

## Geometrical interpretation of the Casimir effect

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- An object or objects immersed in vacuum modify quantum fluctuations of the EM field: there is measurable fluctuation-induced force (Casimir, 1948).
- Also true for any ground state and other fields - analog Casimir effects in condensed matter.
- There also could be topological Casimir/Aharonov-Bohm effects.
- ✓ How can we calculate it efficiently?
- ✓ How to make sense of the (sometime) infinite results?

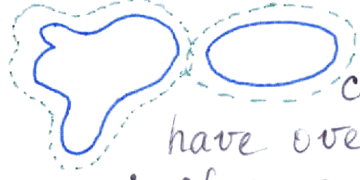
Only  $T=0$  limit, sharp boundaries and pure vacuum will be considered.

- The EM field has zero-point energy

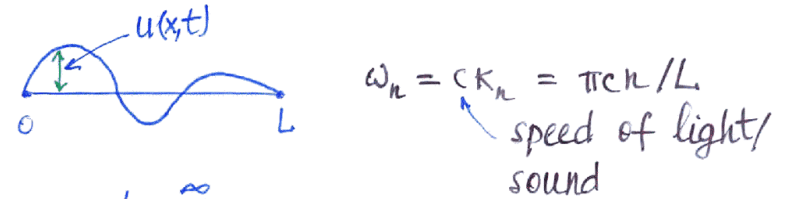
$$E = \sum_{\nu} \frac{1}{2} \hbar \omega_{\nu} \quad \text{which is infinite}$$

← modes

- An object  modifies the spectrum. This gives a self-energy relative to the vacuum. This is also infinite.

-  Objects that are close to each other have overlapping influence. This gives a finite change in the self-energy. This is the Casimir interaction (forces and torques).

Example: a scalar field  $u(x,t)$  on a one-dimensional Dirichlet interval:



$$E = \frac{\pi \hbar c}{2L} \sum_{n=1}^{\infty} n \quad \text{diverges}$$

But real materials become transparent at short wavelengths. So we can cut off the sum at  $n=N$ . Then

$$E = \frac{\pi \hbar c}{4L} (N^2 + N + 0) \quad \text{cutoff frequency}$$

← sharp

The upper limit is of order  $N = \frac{\omega_0 L}{c}$ :

$$E \approx \frac{\hbar \omega_0^2}{c} L + \hbar \omega_0 + (\#) \frac{\hbar c}{L}$$

bulk      ends smooth      finite-size

Insert another Dirichlet partition at  $x=a$ .

$$E(L) = \frac{\hbar\omega_0^2}{c}L + \hbar\omega_0 + (\#) \frac{\hbar c}{L}$$

$$\mathcal{E} = E(a) + E(L-a) =$$

$$= \underbrace{\# \hbar c \left( \frac{1}{a} + \frac{1}{L-a} \right)}_{\text{universal}} + \underbrace{\frac{\hbar\omega_0^2}{c} (a+L-a) + \hbar\omega_0 (1+1)}_{\text{a, L-independent}}$$

- universal
  - size determined by  $\hbar, c$  and macroscopic length scales.
  - if  $L \rightarrow \infty$ , then  $\mathcal{E} \approx \hbar c/a$  is uniquely determined by dimensional analysis.
  - although it is electromagnetic in origin, the charge quantum  $e$  does not appear.
  - determination of  $\#$  requires smooth cutoff; the sign determines if it is attractive or repulsive.
- $a$ -independent
- $\omega_0 \rightarrow \infty$  limit - infinities are subtracted
- non-universal
- $a, L$ -independent

### Determining the numerical prefactor.

• Assume a smooth cutoff function  $F(\frac{\hbar}{N})$

$$E = \frac{\pi \hbar c}{2L} \sum_{n=1}^{\infty} n e^{-n/N} = -\frac{\pi \hbar c}{2L} \frac{\partial}{\partial (1/N)} \left( \sum_{n=1}^{\infty} e^{-n/N} \right) =$$

$$= \frac{\pi \hbar c}{2L} \frac{e^{-\frac{1}{N}}}{(1 - e^{-1/N})^2} \xrightarrow{N \gg 1} \frac{\pi \hbar c}{2L} \left( N^2 \left( \frac{1}{12} \right) \right)$$

← any  $F(y)$

So  $\mathcal{E} = -\frac{\pi \hbar c}{24} \left( \frac{1}{a} + \frac{1}{L-a} \right)$  - attractive

- Another route: the Riemann  $\zeta$ -function:
- $\zeta(s) = \sum_{n=1}^{\infty} n^{-s}$  - defined for  $s > 1$ . There is a singularity at  $s=1$ , but it is possible to continue analytically past the singularity. Then

$$E \stackrel{R}{=} \frac{\pi \hbar c}{2L} \zeta(-1). \text{ It turns out } \zeta(-1) = -\frac{1}{12}$$

which gives same result as above.

This "works" most of the time but there are exceptions when "infinities remain".

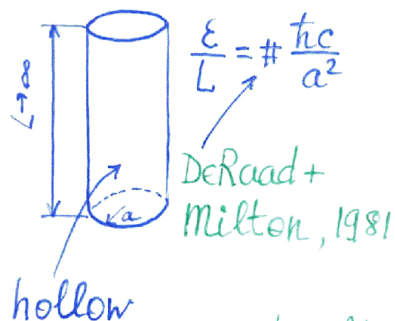
## Exceptions

Always for curved solitary boundaries:

- For a spherical shell in  $d$  spatial dimensions the Casimir pressure is infinite for even  $d$  (Bender+Milton, 1994)

Does it mean that conductive ring in two dimensions is unstable?

- The Casimir energy is finite for a cylinder, but infinite for half-cylinder!



Nesterenko + Lambiase + Scarpetta, 2001



## Formalism

We have found an efficient way to calculate the Casimir energy that does not require an explicit knowledge of the spectrum.

- It is based on the path-integral representation of the zero-point energy
- It works for a scalar field in the presence of arbitrary number of Dirichlet and/or Neumann boundaries.
- It relates two ideas that the Casimir effect can be viewed either as due to zero-point motion of the vacuum or due to fluctuations at the boundaries.

Only the case of single Dirichlet boundary is considered.

Imaginary time action for a scalar field is

$$S[w] = \frac{1}{2} \int_0^{\hbar/T} d\tau \int d^d x \left( \frac{1}{c^2} \left( \frac{\partial w}{\partial \tau} \right)^2 + (\nabla w)^2 \right)$$

imag. time

$w(\vec{r}, 0) = w(\vec{r}, \frac{\hbar}{T})$  - periodicity on the "partition function" Matsubara circle.

Calculate  $Z[w] = \int D w(\vec{r}, \tau) \exp(-S[w]/\hbar)$

over all possible forms for  $w(\vec{r}, \tau)$  satisfying various boundary conditions. Functions slowly varying in space and time contribute most.

The zero-point energy is the "free energy" per unit "length" in the imaginary time direction.

$$E = -\hbar \ln Z[w] / (\hbar/T) = -T \ln Z[w]$$

Introduce a new Dirichlet boundary.

This will constrain the field:

$$w = u + v$$

random = Laplace + constrained random

There is a unique way to associate the unconstrained field  $w$  with a constrained field  $v$  (satisfying new boundary condition):

$$w(\vec{r}, \tau) = v(\vec{r}, \tau) + u(\vec{r}, \tau)$$

Solution to  $\left( \frac{\partial^2}{c^2 \partial \tau^2} + \Delta \right) u = 0$  agreeing with  $w$  on the boundary.

Then  $S[w] = S[v] + S[u] \rightarrow Z[w] = Z[v] Z[u]$ .

$$E_{\text{casimir}} = E_{\text{constrained}} - E_{\text{unconstrained}} = T \ln Z[u]$$

The rule.

In words: the Casimir energy due to a Dirichlet boundary is negative of the zero-point energy of the modes suppressed by this boundary.

In symbols: need to solve the boundary value Laplace problem:

$$\left(\frac{\partial^2}{c^2 \partial t^2} + \Delta\right)u = 0, \quad u|_{\text{boundary}} = f(\vec{r}, t)$$

boundary ← dynamical field

After Fourier expansion

$$u(\vec{r}, t) = \sum_{\omega} u_{\omega}(\vec{r}) \exp i\omega t \rightarrow \text{Helmholtz}$$

$$\left(\Delta - \frac{\omega^2}{c^2}\right)u_{\omega} = 0, \quad u_{\omega}|_{\text{boundary}} = f_{\omega}(\vec{r}) \text{ - put into } S[u]:$$

$$S\{u[f]\} = \frac{1}{2} \int_0^{t/T} dt \int [u \nabla u] d\vec{\sigma} = \frac{\hbar}{2T} \sum_{\omega} \int f_{\omega} [\nabla u_{\omega}] d\vec{\sigma} =$$

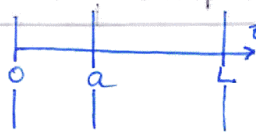
discontinuity      discontinuous

$$u_{\omega} \propto f_{\omega} \frac{\hbar}{2T} \sum_{\omega, \nu} \frac{|f_{\omega \nu}|^2}{\lambda_{\nu} (\omega/c)} \text{ - Gaussian}$$

modes      geometry

$$\mathcal{E} = \frac{I}{2} \sum_{\omega, \nu} \ln(2\pi T \lambda_{\nu} (\omega/c) / \hbar) \xrightarrow{T \rightarrow 0} \frac{\hbar c}{2T} \sum_{\nu} \int_0^{\infty} dk \ln \lambda_{\nu}(k)$$

An example: three planes at  $z=0, a, L$  in  $d$  dimensions



What is the Casimir pressure on the plane at  $z=a$ ?

$$u_{\omega}(\vec{r}) = \sum_{\vec{q}} u_{\omega \vec{q}}(z) \exp i\vec{q} \cdot \vec{r}_{\perp}$$

to  $z$

$$\left(\frac{d^2}{dz^2} - q^2 - \frac{\omega^2}{c^2}\right)u_{\omega \vec{q}} = 0, \quad u_{\omega \vec{q}}|_{0,L} = 0, \quad u_{\omega \vec{q}}|_a = f_{\omega \vec{q}}$$

The particular solution

$$u_{\omega \vec{q}}(z) = f_{\omega \vec{q}} \frac{\sinh(|x|z)}{\sinh(|x|a)}, \quad 0 \leq z \leq a, \quad x^2 = q^2 + \frac{\omega^2}{c^2}$$

$$u_{\omega \vec{q}}(z) = f_{\omega \vec{q}} \frac{\sinh(|x|(L-z))}{\sinh(|x|(L-a))}, \quad a < z \leq L$$

becomes localized for large  $|x|$

$$S = \frac{\hbar A}{2T} \sum_{\omega, \vec{q}} |x| (\coth|x|a + \coth|x|(L-a)) |f_{\omega \vec{q}}|^2$$

area      Gaussian

$$\frac{\mathcal{E}}{A} \rightarrow -\frac{\hbar}{2} \int \frac{d\omega d^d q}{(2\pi)^d} \ln(\dots) \text{ - divergent}$$

$$P = -\frac{\partial(\mathcal{E}/A)}{\partial a} = \# \hbar c [(L-a)^{-d-1} - a^{-d-1}] \text{ - finite}$$

known  $d$ -dependence, Ambjørn + Wolfram, 1983

An example: the circle of radius  $a$

$$\left( \frac{1}{\rho} \frac{\partial}{\partial \rho} \left( \rho \frac{\partial}{\partial \rho} \right) + \frac{1}{\rho^2} \frac{\partial^2}{\partial \varphi^2} - \frac{\omega^2}{c^2} \right) u_\omega = 0, \quad u_\omega|_a = f_\omega(\varphi)$$

The solution is

$$\left\{ \begin{aligned} u_\omega &= \sum_{n=-\infty}^{\infty} \frac{I_n(\omega \rho / c)}{I_n(\omega a / c)} f_\omega e^{in\varphi}, & \rho \leq a \\ u_\omega &= \sum_{n=-\infty}^{\infty} \frac{K_n(\omega \rho / c)}{K_n(\omega a / c)} f_\omega e^{in\varphi}, & \rho > a \end{aligned} \right.$$

where  $I_n(z)$  and  $K_n(z)$  are modified Bessel functions

$$S = \frac{\pi \hbar}{T} \sum_{\omega, n} \frac{|f_\omega|^2}{I_n(\omega a / c) K_n(\omega a / c)} \quad \text{- Gaussian}$$

$$\mathcal{E} = \frac{\hbar c}{2\pi} \sum_{n=-\infty}^{\infty} \int_0^\infty dx \ln(I_n(x) K_n(x)) \quad \leftarrow \text{divergent.}$$

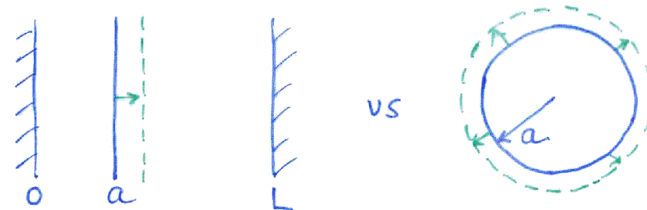
$$P = -\frac{\partial \mathcal{E}}{2\pi a \partial a} = -\frac{\hbar c}{4\pi^2 a^3} \sum_{n=-\infty}^{\infty} \int_0^\infty x dx \frac{d}{dx} \ln(I_n(x) K_n(x))$$

Sen, 1981

Divergences exist in any curved geometry!

Why? Largely (but not entirely) due to infinite self-energy (per unit area) of a boundary (Deutsch + Candelas, 1979; Jaffe et al., 2002+, Barton, 2004). How come?

The force is the energy change per infinitesimal displacement:



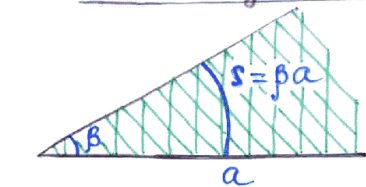
planar: the area does not change, the force is finite

curved: the area changes, the force is infinite

Our contention: geometry-dependent "infinities", sometime "removed" and sometime not, are finite non-universal pieces of the effect having geometrical origin and measurable consequences.

Test: need a curved geometry where various geometrical and universal pieces of the effect can be unambiguously separated; ideally a system characterized by more than one length scale so that the area of the boundary is independent of its curvature.

The wedge-arc geometry in 2d



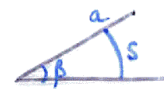
Circular Dirichlet arc with Dirichlet  $u(\rho, 0) = u(\rho, \beta) = 0$ , or periodic,  $u(\rho, \varphi) = u(\rho, \varphi + \beta)$ , edges.

- close relative of the circle problem; periodic version includes the circle as a special case  $\beta = 2\pi$
- if  $s = \beta a$  is kept fixed, the answer is expected to be universal (almost true)
- Both the force and torque exist here.

Solution employs:

- general cutoff function reflecting material's transparency to high-energy modes
- extremely useful and accurate Debye expansion (Milton+DeRaad+Schwinger, 1978).

$\beta = s/a \ll 1$



line tension  $\rightarrow$  universal  $\rightarrow$  log-accuracy

Dirichlet:  $E = d \cdot s - \frac{\hbar c s}{256 \pi a^2} \ln \frac{\omega_0 s}{c} + \frac{\pi \hbar c}{48 s}$

Labels:  $d \cdot s$  is non-universal;  $\frac{\hbar c s}{256 \pi a^2} \ln \frac{\omega_0 s}{c}$  is nearly-universal;  $\frac{\pi \hbar c}{48 s}$  is universal.

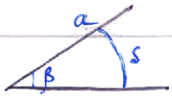
Periodic:  $E = d \cdot s - \frac{\hbar c s}{256 \pi a^2} \ln \frac{\omega_0 s}{c} + \frac{\pi \hbar c}{12 s} - \frac{0.08808 \hbar c}{2 \pi a}$

$d = \frac{\hbar \omega_0}{c / \omega_0} = \frac{\hbar \omega_0^2}{c}$   
 can be given exactly

for  $s = 2\pi a$  and exponential cutoff function this was first given by Sen, 1981



Solution again:



$$\text{D: } \mathcal{E} = \alpha \cdot s - \frac{\hbar c s}{256 \pi a^2} \ln \frac{\omega_0 s}{c} + \frac{\pi \hbar c}{48 s}$$

$$\text{P: } \mathcal{E} = \alpha \cdot s - \frac{\hbar c s}{256 \pi a^2} \ln \frac{\omega_0 s}{c} + \frac{\pi \hbar c}{12 s} - \frac{0.08808 \hbar c}{2 \pi a}$$

Summary of properties

Property Contribution	sensitive to topology	additive
non-universal	no	yes
nearly-universal	no	nearly
universal	yes	no

Why are the non-universal terms insensitive to topology and (largely) additive?

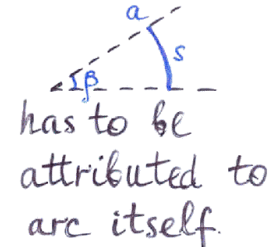
Because they are dominated by high energy localized modes - they are (almost) local properties of the boundary.

Geometrical interpretation of non-universal contributions

$$\mathcal{E}(\omega_0) = \underbrace{\alpha \cdot s}_{\text{line energy}} - \frac{\hbar c s}{256 \pi a^2} \ln \frac{\omega_0 s}{c}$$

line energy

How to understand this?



- Finite-size correction to line tension  
 $\alpha \rightarrow \alpha - \frac{\hbar c}{256 \pi a^2} \ln \frac{\omega_0 s}{c}$  or
- Marginally non-local contribution to curvature (bending) rigidity  
 $K = \frac{1}{a}$  - local curvature  
 $\left(-\frac{\hbar c}{256 \pi a^2} \ln \frac{\omega_0 s}{c}\right) s = \left(-\frac{\hbar c}{256 \pi} K^2 \ln \frac{\omega_0 s}{c}\right) s$

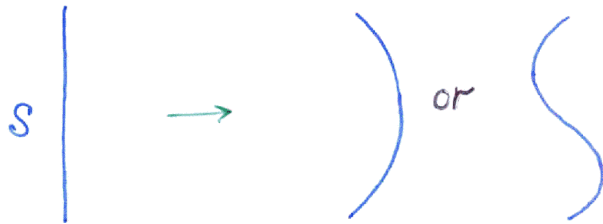
### Arbitrary smooth boundary

$$\mathcal{E}(\omega_0) = \alpha \cdot s - \frac{\hbar c s}{256\pi} \kappa^2 \ln \frac{\omega_0 s}{c}$$

For infinitesimally small  $s$  this looks almost like a differential of  $\mathcal{E}(\omega_0)$ .  
Integrate over the boundary and fix the log by hand:

$$\mathcal{E}(\omega_0) = \alpha \cdot S - \underbrace{\left( \frac{\hbar c}{256\pi} \int \kappa^2(s) ds \right)}_{\int ds} \ln \frac{\omega_0 S}{c} \quad (\text{Sen, 1981})$$

Casimir wrinkling of sufficiently long straight fixed length boundary



### Conclusions

- Casimir energy of a smooth boundary is generally non-universal.
- "Infinities" removed by formal methods are either irrelevant geometry-independent constants or measurable local contributions into elastic properties of the boundary.
- "Infinities" that cannot be removed by formal techniques represent non-local (logarithmic in  $\omega_0$ ) contributions into boundary's elasticity. This happens for a spherical shell in even dimensions (Sen almost mentions this as a possibility in 2d).
- These issues do not exist when the Casimir force is not due to shape change.

How to see that even dimensions  
are special?

For a sphere of radius  $a$  the non-universal part of the effect must have the form *even function of curvature*

$$\begin{aligned} E(\omega_0) &= \int E(k) d\sigma = \int \sum_{n=0}^N \gamma_n k^{2n} d\sigma \approx \\ &\approx \sum_{n=0}^N \gamma_n a^{-2n} a^{d-1} = \sum_{n=0}^N \gamma_n a^{d-1-2n} \end{aligned}$$

The rigidity constants  $\gamma_n$  have dimension of energy/(length)<sup>d-1-2n</sup> and they are determined by microscopic physics, i.e. by  $\hbar\omega_0$  and  $c/\omega_0$  (microscopic energy + length scales)

$$\gamma_n \approx \frac{\hbar\omega_0}{\left(\frac{c}{\omega_0}\right)^{d-1-2n}} = \hbar c \left(\frac{\omega_0}{c}\right)^{d-2n}$$

$$E(\omega_0) \approx \hbar c \sum_{n=0}^N \left(\frac{\omega_0}{c}\right)^{d-2n} a^{d-1-2n} \left. \begin{array}{l} \text{more restrictive} \\ \text{than dimensional} \\ \text{analysis implies} \end{array} \right\}$$

$d-2N \geq 0$  fixes the # of terms.

For  $d$  even there are  $\frac{d}{2}+1$  terms and the last one is  $\omega_0$ -independent  $\rightarrow$  marginal case  $\rightarrow$  deviation from locality.