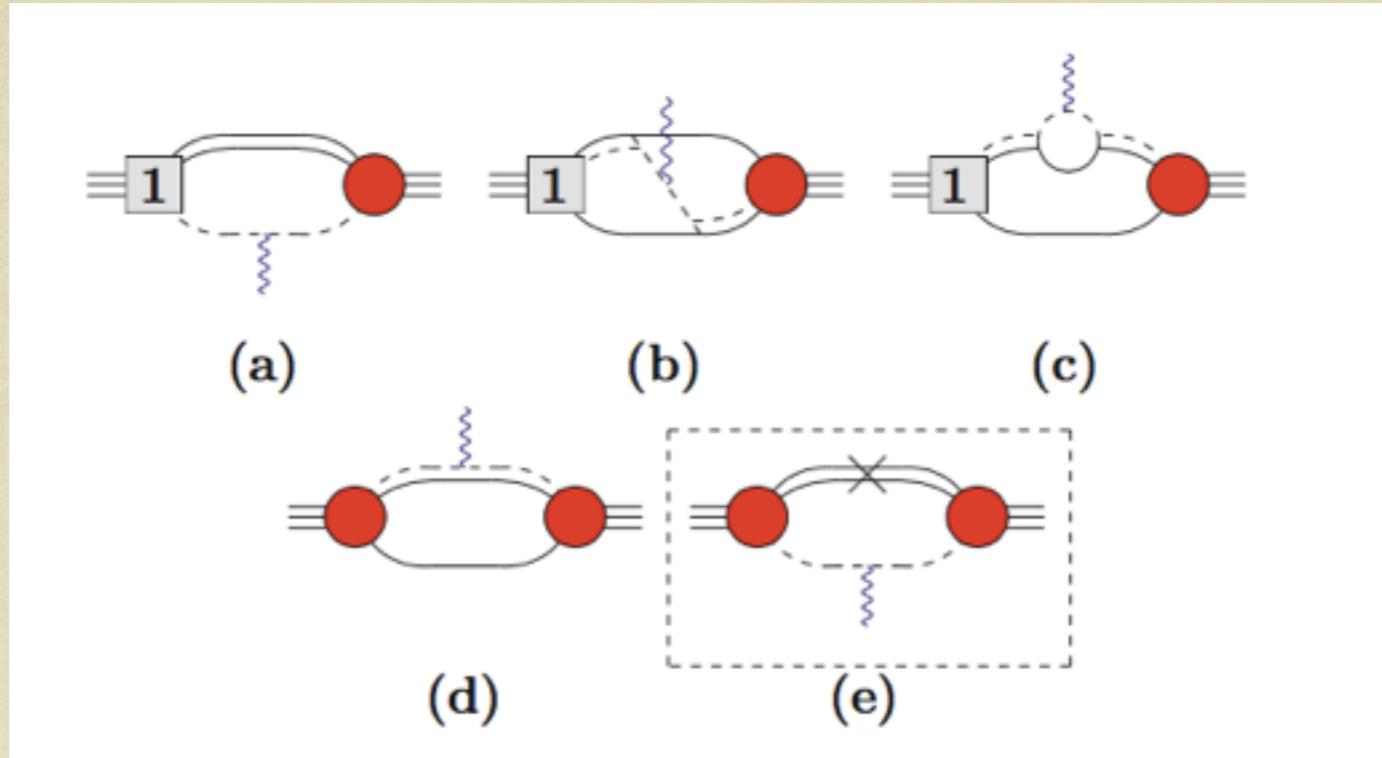
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LO NLO NNLO

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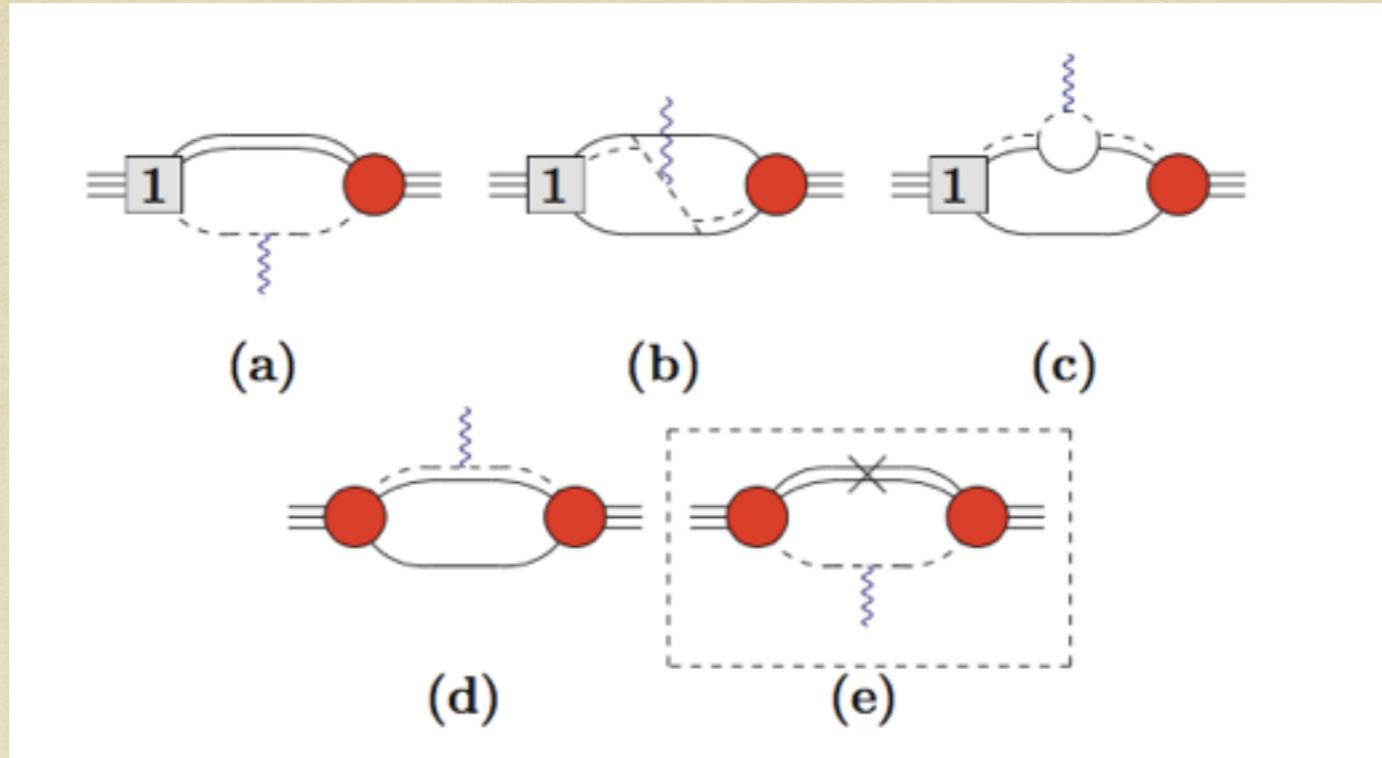
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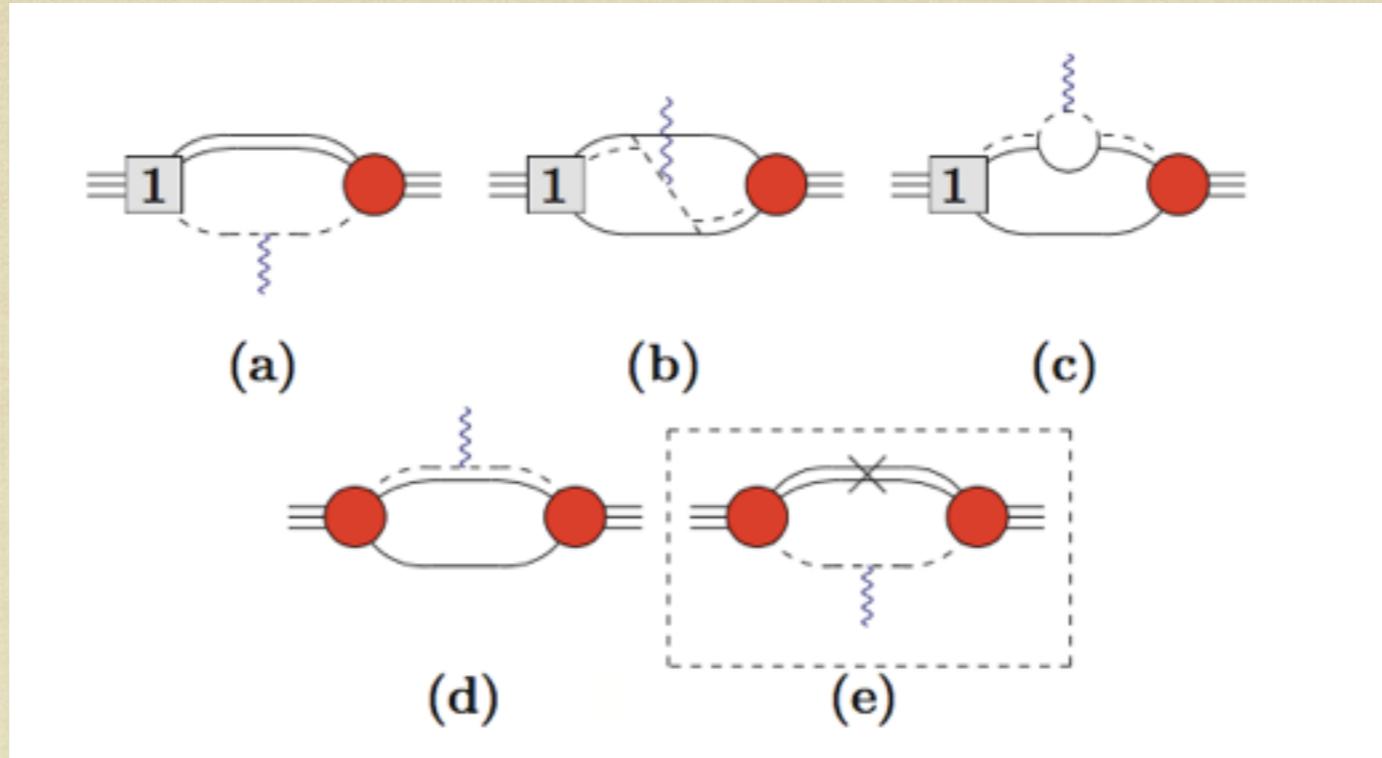
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Vanasse & DP, Few-body Systems, to appear

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- SU(4) limit gives similarly good description of radii for ${}^3\text{H}$ and ${}^3\text{He}$
- Need to know ranges for ${}^{11}\text{Li}$, ${}^{14}\text{Be}$, ${}^{22}\text{C}$; estimates mostly move EFT prediction closer to data.

Vanasse & DP, Few-body Systems, to appear

Photodissociation of trimers

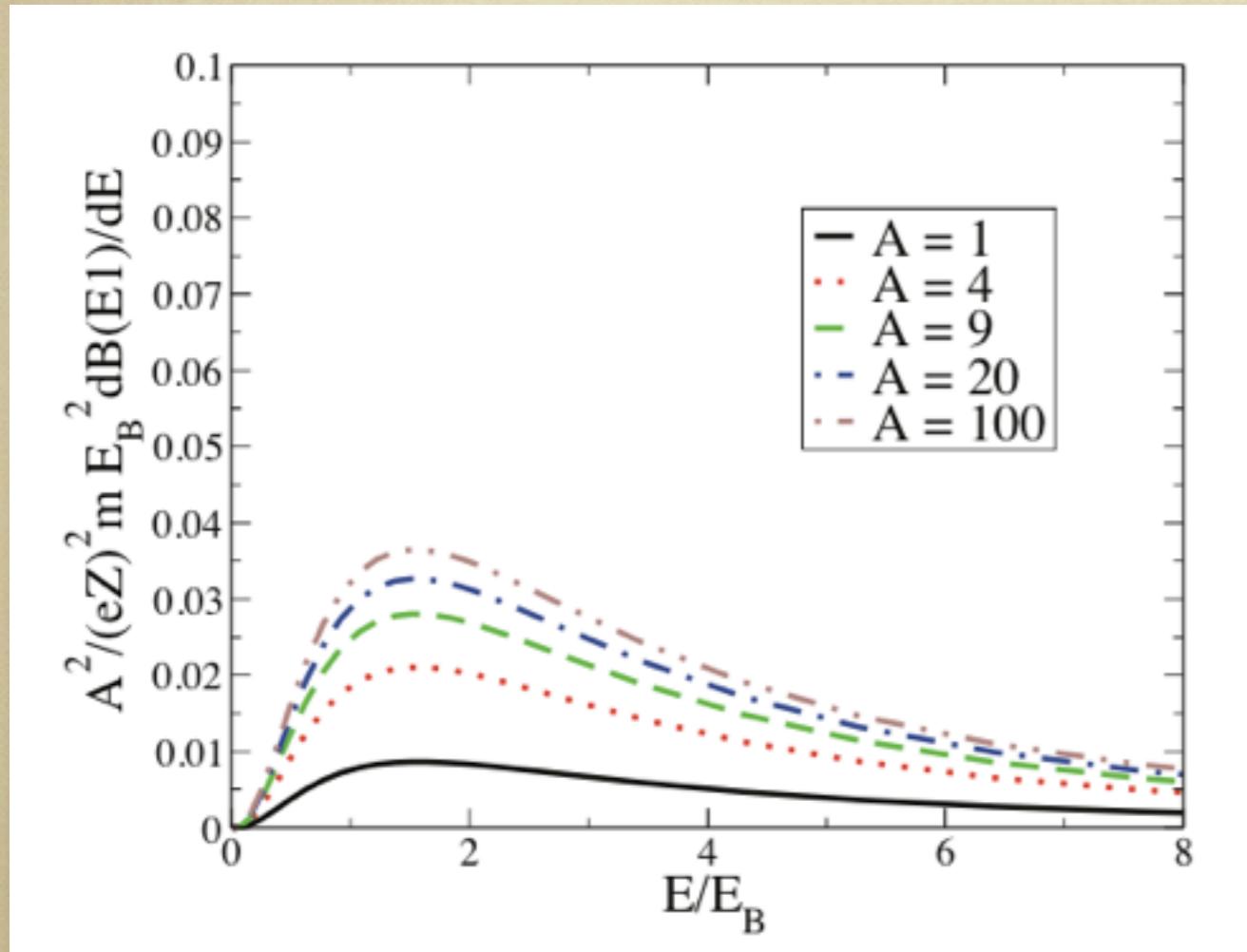
Acharya, Hagen, Hammer, Phillips , in preparation

- Look at E1 dissociation of Borromean core-neutron-neutron system into three particles
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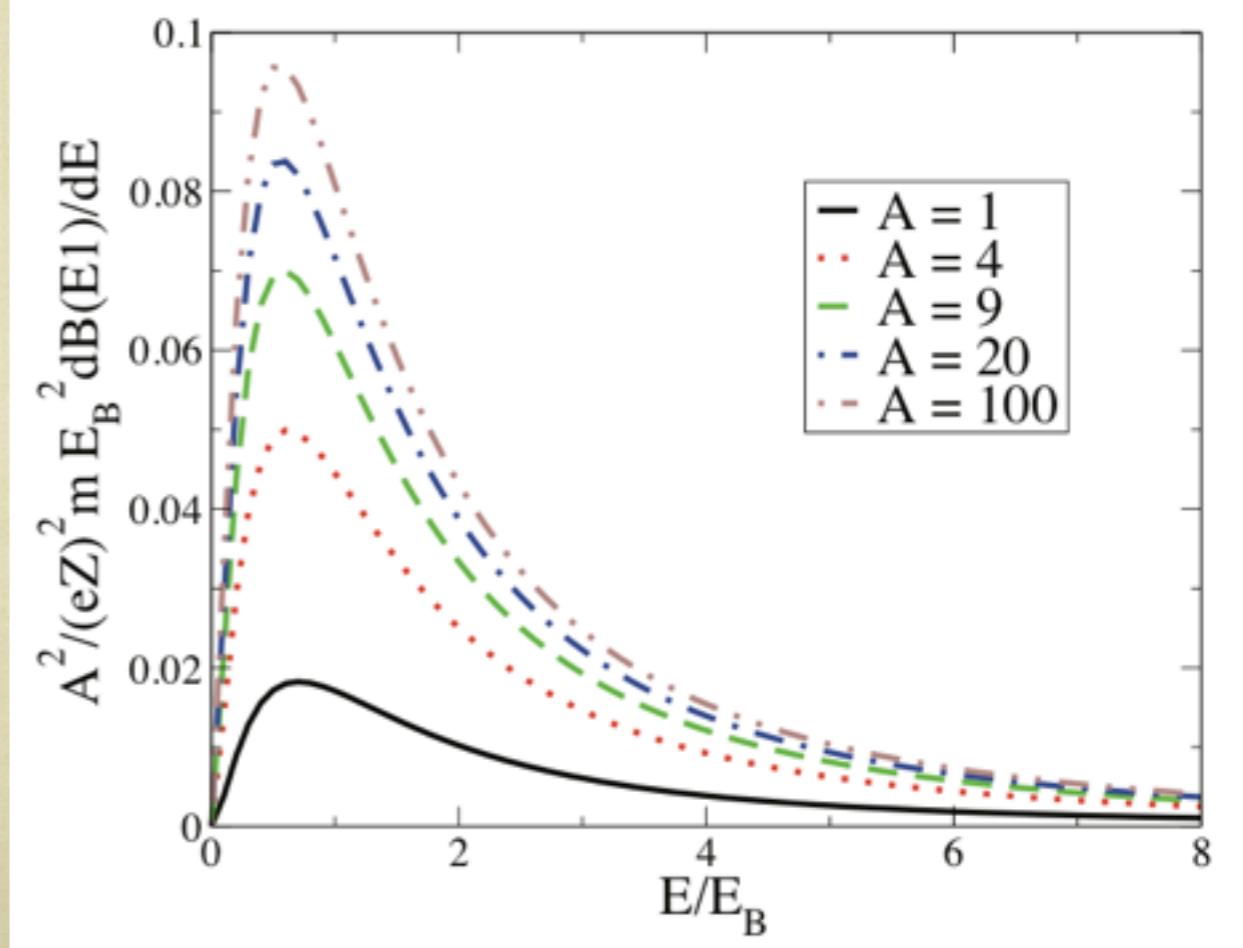
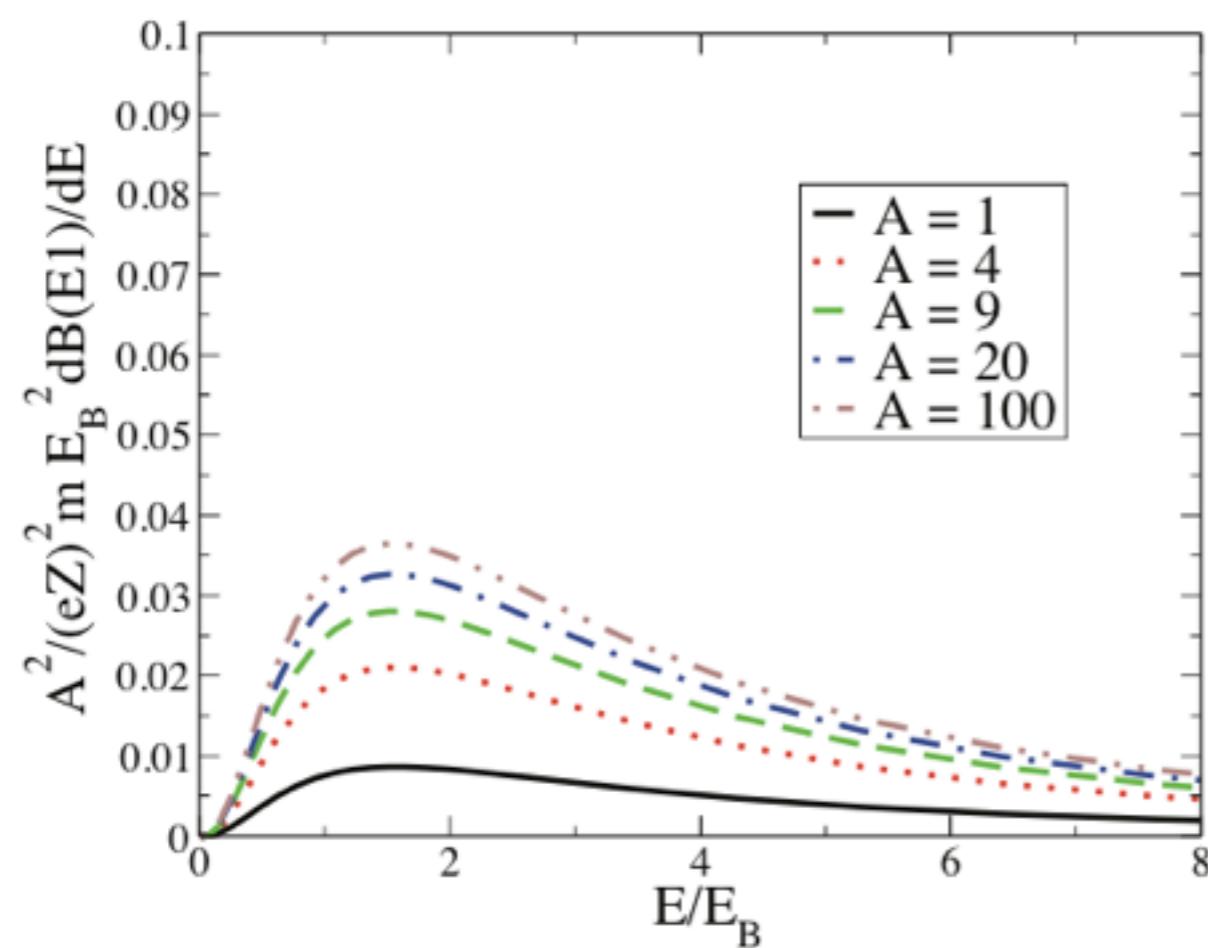
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Acharya, Hagen, Hammer, Phillips , in preparation

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- Go to unitary limit $E_{nn}=E_{nc}=0$. Only scales are B and energy of outgoing particles



Photodissociation of ^{11}Li

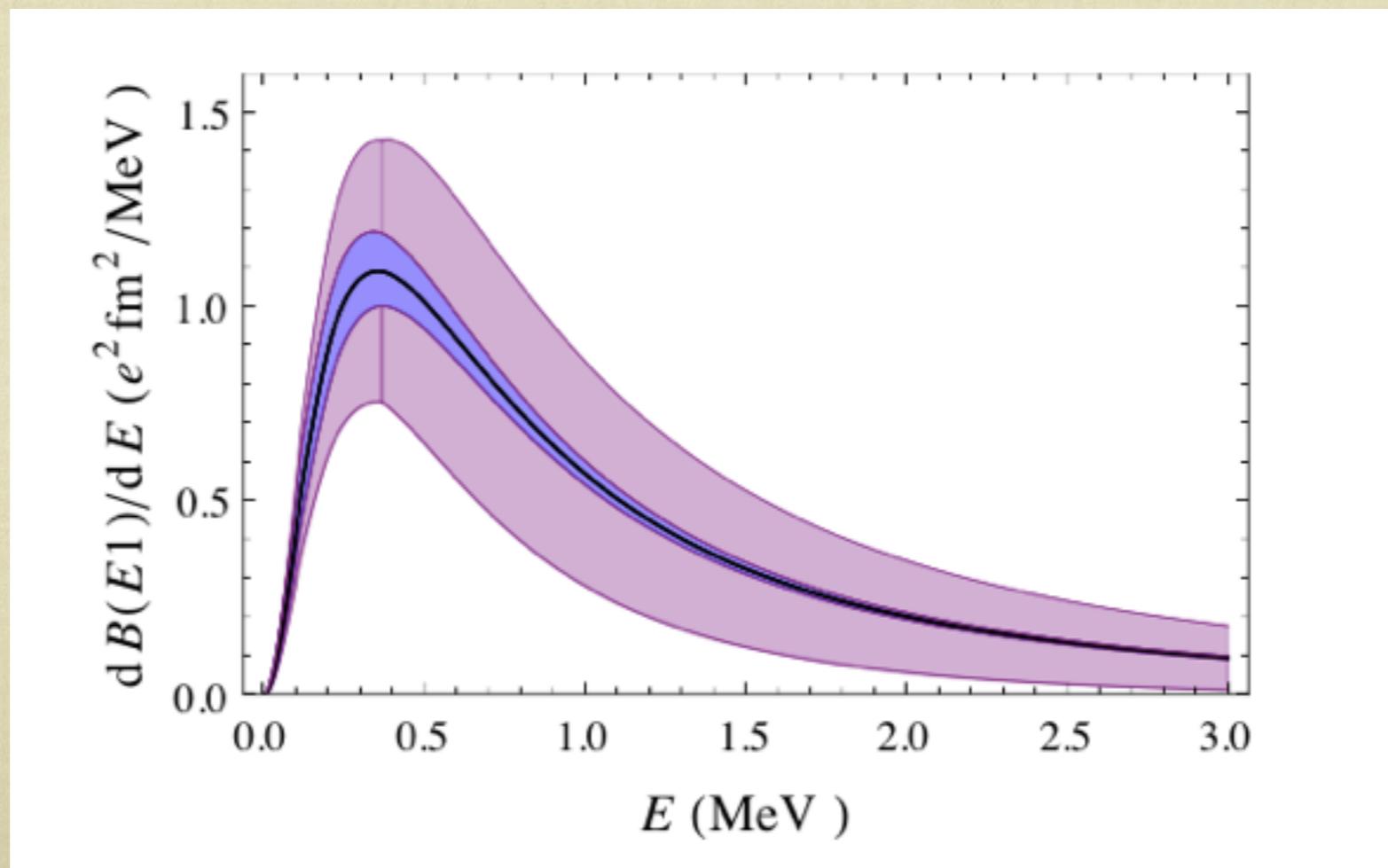
Acharya, Phillips , in preparation

- ^{11}Li as Borromean system with S-wave interactions.
- $E_{\text{nc}}=26 \text{ keV}$, $E_{\text{nn}}=118 \text{ keV}$, $B=369 \text{ keV}$. Breakdown $\approx 70 \text{ MeV}$

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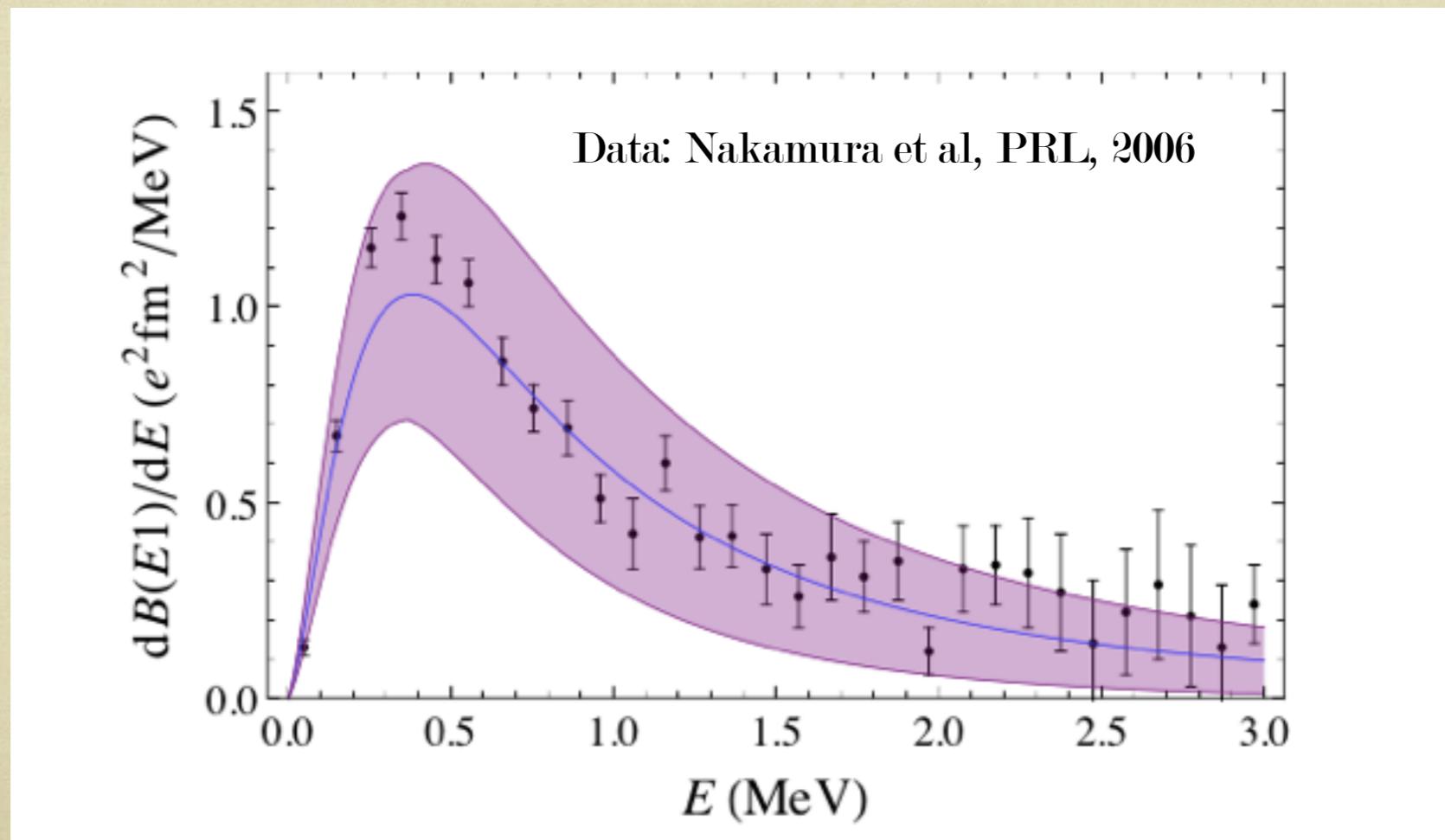
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Conclusion

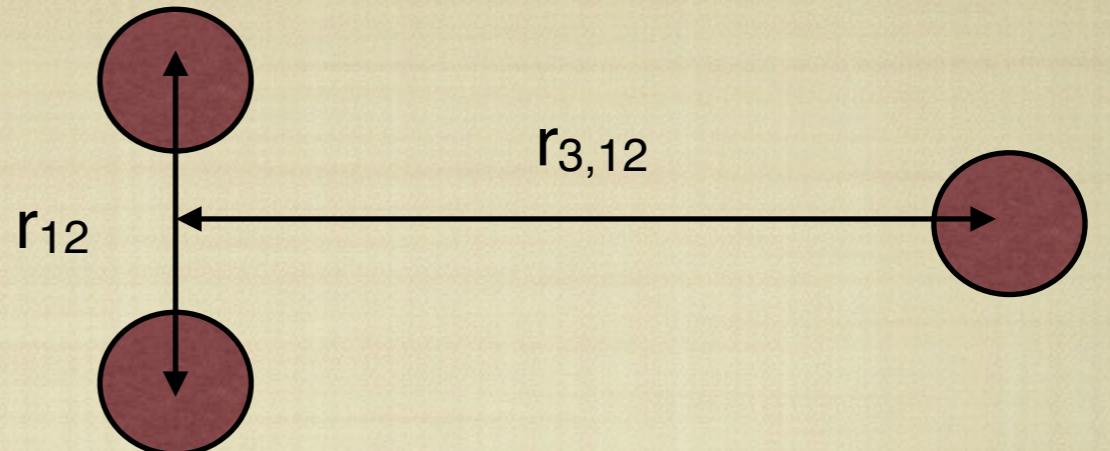
- Universality: quantum few-body systems with $R \ll l_{\text{al}}$ that differ in scale by orders of magnitude exhibit the same correlations
Correlations between three- and four-body systems...and beyond
- Short-range EFT: expand observables in $r/a, kr$
- Two-body: compute matter radii, photodissociation cross sections
- Three-body: halo radii in terms of B, E_{nc}, E_{nn} : ^3H (NNLO), ^{22}C (LO)
- Photodissociation (aka Coulomb excitation): do halo nuclei approach the E1 response of universal trimers?
- Range effects are sizable: do such systems still exhibit universality?
- p-waves are another (controllable) source of universality violation

Backup slides: Efimov effect

Efimov effect

Efimov, 1970

Consider three-body problem in
limit $R \rightarrow 0$, $|a| \rightarrow \infty$

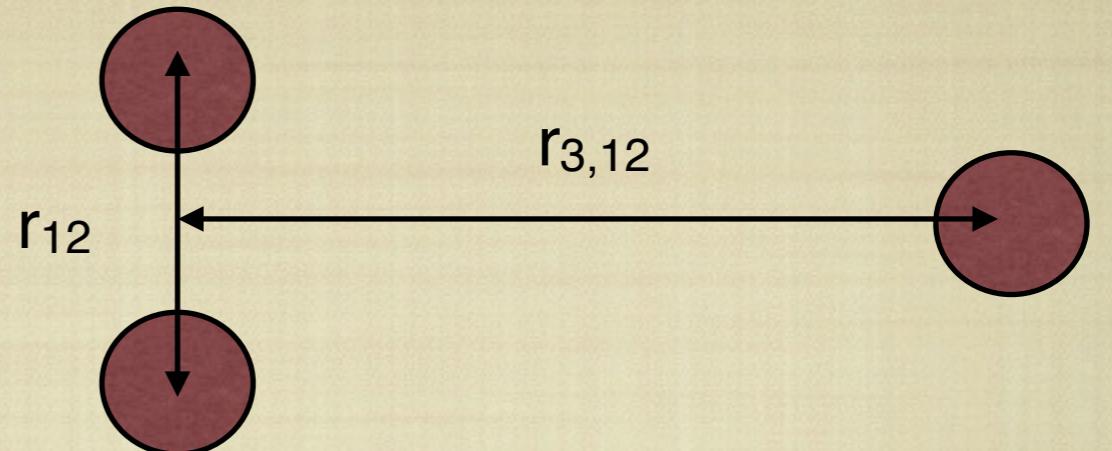


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Consider three-body problem in
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- Spectrum unbounded from below



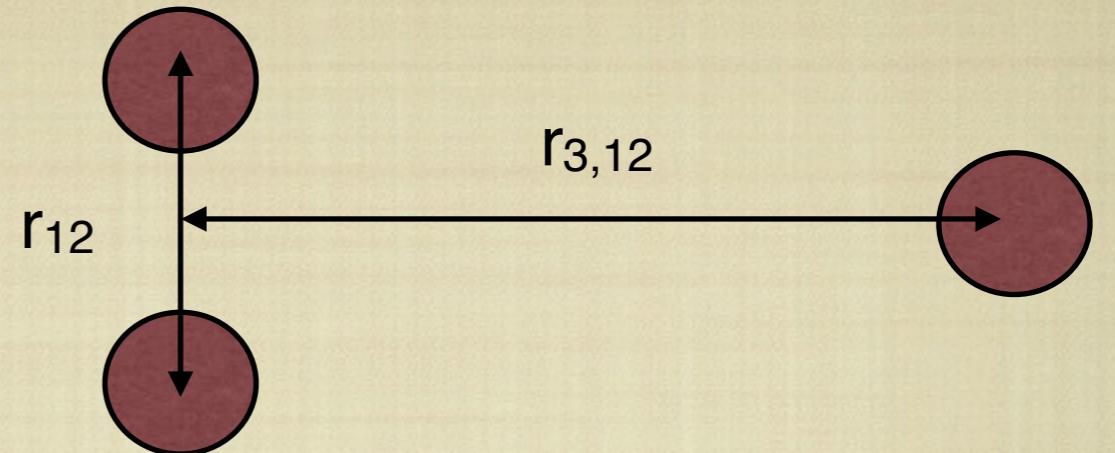
Thomas, 1935

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Thomas, 1935

$$\rho = (r_{12}^2 + r_{3,12}^2)^{1/2}$$

$$\tan \alpha = r_{3,12} / r_{12}$$

3B Schrödinger equation:

$$-\frac{\hbar^2}{2m} \frac{d^2\Psi}{d\rho^2} + V_{\text{eff}}(\rho)\Psi(\rho) = E\Psi(\rho)$$

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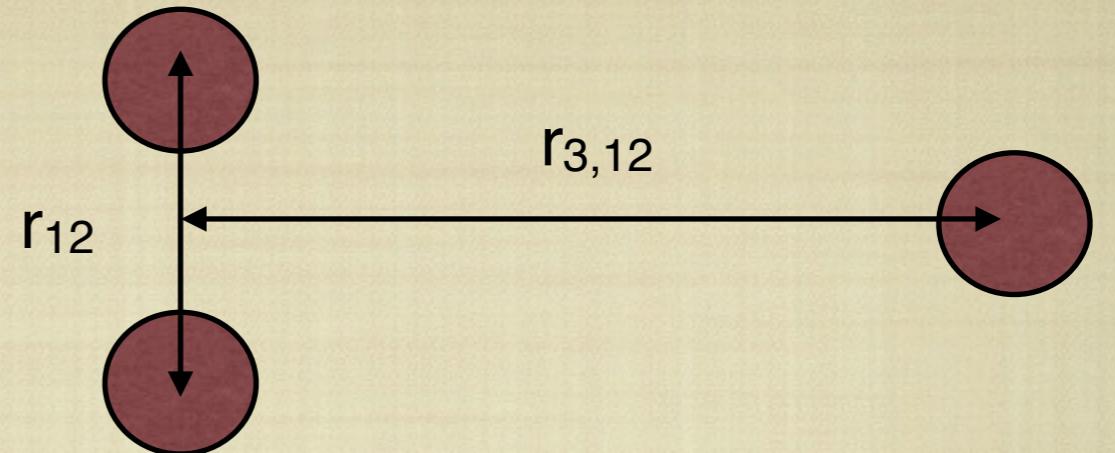
- Spectrum unbounded from below

Thomas, 1935

- $V_{\text{eff}}(\rho) = -s_0(s_0 + 1)/\rho^2$ for $R \ll \rho \ll |a|$

$$\rho = (r_{12}^2 + r_{3,12}^2)^{1/2}$$

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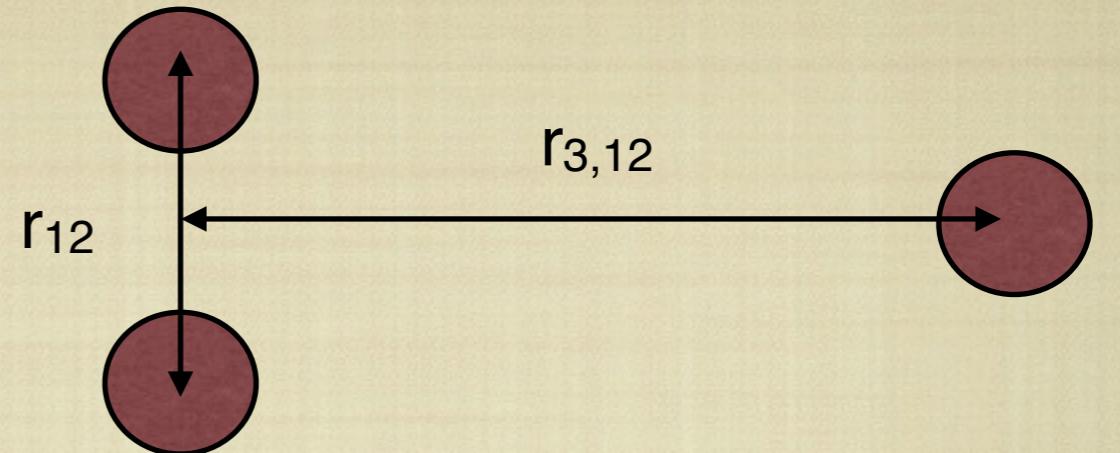
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- Value $s_0 = 1.0062$ by matching at small a

$$\tan a = r_{3,12}/r_{12}$$

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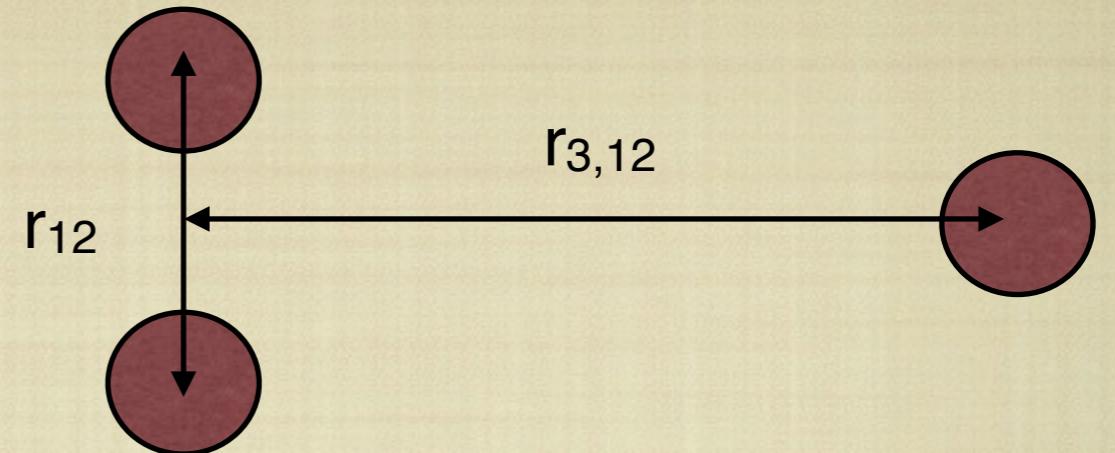
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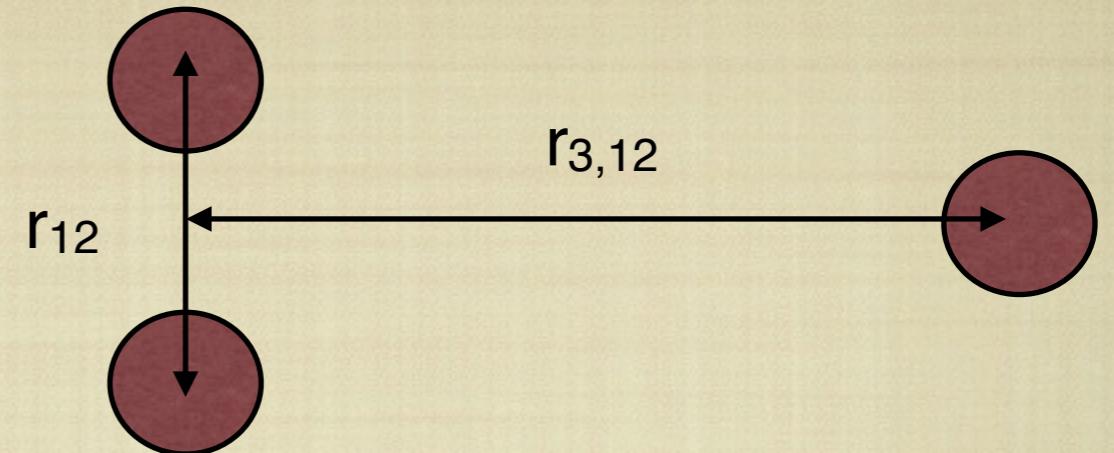
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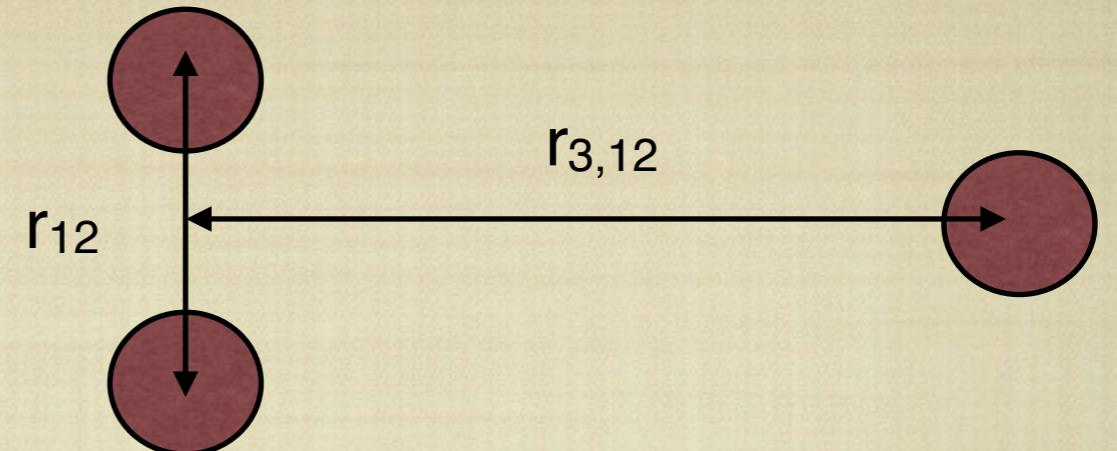
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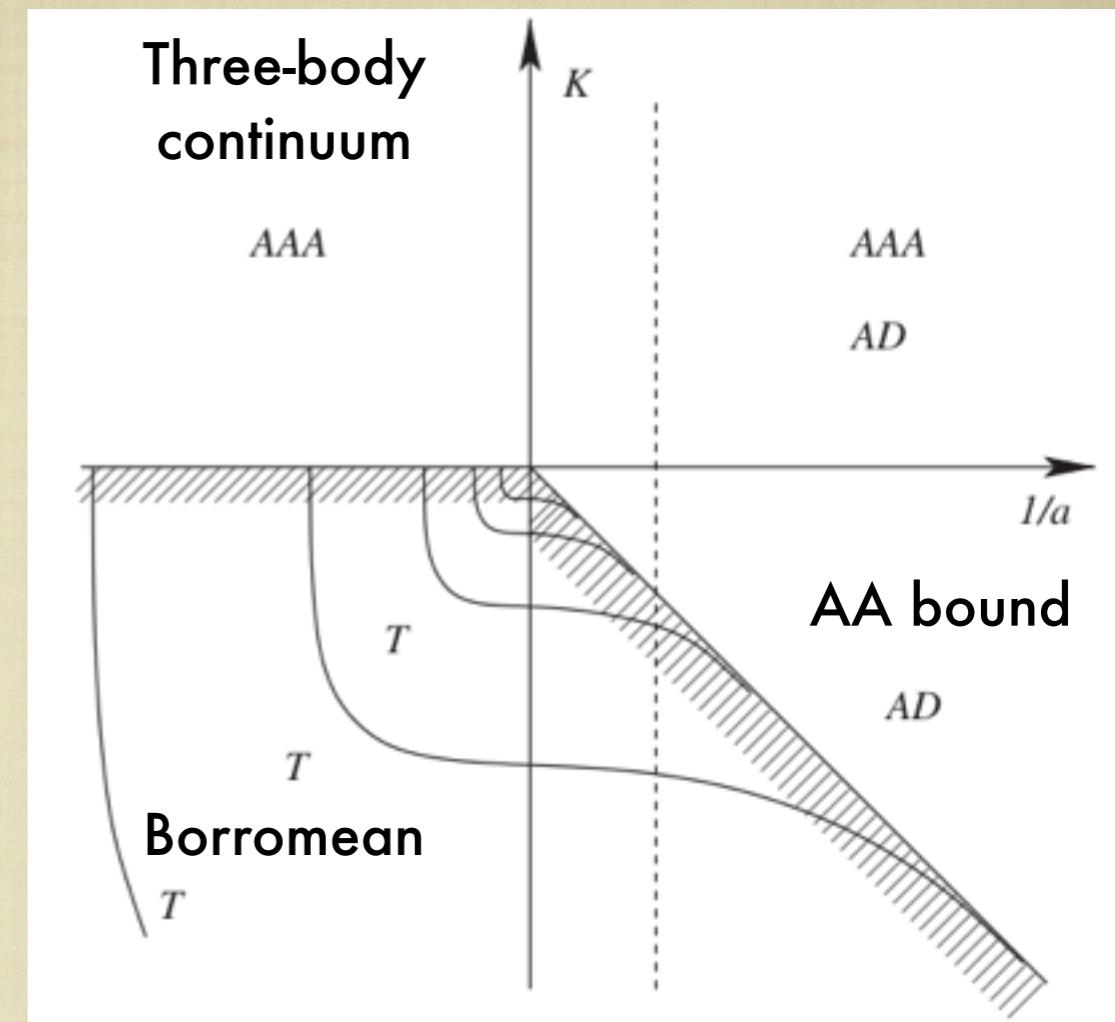
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- Energy of lowest state set by short-distance dynamics $\sim \frac{\hbar^2}{mR^2}$



The Efimov spectrum

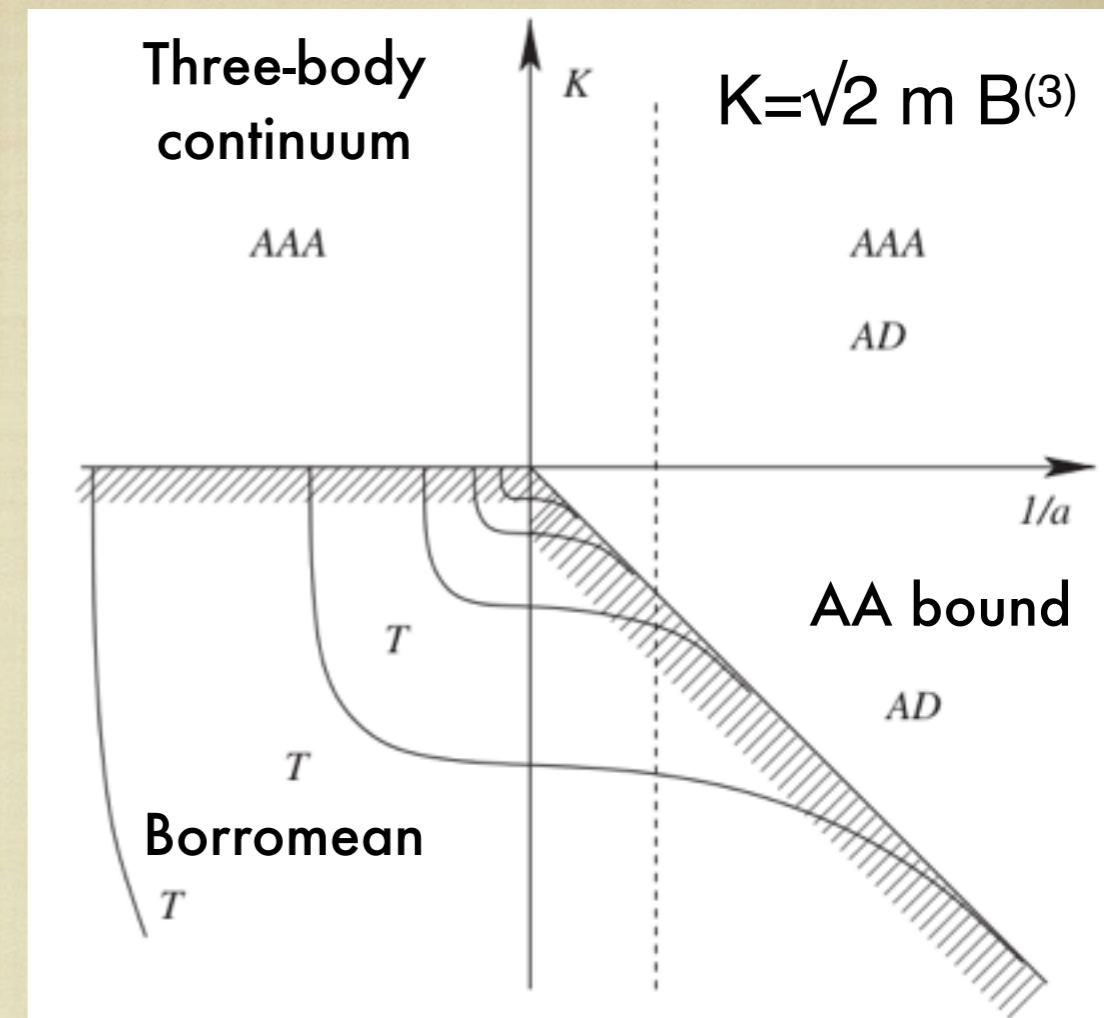


The Efimov spectrum

- $1/a=0$: infinite set of bound states, related by $K_{n+1}=K_n e^{\pi/s_0} = K_n$ (22.7)

Efimov, Yad. Fiz., 1970

Braaten & Hammer, Phys. Rep., 2003

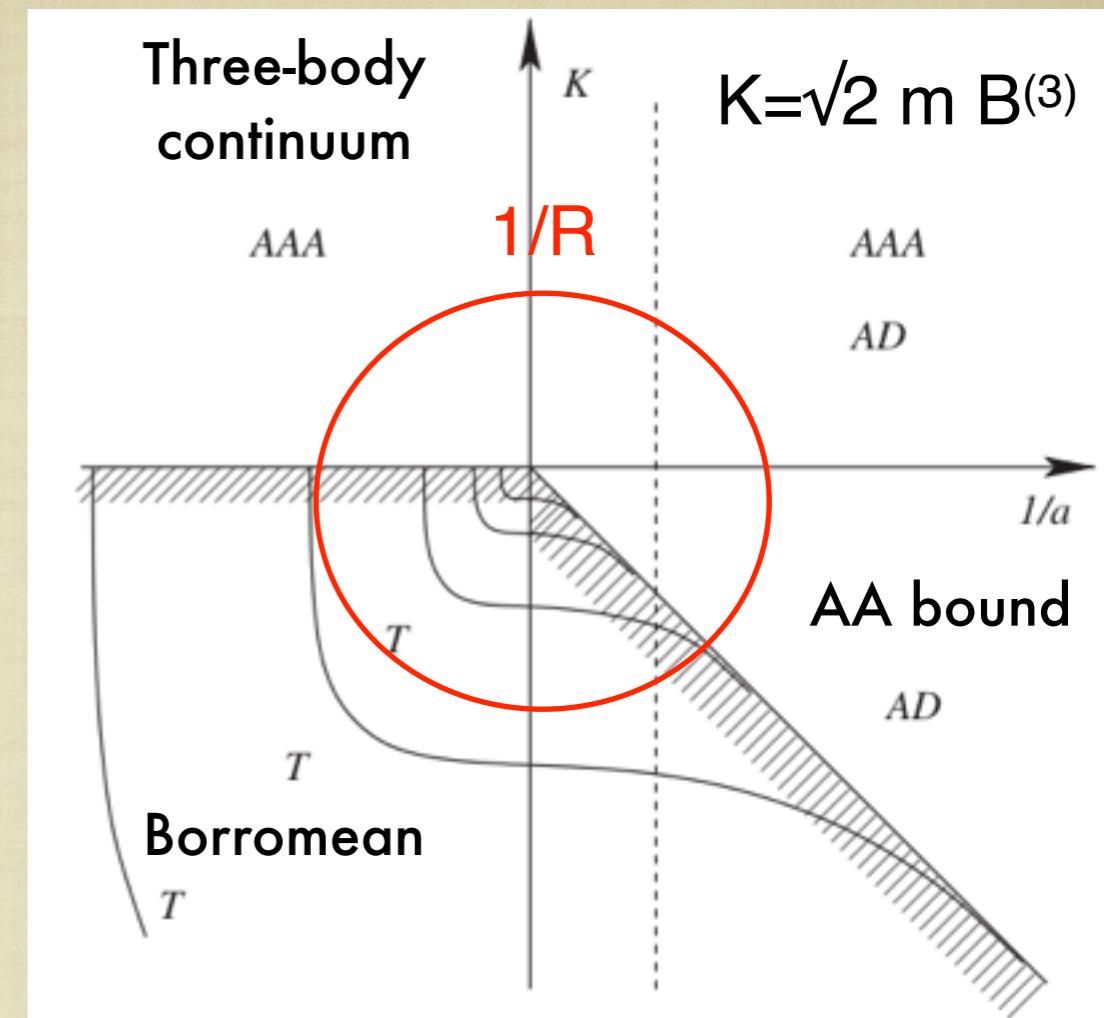


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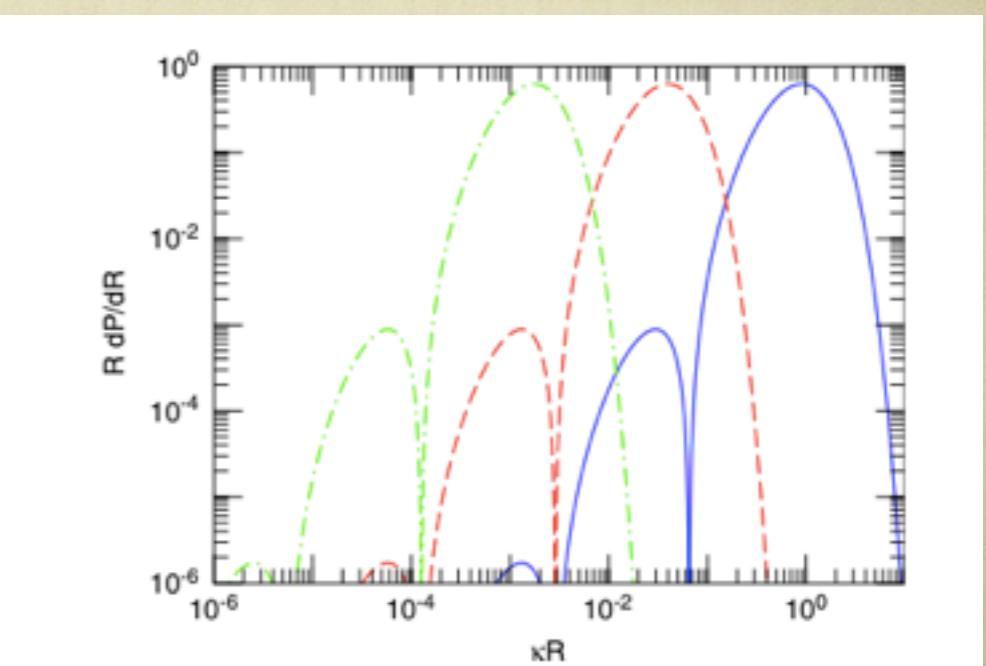
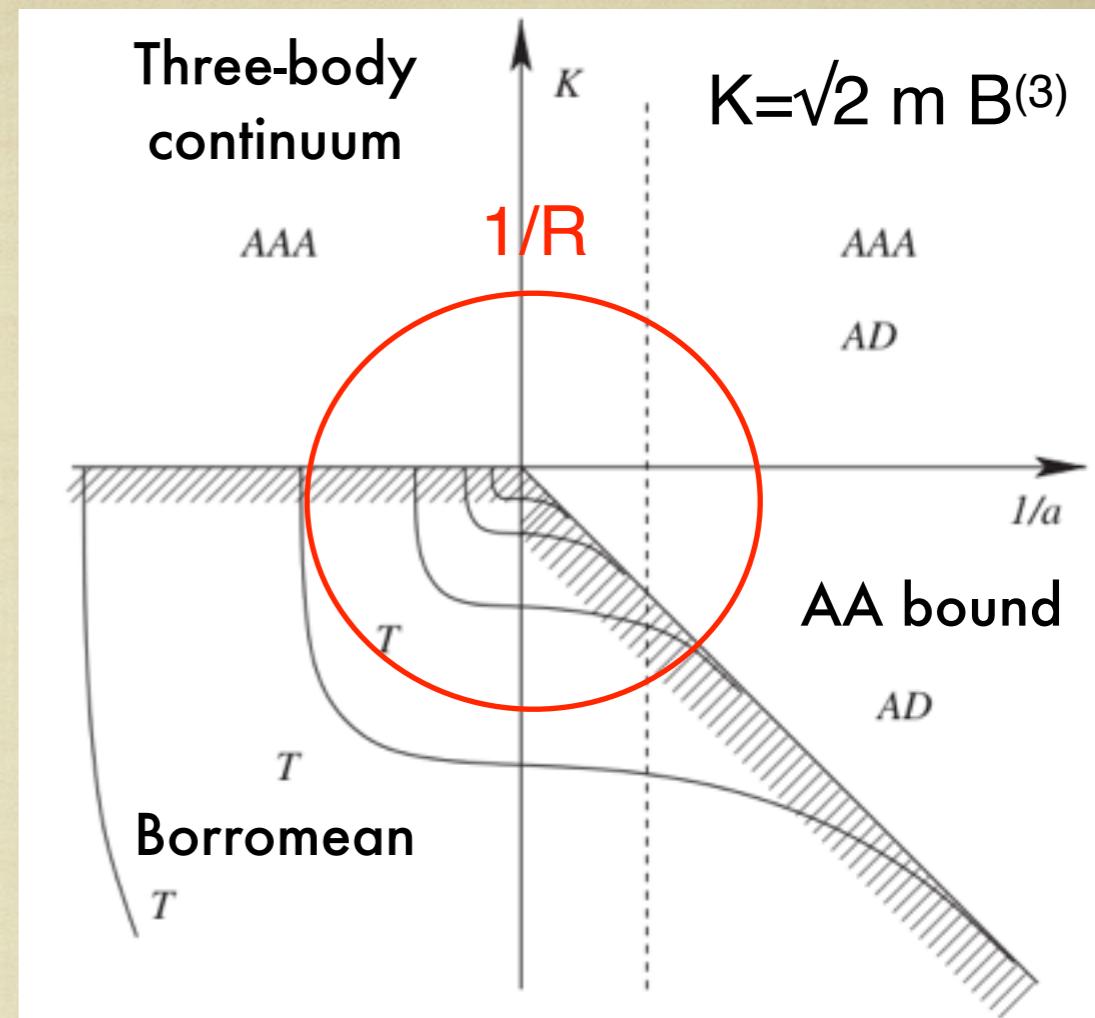


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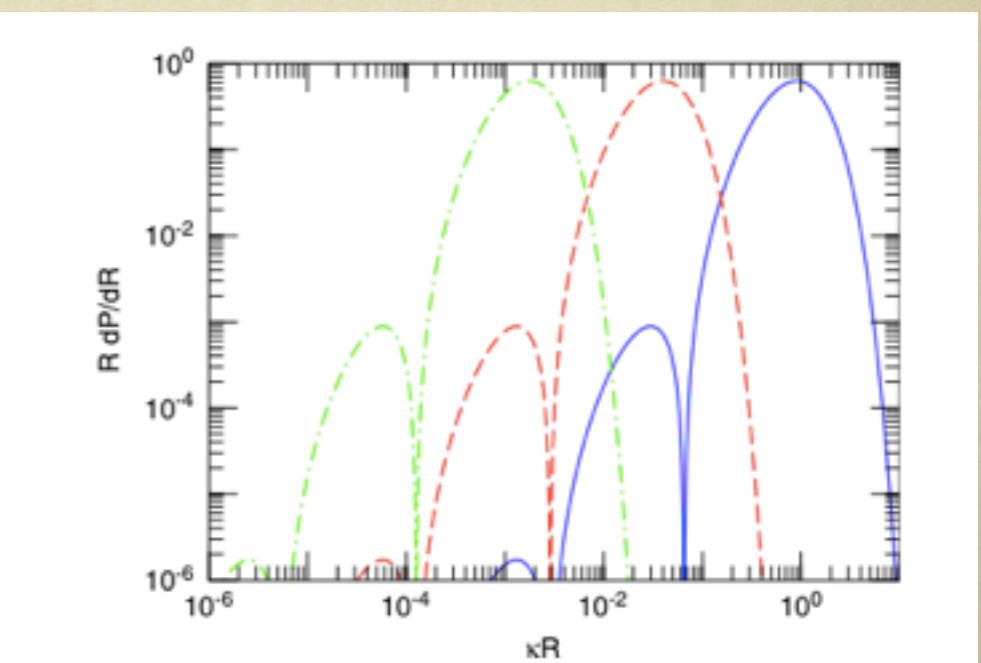
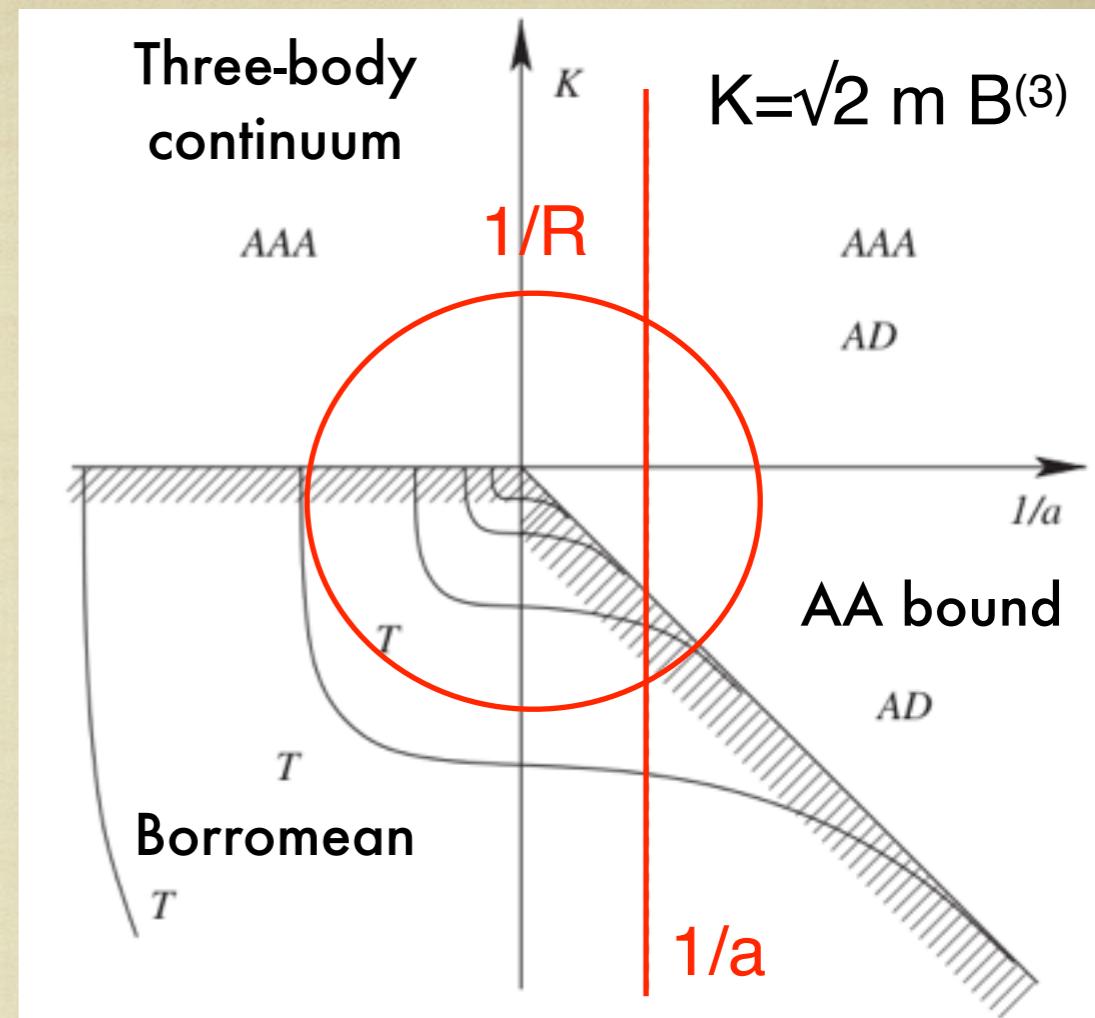


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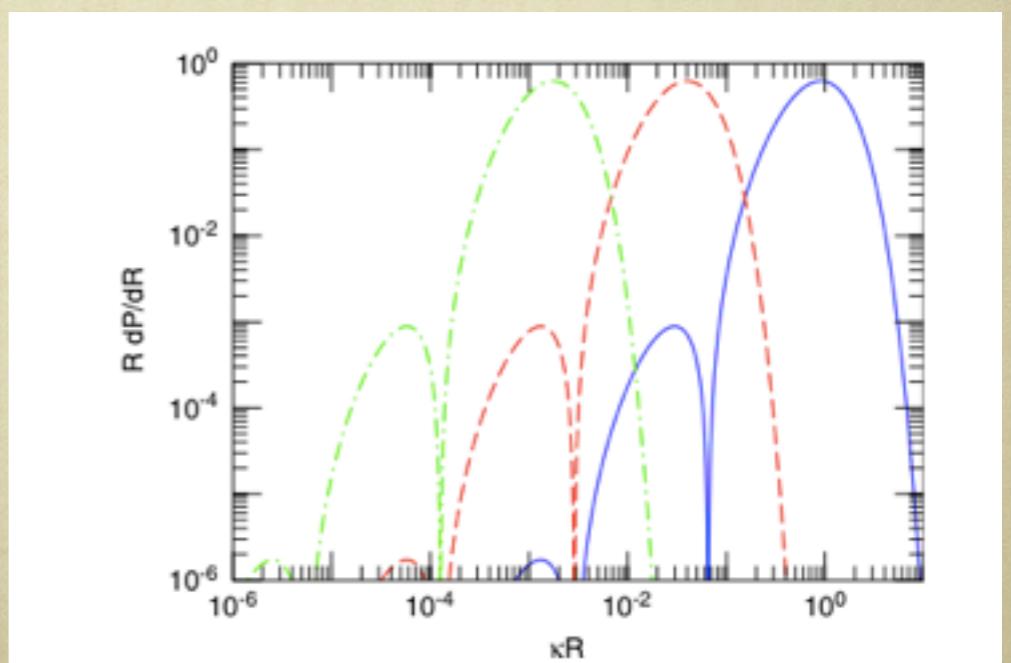
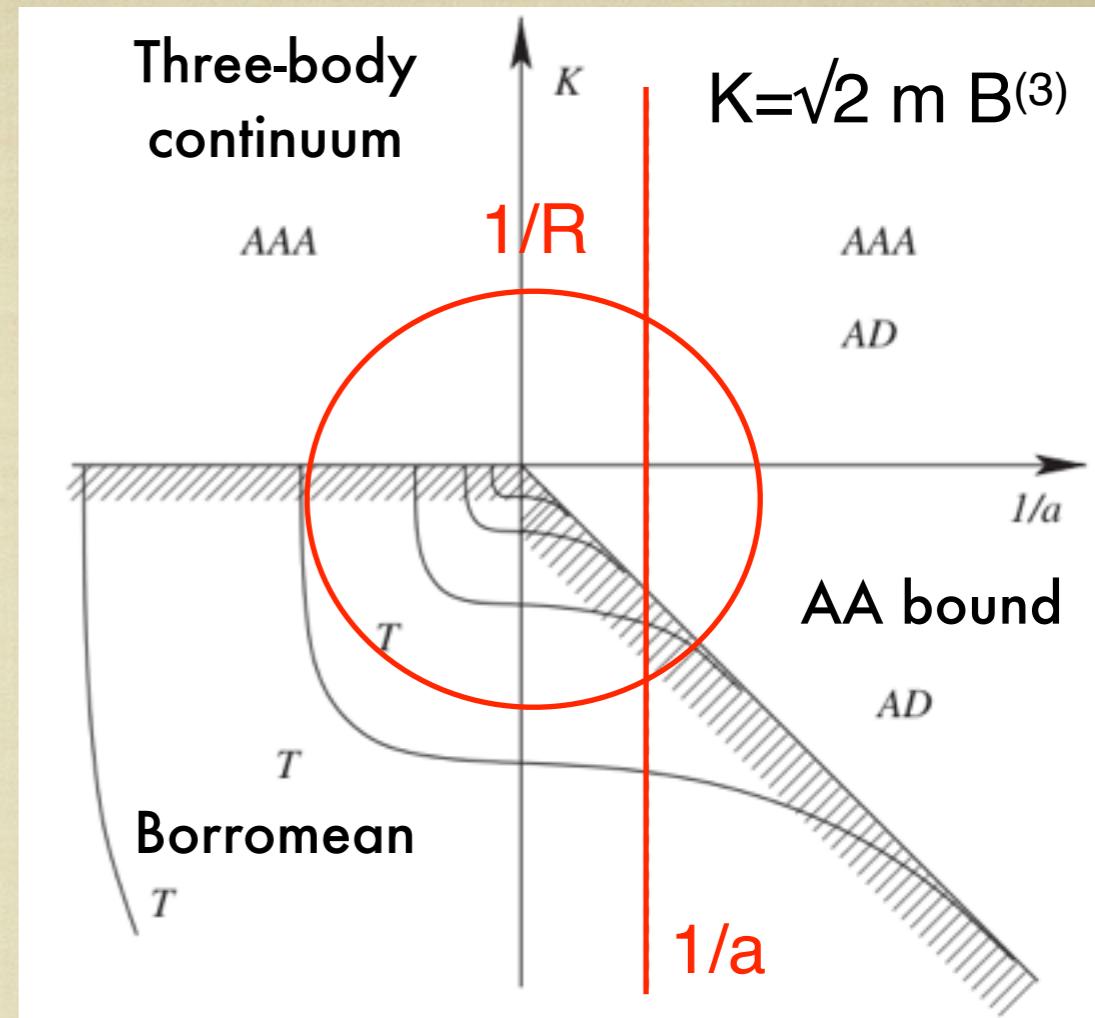


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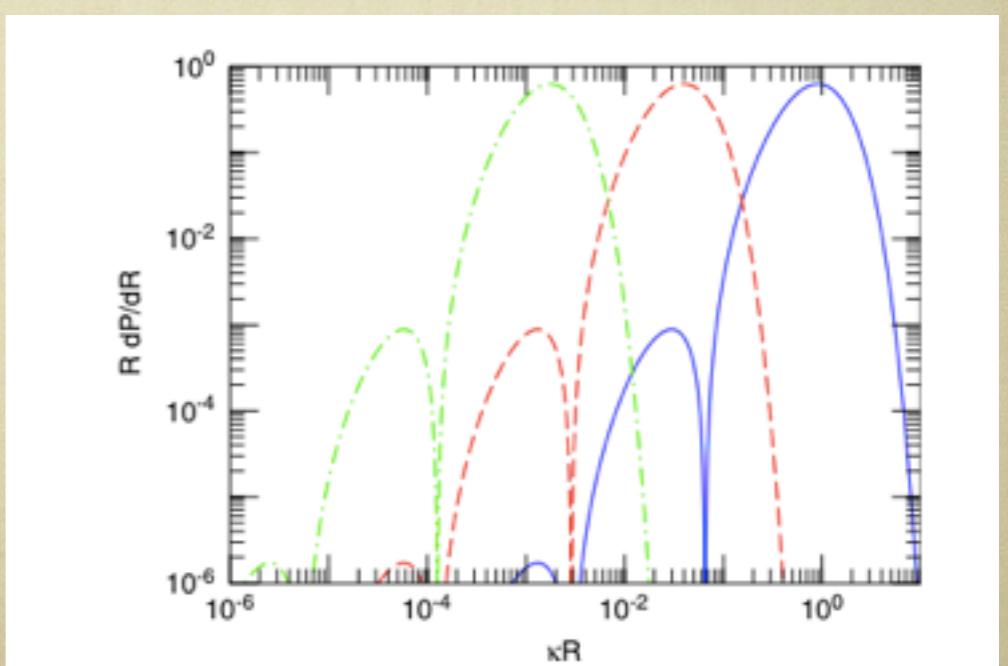
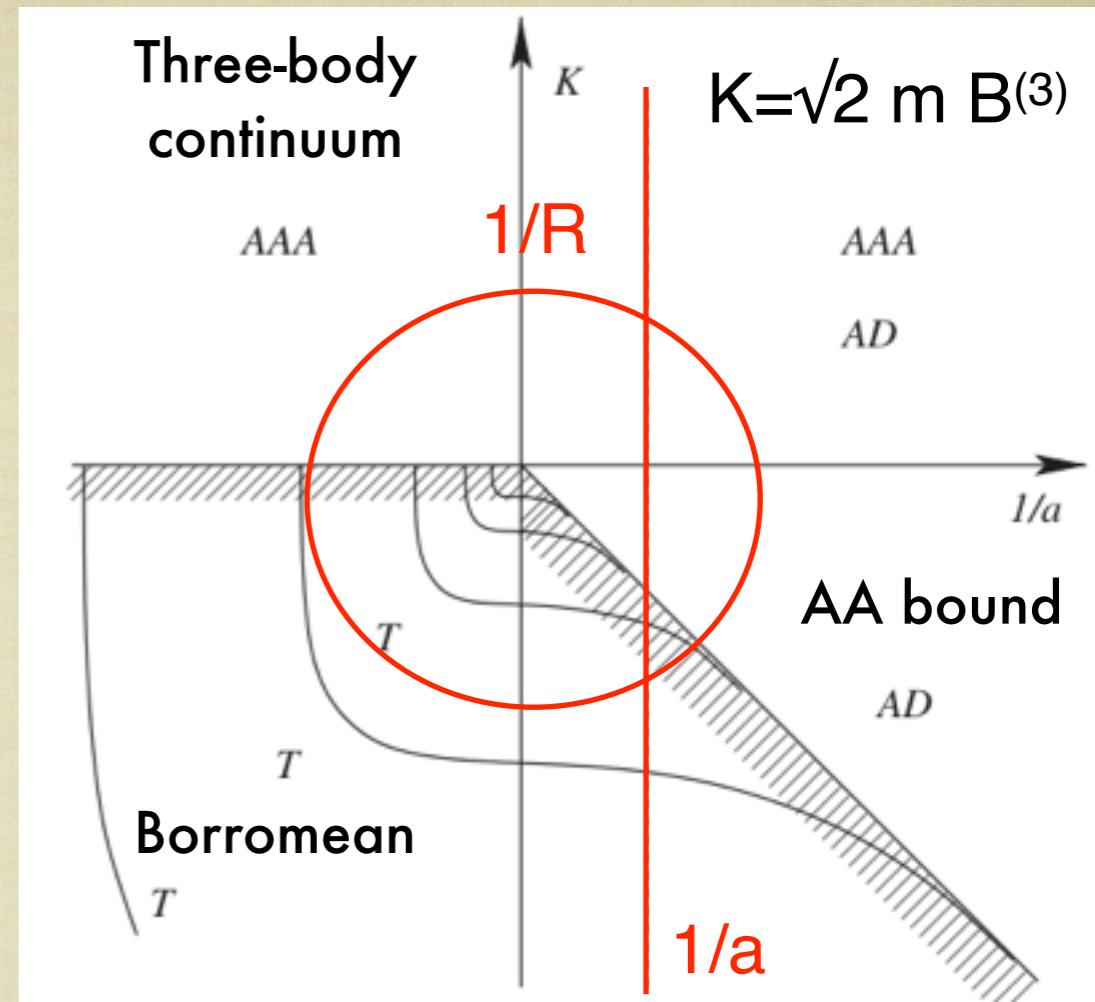
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Observed in Innsbruck experiment, 2014

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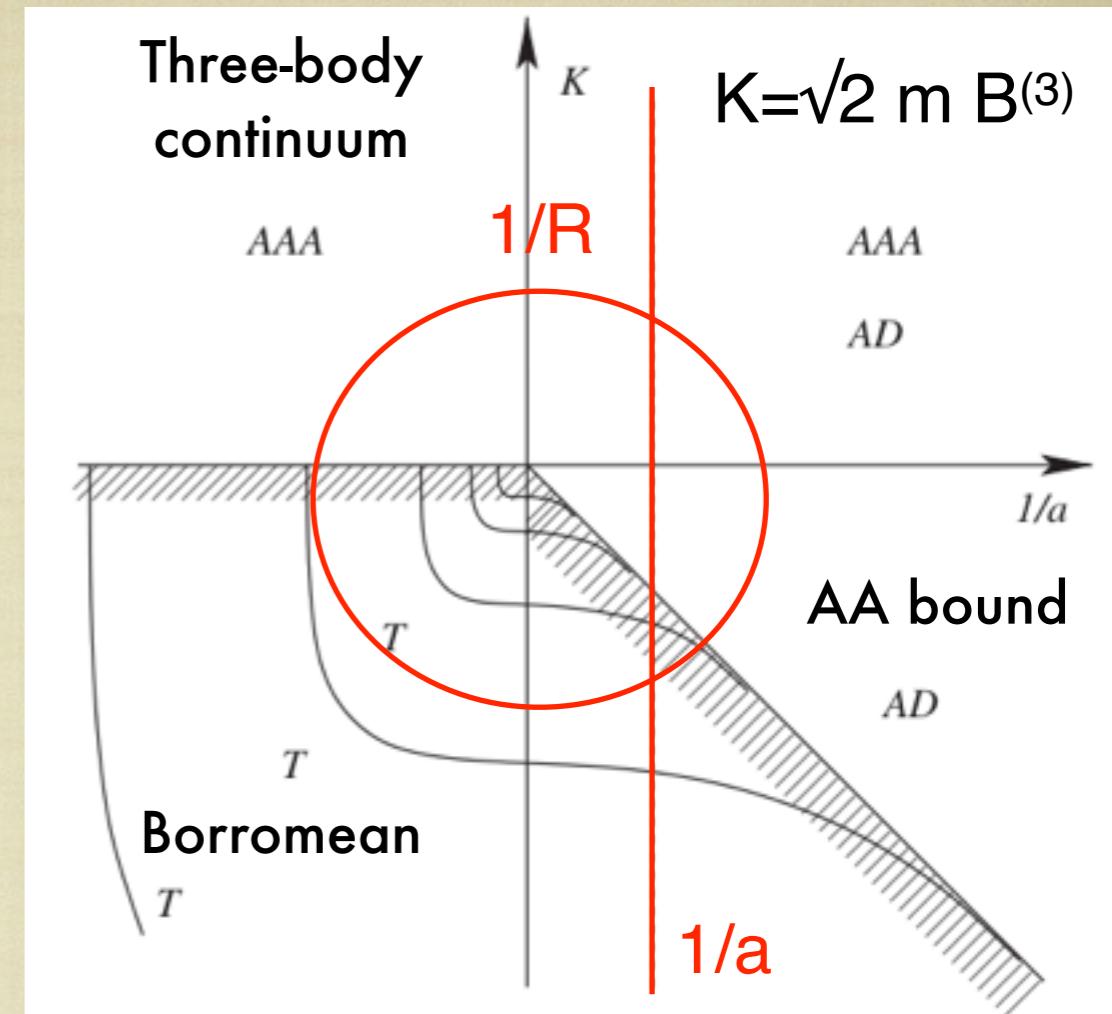


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Observed in Innsbruck experiment, 2014

- Correlations between different recombination features on an Efimov branch (or different branches): universal relations

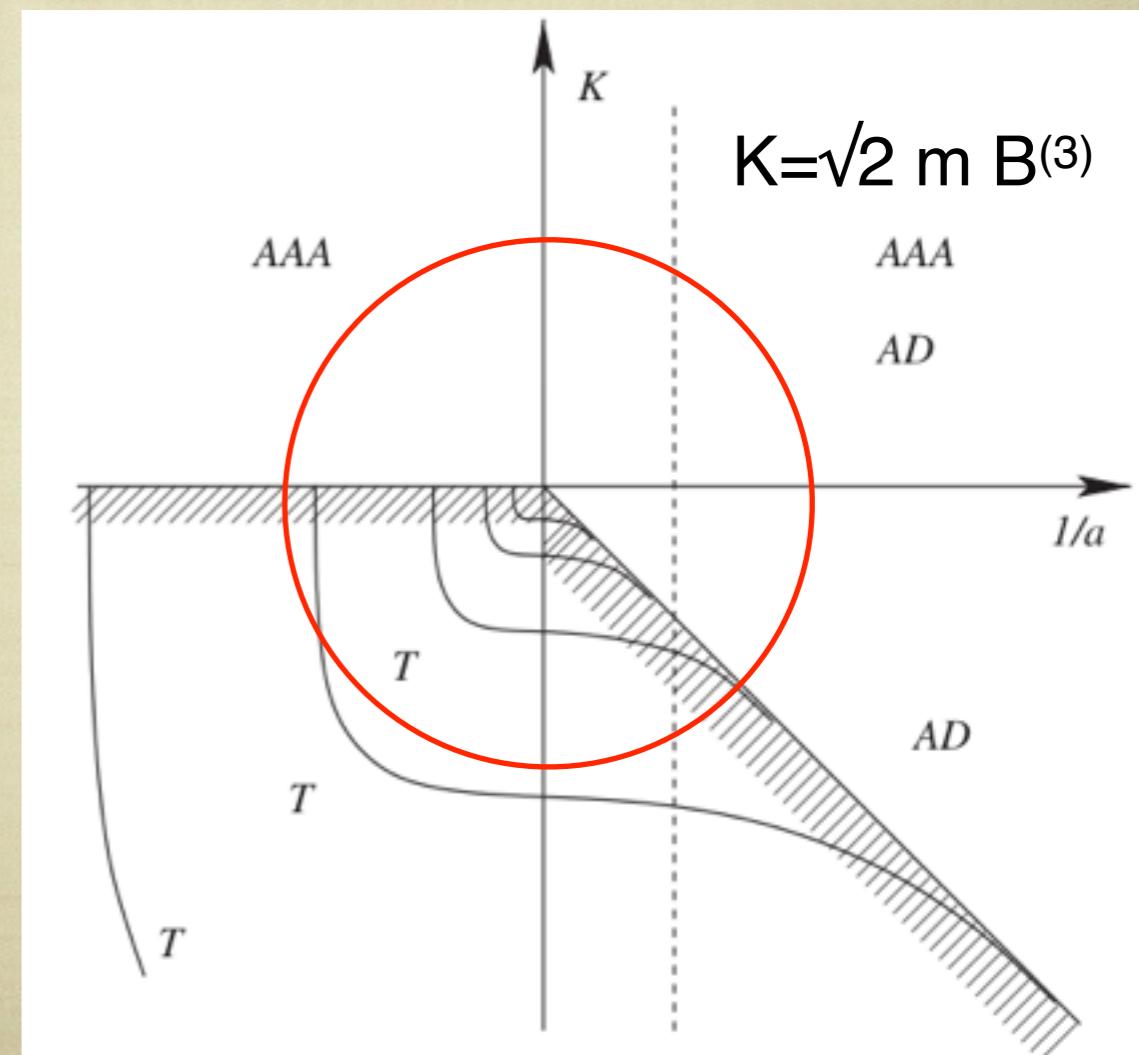
Efimov states in halo nuclei?

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- Candidates: ^{20}C ; ^{11}Li , ^{22}C , ^{14}Be .

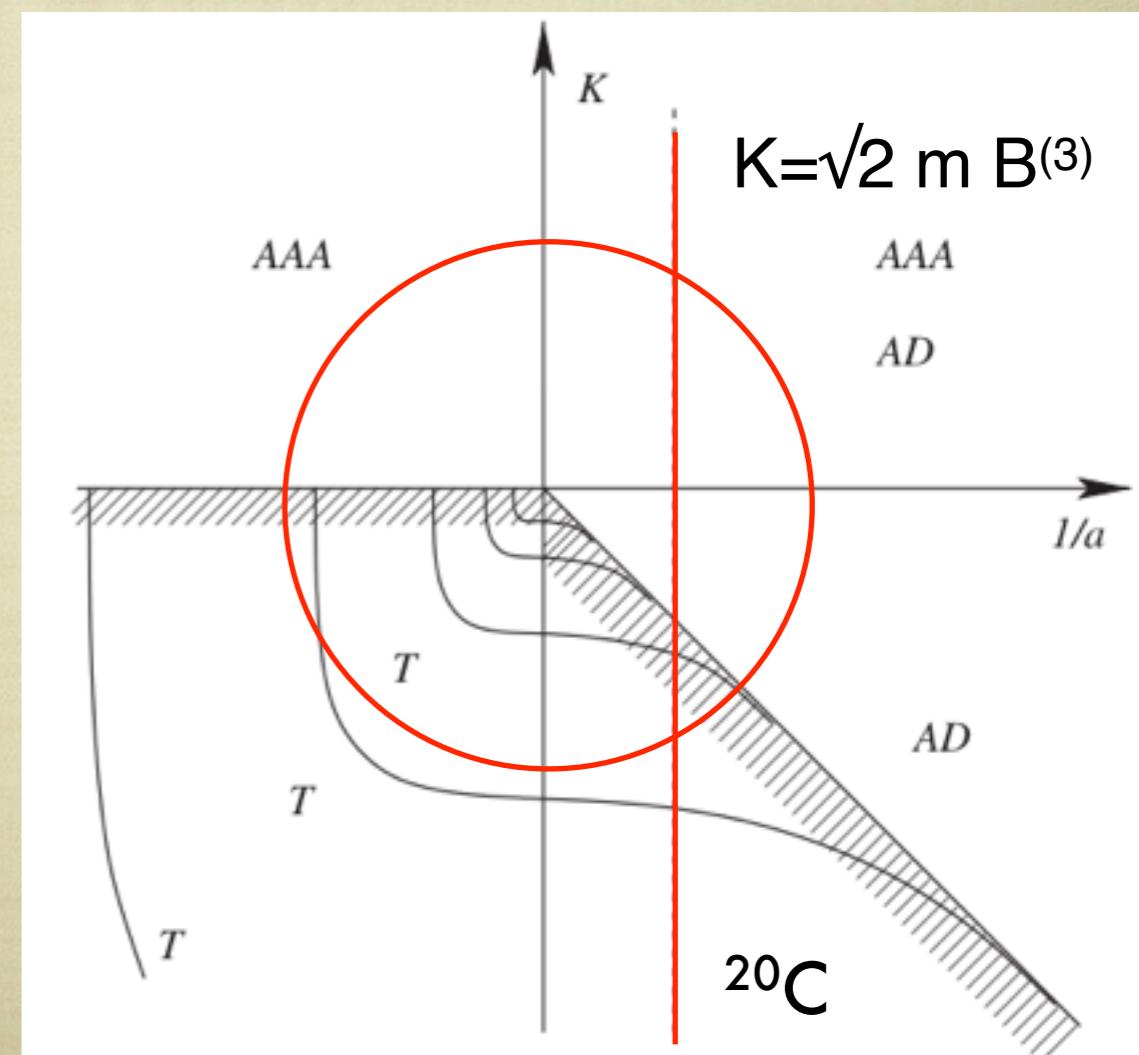
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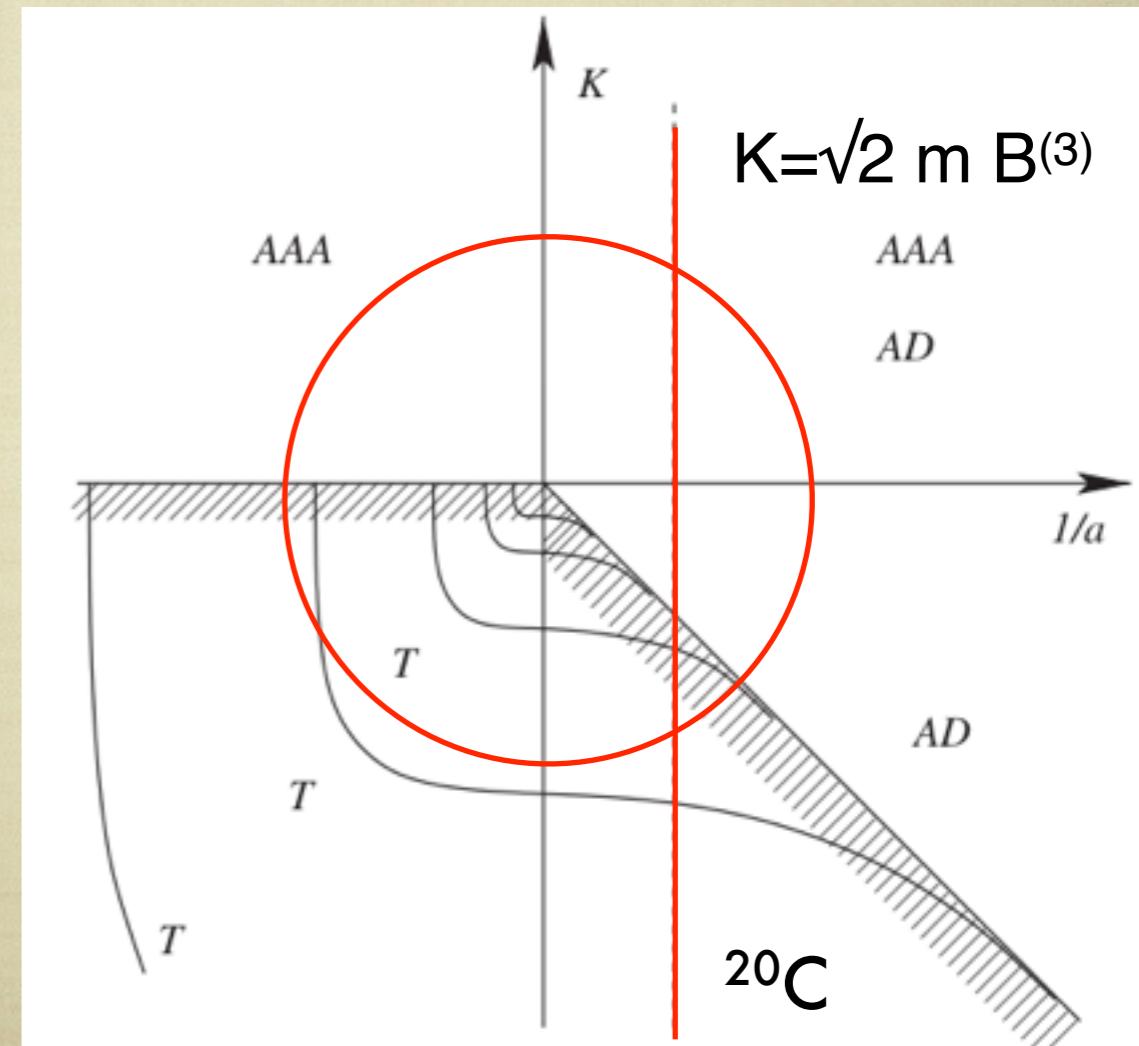
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- For Efimovian bound states need both $K_0 a$ and a/R large

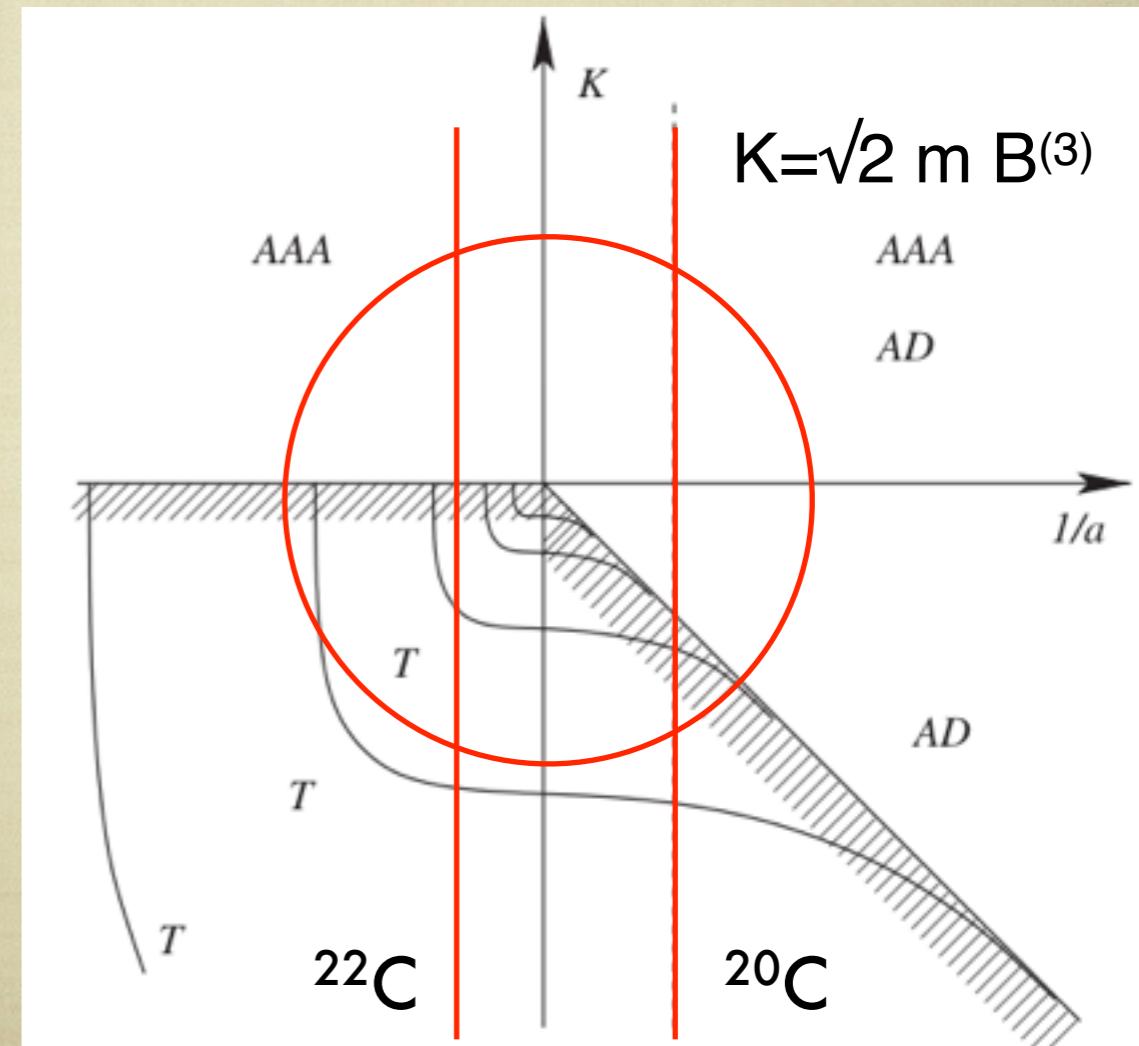
Canham & Hammer, EPJA (2008)



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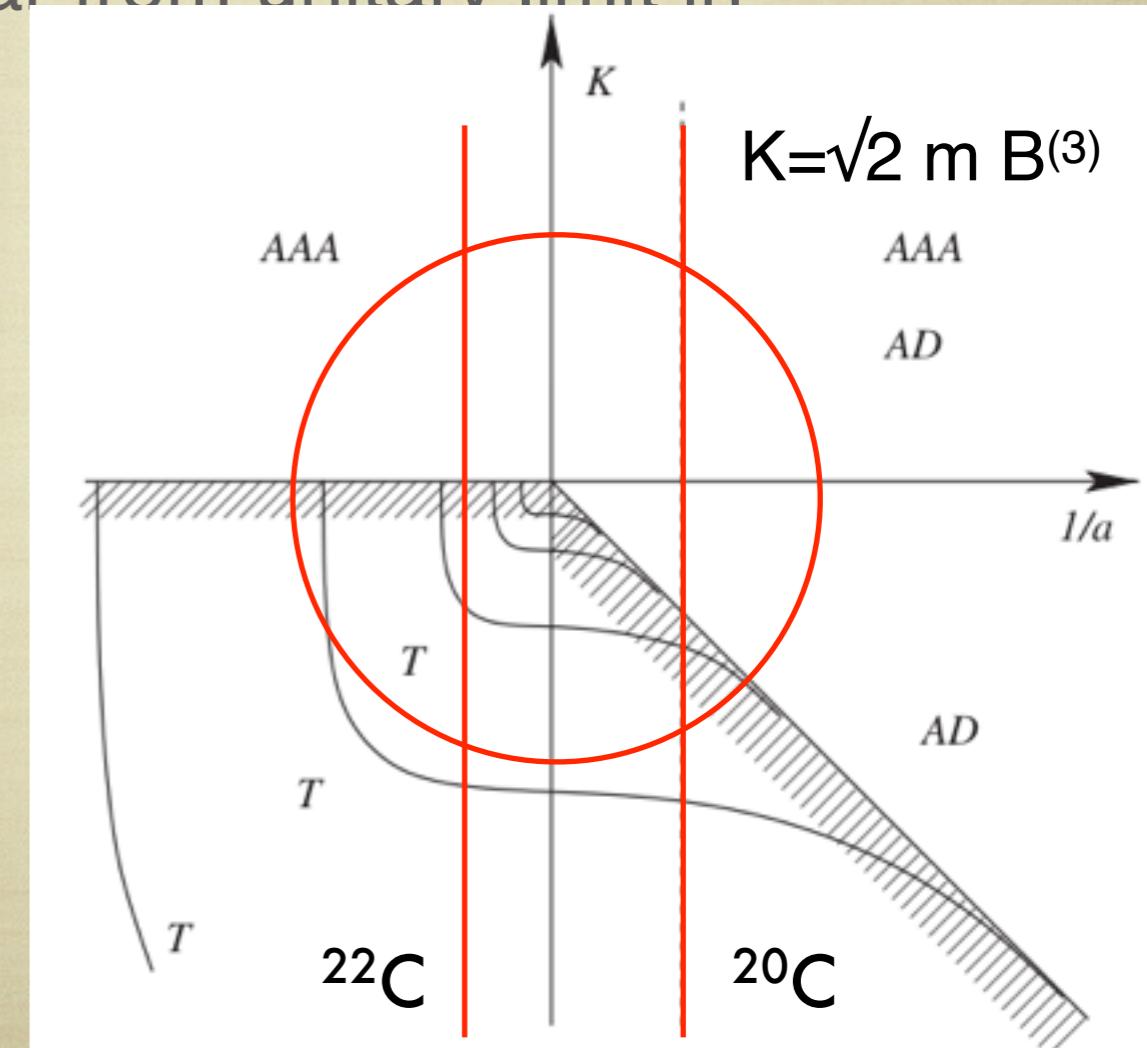
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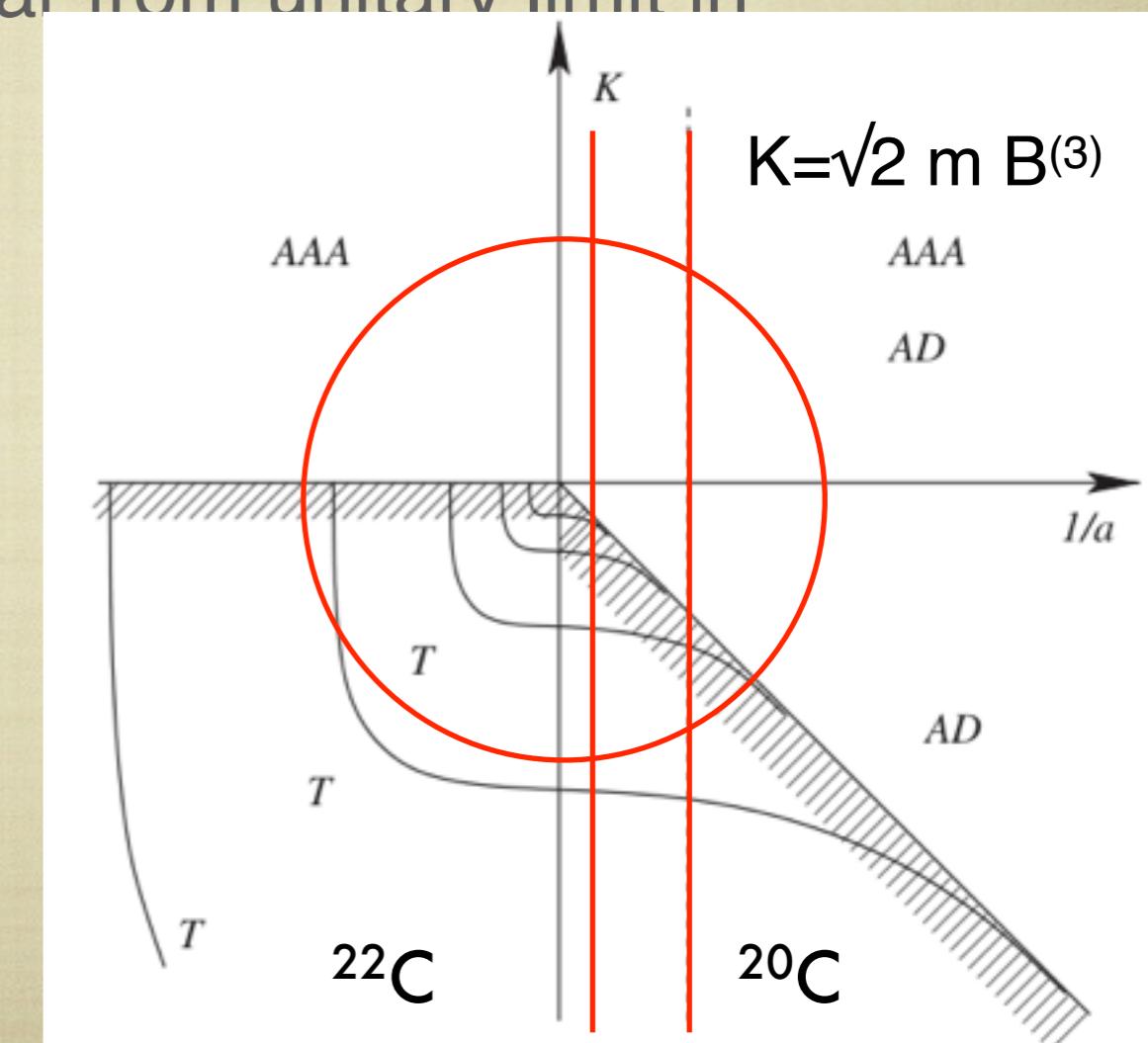
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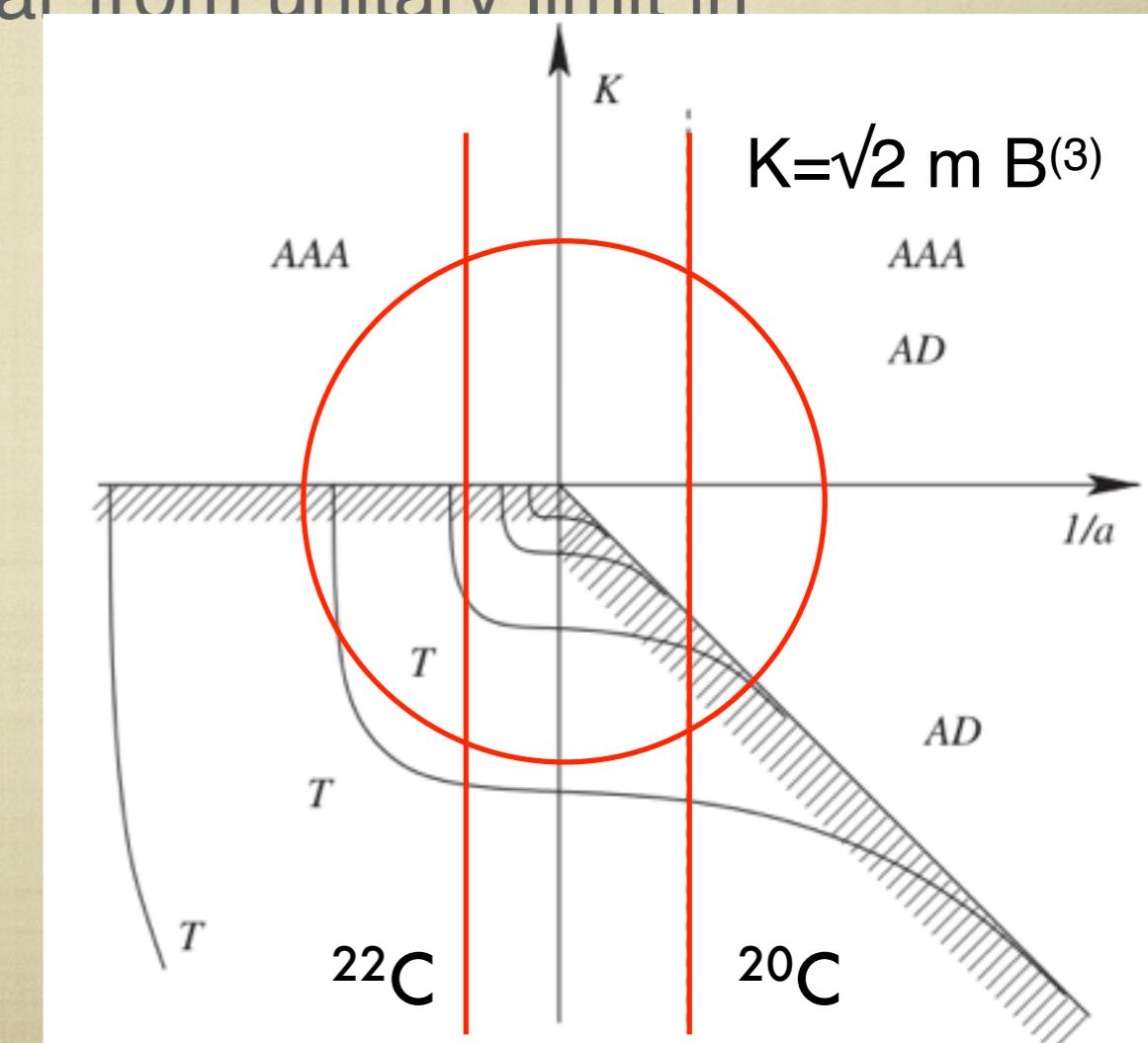
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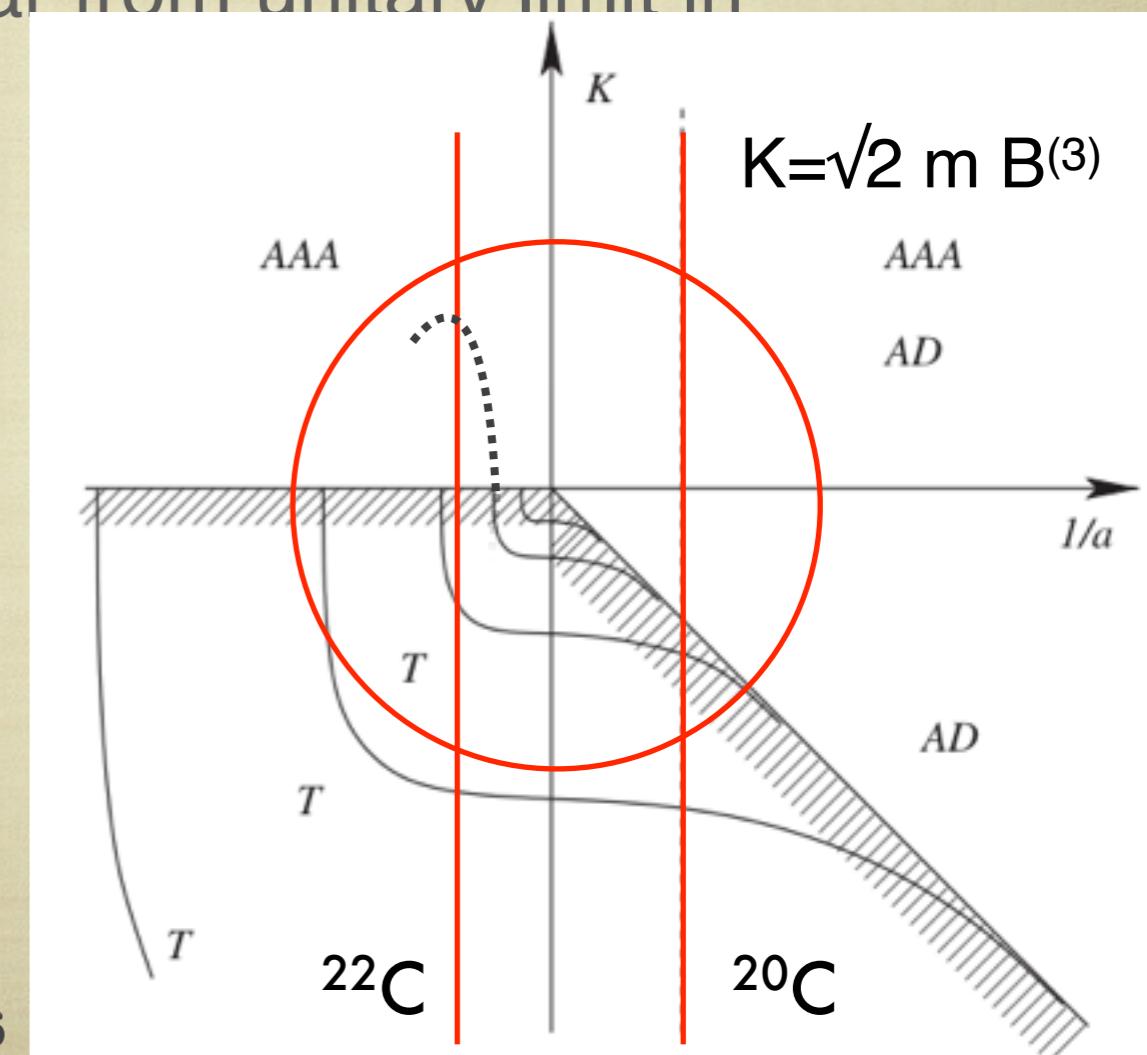
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- $\gamma + ^{22}\text{C} \rightarrow ^{20}\text{C} + \text{n} + \text{n}$



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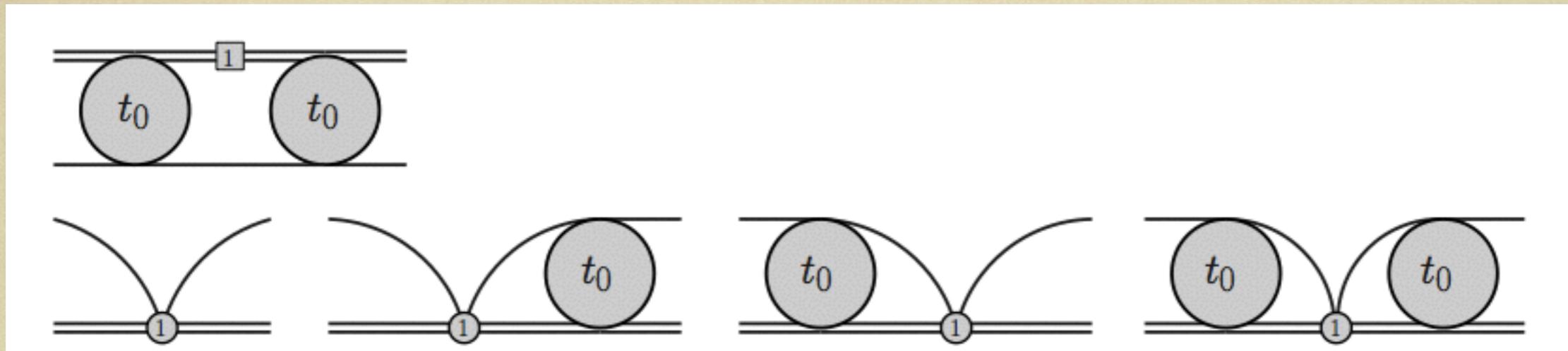


Backup slides: NLO and NNLO
in three-body systems

Perturbation theory at NLO

Hammer, Mehen (2001); Ji, Phillips, Platter (2009, 2010)

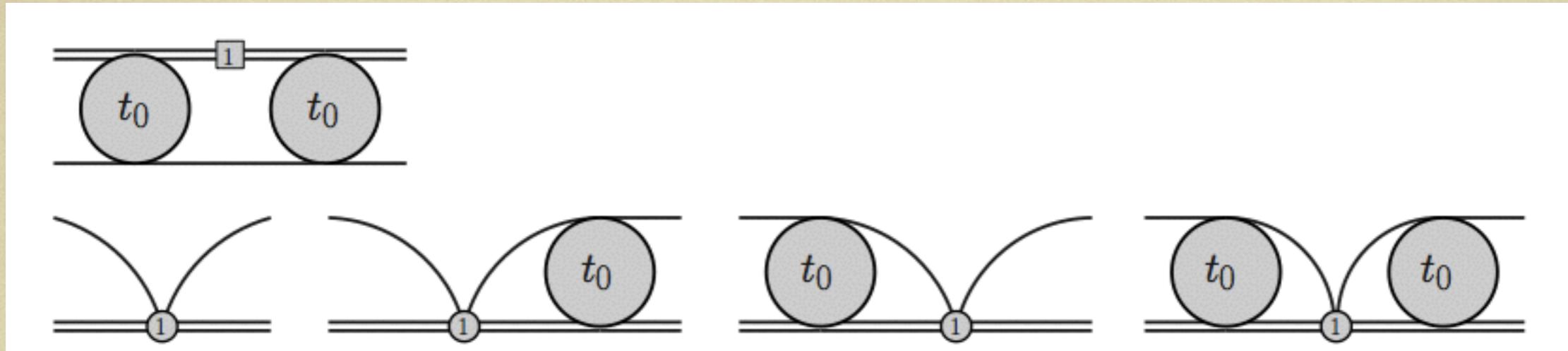
- Insert t_1^{2B} in first-order perturbation theory between LO wfs



Perturbation theory at NLO

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- Insert $t_1^{(2B)}$ in first-order perturbation theory between LO wfs



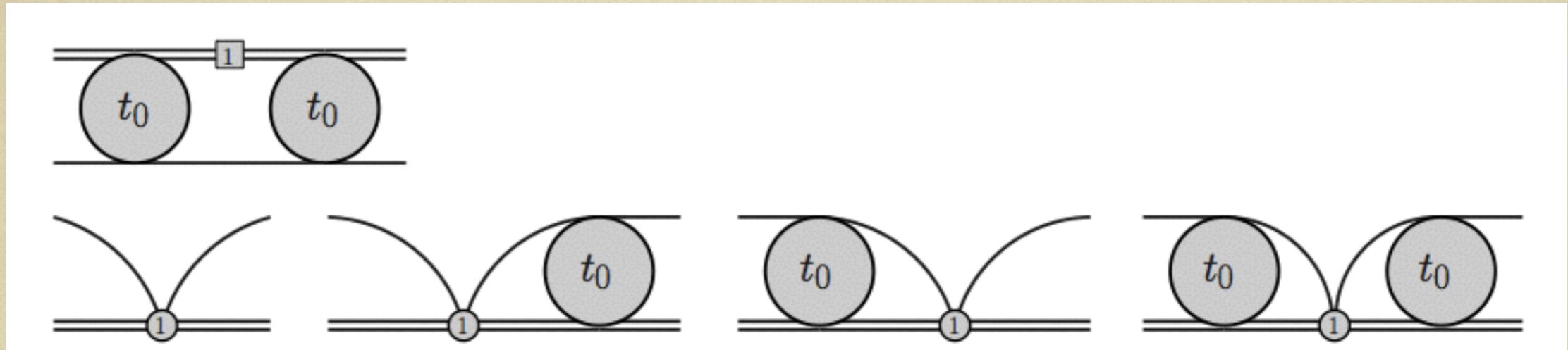
$$t_1^{(a)}(k, k; E) = \frac{r}{\pi} \int dq q^2 \frac{1/a + \sqrt{3/4q^2 - mE}}{-1/a + \sqrt{3/4q^2 - mE}} t_0^2(q, k; E)$$

$$t_1^{(b)}(k, k; E) = \frac{2H_1(\Lambda)}{\Lambda^2} \left[1 + \frac{2}{\pi} \int dq \frac{q^2}{-1/a + \sqrt{3q^2/4 - mE}} t_0(k, q) \right]^2$$

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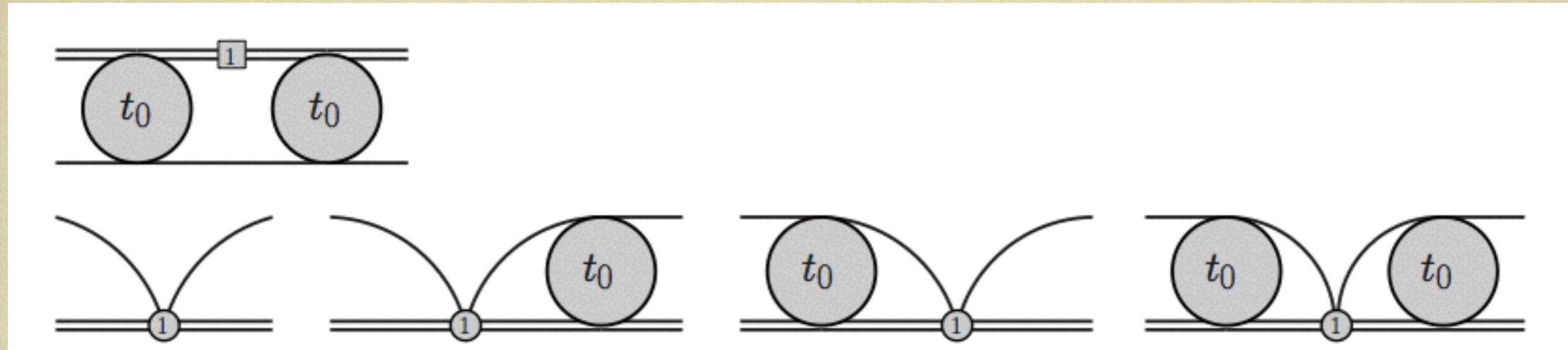
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$t_0 \sim 1/q \Rightarrow$ linear divergence $\sim r\Lambda$. Can be absorbed in H_1 .

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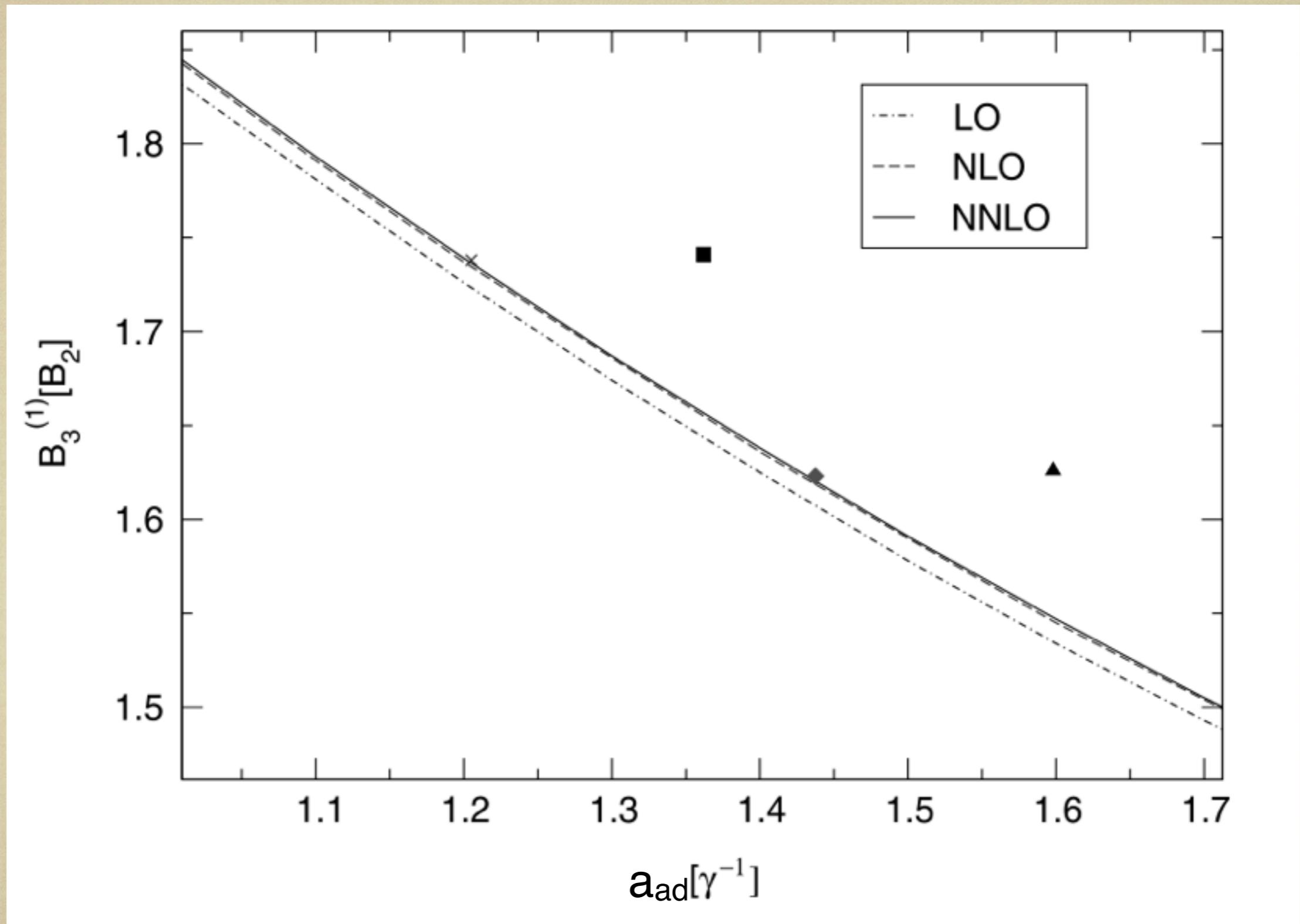
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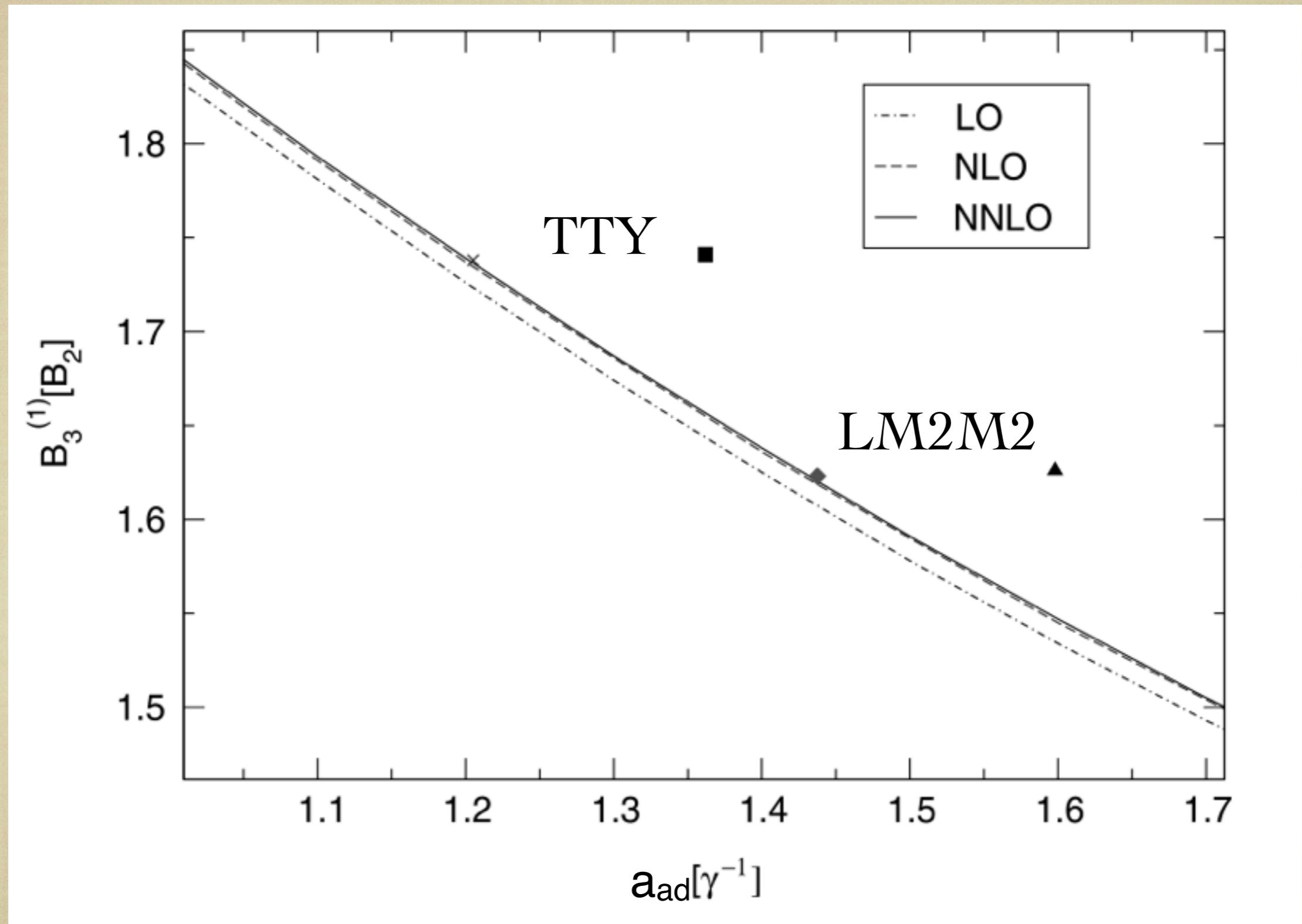
No new 3B datum needed at NLO **at fixed a** .

Corrections to universality



Platter, Phillips (2006)

Corrections to universality



He-4 trimers at NLO

Platter, Phillips (2006); Ji, Phillips (2012)

He-4 trimers at NLO

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- Experimentally: $a = 104^{+8}_{-18} \text{ \AA}$; $B_d = 1.1^{+0.3}_{-0.2} \text{ mK} \Rightarrow r \sim 10 \text{ \AA}$,
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	$B_t^{(1)} [B_d]$	$B_t^{(0)} [B_d]$	$a_{ad} [\gamma]$	$r_{ad} [1/\gamma]$
TTY	1.738	96.33	1.205	?

He-4 trimers at NLO

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	$B_t^{(1)} [B_d]$	$B_t^{(0)} [B_d]$	$a_{ad} [\gamma]$	$r_{ad} [1/\gamma]$
TTY	1.738	96.33	1.205	?
LO, a_{ad}	1.723	97.12	1.205	0.8352
NLO, a_{ad}	1.736	89.72	1.205	0.9049

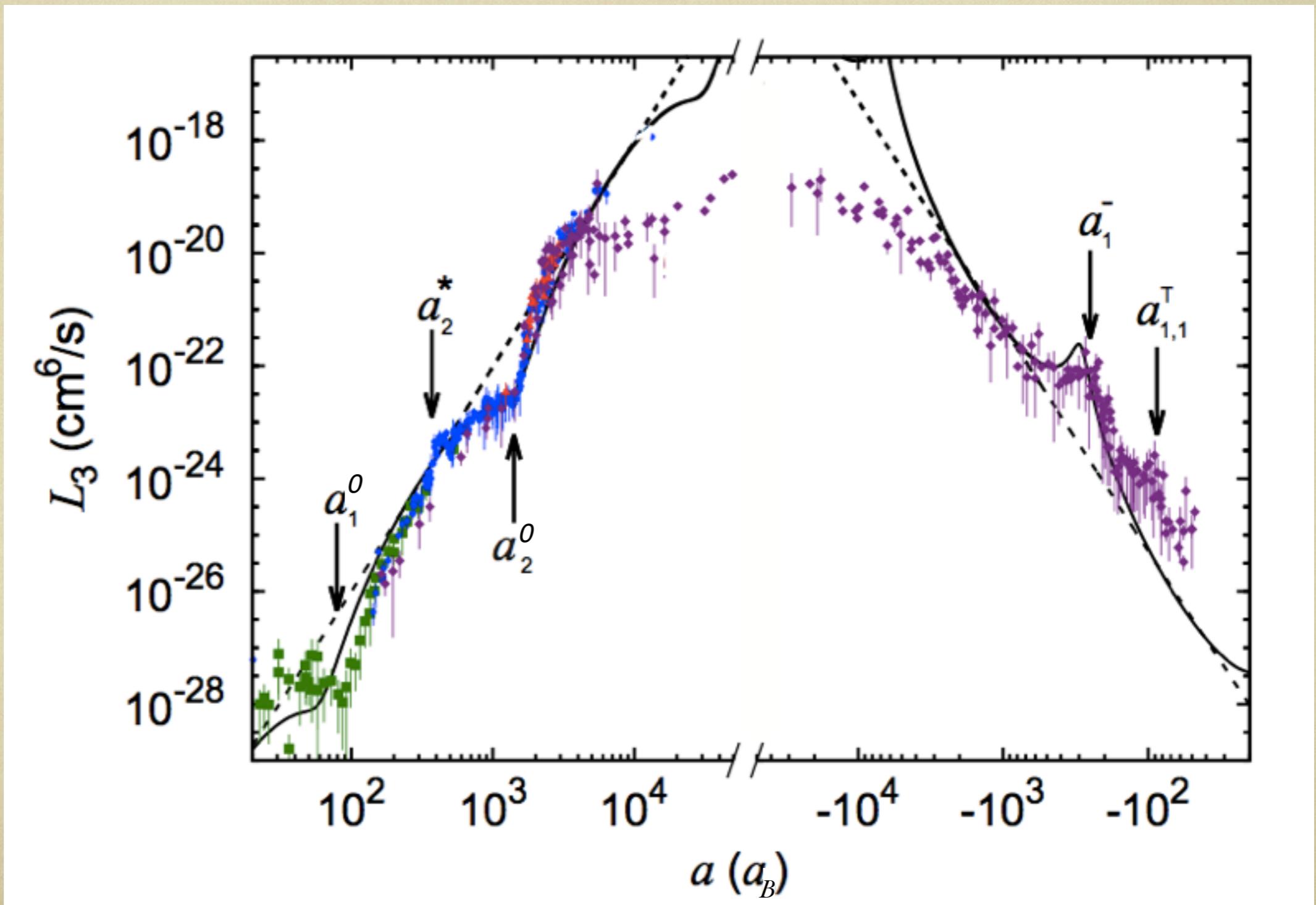
He-4 trimers at NLO

Platter, Phillips (2006); Ji, Phillips (2012)

- Experimentally: $a=104^{+8}_{-18}$ Å; $B_d=1.1^{+0.3}_{-0.2}$ mK $\Rightarrow r \sim 10$ Å, and trimer observed, but no measurement of B_t or a_{ad}
- To make quantitative EFT predictions we take as input

	$B_t^{(1)} [B_d]$	$B_t^{(0)} [B_d]$	$a_{ad} [\gamma]$	$r_{ad} [1/\gamma]$
TTY	1.738	96.33	1.205	?
LO, a_{ad}	1.723	97.12	1.205	0.8352
NLO, a_{ad}	1.736	89.72	1.205	0.9049
LO, $B_t^{(1)}$	1.738	99.37	1.178	0.8752
NLO, $B_t^{(1)}$	1.738	89.77	1.201	0.913

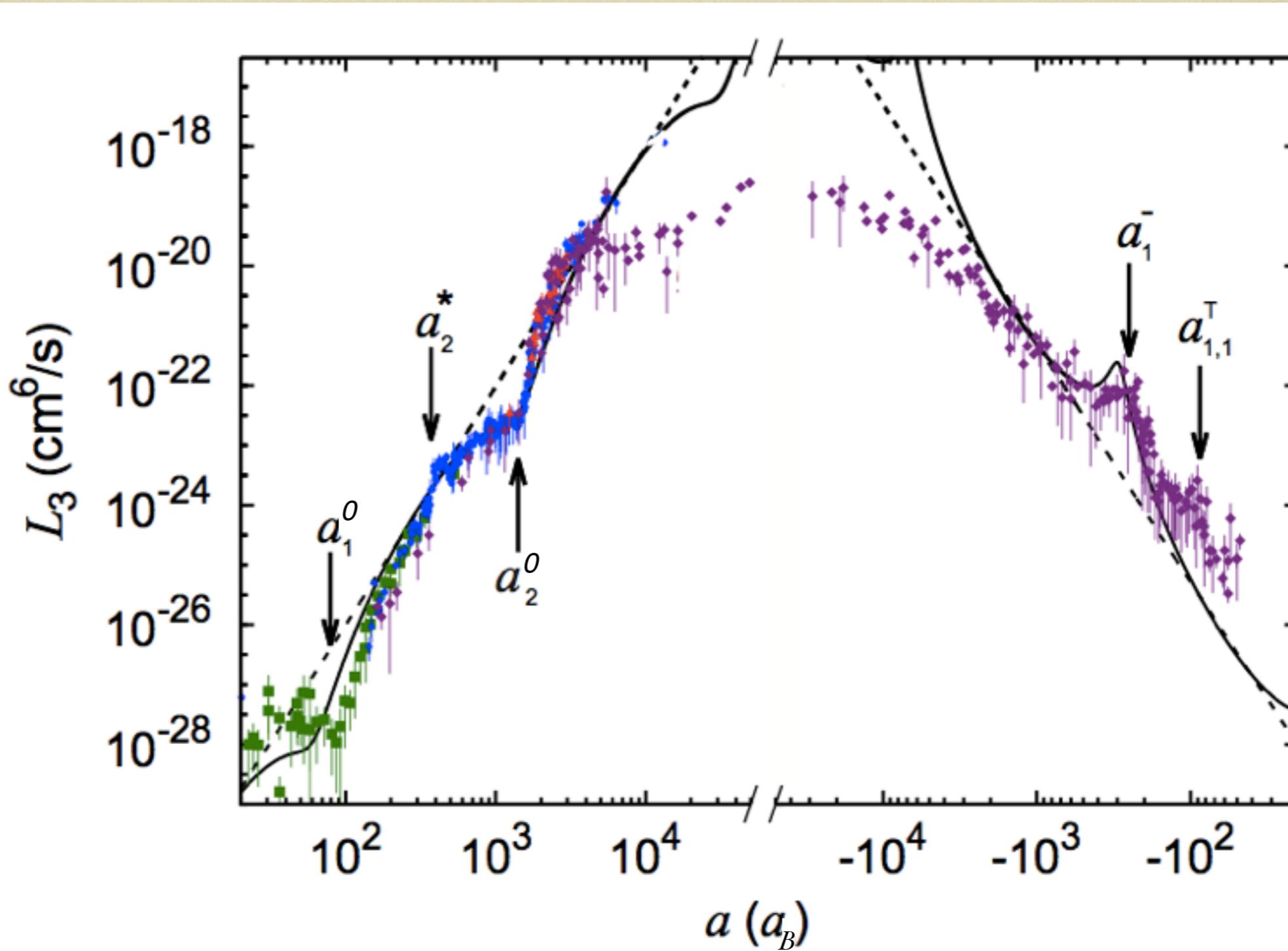
Semi-universal relations in ${}^7\text{Li}$



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Dyke, Pollack,
Hulet (2013)

$F=1$,
 $m_F=1$

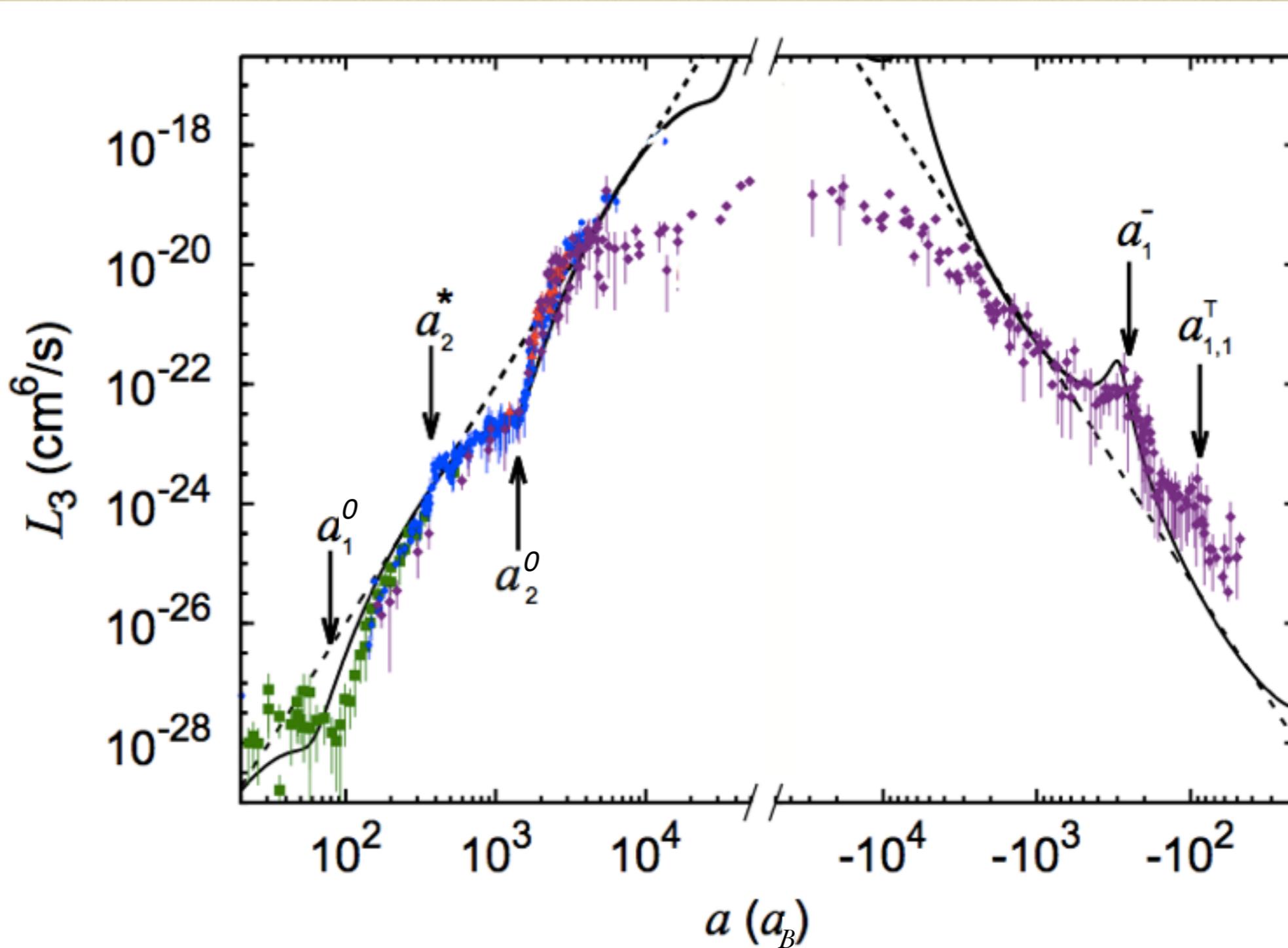


Semi-universal relations in ${}^7\text{Li}$

$$\frac{1}{a_*} = -0.939\gamma_-^{(-)} - 0.309r(a_*)\frac{1}{a_-^{(-)}} + 7.17\frac{r(a_*)}{r(a_0)} \left(\frac{1}{a_0} + 0.210\frac{1}{a_-^{(-)}} \right)$$

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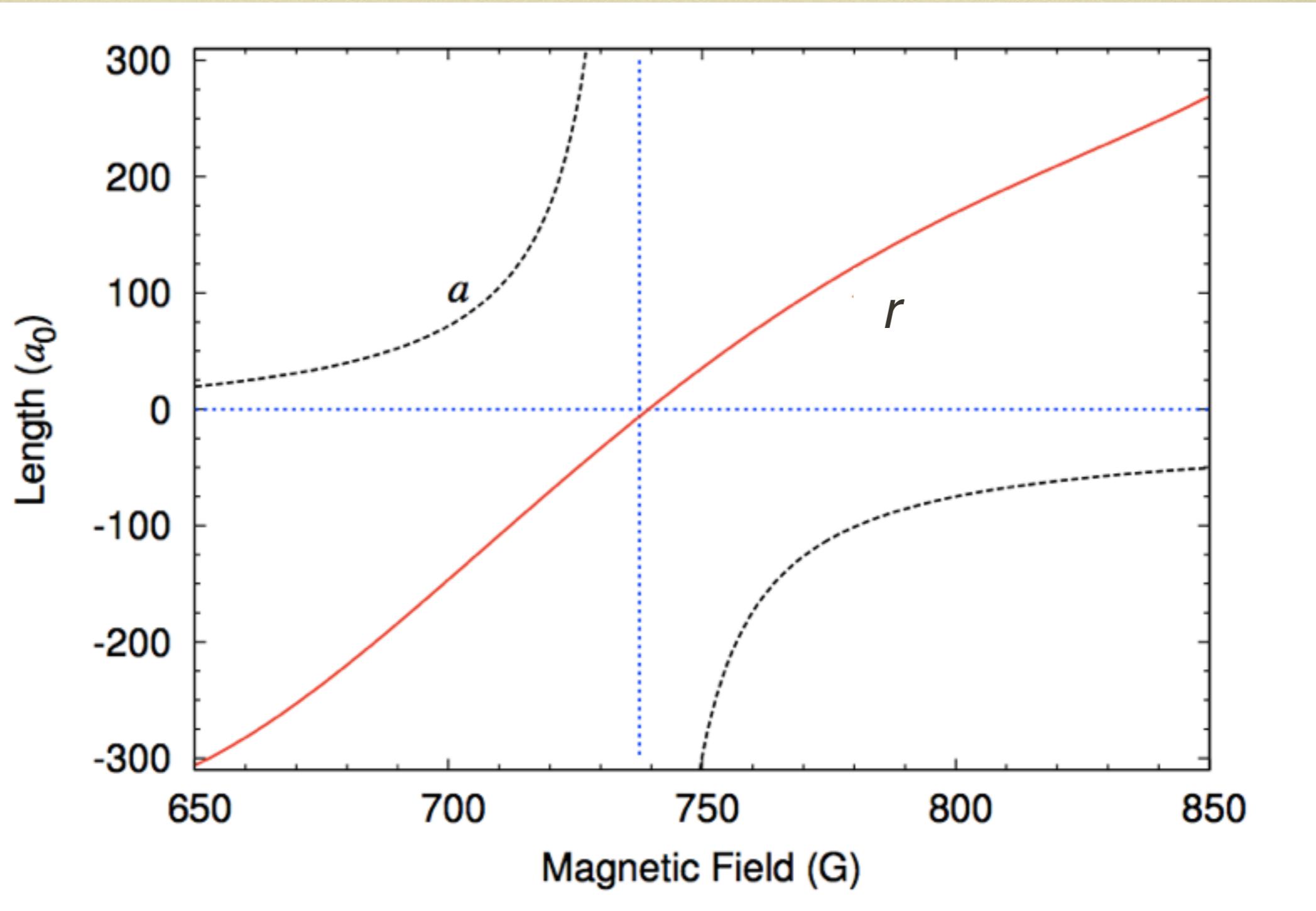


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Gross et al. (2010)

Minima and maxima at: $a_0=1160$ a_B; $a_-^{(-)}=-264$ a_B

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$m_F=0$ $a^*=(271 - 108 + \dots)$ a_B with $a_-^{(-)}$ as LO input

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Averaging: $a^*=(210 \pm 44)$ a_B

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Gross et al. (2012)

Experiment: $a^*=(196 \pm 4)$ a_B

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Dyke, Pollack,
Hulet (2013)

Minima and maxima at: $a_0=1402$ a_B; $a_-^{(-)}=-241$ a_B

Effective ranges: $r(a_0)=-14.8$ a_B; $r(a^*)=-33.9$ a_B; $r(a_-^{(-)})=45.8$ a_B

$F=1$, **SREFT prediction:**

$m_F=1$

Semi-universal relations in ${}^7\text{Li}$

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Averaging: $a^*=(339 \pm 65)$ a_B

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Averaging: $a^*=(339 \pm 65)$ a_B

Experiment: $a^*=(426 \pm 20)$ a_B

N^2LO results using STM equation

Bedaque, Greisshammer, Hammer, Rupak (2002)

- Insert $t_0^{2B} + t_1^{2B} + t_2^{2B}$ in STM eqn, solve to get $t_0 + t_1 + t_2$

Only reliable for $\Lambda\ell \ll 1$

see Platter & Phillips (2005), Platter (2006)

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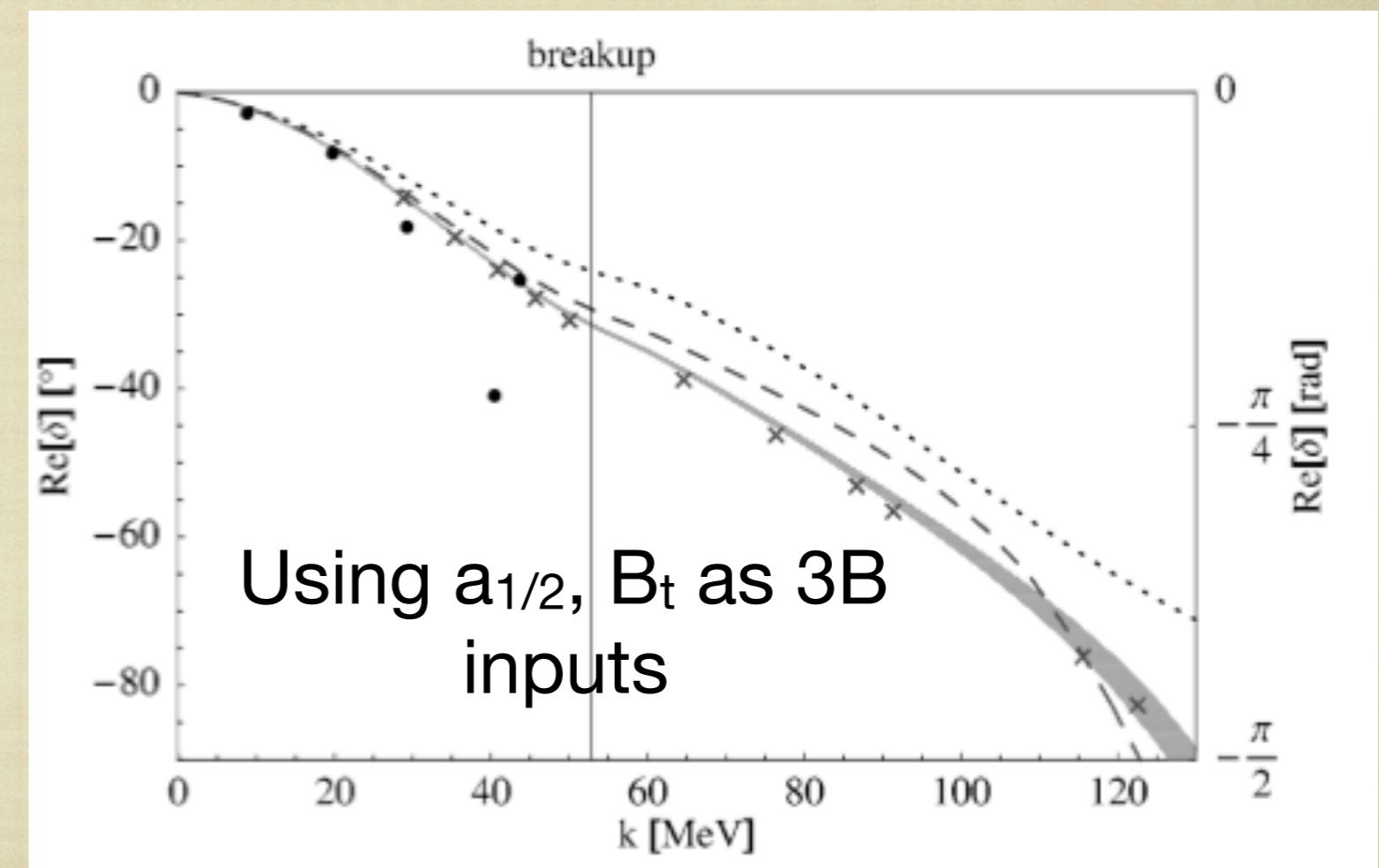
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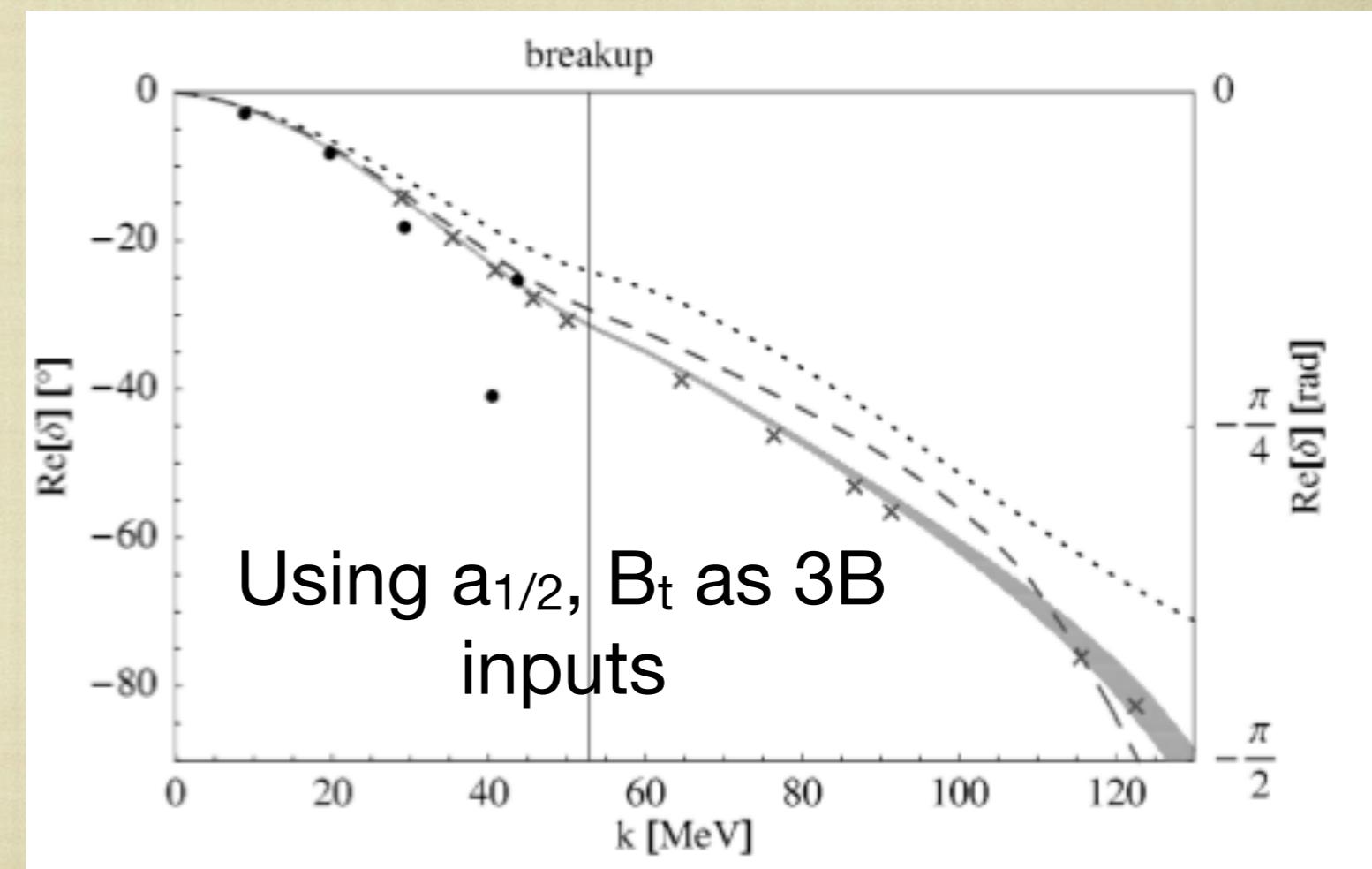
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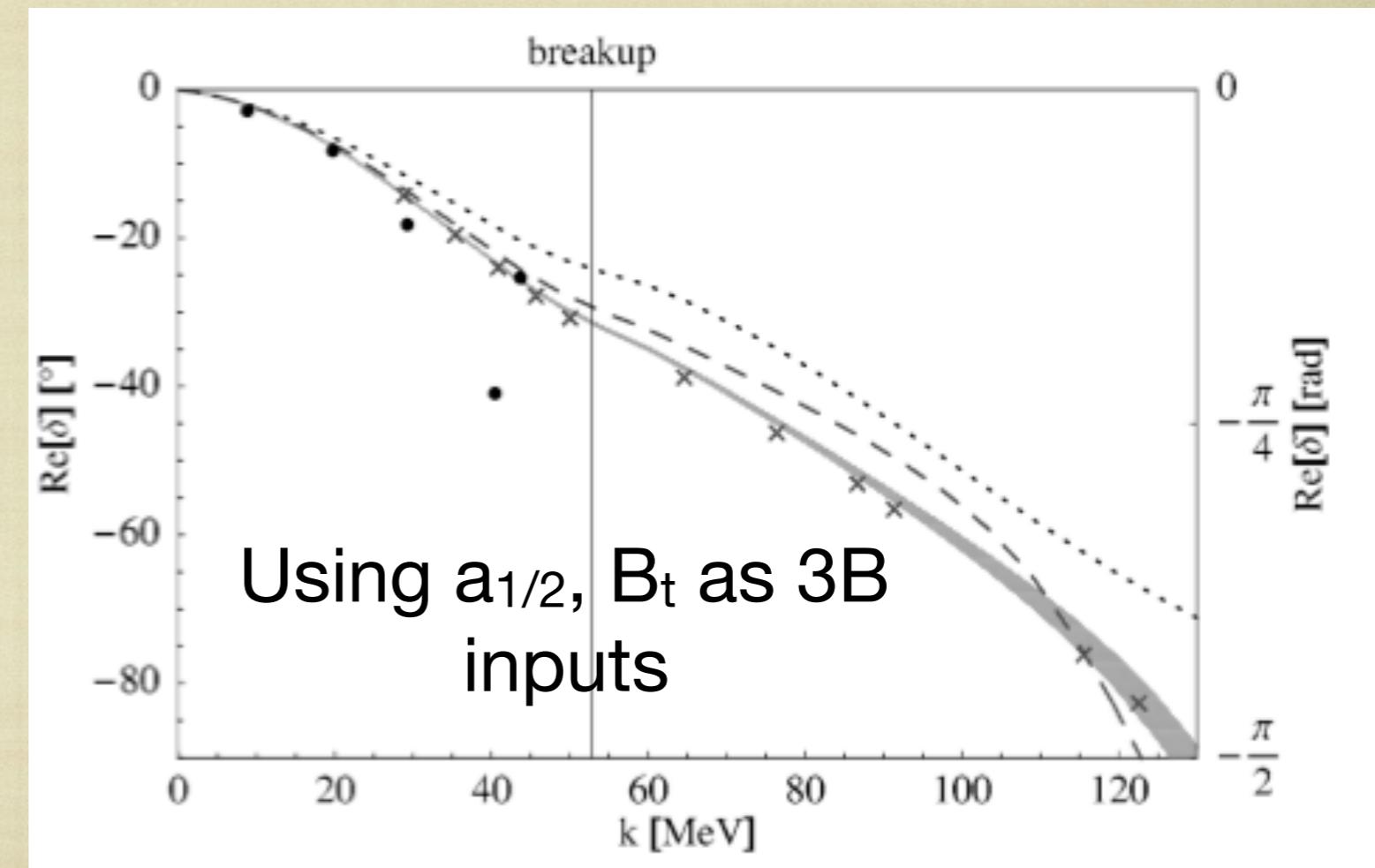
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- Alternative: replace $-\frac{1}{a} \rightarrow -\frac{1}{a} + \frac{1}{2}rk^2$ in t^{2B} , solve integral equation. Caution required!

Bedaque, van Kolck (1999)

Results at N²LO

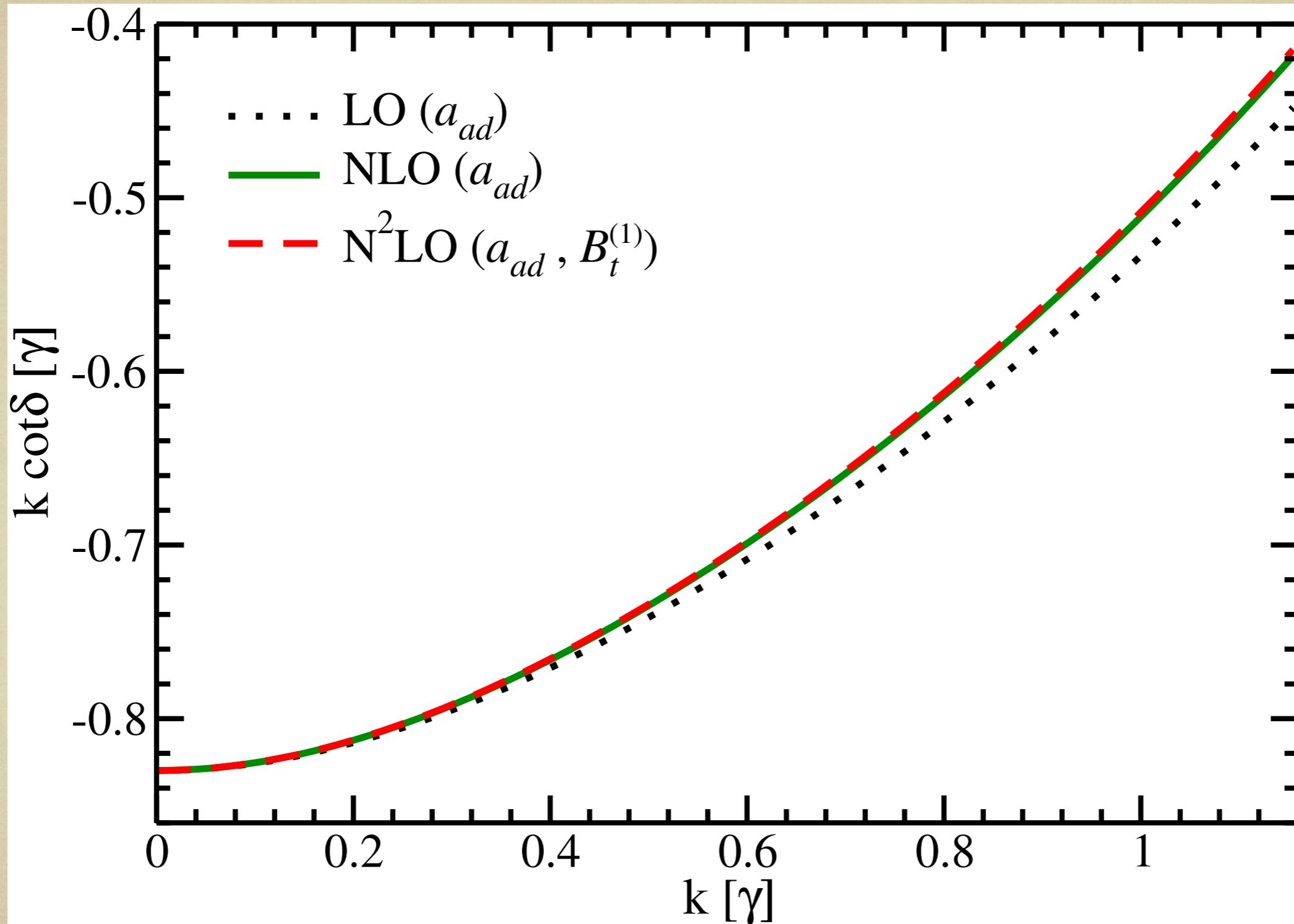
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Results at N²LO



Phase shifts predicted to better than 0.2% at N²LO

For application to n-d system see: Bedaque, Greisshammer, Hammer, Rupak (2002)