

Holography for $N=2^*$ on S^4

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N=4 SYM

N=2* theory in flat space
massless vector mult.
& massive hypermult.

gravitational dual

Pilch-Warner flow w/ flat-sliced domain walls.

Pilch-Warner (2000)

Buchel-Peet-Polchinski (2000)

Evans-Johnson-Petrini (2000)

Localization

N=2* theory on S^4

Free energy F

Pestun (2007)

Buchel-Russo-Zarembo (2013)

gravitational dual?

On-shell action $S_{\text{on-shell}}$

Localization

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large N and large 't Hooft coupling $\lambda = g_{\text{YM}}^2 N$

$$F_{S^4} = -\frac{N^2}{2}(1 + m^2 a^2) \log \frac{\lambda(1 + m^2 a^2) e^{2\gamma + \frac{1}{2}}}{16\pi^2}$$

Diagram illustrating the localization formula for the free energy F_{S^4} on S^4 . The formula is:

$$F_{S^4} = -\frac{N^2}{2}(1 + m^2 a^2) \log \frac{\lambda(1 + m^2 a^2) e^{2\gamma + \frac{1}{2}}}{16\pi^2}$$

The terms are labeled as follows:

- m : mass of hyper (indicated by an arrow pointing to $m^2 a^2$)
- a : radius of S^4 (indicated by an arrow pointing to a^2)
- γ : Euler const (indicated by an arrow pointing to γ)

Ambiguities due to UV subtractions are eliminated in 3rd derivative:

$$\frac{d^3 F_{S^4}}{d(ma)^3} = -2N^2 \frac{ma(m^2 a^2 + 3)}{(m^2 a^2 + 1)^2}$$

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$$\begin{aligned}\frac{d^3 F_{S^4}}{d(ma)^3} &= -2N^2 \frac{ma(m^2 a^2 + 3)}{(m^2 a^2 + 1)^2} \\ &= N^2 (-6ma + 10m^3 a^3 + \dots) \\ &\quad \text{for } ma \ll 1\end{aligned}$$

Buchel initiated holographic study.

Showed that $m^3 a^3$ -term in the free energy was **NOT** matched by S^4 -sliced flow in the 5d Pilch-Warner supergravity model. Buchel (2013)

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In this talk:

- Explain why Pilch-Warner does not capture flow of $N=2^*$ theory on S^4 .
- Construct flow in suitable truncation of 5d $N=8$ supergravity.
- Use holographic renormalization *to exactly match full functional form of the universal part of the free energy*.

This offers a precision-test of the gauge-gravity duality in the context of a Euclidean non-conformal setting.

Plan

- $N=2^*$ on S^4
- Gravity dual
- Holographic renormalization
- Comments

*Based on 1311.1508 with
Nikolay Bobev, Dan Freedman, and Silviu Pufu*

Review of $N=2^*$: setup in flat space

$N=4$ SYM

$$A_\mu, \quad \lambda_1, \lambda_2, \lambda_3, \lambda_4, \quad X_1, X_2, X_3, X_4, X_5, X_6$$

4 gluinos 6 scalars

Global R-symmetry: $SU(4)_R \sim SO(6)_R$

In $N=2$ formulation:

$N=2$ vector multiplet: $A_\mu, \quad \begin{matrix} \psi_1 = \lambda_3 \\ \psi_2 = \lambda_4 \end{matrix}, \quad Z_3 = \frac{1}{\sqrt{2}}(X_3 + iX_6)$

$N=2$ hypermultiplet: $\begin{matrix} \chi_1 = \lambda_1 \\ \chi_2 = \lambda_2 \end{matrix}, \quad \begin{matrix} Z_1 = \frac{1}{\sqrt{2}}(X_1 + iX_4) \\ Z_2 = \frac{1}{\sqrt{2}}(X_2 + iX_5) \end{matrix}$

Add mass for hyper $\longrightarrow N=2^*$ theory

Review of $N=2^*$: global symmetries

$SU(4)_R \sim SO(6)_R \xrightarrow{N=2}$

	$SU(2)_V$	$SU(2)_H$	$U(1)_R$
A_μ	0	0	0
Φ	0	0	+2
$\psi_{1,2}$	1/2	0	+1
$\tilde{\psi}_{1,2}$	1/2	0	-1
$Z_{1,2}$	$1/2^\dagger$	1/2	0
$\chi_{1,2}$	0	1/2	-1
$\tilde{\chi}_{1,2}$	0	1/2	+1

Mass term for $N=2^*$ in flat space

$$\mathcal{L}_m^{\mathbb{R}^4} = m^2 \text{tr} (|Z_1|^2 + |Z_2|^2) + m \text{tr} (\chi_1 \chi_1 + \chi_2 \chi_2 + \text{h.c.})$$

breaks $SU(2)_V \times SU(2)_H \times U(1)_R \longrightarrow SU(2)_V \times U(1)_H$.

Review of N=2*: holography

Mass term for N=2* in flat space $SU(2)_V \times U(1)_H$,

$$\mathcal{L}_m^{\mathbb{R}^4} = m^2 \operatorname{tr} (|Z_1|^2 + |Z_2|^2) + m \operatorname{tr} (\chi_1 \chi_1 + \chi_2 \chi_2 + \text{h.c.})$$



dimension 2 operator



dual scalar ϕ



dimension 3 operator



dual scalar ψ

5d Pilch-Warner flow has flat domain walls

$$ds^2 = dr^2 + e^{2A(r)} \delta_{ij} dx^i dx^j$$

and non-trivial radial profiles for the two scalars

Type IIB lift: $SU(2)_V \times U(1)_H \times U(1)_Y$

Now put the theory on S^4

Euclidean formalism for supergravity

Festuccia and Seiberg (2011)
“Rigid SUSY in curved spacetime”

Freedman and Pufu (2013)
“Holography of F-maximization”

On S^4 : $N=4$ SYM

$N=4$ SYM is conformal,

so just need conformal coupling for the scalars:

$$\mathcal{L}_{\mathcal{N}=4}^{S^4} = \mathcal{L}_{\mathcal{N}=4}^{\mathbb{R}^4} \Big|_{\eta_{\mu\nu} \rightarrow g_{\mu\nu}} + \frac{2}{a^2} \text{tr} \left(|Z_1|^2 + |Z_2|^2 + |Z_3|^2 \right)$$

where a is the radius of the sphere.

On S^4 : $N=2^*$ SYM

$N=2^*$ theory is ***NOT*** conformal,

so in addition to conformal coupling for the scalars,

$$\frac{2}{a^2} \text{tr}(|Z_1|^2 + |Z_2|^2 + |Z_3|^2)$$

the presence of the mass terms

$$m^2 \text{tr}(|Z_1|^2 + |Z_2|^2) + m \text{tr}(\chi_1 \chi_1 + \chi_2 \chi_2 + \text{h.c.})$$

requires another term

$$\frac{im}{2a} \text{tr}(Z_1^2 + Z_2^2 + \text{h.c.})$$

in order for supersymmetry to be preserved.

$$\text{SUSY transf w/ } S^4 \text{ Killing spinors} \quad \nabla_\mu \epsilon_\pm = \pm \frac{i}{2a} \sigma_\mu \tilde{\epsilon}_\pm$$

On S^4 : $N=2^*$ SYM

Consequences of $\frac{im}{2a} \text{tr} (Z_1^2 + Z_2^2 + \text{h.c.})$

1) $SU(2)_V \times U(1)_H$ is broken to $U(1)_V \times U(1)_H$

2) The gravity dual can be expected to involve one more scalar dual to this dimension 2 operator.



dual scalar χ

The third scalar turns out to be necessary for the gravitational flow dual to $N=2^*$ theory on S^4 .

This is why the two-scalar Pilch-Warner model does not capture this flow on S^4 .

Holographic dual of N=2* SYM on S⁴

Fields $g_{\mu\nu}$, ϕ , ψ , and χ

with the three scalars dual to the three operators

$$\mathcal{O}_\phi = \text{tr}(|Z_1|^2 + |Z_2|^2), \quad \mathcal{O}_\psi = \text{tr}(\chi_1\chi_1 + \chi_2\chi_2 + \text{h.c.}), \quad \mathcal{O}_\chi = \text{tr}(Z_1^2 + Z_2^2 + \text{h.c.}).$$

symmetry $U(1)_V \times U(1)_H \times U(1)_Y$

Intriligator (1998)

Buchel-Peet-Polchinski (2000)

bonus symmetry at large-N

Truncation of N=8 gauged supergravity in 5d:

$$\mathcal{L} = \frac{1}{2\kappa^2} \left[-R + \frac{12\partial_\mu\eta\partial^\mu\eta}{\eta^2} + \frac{4\partial_\mu z\partial^\mu\tilde{z}}{(1-z\tilde{z})^2} + V \right],$$

$$V \equiv -\frac{4}{L^2} \left(\frac{1}{\eta^4} + 2\eta^2 \frac{1+z\tilde{z}}{1-z\tilde{z}} + \frac{\eta^8}{4} \frac{(z-\tilde{z})^2}{(1-z\tilde{z})^2} \right).$$

$$\eta = e^{\phi/\sqrt{6}}$$

$$z = \frac{1}{\sqrt{2}}(\chi + i\psi)$$

$$\tilde{z} = \frac{1}{\sqrt{2}}(\chi - i\psi)$$

Holographic dual of $N=2^*$ SYM on S^4

1) Bulk theory:

$$V = -3 - 2\phi^2 - 2\chi^2 - \frac{3}{2}\psi^2 + \dots$$

Scale dimension $\Delta = 2 + \sqrt{4 + m^2}$.

Two scaling dimension 2, one 3.

2) Truncation $z = -\tilde{z}$

gives Pilch-Warner model with flat-sliced domain wall solutions.

$$\eta = e^{\phi/\sqrt{6}}$$

$$z = \frac{1}{\sqrt{2}}(\chi + i\psi)$$

$$\tilde{z} = \frac{1}{\sqrt{2}}(\chi - i\psi)$$

Holographic dual of N=2* SYM on S⁴

3) We want S⁴-sliced domain wall solutions. Ansatz:

$$ds^2 = L^2 e^{2A(r)} ds_{S^4}^2 + dr^2$$

$$\eta = \eta(r) \quad z = z(r) \quad \tilde{z} = \tilde{z}(r)$$

Note: Euclidean solution, z and \tilde{z} are indep!

BPS equations:

$$z' = \frac{3\eta'(z\tilde{z} - 1) [2(z + \tilde{z}) + \eta^6(z - \tilde{z})]}{2\eta [\eta^6(\tilde{z}^2 - 1) + \tilde{z}^2 + 1]},$$

$$\tilde{z}' = \frac{3\eta'(z\tilde{z} - 1) [2(z + \tilde{z}) - \eta^6(z - \tilde{z})]}{2\eta [\eta^6(z^2 - 1) + z^2 + 1]},$$

$$(\eta')^2 = \frac{[\eta^6(z^2 - 1) + z^2 + 1] [\eta^6(\tilde{z}^2 - 1) + \tilde{z}^2 + 1]}{9L^2\eta^2(z\tilde{z} - 1)^2}$$

$$e^{2A} = \frac{(z\tilde{z} - 1)^2 [\eta^6(z^2 - 1) + z^2 + 1] [\eta^6(\tilde{z}^2 - 1) + \tilde{z}^2 + 1]}{\eta^8(z^2 - \tilde{z}^2)^2}$$

(imply EOM)

Holographic dual of $N=2^*$ SYM on S^4

Have not found analytic solution to BPS eqs, but can analyze UV and IR behavior:

UV behavior: $r \rightarrow \infty$.

Solution approaches Euclidean AdS_5 (scalars $\rightarrow 0$)

$$ds_5^2 = dr^2 + L^2 \sinh^2 \left(\frac{r}{L} \right) ds_{S^4}^2$$

Solving the BPS eqs iteratively order by order gives

$$e^{2A} = \frac{e^{2r}}{4} + \frac{1}{6}(\mu^2 - 3) + \dots \quad \eta = 1 + e^{-2r} \left[\frac{2\mu^2}{3}r + \frac{\mu(\mu + v)}{3} \right] + \dots$$

$$\frac{1}{2}(z + \tilde{z}) = e^{-2r} [2\mu r + v] + \dots \quad \frac{1}{2}(z - \tilde{z}) = \mp \mu e^{-r} \mp \dots$$

Two parameters: μ mass, v vev

Holographic dual of $N=2^*$ SYM on S^4

The holographic dual of $N=2^*$ SYM should depend on
Just one dimensionless parameter: ma

IR behavior:

Solution approaches Euclidean flat space smoothly as
 S^4 shrinks to zero size and scalars \rightarrow constants.

Solve BPS equations iteratively as $r \rightarrow 0$

Smoothness condition gives 1-parameter family

$$\begin{aligned}\eta &= \eta_0 + O(r^2) \\ e^{2A} &= O(r^2) \\ \frac{1}{2}(z + \tilde{z}) &= \sqrt{\frac{\eta_0^6 - 1}{\eta_0^6 + 1}} \frac{\eta_0^6}{\eta_0^6 + 2} + O(r^2) \\ &\text{etc}\end{aligned}$$

Holographic dual of $N=2^*$ SYM on S^4

Now match

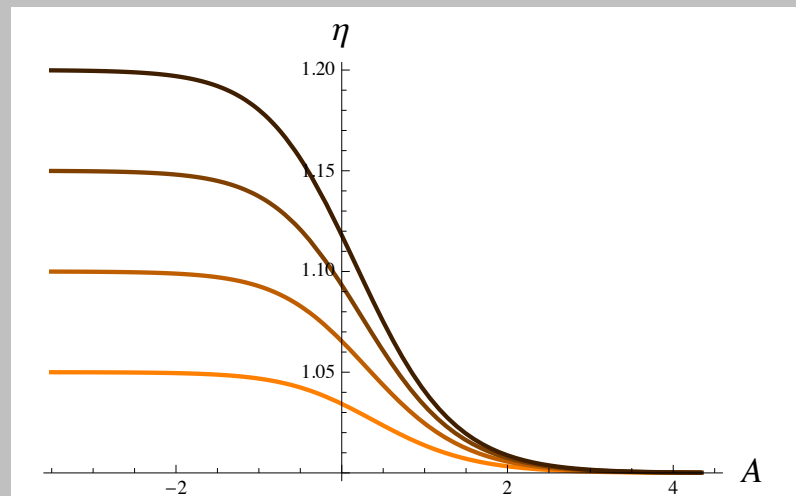
UV and IR behavior

Numerical solution
interpolates between UV
and IR region: fixes the
two UV parameters in terms of the single IR parameter:

$$\mu(\eta_0) \quad \text{and} \quad v(\eta_0)$$

Extract from the numerics

$$v(\mu) = -2\mu - \mu \log(1 - \mu^2)$$





Official Old San Marcos Pass. 3 miles, 1165 ft.
[JoeP](#) Personal Record 23:02min. Feb 26, 2014.

Recap:

We have constructed a candidate dual for the RG flow of $N=2^*$ on S^4

Next is calculation of the *on-shell action*.

Recap:

We have constructed a candidate dual for the RG flow of $N=2^*$ on S^4

Next is calculation of the *on-shell action*.

Divergent, but standard systematic technique for handling it with ***infinite counterterms***:

Holographic renormalization

Leaves ambiguity of ***finite counterterms***

Holographic renormalization

Finite counterterms tricky.

For *flat-sliced* domain walls, one can use the **Bogomolnyi trick** to determine the counterterms.

Idea:

Supergravity theory with several scalars, Kahler potential, and scalar potential given in terms of superpotential as

$$V = \frac{1}{2} K^{ij} \partial_i W \partial_j W - \frac{4}{3} W^2$$

gives BPS eqs

the action can be rearranged to sum of squares:

$$S = \int d^4x \int^{r_0} \left(e^{4A} \left[-3 \left(A' - \frac{2}{3} W \right)^2 + \frac{1}{2} K_{ij} (\phi^{i'} - K^{il} \partial_l W) (\phi^{j'} - K^{jm} \partial_m W) \right] - \frac{d}{dr} (e^{4A} W) \right)$$

Bdr counterterms thus fixed by SUSY:

$$S_W = \int d^4x e^{4A(r_0)} W(\phi_i(r_0))$$

Holographic renormalization

- Our flows have S^4 -slicing.
The flat-sliced limit is only consistent for $\tilde{z} = \pm z$
- There is no superpotential W for our scalar potential.

No Bogo. for us!!?

Holographic renormalization

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The flat-sliced limit is only consistent for $\tilde{z} = \pm z$
- There is no superpotential W for our scalar potential.

No Bogo. for us!!?

Idea:

Our scalar potential has an approximate superpotential, valid near the boundary:

$$W = \frac{3}{2} + \phi^2 + \frac{1}{2}\psi^2 + \frac{1}{2}\chi^2 + \sqrt{\frac{2}{3}}\phi\psi^2 + \frac{1}{4}\psi^4$$

Use it to determine finite counterterms for flat-sliced case.

Call upon *counterterm universality!*

Satisfies multiple checks & truncations

$$S_{\text{susy}} = \int d^4x \sqrt{\gamma} W$$

On-shell action & the free energy

$$S_{\text{ren}} = S_{5\text{D}} + S_{\text{GH}} + S_{\text{susy}}, \quad \text{with} \quad S_{\text{susy}} = \int d^4x \sqrt{\gamma} W,$$

Show that the derivative of the free energy with respect to the source parameter is given in terms of 1-point functions:

$$\frac{dF}{d\mu} = \frac{N^2}{2\pi^2} \int d^4x \sqrt{g_0} \left(\langle \mathcal{O}_\psi \rangle \frac{\partial \psi_0}{\partial \mu} + \langle \mathcal{O}_\phi \rangle \frac{\partial \phi_0}{\partial \mu} + \langle \mathcal{O}_\chi \rangle \frac{\partial \chi_0}{\partial \mu} \right)$$

where

$$\langle \mathcal{O}_\psi \rangle = \lim_{\epsilon \rightarrow 0} \frac{1}{\epsilon^{3/2}} \frac{1}{\sqrt{\gamma}} \frac{\delta S_{\text{ren}}}{\delta \psi} \quad \text{and} \quad \langle \mathcal{O}_\phi \rangle = \lim_{\epsilon \rightarrow 0} \frac{\log \epsilon}{\epsilon} \frac{1}{\sqrt{\gamma}} \frac{\delta S_{\text{ren}}}{\delta \phi}$$

and we have used $1/4\pi G_5 = N^2/2\pi^2$,

$$\frac{dF}{d\mu} = \frac{N^2}{2\pi^2} \text{vol}_0(S^4) \left(4\mu - 12v(\mu) \right) = N^2 \left(\frac{1}{3}\mu - v(\mu) \right)$$

On-shell action & the free energy

Take two more derivatives:

$$\frac{d^3 F}{d\mu^3} = -N^2 v''(\mu) = -2N^2 \frac{\mu(3 - \mu^2)}{(1 - \mu^2)^2}$$

using

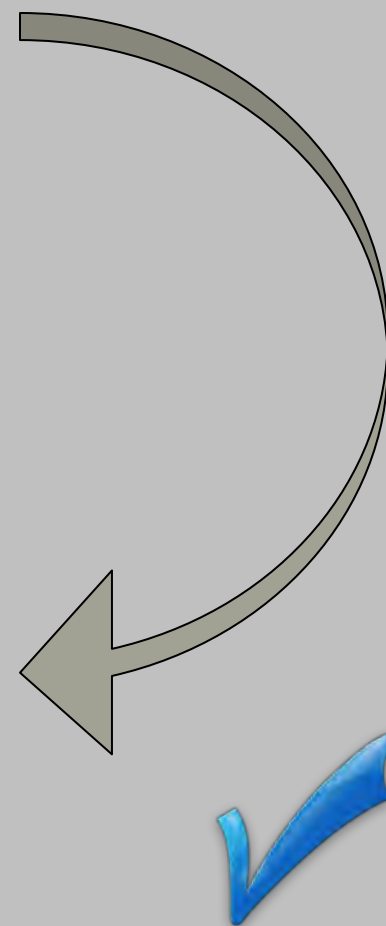
$$v(\mu) = -2\mu - \mu \log(1 - \mu^2)$$

from the UV/IR match in the BPS flow solution.

Compare with the field theory result:

$$\frac{d^3 F_{S^4}}{d(ma)^3} = -2N^2 \frac{ma(m^2 a^2 + 3)}{(m^2 a^2 + 1)^2}$$

Perfect match after identification $\mu = \pm ima$.



Comments

1) Finite counterterms were key for cancelation of μ^3 terms

Could we have found a match without using supersymmetry and universality argument to fix finite counterterms?

Yes.

One can list the possible ambiguous finite counterterms, such as

$$R[\gamma] \psi^2, \quad (\log \epsilon)^{-1} \psi^2 \phi, \quad \psi^4 \quad \text{etc}$$

and calculate their potential contribution to $\frac{dF}{d\mu}$

Turns out that they contribute only μ or μ^3 , but never $v(\mu)$

So all ambiguity of finite counterterms eliminated in $\frac{d^5 F}{d\mu^5}$

➡ Full functional match with field theory result.

2) Why $\frac{d^3 F}{d\mu^3}$?

Superconformal theory on S^4 has free energy of the form

$$F = \alpha_2 \frac{a^2}{\epsilon^2} + \alpha_0 - a_{\text{anom}} \log \frac{a}{\epsilon} + \mathcal{O}(\epsilon/a)$$

small distance cutoff ϵ

For $N=2^*$ theory on S^4 , F can also depend on dim'less $m^2 \epsilon^2$

For small $m^2 \epsilon^2$, the coefficients of the non-universal terms can be expressed as

$$\alpha_2 = \tilde{\alpha}_2 + m^2 \epsilon^2 \beta_2 + O(m^4 \epsilon^4) \quad \text{and} \quad \alpha_0 = \tilde{\alpha}_0 + O(m^2 \epsilon^2)$$

So the non-universal contributions $\tilde{\alpha}_2 \frac{a^2}{\epsilon^2} + \tilde{\alpha}_0 + \beta_2 m^2 a^2$

are eliminated in $\frac{d^3 F}{d(ma)^3}$

$$\mathbf{3)} \quad \frac{d^3 F_{S^4}}{d(ma)^3} = -2N^2 \frac{ma(m^2 a^2 + 3)}{(m^2 a^2 + 1)^2}$$

**What is special
about $m^2 a^2 = -1$?**

Recall the mass terms in N=2* on S⁴:

$$\frac{2}{a^2} \text{tr}(|Z_1|^2 + |Z_2|^2 + |Z_3|^2) + m^2 \text{tr}(|Z_1|^2 + |Z_2|^2) + \frac{im}{2a} \text{tr}(Z_1^2 + Z_2^2 + \text{h.c.})$$

Write $Z = \frac{1}{\sqrt{2}}(A + iB)$ to find

$$\begin{aligned} \mathcal{L}_{\text{scalar}} &= \frac{1}{2} \left[\left(\frac{2}{a^2} + i\frac{m}{a} + m^2 \right) A^2 + \left(\frac{2}{a^2} - i\frac{m}{a} + m^2 \right) B^2 \right] \\ &= \frac{1}{2a^2} [(1 + ima)(2 - ima)A^2 + (1 - ima)(2 + ima)B^2] \\ &\quad \quad \quad -1 < ima < 2 \qquad \qquad \quad -2 < ima < 1 \end{aligned}$$

So $ma = \pm i$ is the tachyon threshold!

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Happy Birthday, Joe!



thank you
for all the good you have done for us