



ON THE ORIGIN, NATURE, AND MIXING OF  
MULTIPHASE GAS IN ASTROPHYSICS

KITP  
10.15.20

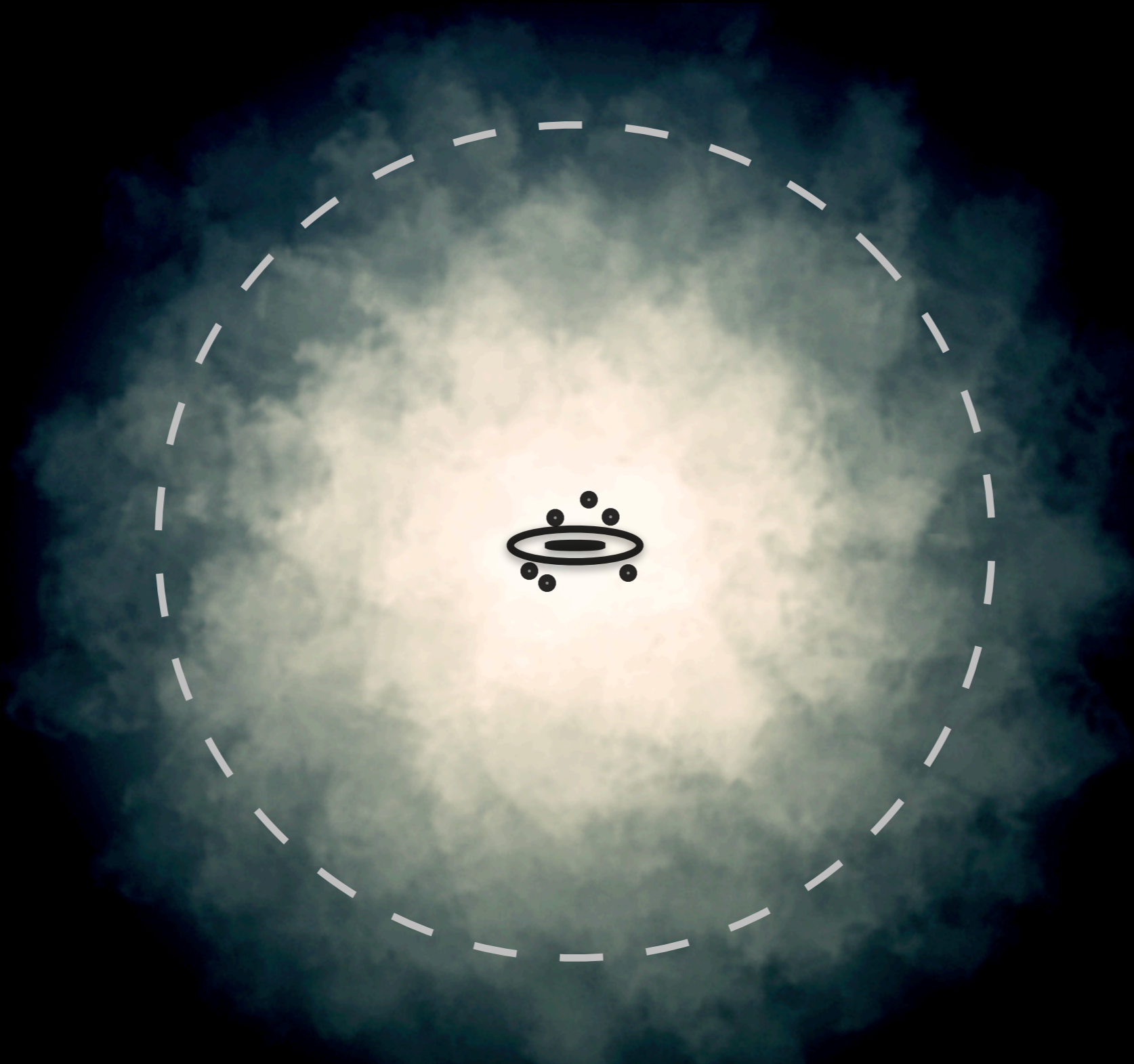
# The Impact of **Cosmic Rays** on **Thermal Instability** in the Circumgalactic Medium

IRYNA BUTSKY

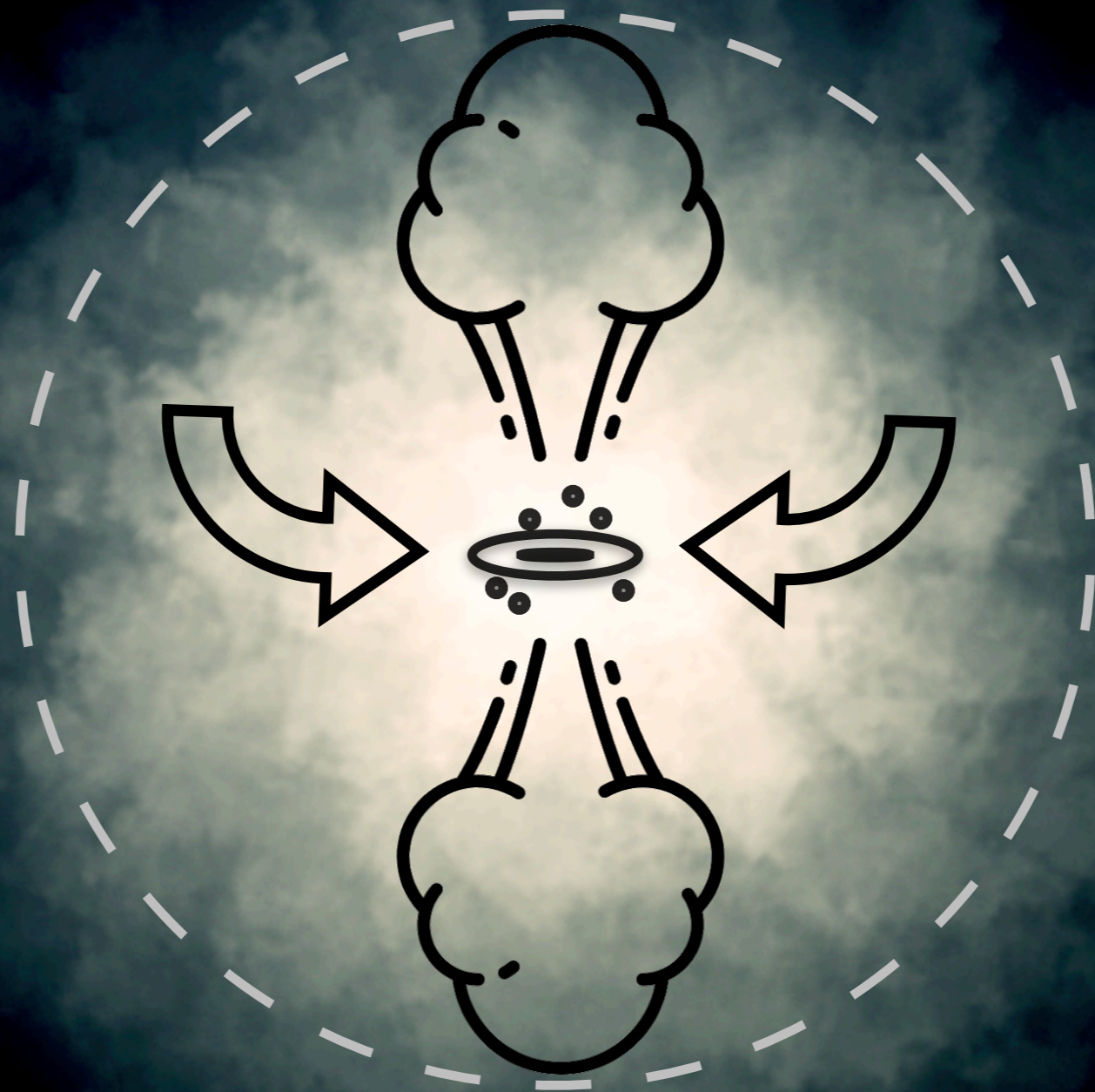
Drummond **Fielding**, Chris **Hayward**,  
Cameron **Hummels**, Tom **Quinn**, Jess **Werk**



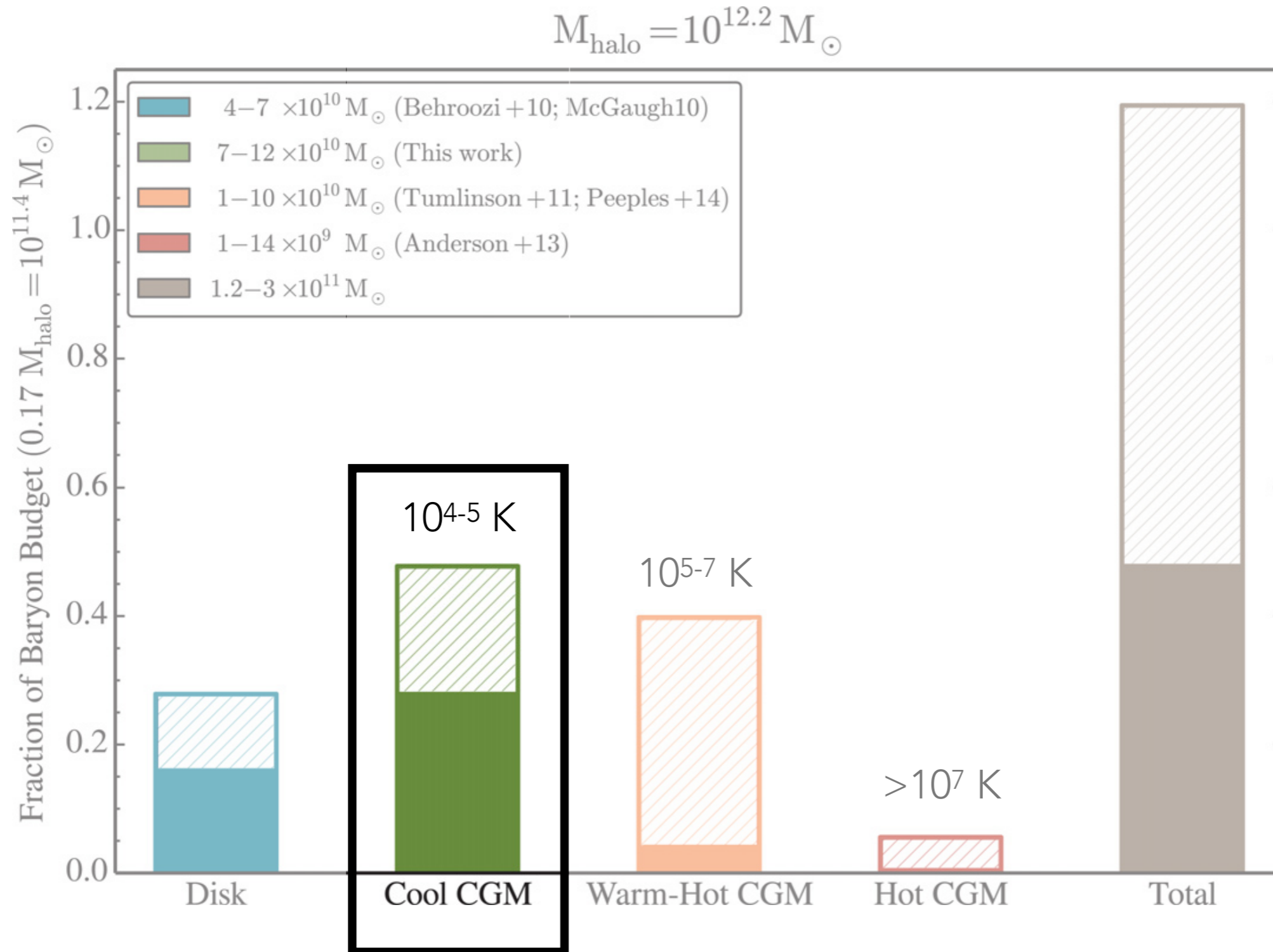
# The Circumgalactic Medium (CGM)



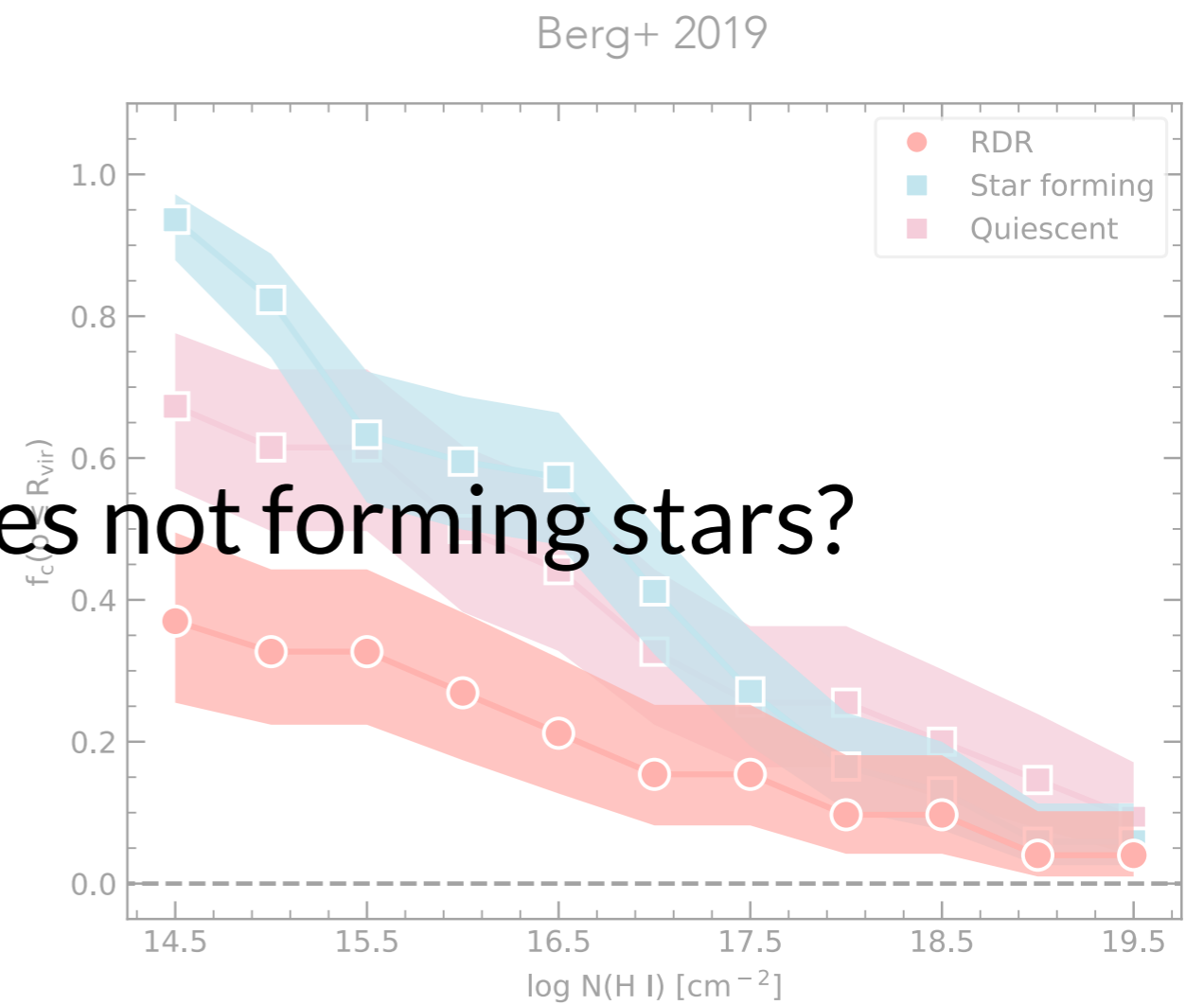
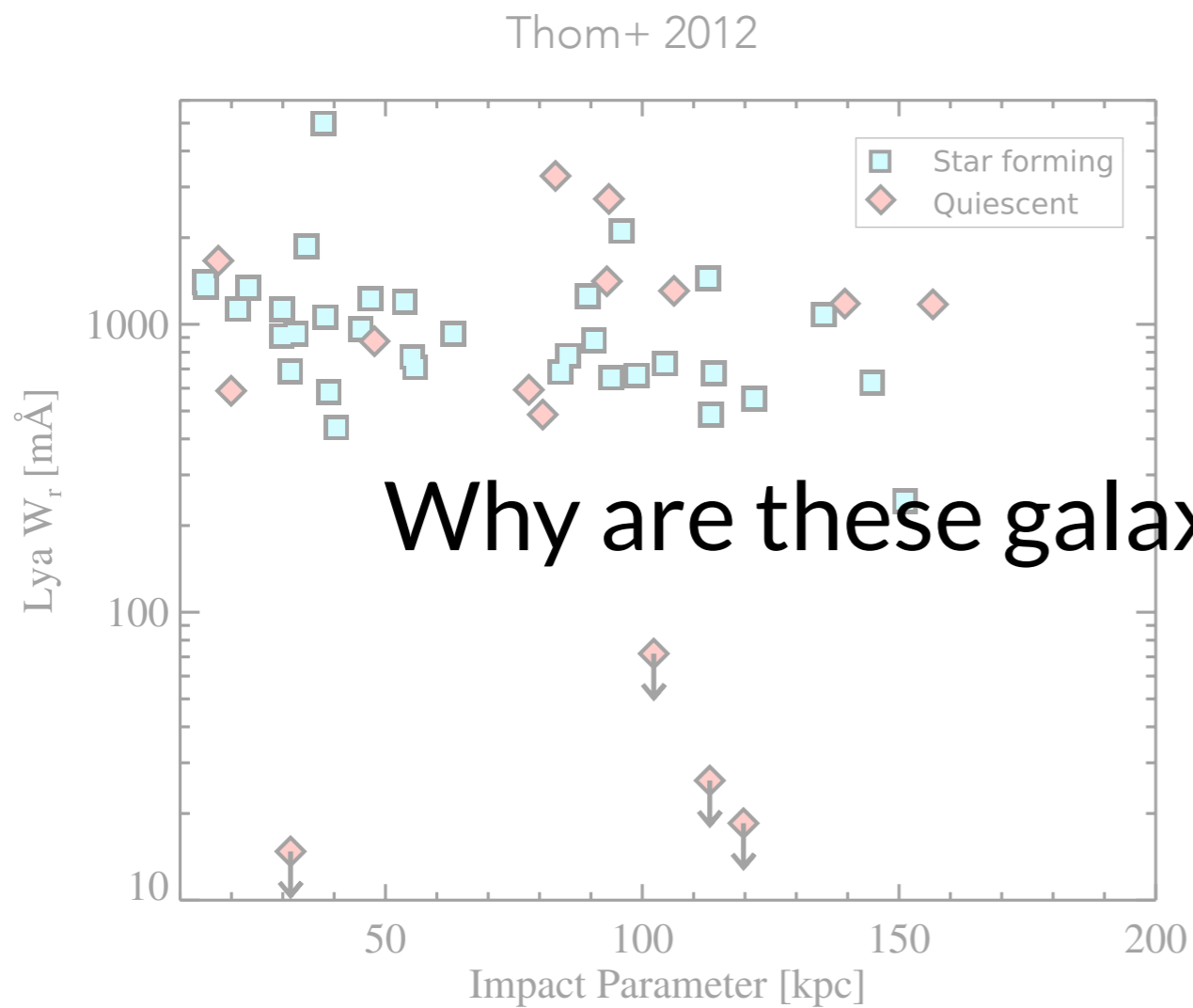
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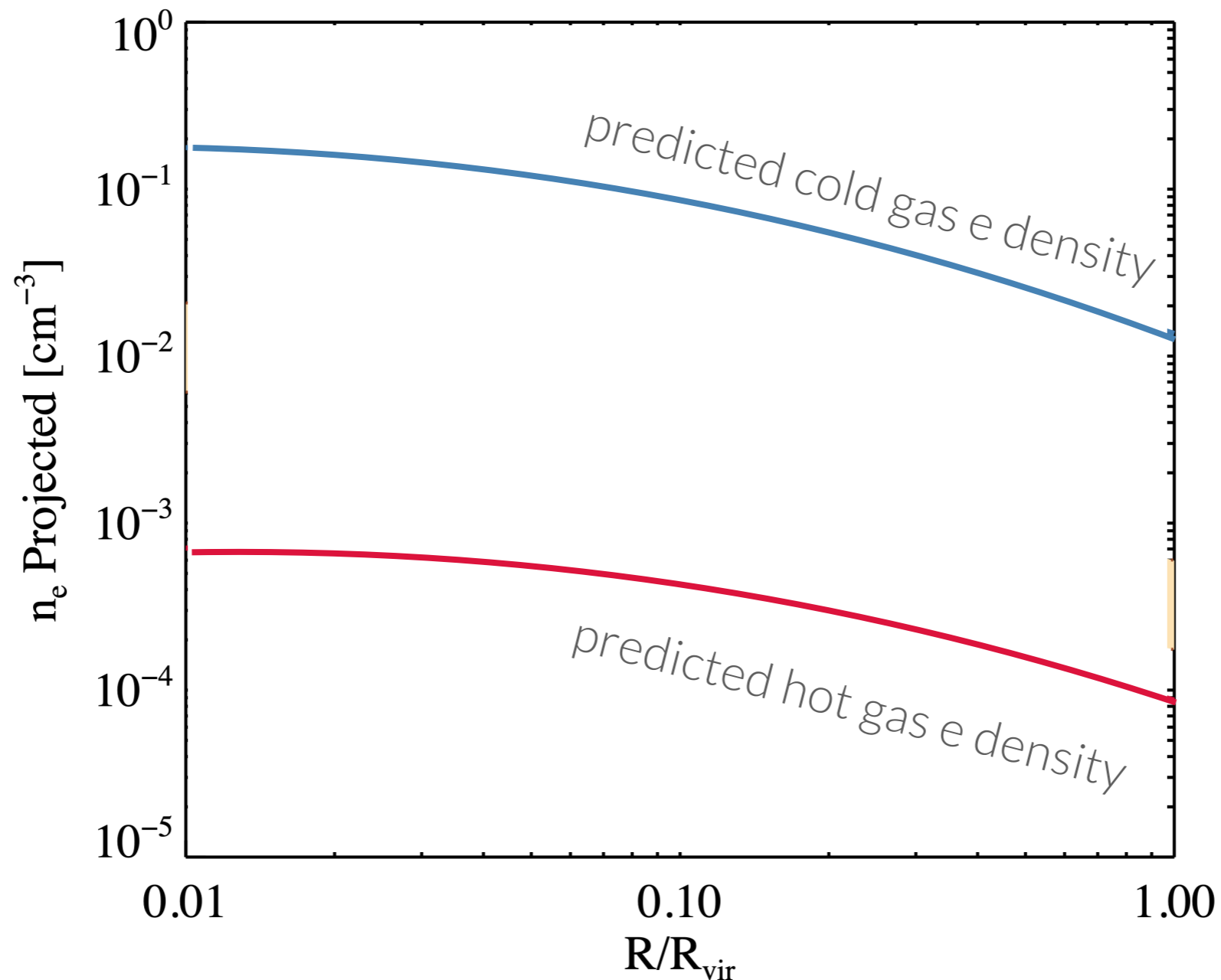
# COOL CGM GAS COMPRISES 25-45% OF GALACTIC BARYONS



# Quiescent galaxies have **massive reservoirs of cold gas** in their CGM



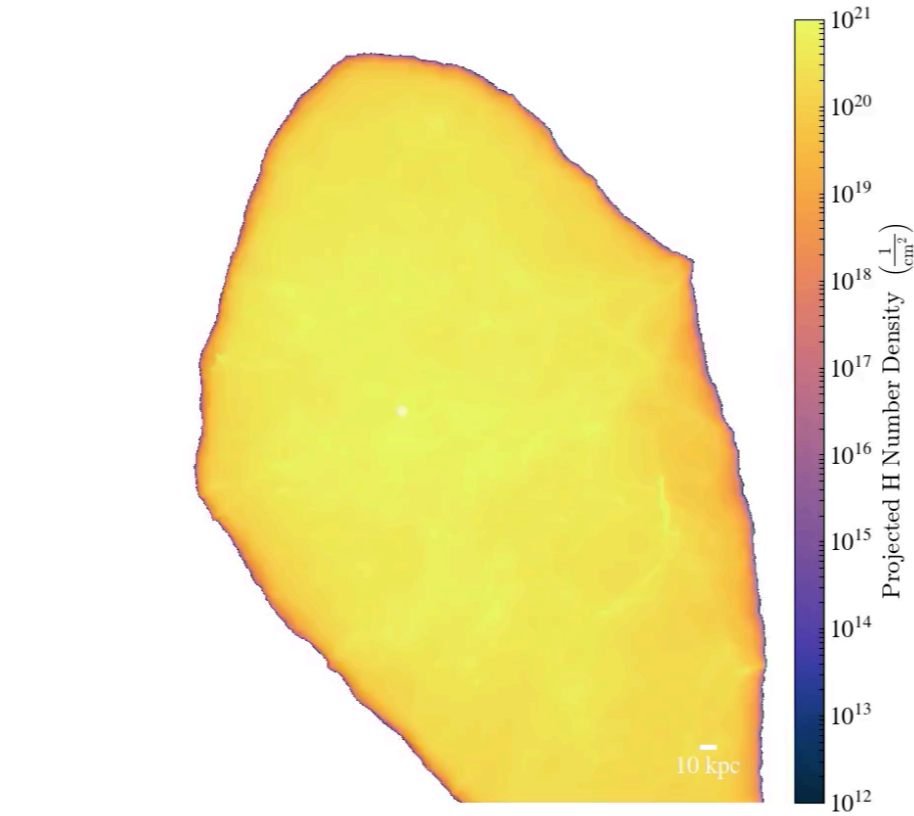
Cold CGM gas appears to be **out of** thermal pressure equilibrium



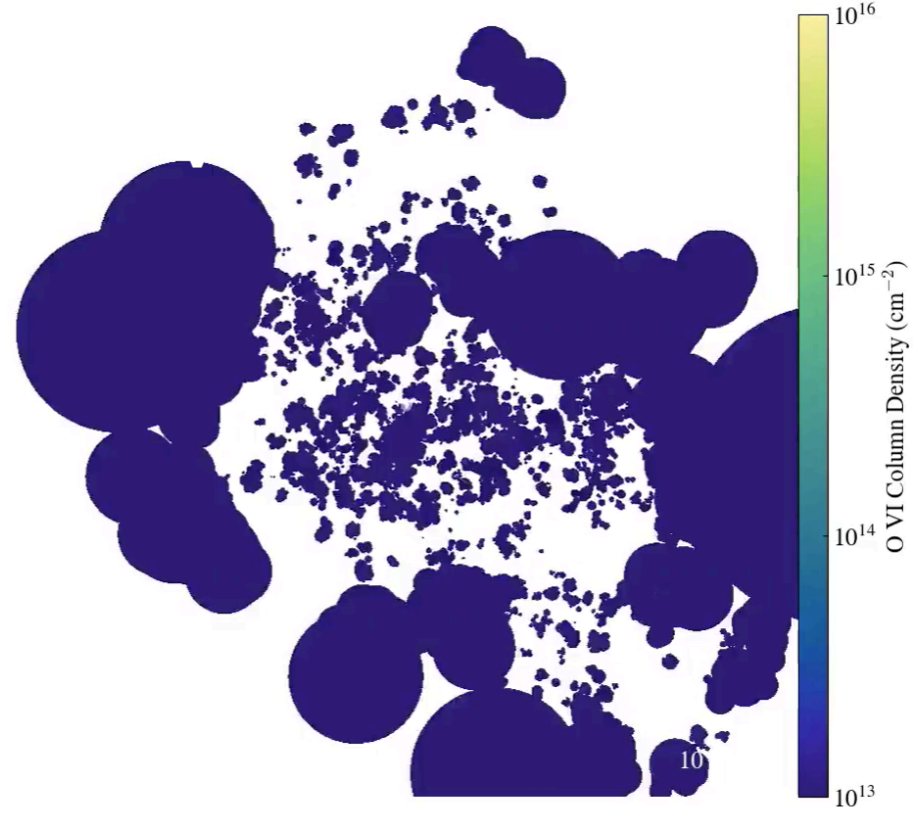
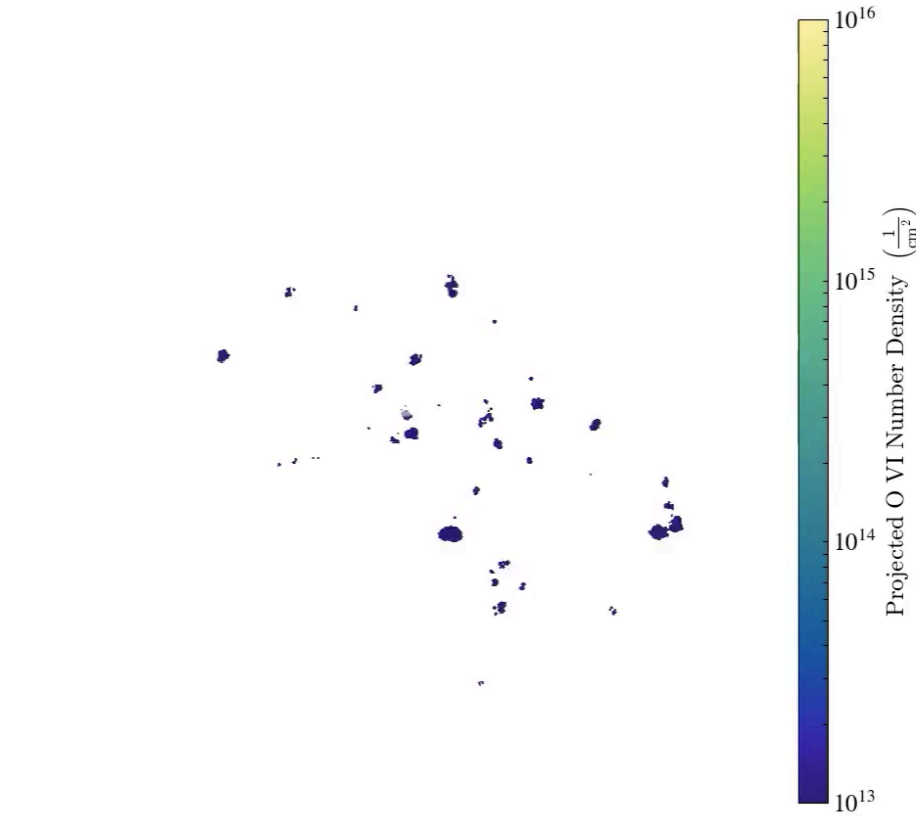
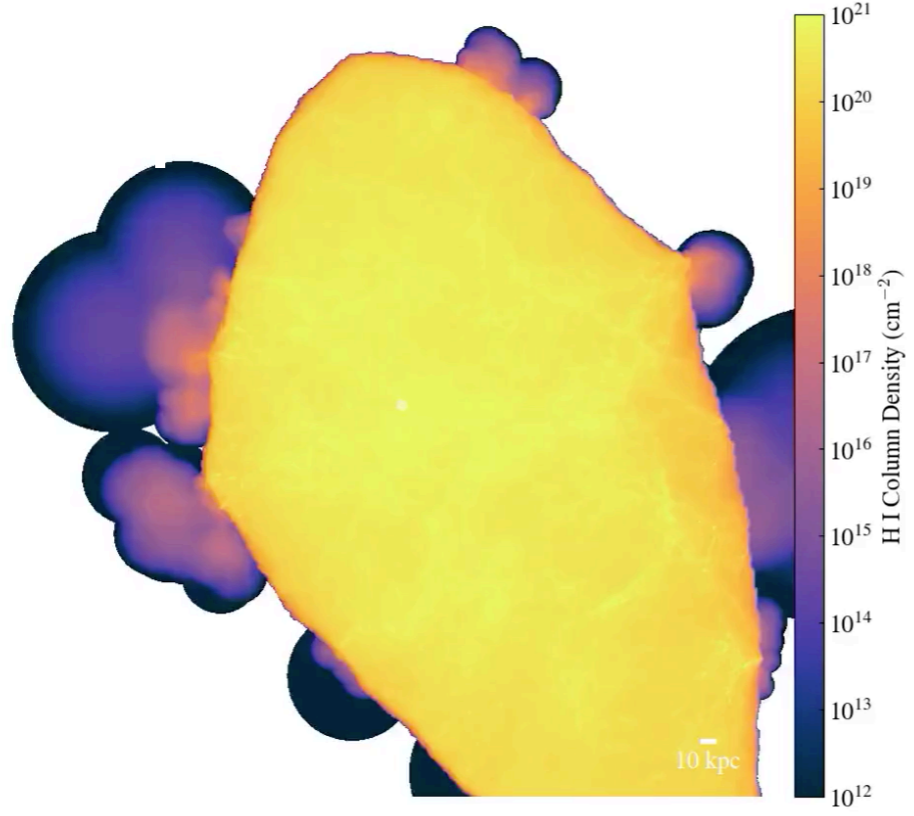
Is cold gas really out of **pressure** equilibrium?

# Cosmic rays **alter ionization state** in the CGM

MHD - only

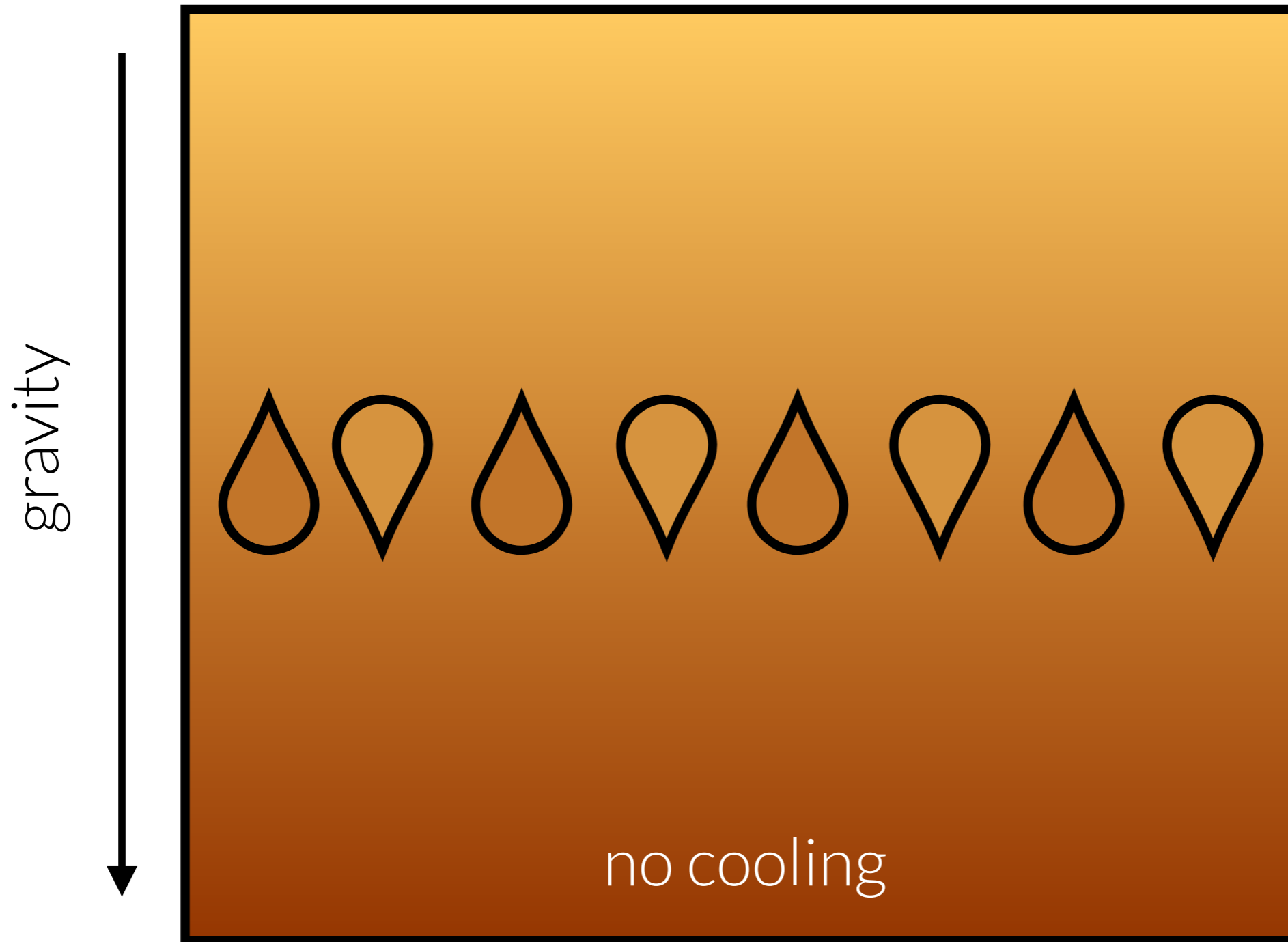


MHD + cosmic rays

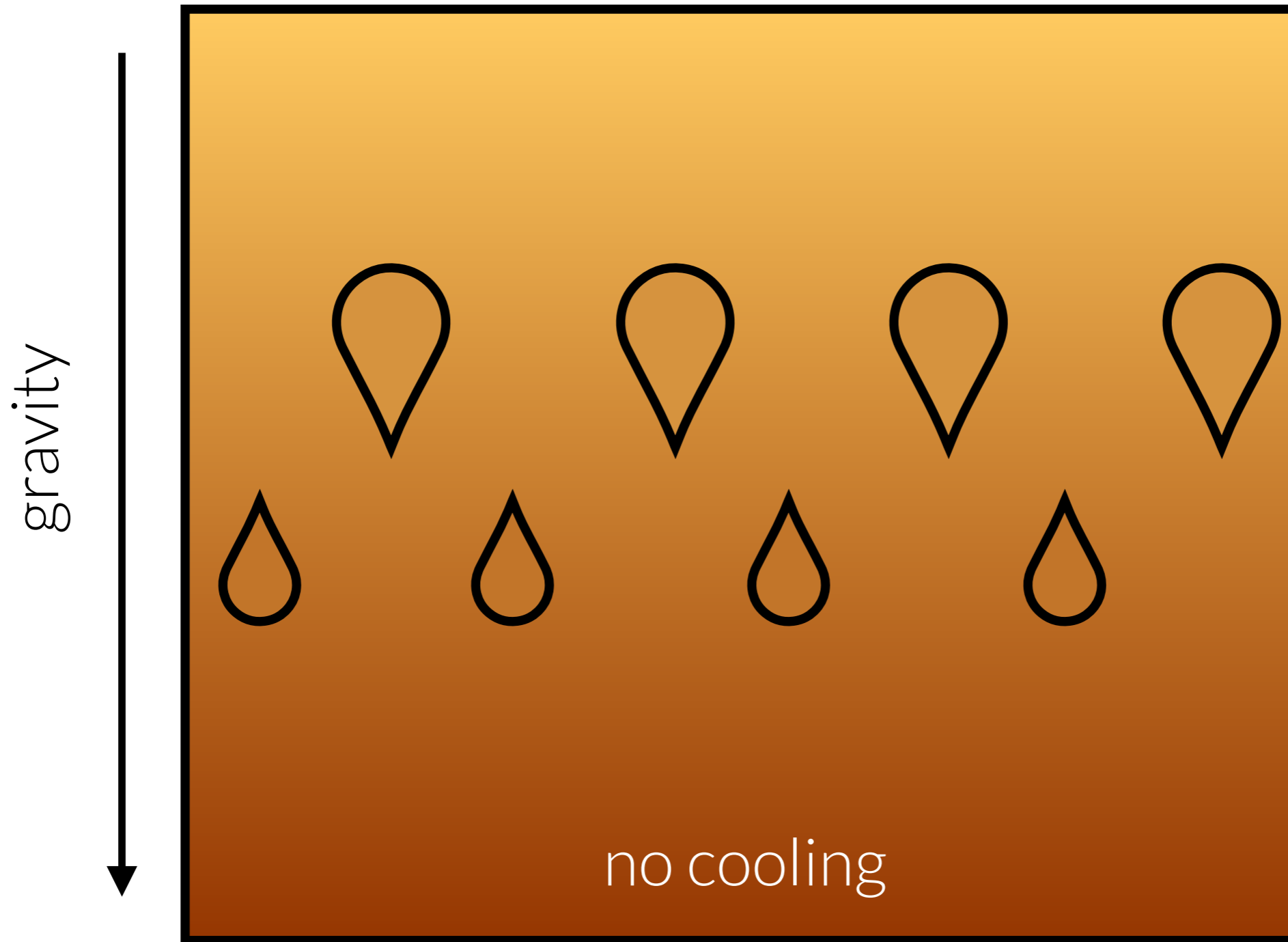




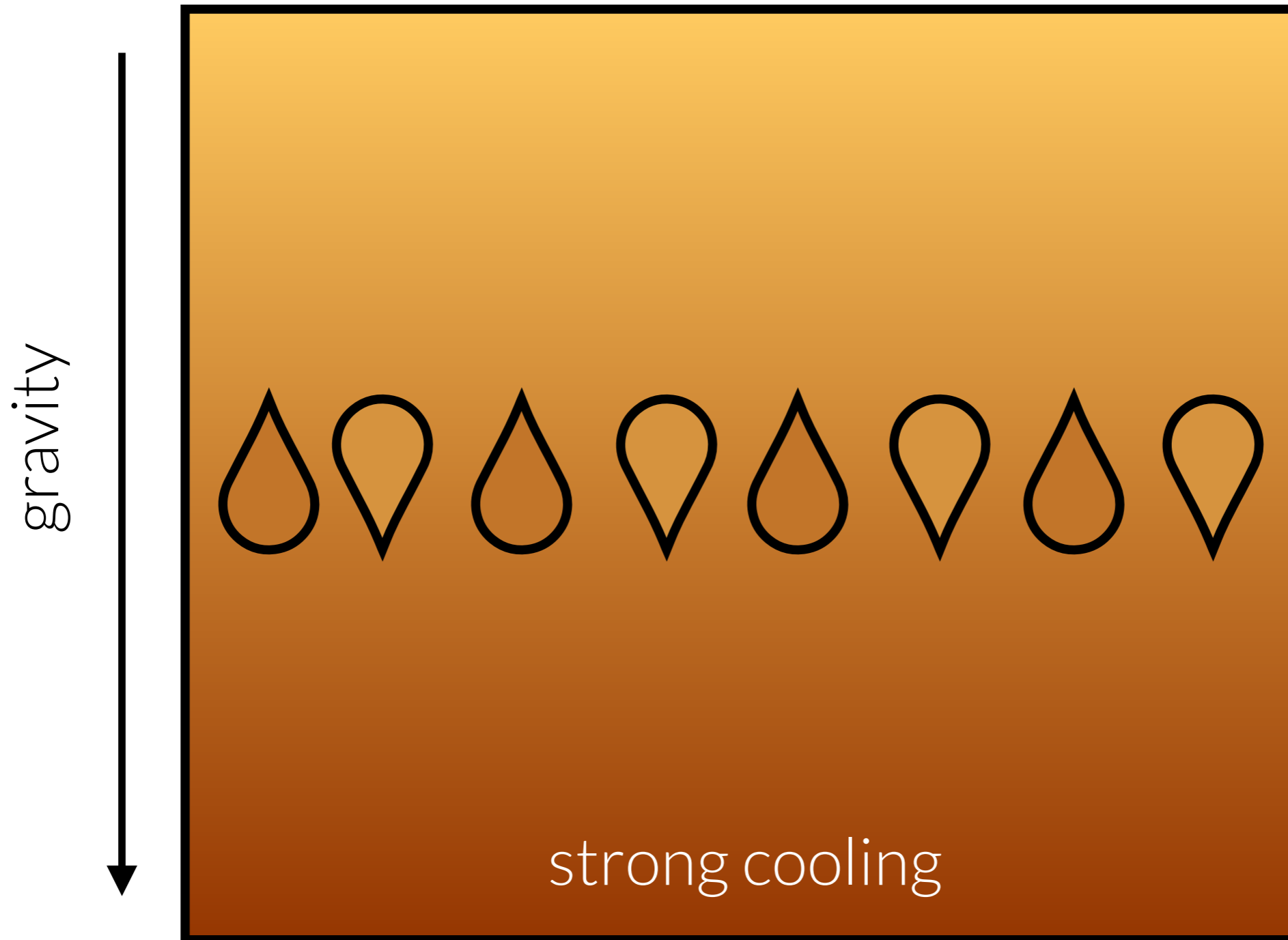
# Thermal instability in a stratified plasma



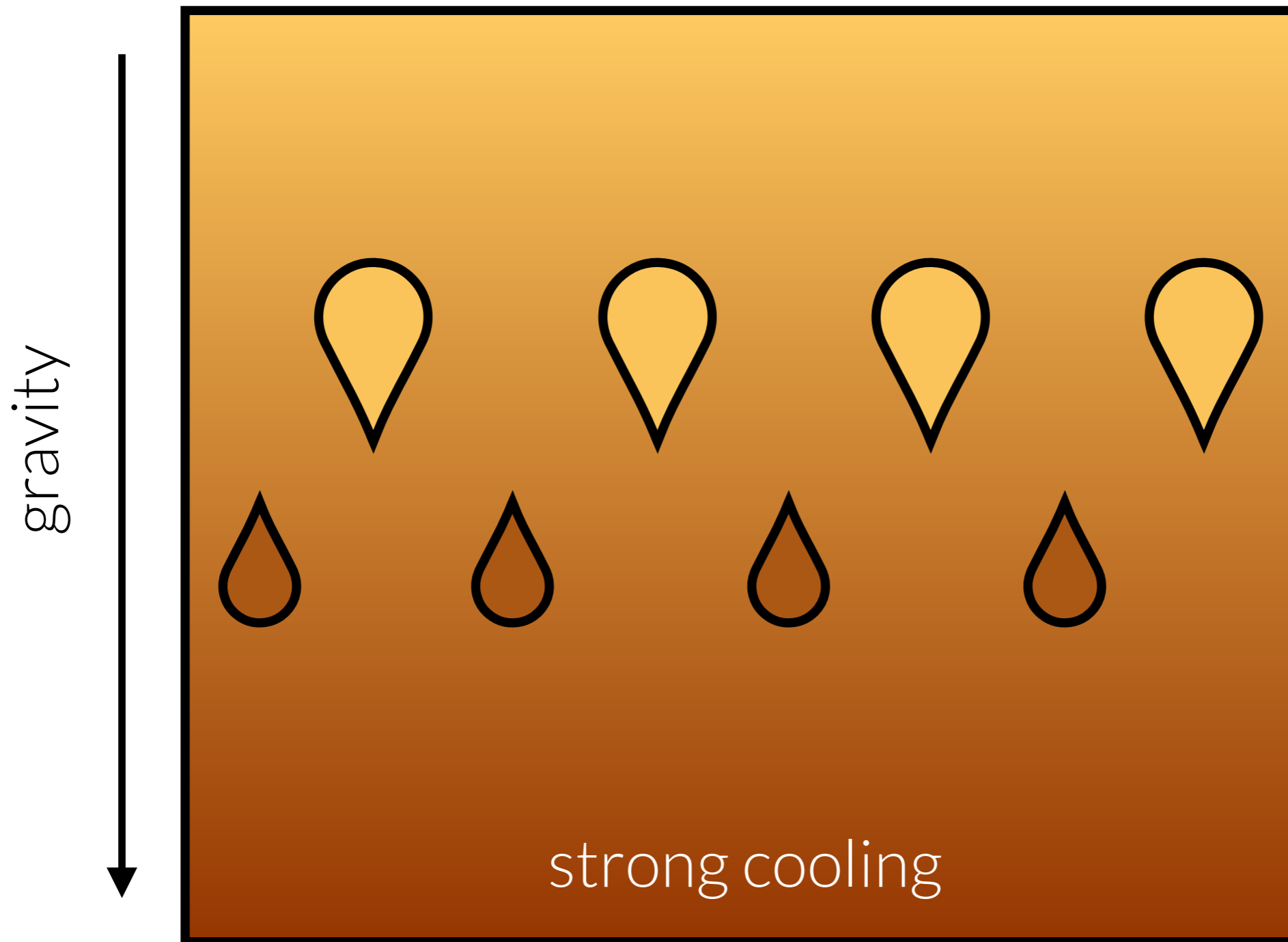
# Thermal instability in a stratified plasma



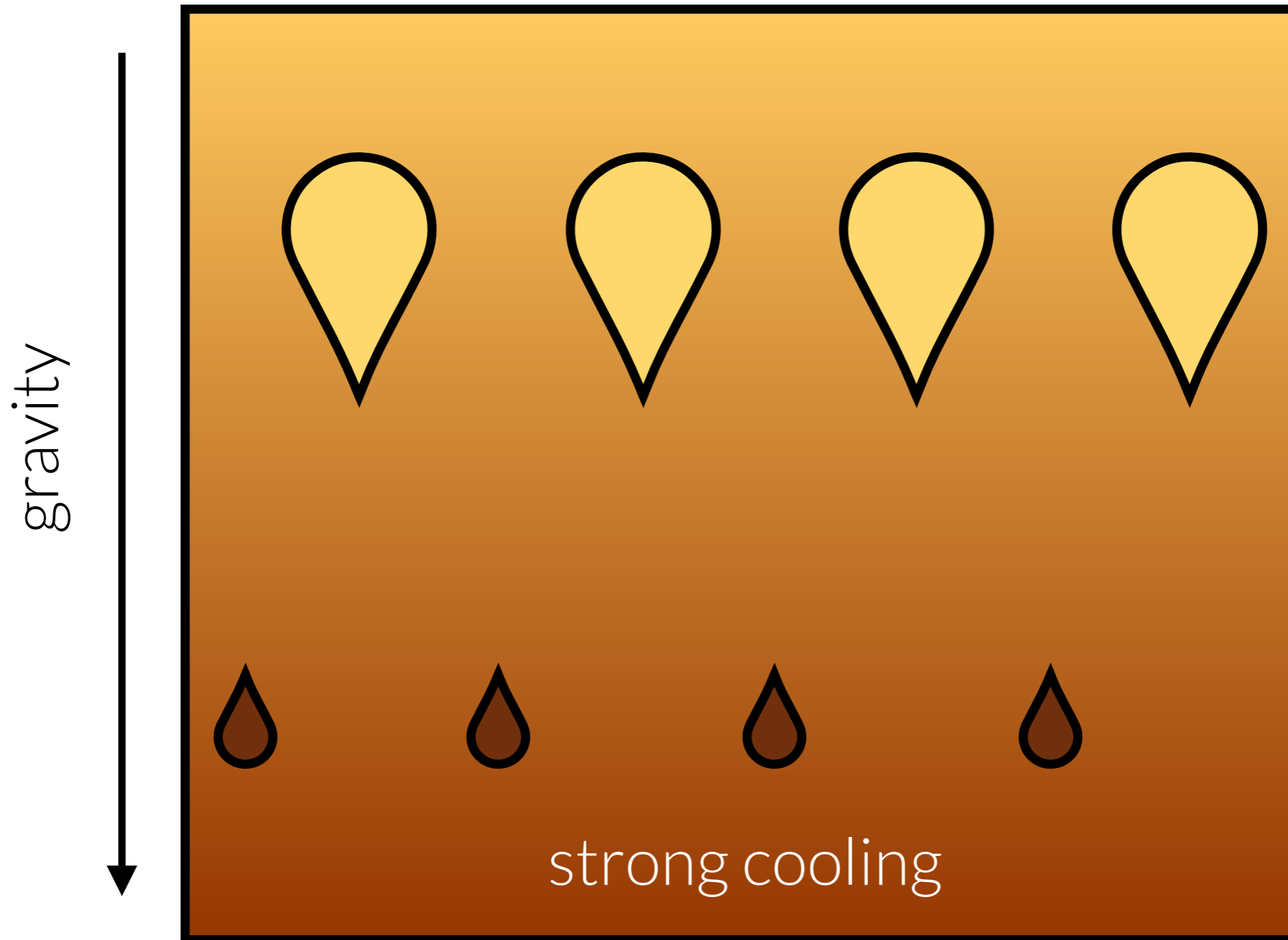
# Thermal instability in a stratified plasma



# Thermal instability in a stratified plasma



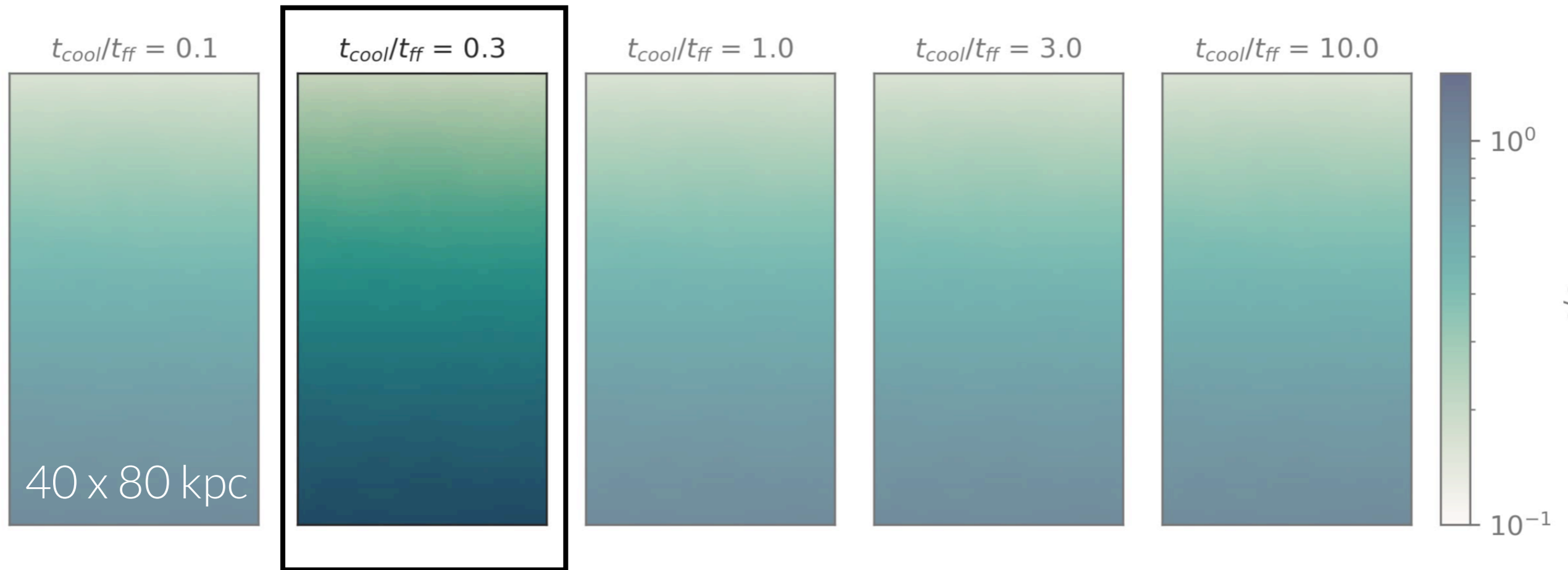
# Thermal instability in a stratified plasma



# Thermal instability **without** cosmic rays

fast cooling

slow cooling



# Cosmic Rays and Thermal Instability

pressure

transport



# Cosmic Rays and Thermal Instability

pressure

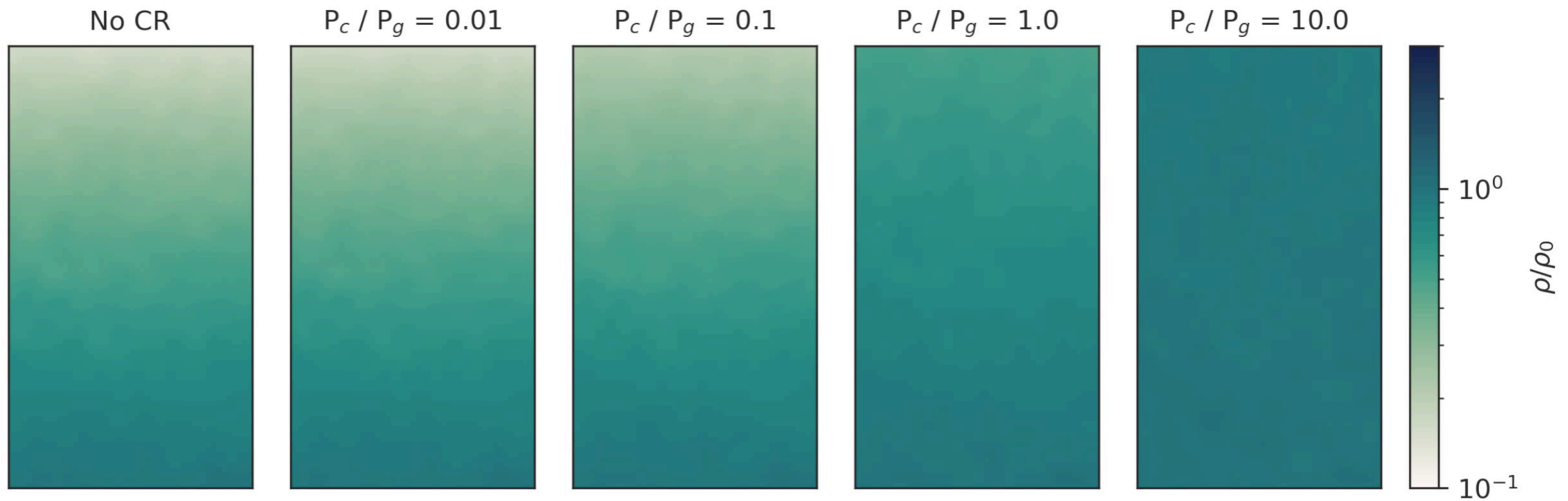


transport

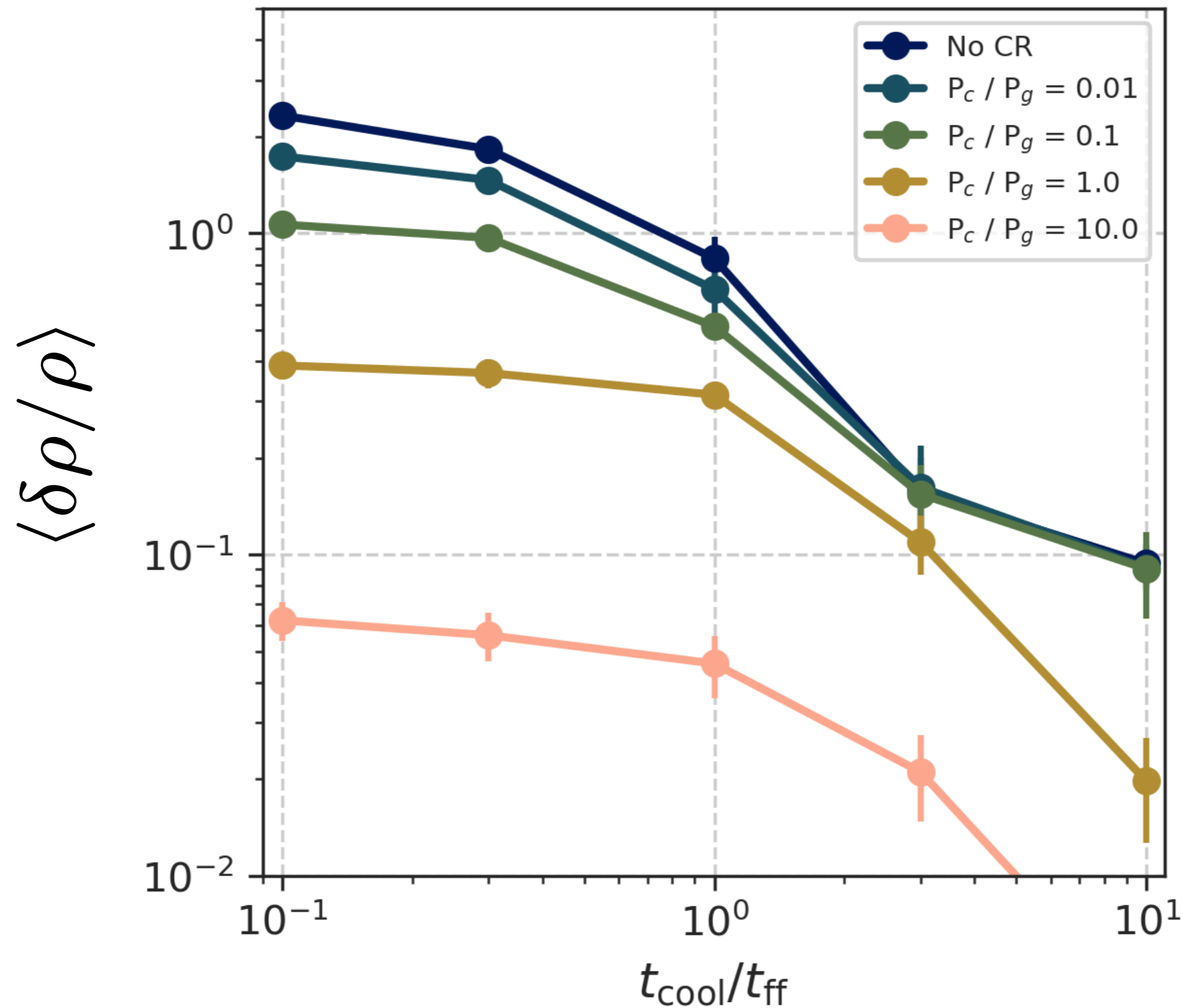




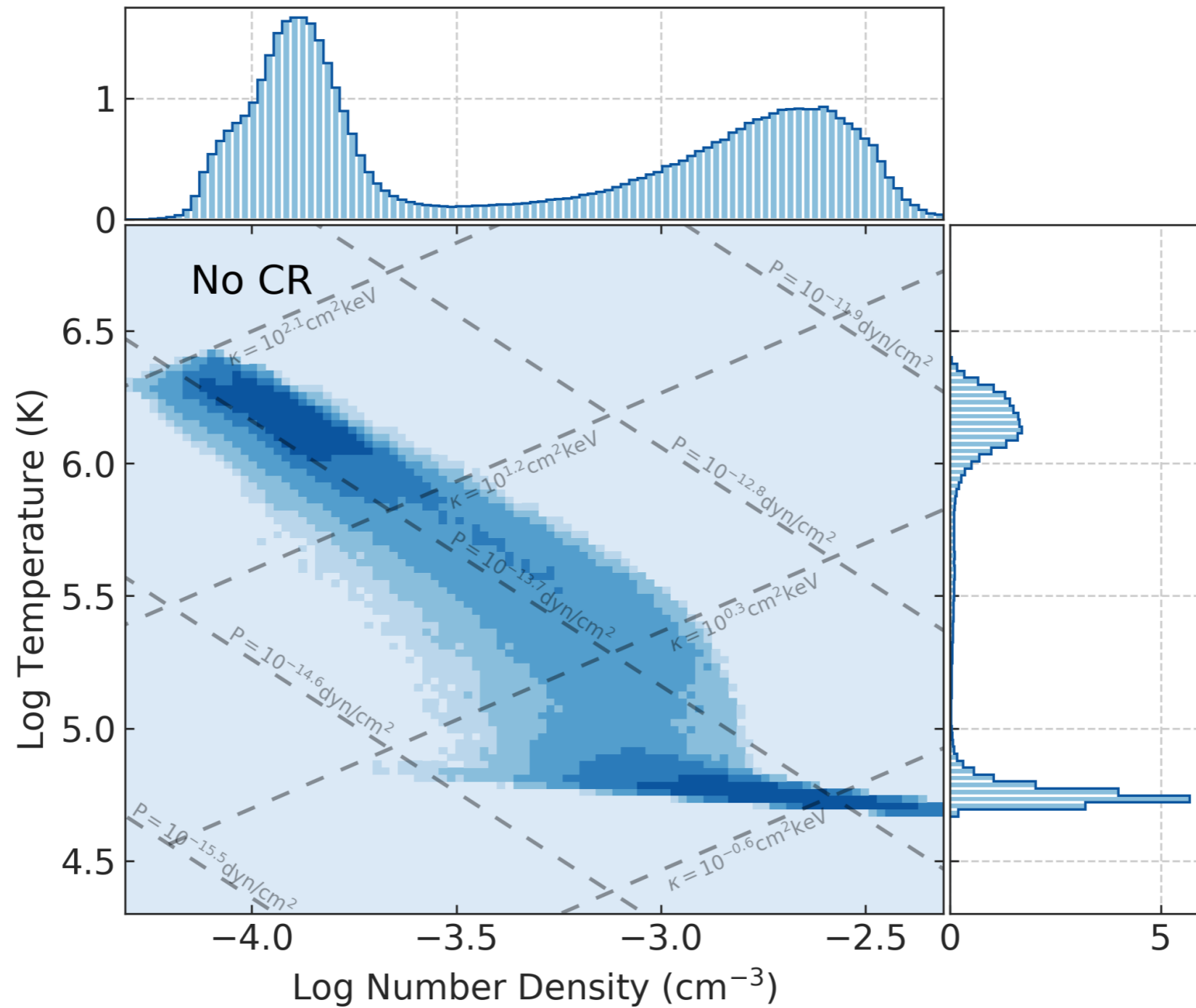
# Cosmic ray pressure decreases **cold gas density**



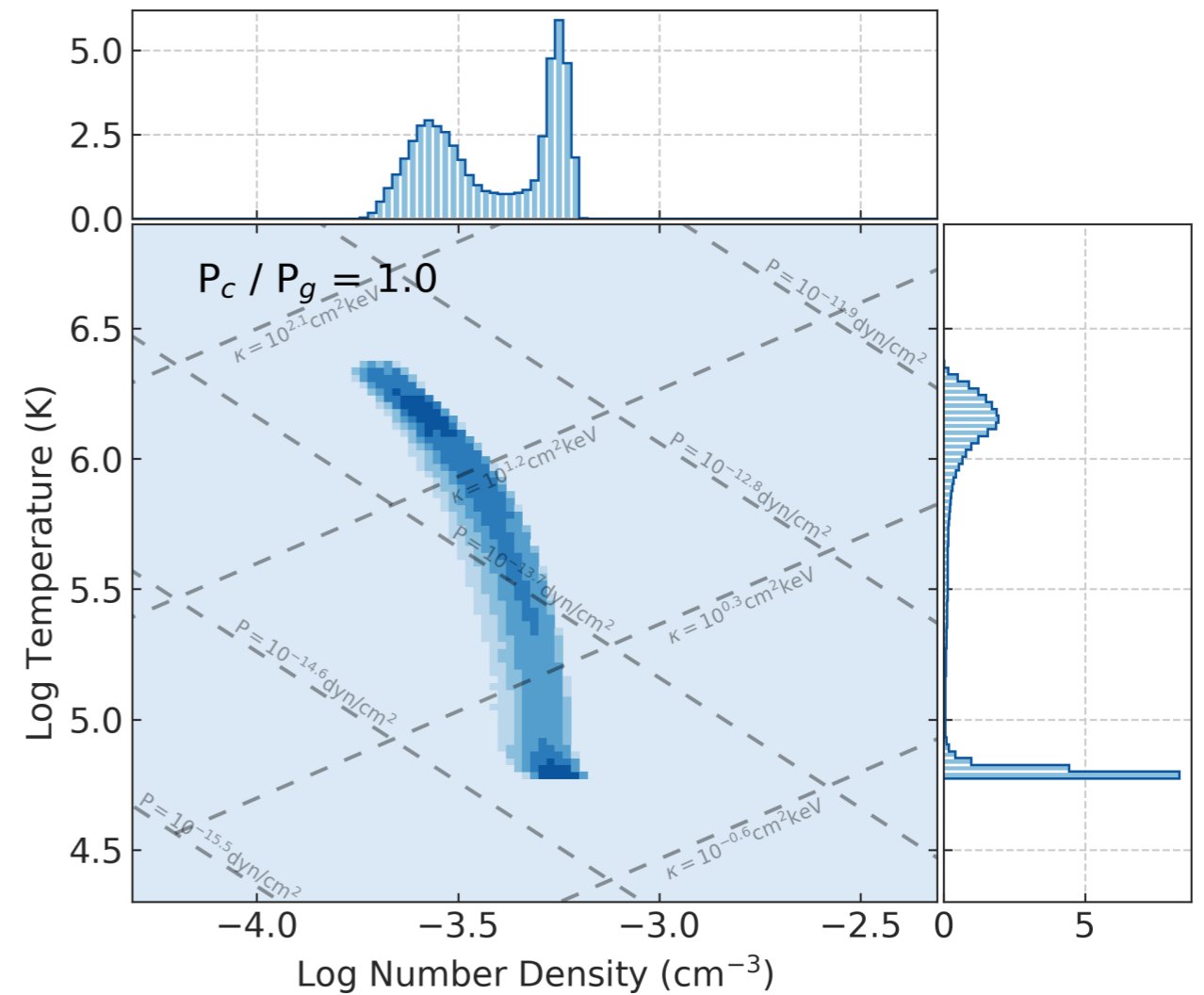
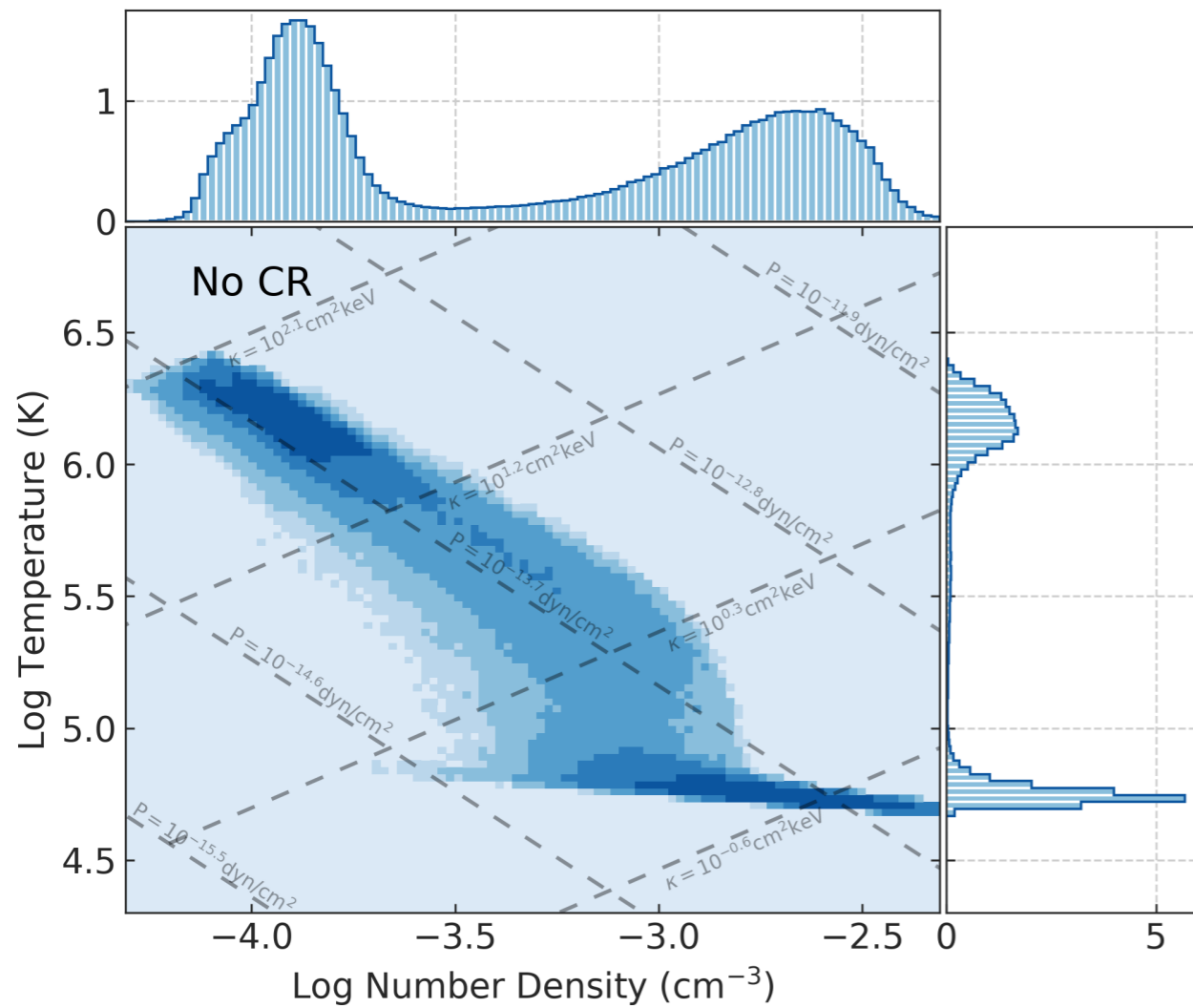
# Cosmic ray pressure washes out **density fluctuation**



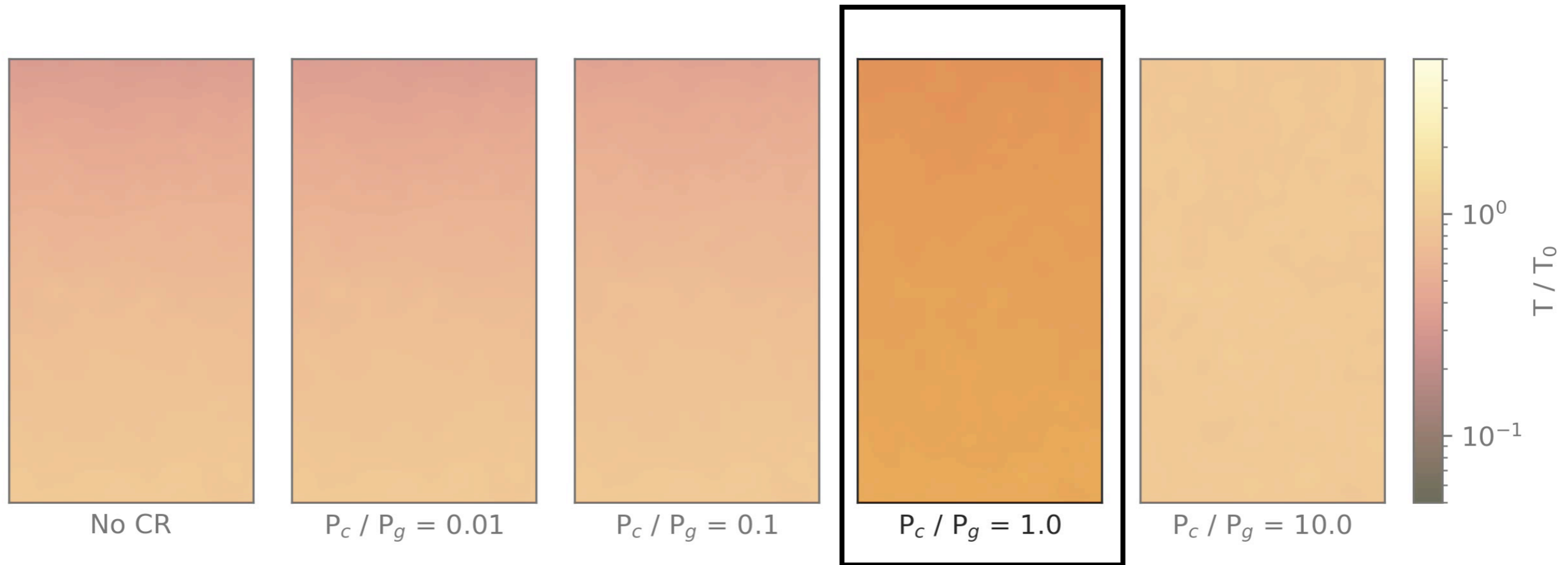
Without cosmic rays, gas cools in thermal pressure equilibrium (**isobarically**)



With high cosmic ray pressure, gas cools at constant density (**isochorically**)

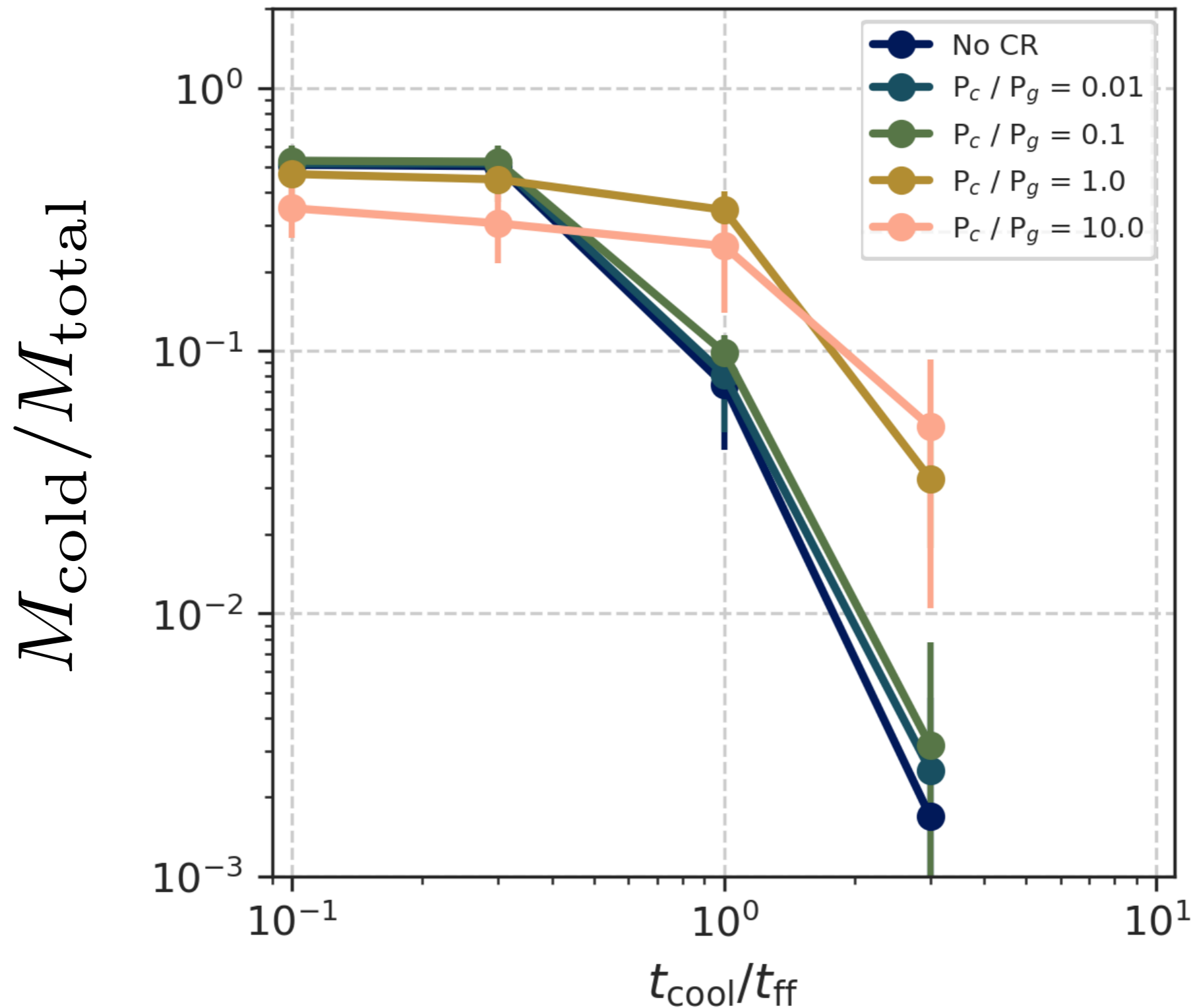


# Cold gas **temperature is independent** of cosmic ray pressure

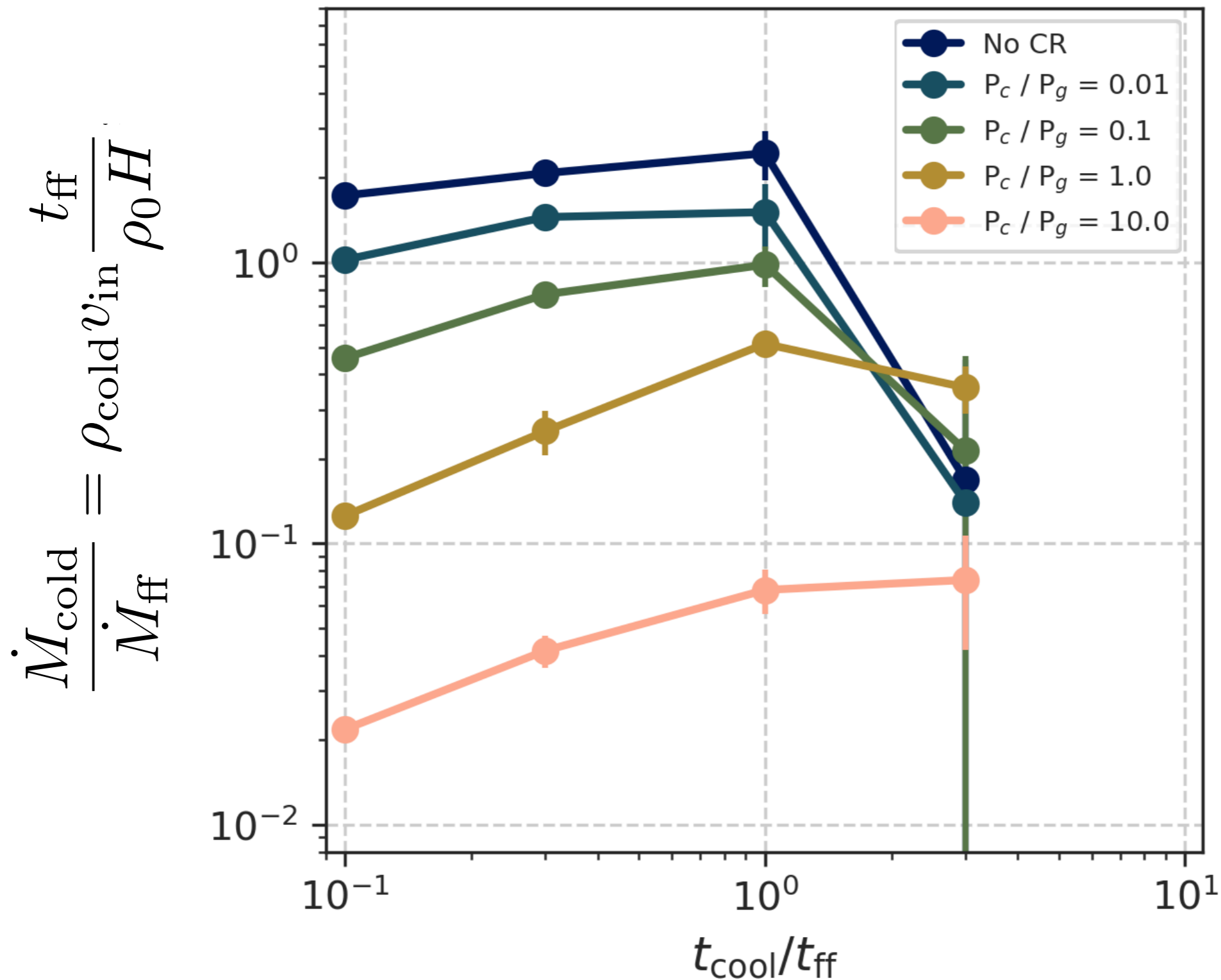


Cold gas temperature is set by the shape of the **cooling curve**, which is an approximation to atomic physics and insensitive of cosmic ray pressure

# Cosmic ray pressure can **increase cold mass fraction**

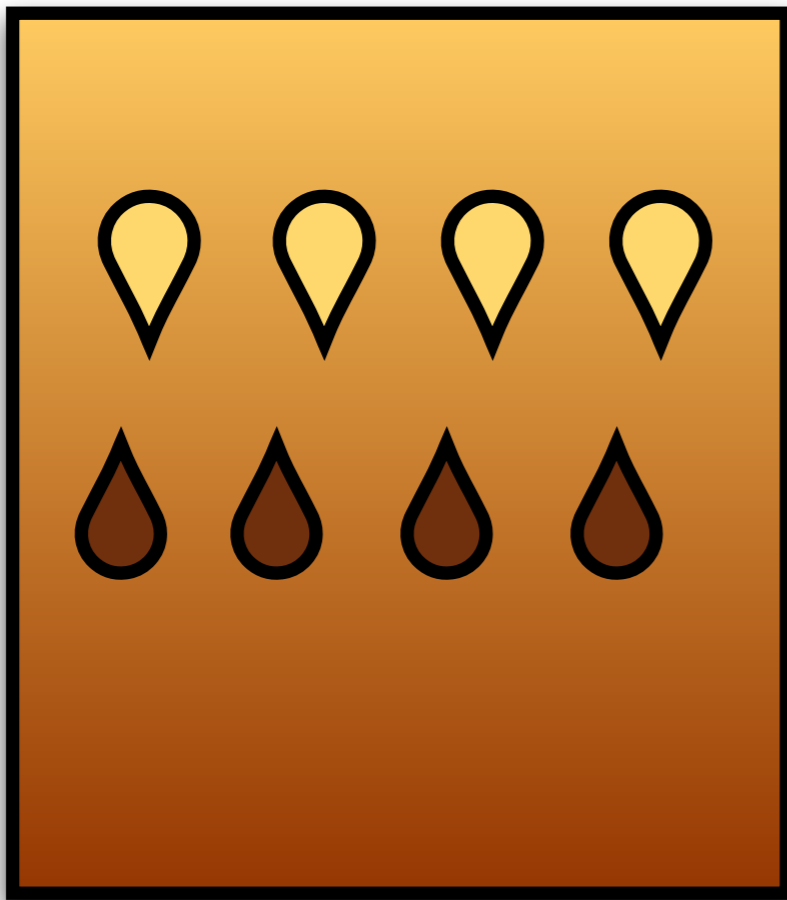


# Cosmic ray pressure **decreases cold mass flux**



# Cosmic Rays and Thermal Instability

pressure



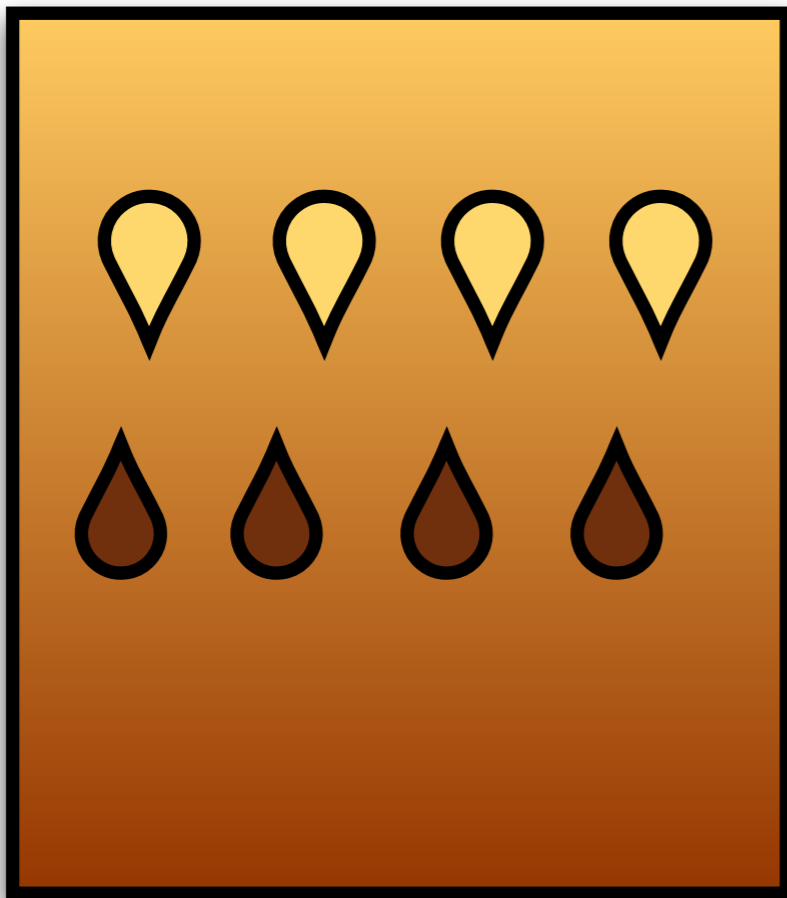
transport





# Cosmic Rays and Thermal Instability

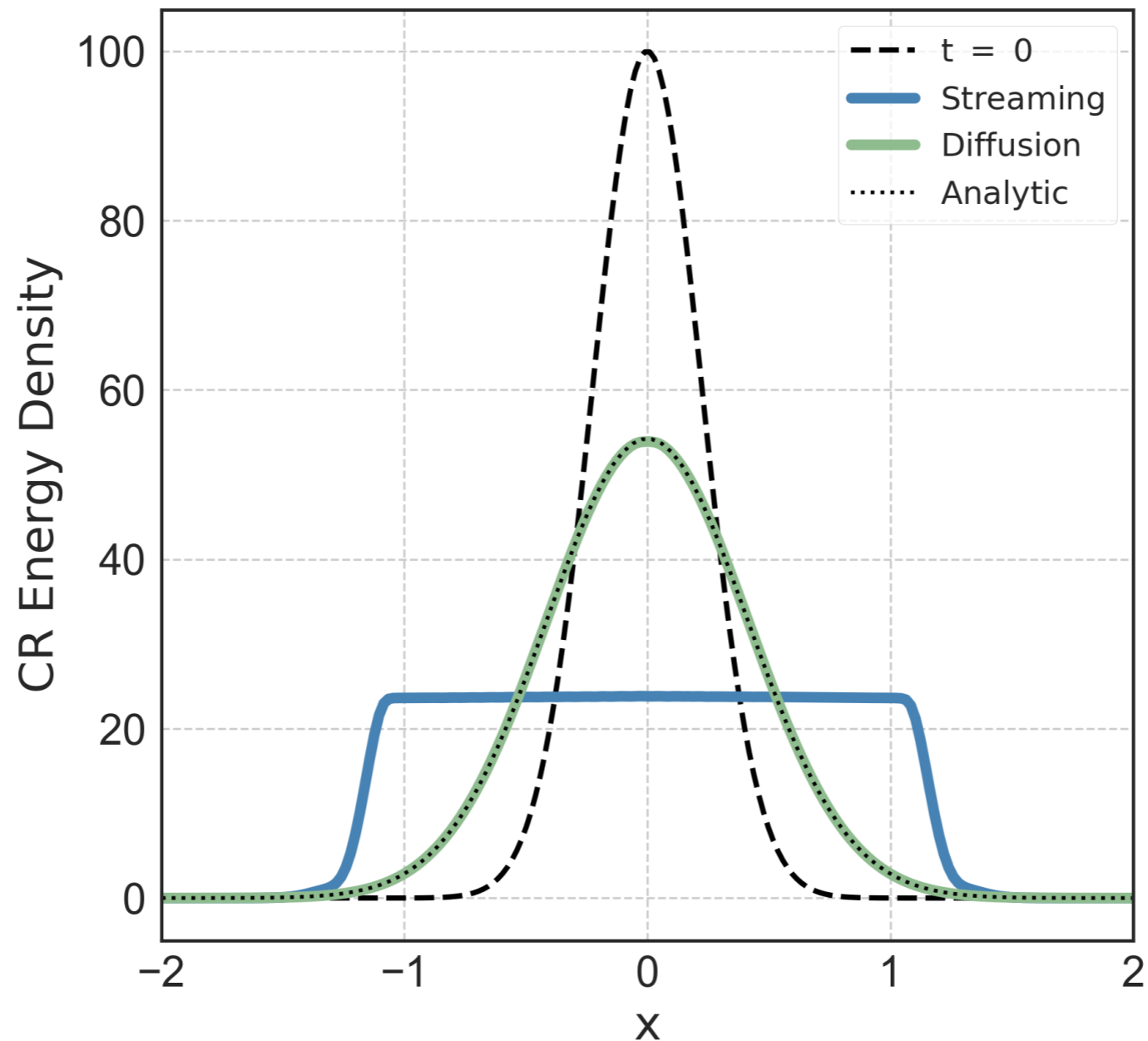
pressure



transport

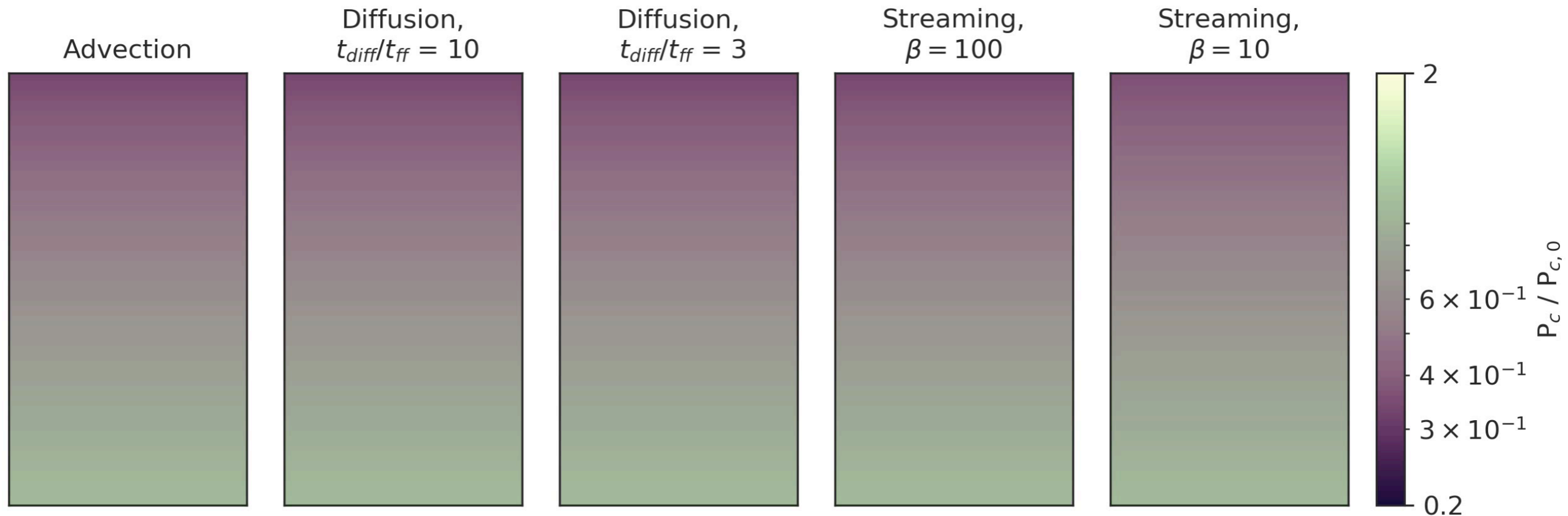


# Cosmic ray transport: **streaming** and **diffusion**



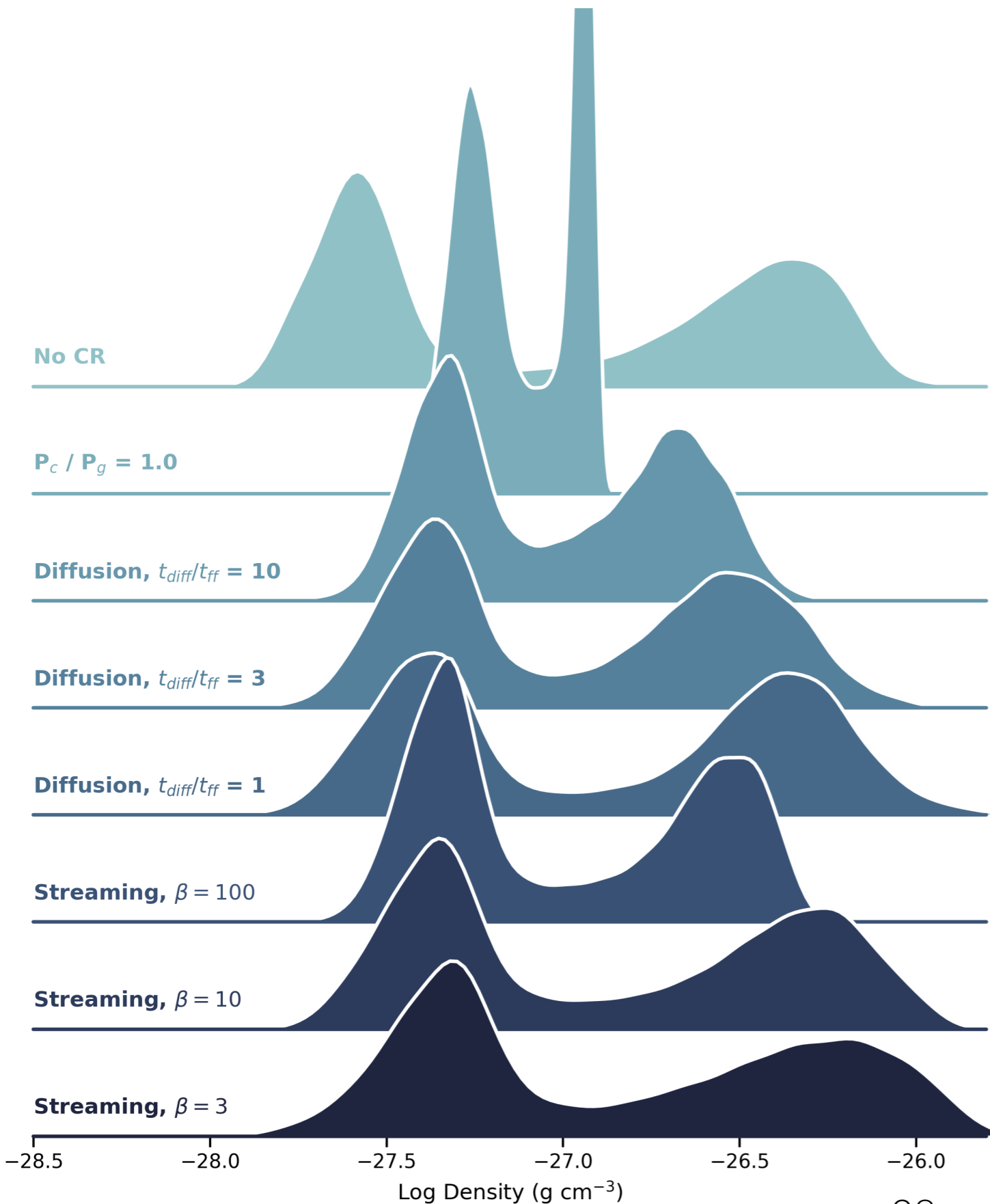
Cosmic rays move **down** their energy gradient, **along** magnetic field lines

# Impact of cosmic ray transport



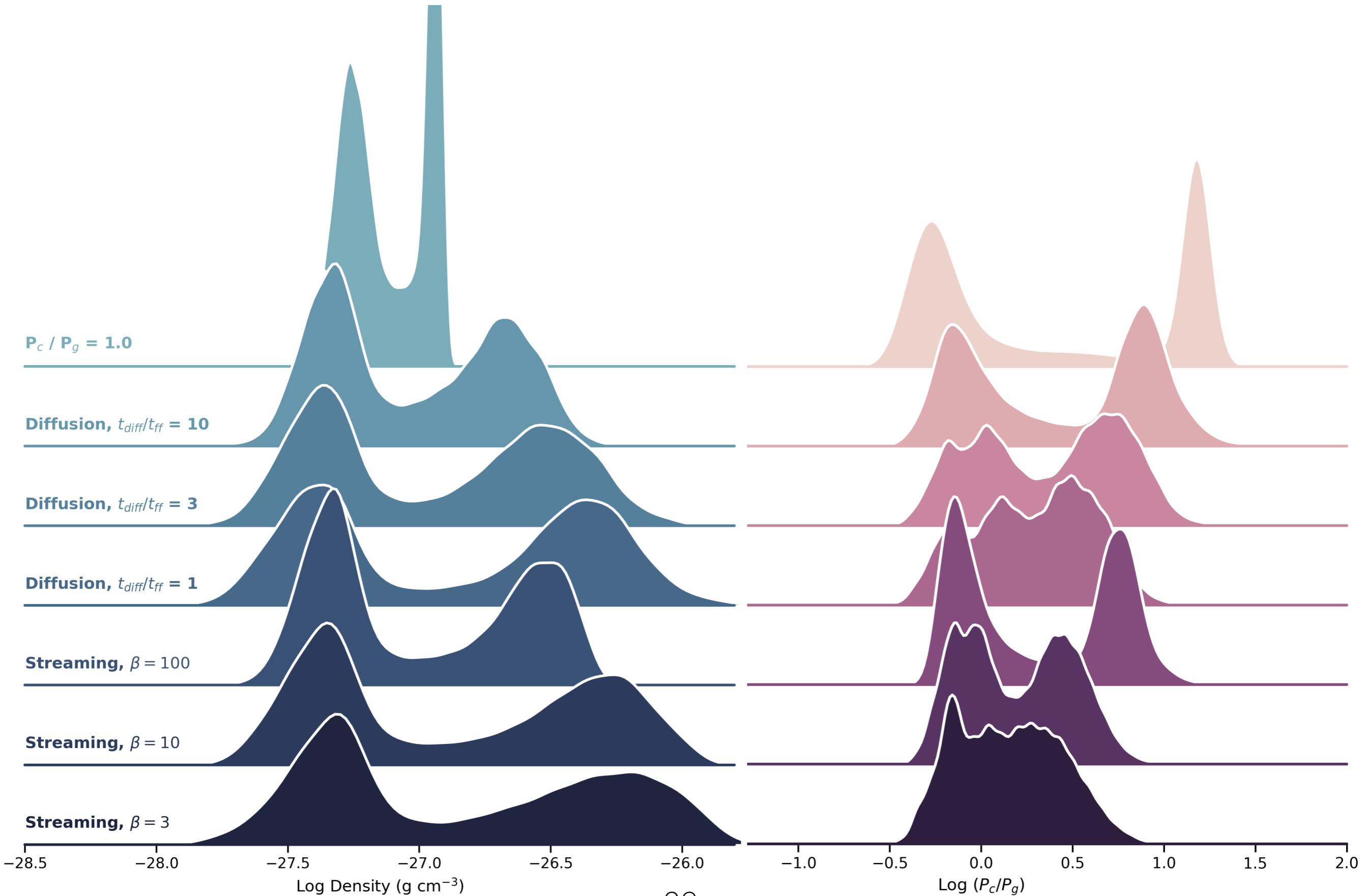
**Cosmic ray transport** redistributes cosmic ray pressure **from** high concentrations (in **cold gas**) **to** low concentrations ( in **hot gas** )

# Impact of cosmic rays on gas **density profiles**

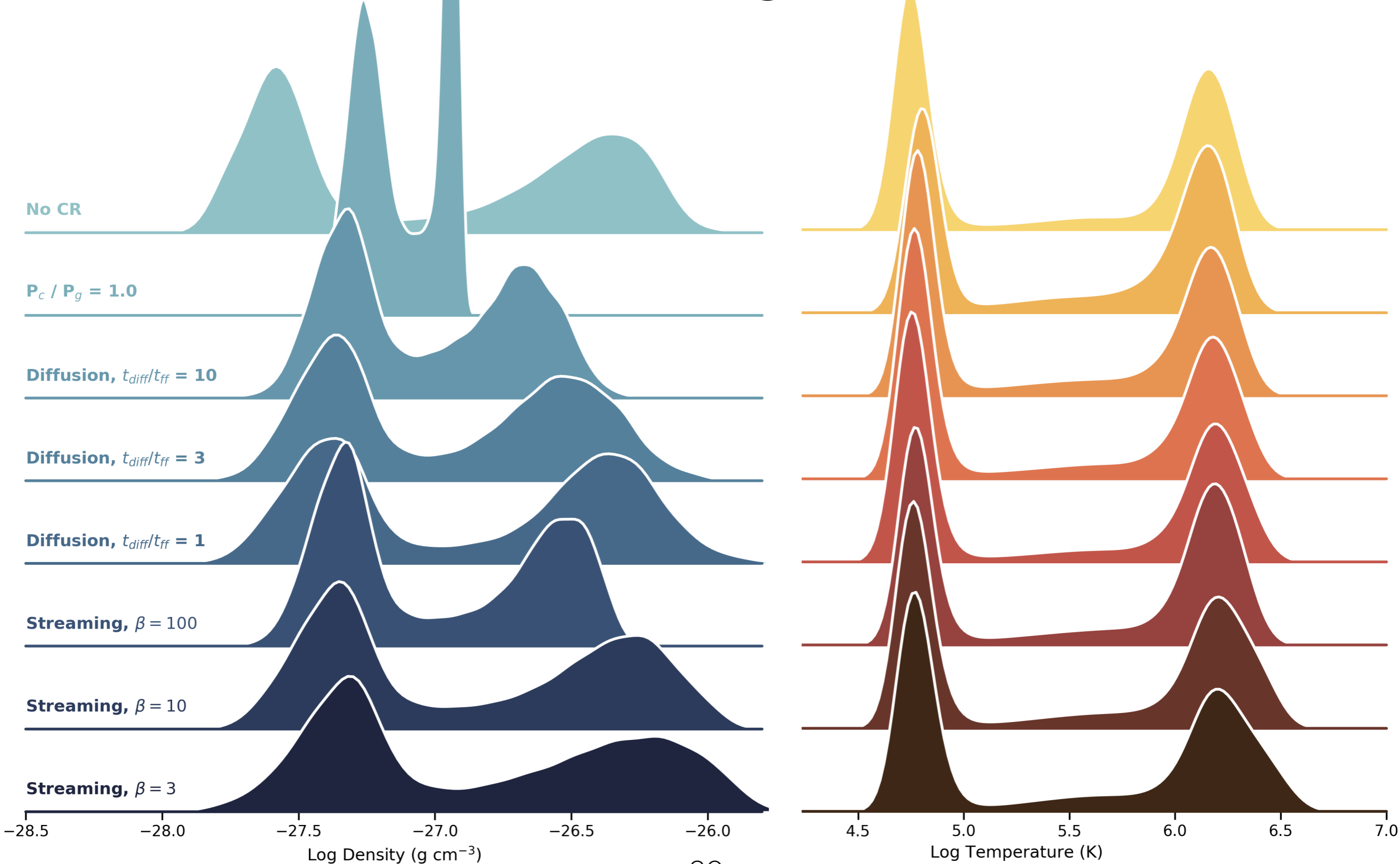


Simulations with cosmic ray transport have density distributions in between those of simulations **without cosmic rays** and simulations with cosmic rays but **without cosmic ray transport**

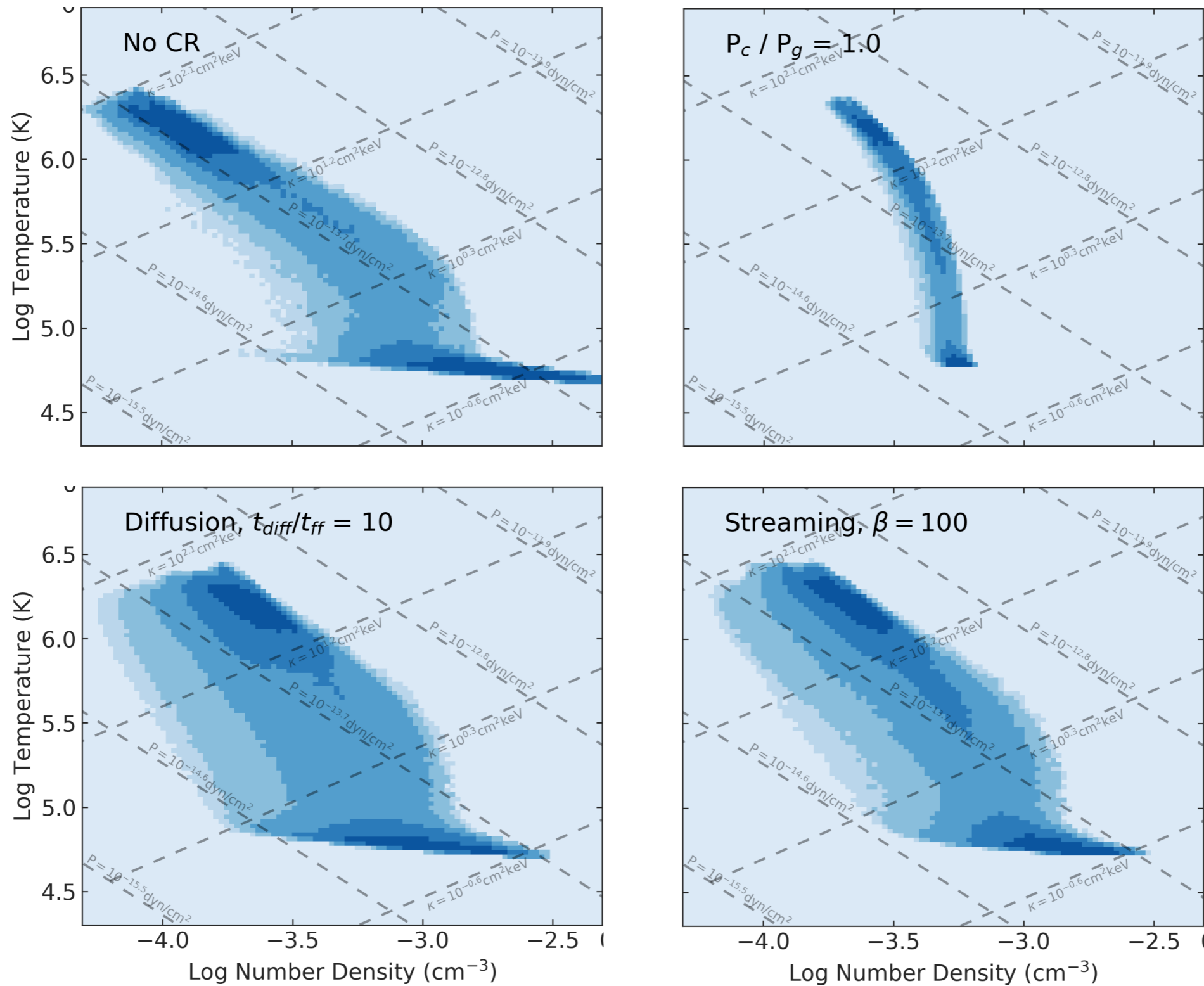
# Gas density pdf depends on **cosmic ray pressure** pdf



# Gas **temperature** distribution remains relatively unchanged

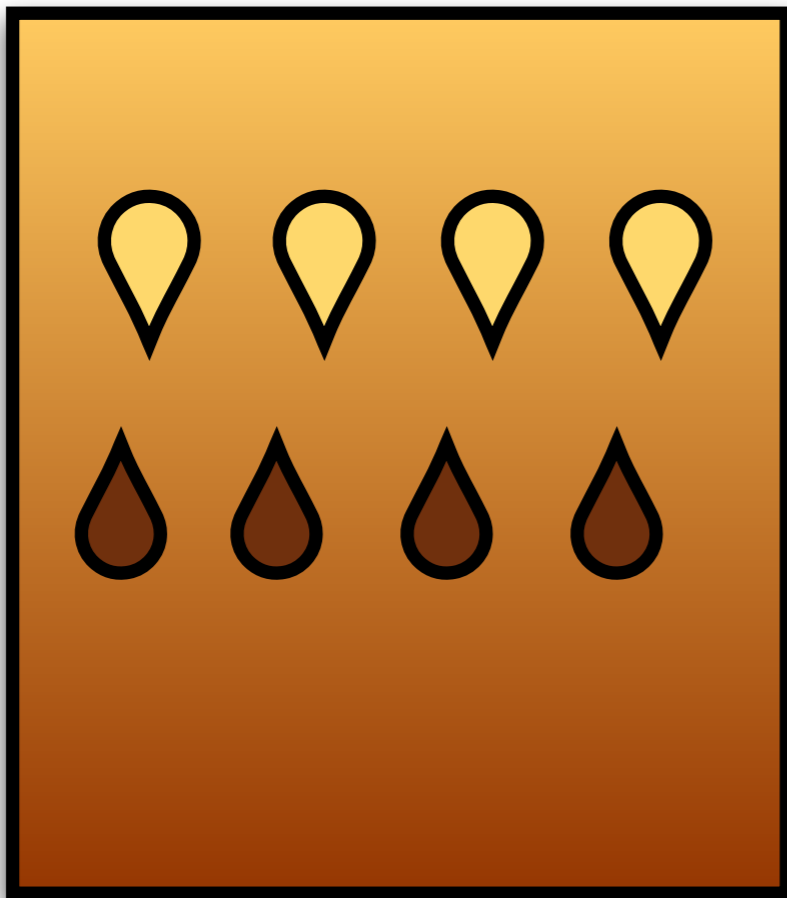


# Cosmic ray transport **fills larger** density-temperature phase space

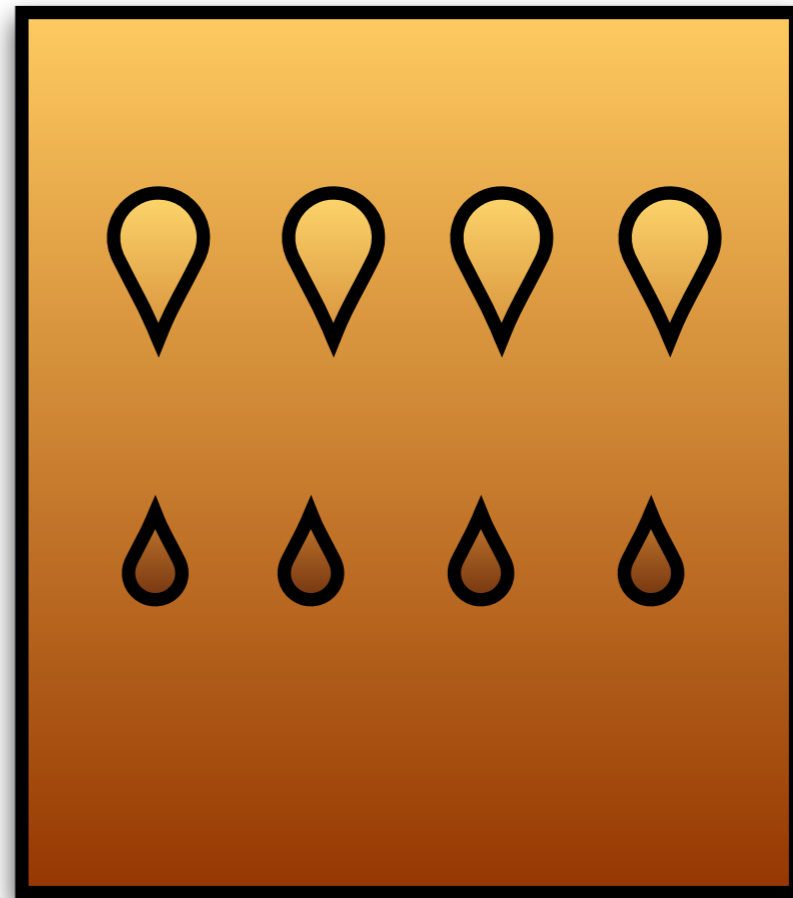


# Cosmic Rays and Thermal Instability

pressure

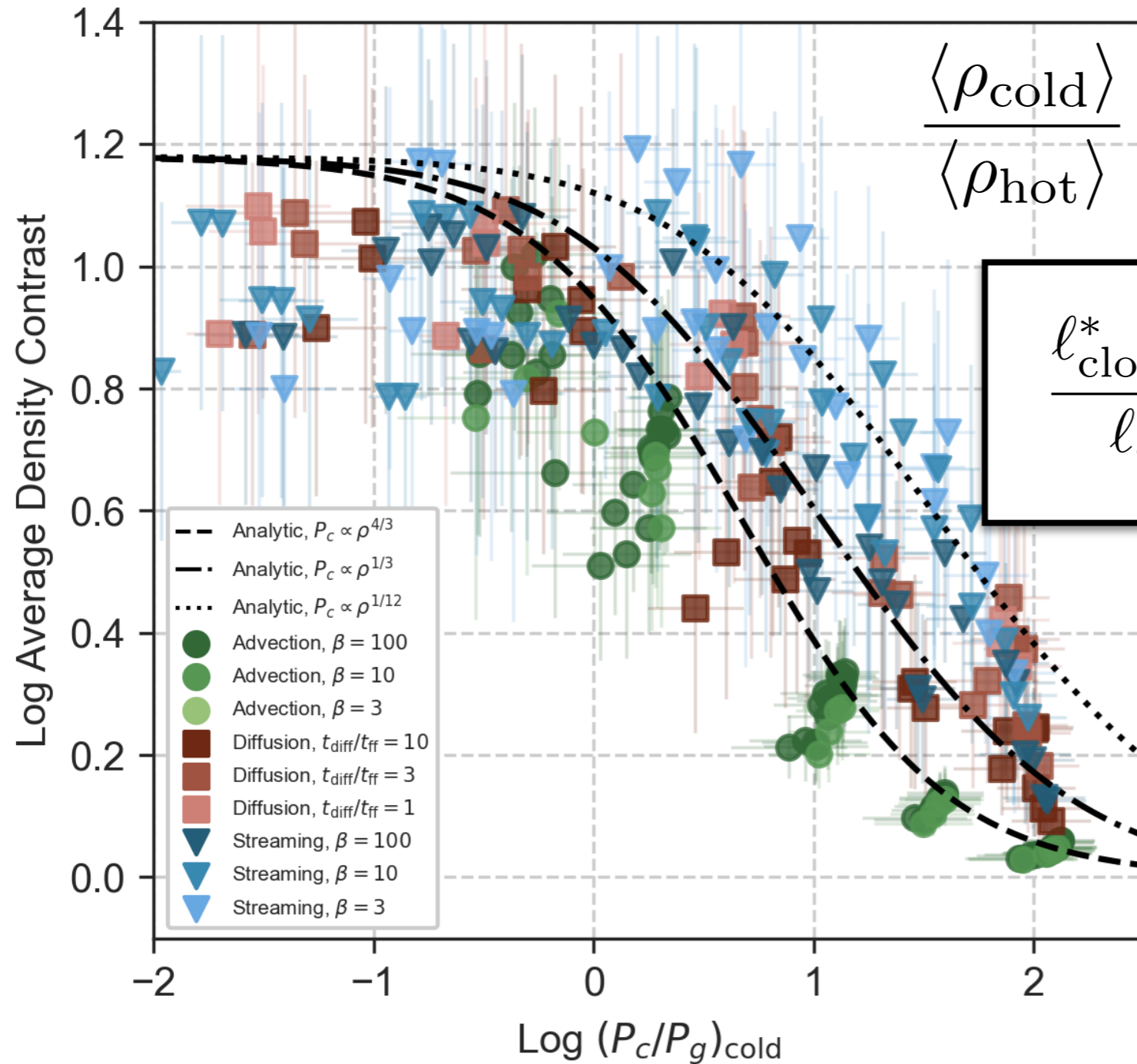


transport





# Non-thermal pressure support **decreases cold gas density** and **increases size of cold clouds**



$$\frac{\langle \rho_{\text{cold}} \rangle}{\langle \rho_{\text{hot}} \rangle}$$

$$l_{\text{cloudlet}} \sim \min(c_s t_{\text{cool}})$$

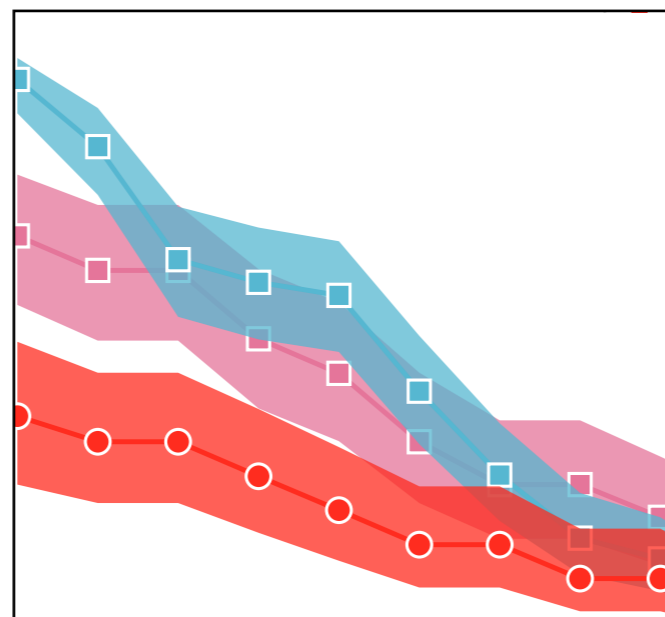
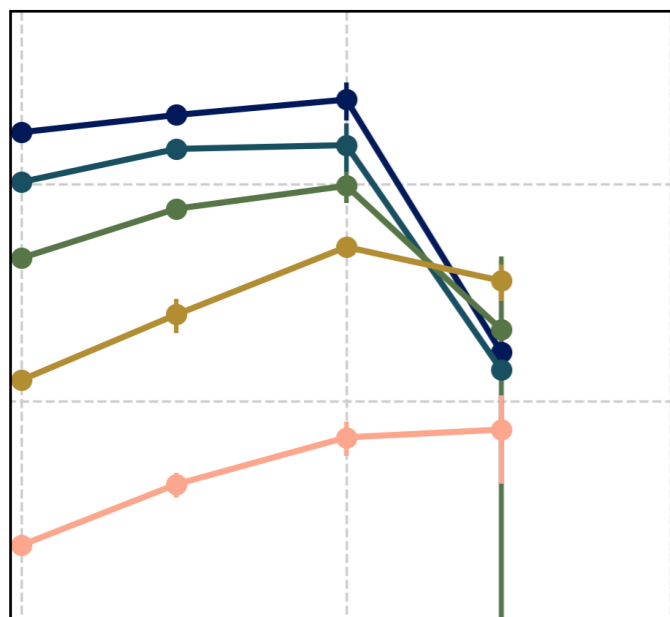
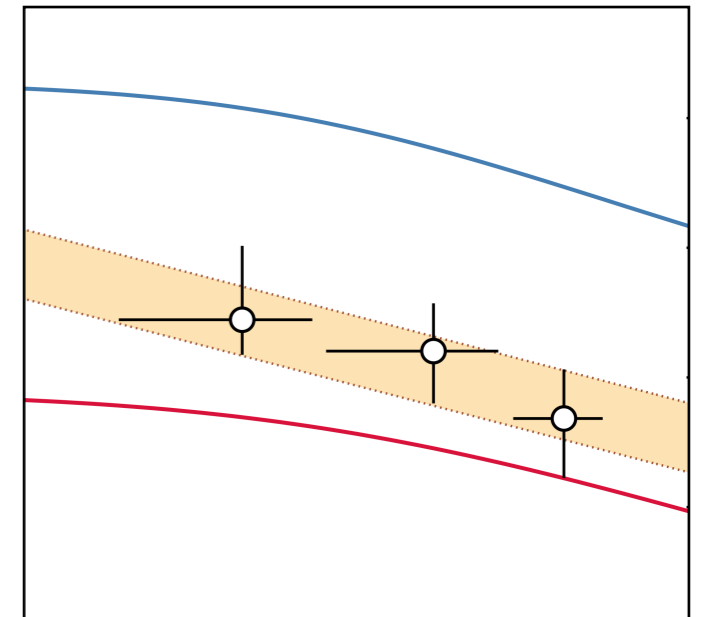
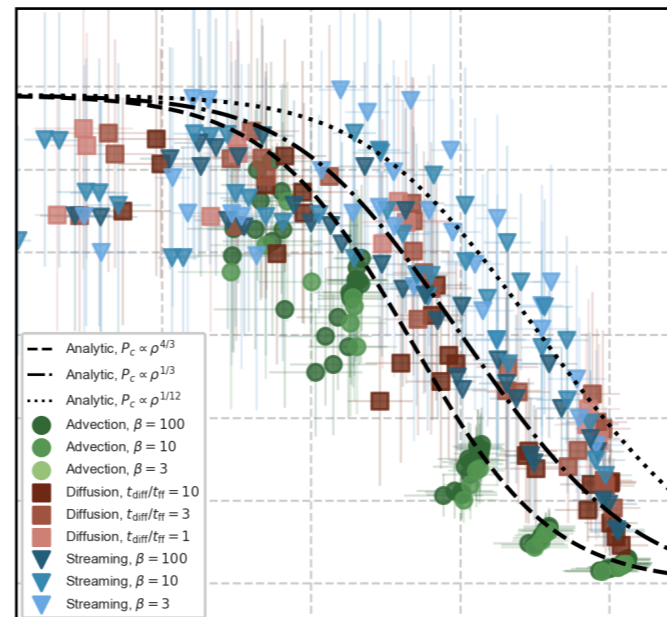
$$\frac{l_{\text{cloudlet}, \eta \gg 1}^*}{l_{\text{cloudlet}}} \approx \Theta \left( \frac{\gamma_{\text{c,eff}}}{\gamma} \eta_{\text{cold}} \right)^{1/2}$$

$$\Theta = T_{\text{hot}}/T_{\text{cold}}$$

$$\eta = P_c/P_g$$

In summary, **accurate interpretations** of CGM observations **need to account for** the presence of **cosmic rays**.

for example, cosmic rays **decrease** cold gas density and **increase** cold cloud sizes



cosmic ray pressure can also **decrease** cold gas **accretion rates**