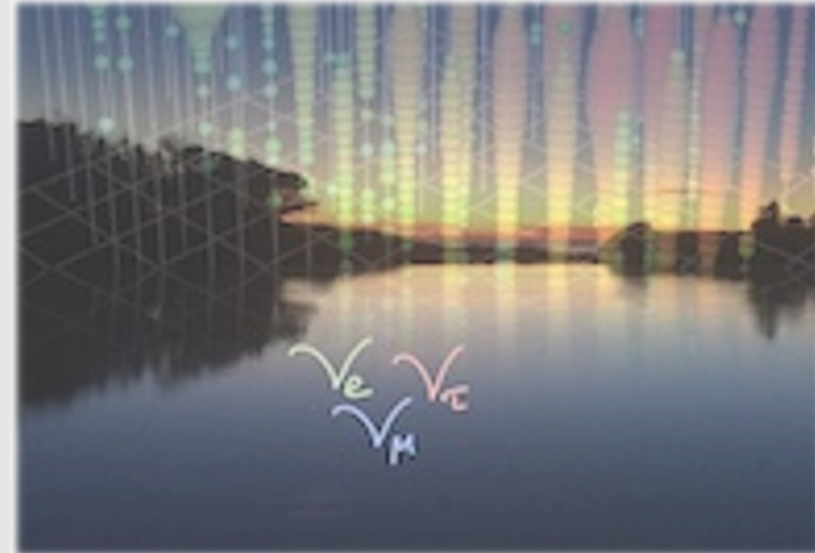


# Neutrino scattering and oscillation measurements with NOvA data

Interdisciplinary Developments  
in Neutrino Physics

@ KITP, UC Santa Barbara



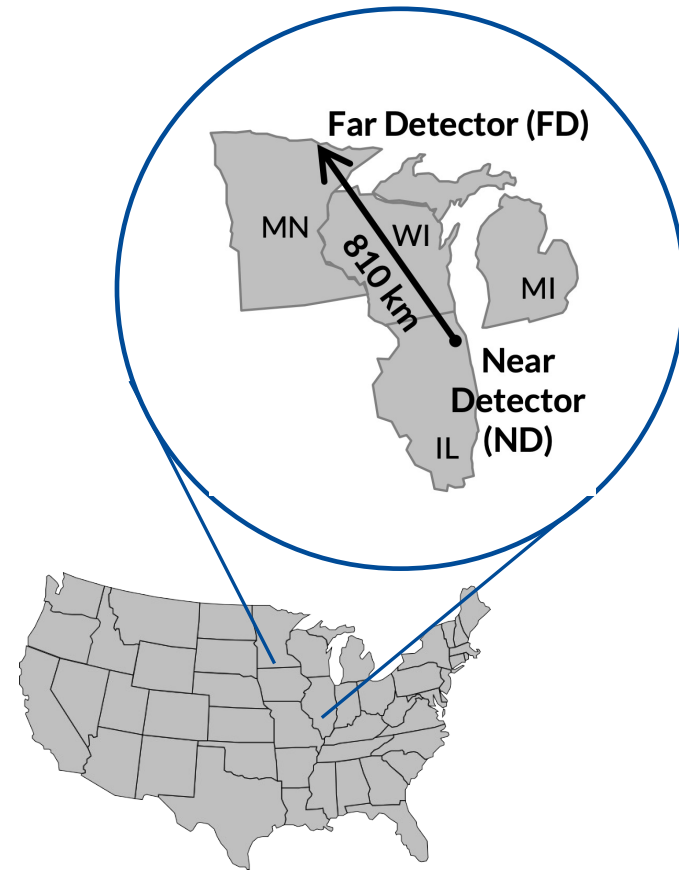
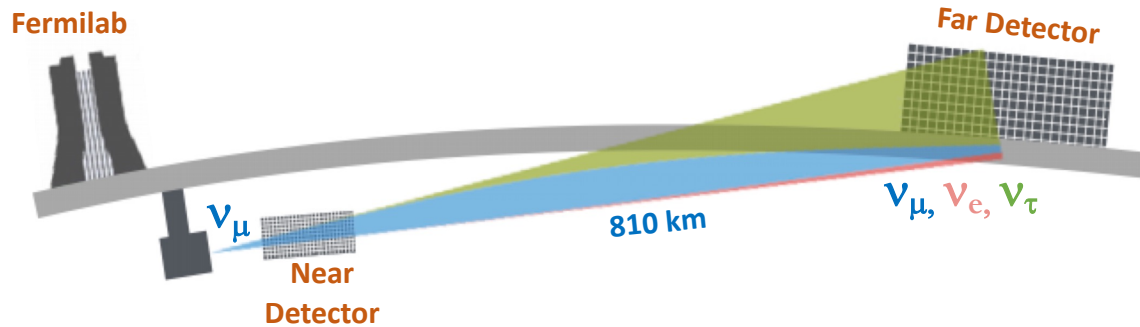
**Zoya Vallari**

on behalf of NOvA collaboration

**28 March 2022**

[ zoya@caltech.edu ]

# The NOvA Experiment

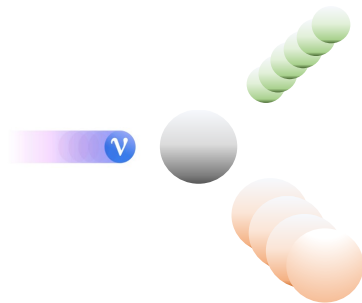
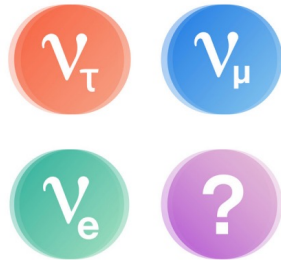
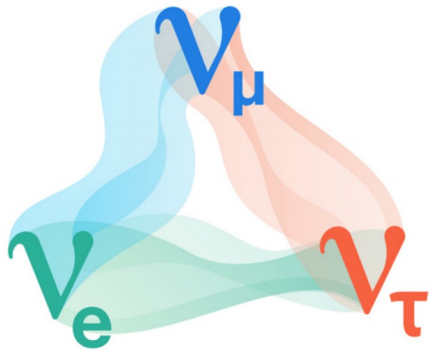


- NOvA is an accelerator based **long-baseline neutrino oscillation** experiment.
- It receives a beam of **muon (anti)neutrinos** peaked at **2 GeV** energy from the NuMI beam facility at Fermilab.

## HUMANS OF NOvA

- ◎ 200 collaborators
- ◎ 50 institutions
- ◎ 8 countries

# NOvA's Physics Program



3-flavor Neutrino Oscillation

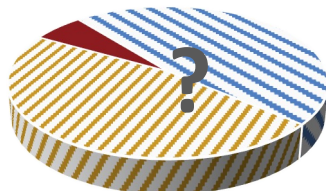
Sterile Neutrino Search

Astrophysical measurements

Neutrino-nucleon interaction

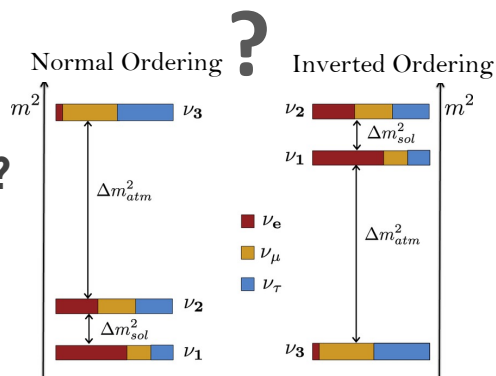
# NOvA's Physics Goals

1.  $\theta_{23}$  : Is the mixing maximal?



3-flavor Neutrino Oscillation

2. Mass Ordering: Normal or Inverted?



Sterile Neutrino Search

3.  $\delta_{CP}$ : Do neutrinos violate CP?

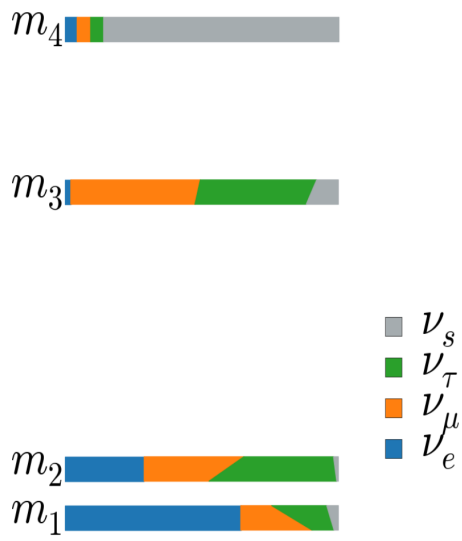


Astrophysical measurements

Neutrino-nucleon interaction

# NOvA's Physics Goals

- Search for 3+1 sterile neutrino models in the neutral current disappearance channel.



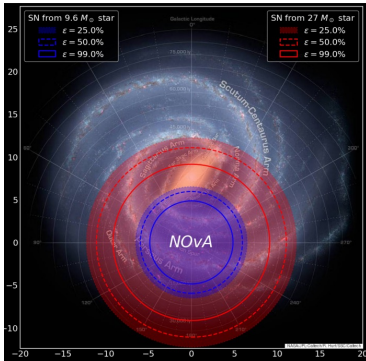
3-flavor Neutrino Oscillation

**Sterile Neutrino Search**

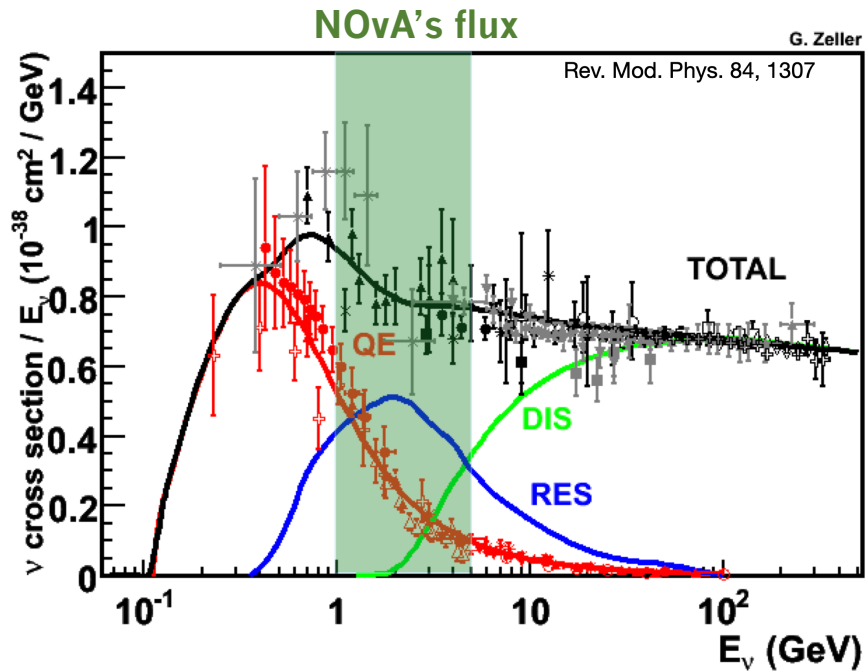
**Astrophysical measurements**

Neutrino-nucleon interaction

- Seasonal Cosmic Ray variations.
- Search for a magnetic monopole component of cosmic rays.
- Detection of supernova neutrinos.



# NOvA's Physics Goals



- Very high statistics environment at the NOvA ND.
- Cross section measurements provide constraints on neutrino interaction models.

3-flavor Neutrino Oscillation

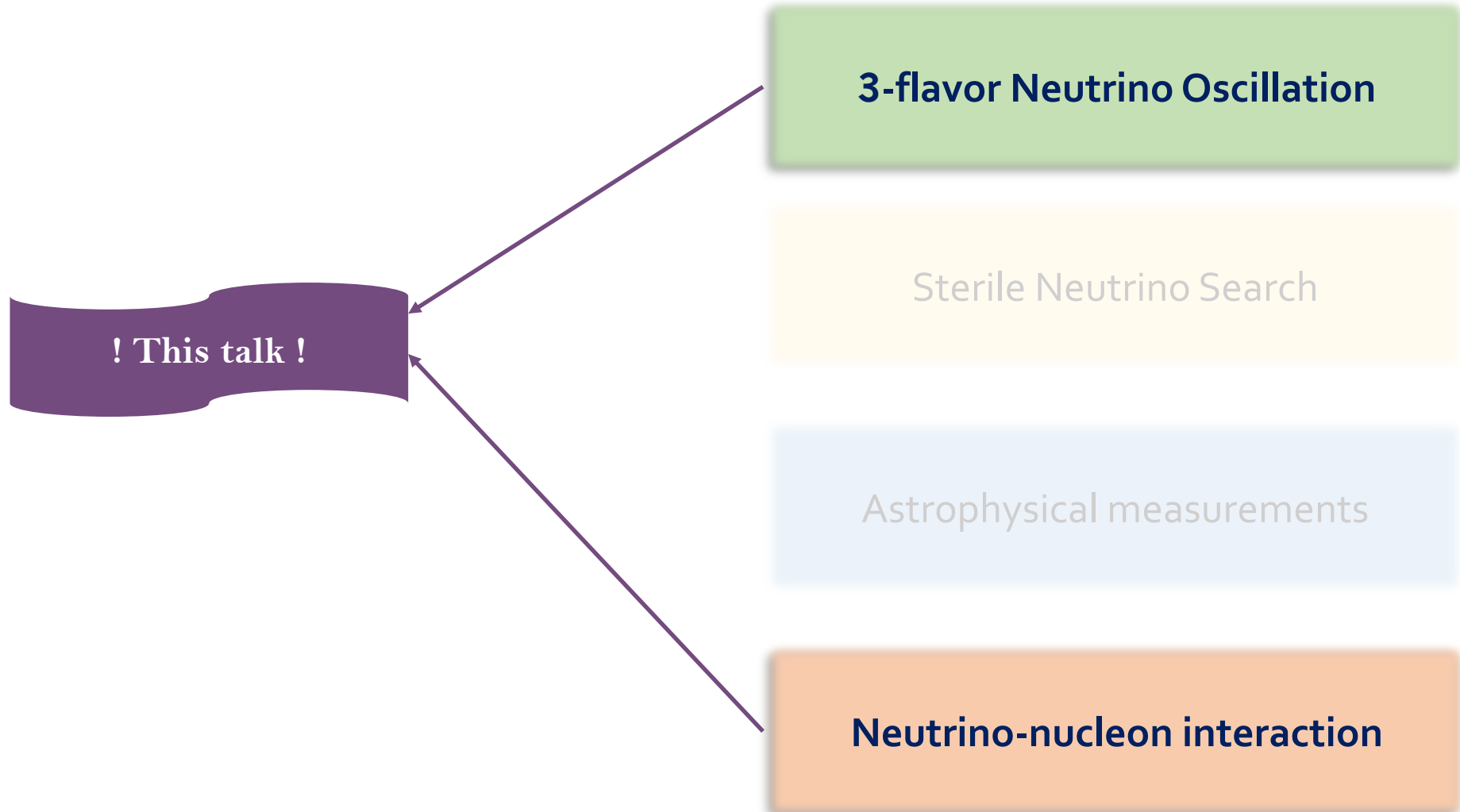
Sterile Neutrino Search

Astrophysical measurements

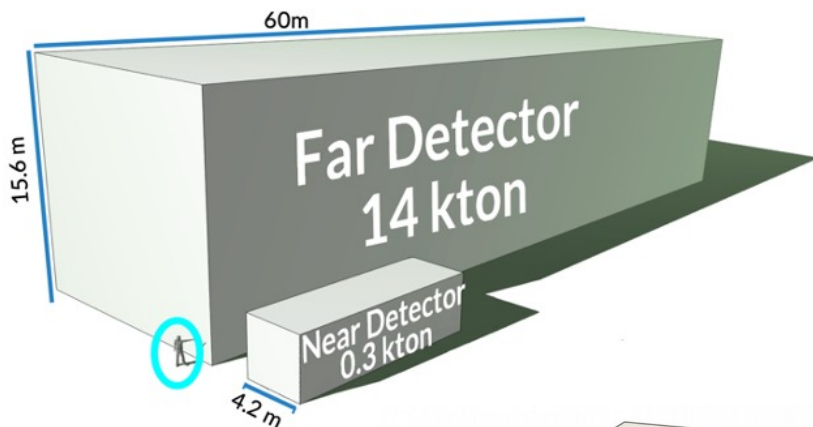
**Neutrino-nucleon interaction**

# NOvA's Physics Program

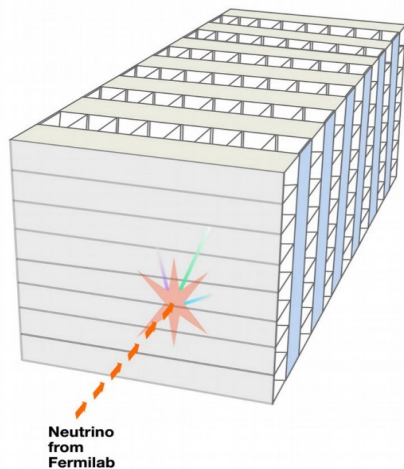
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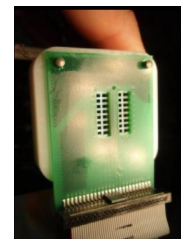
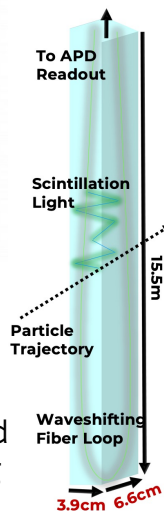
# The NOvA detectors



Alternating layers of scintillator planes which provide two 2D views.



Liquid scintillator contained in PVC cells. Scintillation light is carried by the wavelength shifting fibers to APDs.



Avalanche Photodiode (APD)

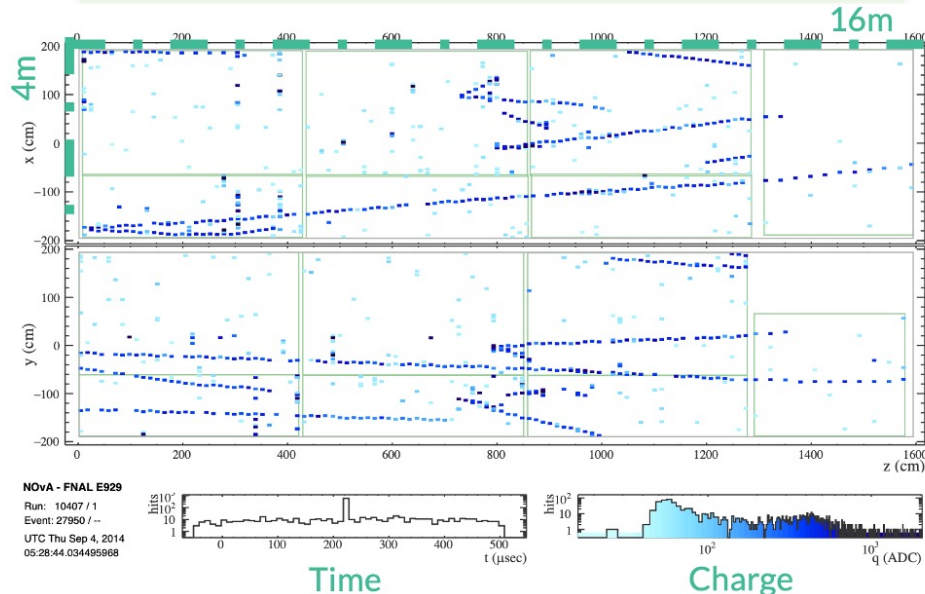
- NOvA's ND and FD are **functionally identical detectors**.
- Segmented liquid scintillator detectors:
  - Particle detection via tracking and calorimetry.
- Optimized for electron showers:
  - **~6 samples per  $X_0$**  (40cm) & **~60%** active volume
- Nuclear Targets: CH<sub>2</sub> (77%), Cl (16%), TiO<sub>2</sub>(6%)
- Spatial resolution: **~few cm** (good rejection of cosmic events)
- Timing resolution: **few ns** (to distinguish pile-ups)



# The NOvA detectors

## Near Detector (ND)

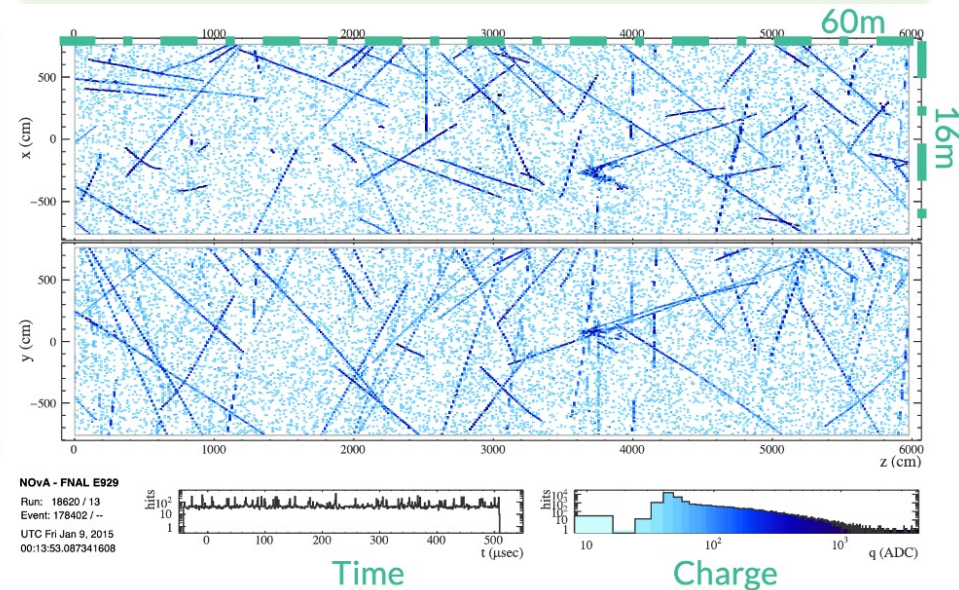
- 290-ton
- ~20k channels
- **1km** from beam target
- 5 contained neutrino events per beam pulse
- **Underground** : negligible cosmic neutrino events



**ND Event Display** : Shows pileup during the 10 $\mu$ s ND beam spill

## Far Detector (FD)

- 14k-ton
- ~344k channels
- **810 km** from beam target
- < 1 neutrino event per day
- **On the ground** : ~150 kHz cosmic neutrino background



**FD Event Display** : full 550 $\mu$ s exposure in FD dominated by cosmics

# Simulation

## Neutrino Flux



- Geant4 based simulation for NuMI beamline : **G4NuMI**
- Hadron production model constrained with external measurements on thin target.

## Neutrino Interaction



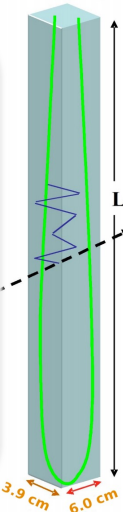
- $\nu$ -A interactions are simulated with the GENIE generator.
- The simulation is tuned to both the external and the internal data from the ND.

Today's results:

1. Neutrino-oscillation : **GENIE v3.0.6**
2. Cross-section – **GENIE v2.12.2**

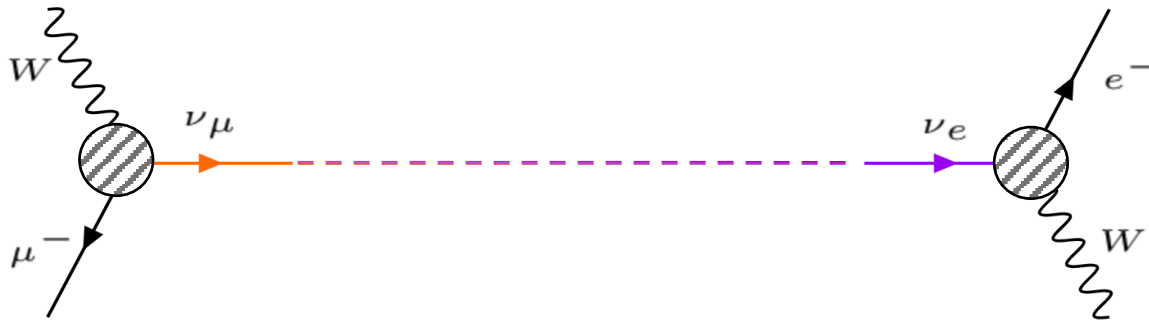
## Detector Response

- Geant4 is used to propagate the final state particles inside the detector volume.
- Custom simulation for light readout and front-end electronics.





## 3-flavor neutrino oscillation



# Measuring Oscillations

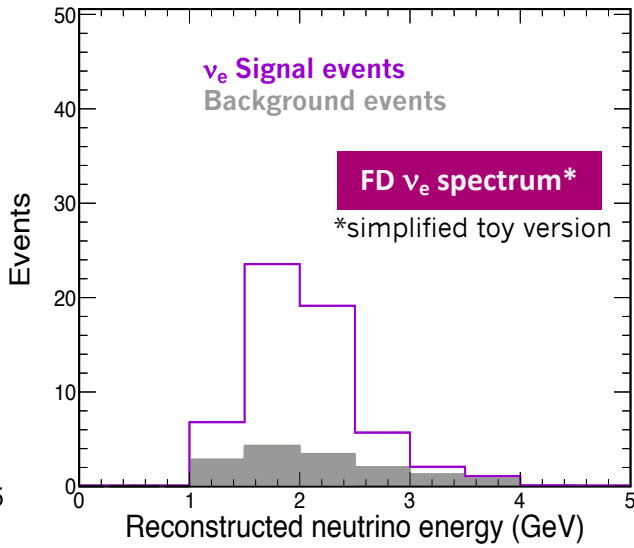
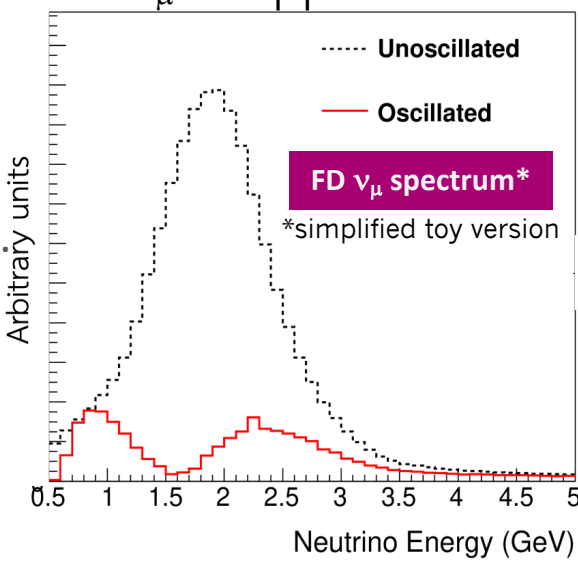
Two Channels

$\bar{\nu}_\mu$  disappearance

$\bar{\nu}_e$  appearance

## Measurement

- 1. Detect neutrino interactions.
- 2. Identify the flavor.
- 3. Reconstruct energy.

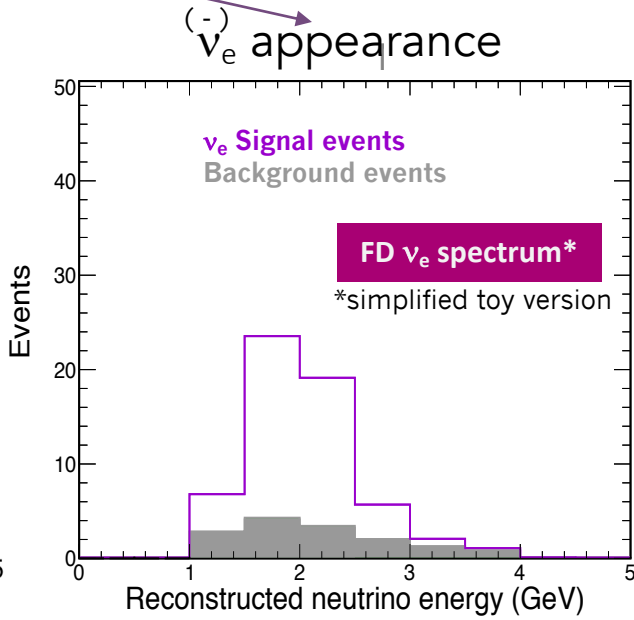
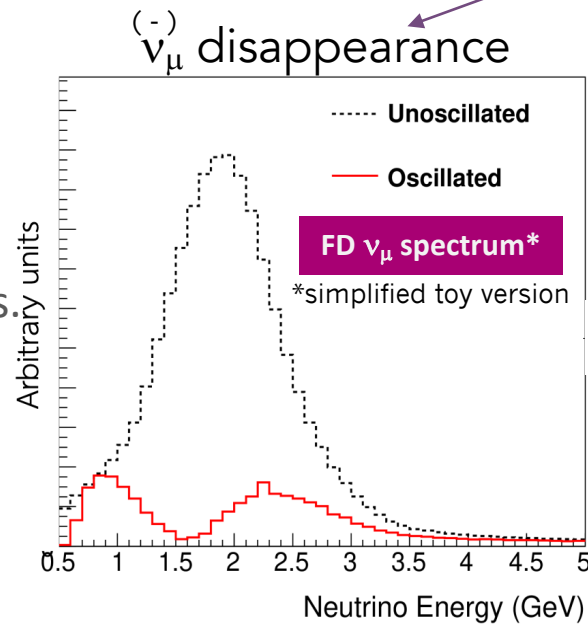


# Measuring Oscillations

Two Channels

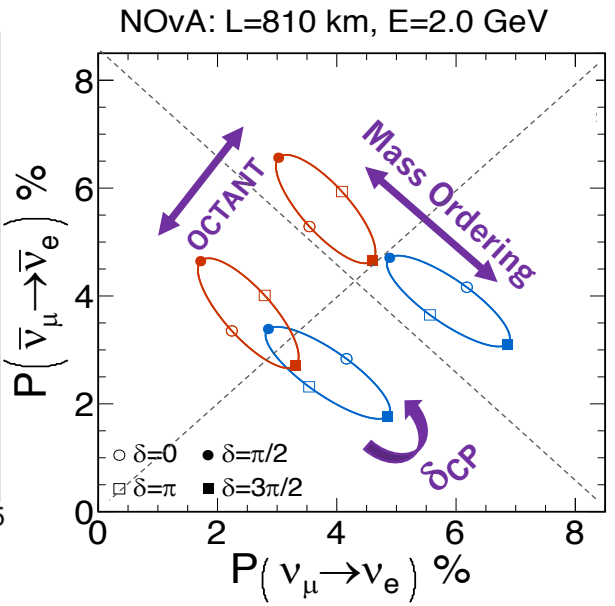
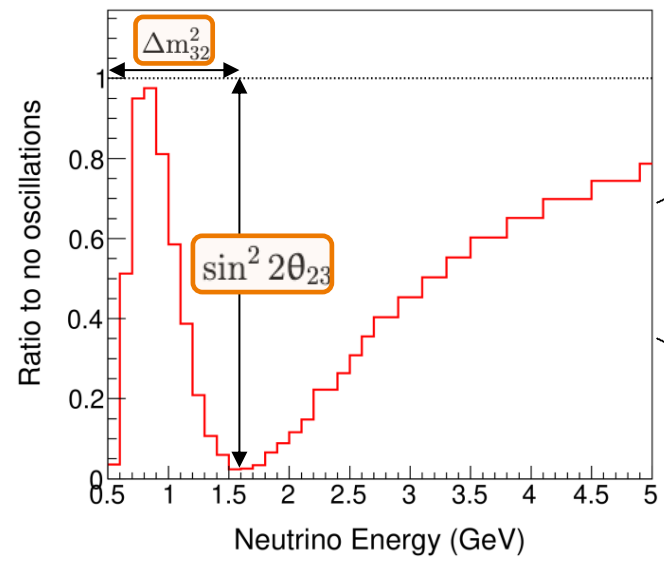
## Measurement

1. Detect neutrino interactions.
2. Identify the flavor.
3. Reconstruct energy.



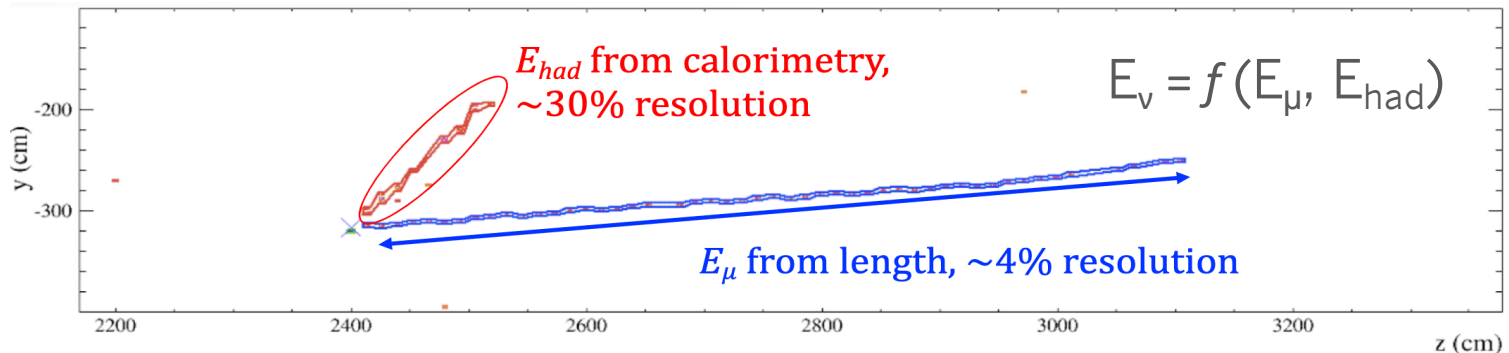
## Inference

1. Compare with the no oscillation prediction.
2. Compare  $\nu_e$  and  $\bar{\nu}_e$  appearance rate.



# Reconstructing Neutrino Energy

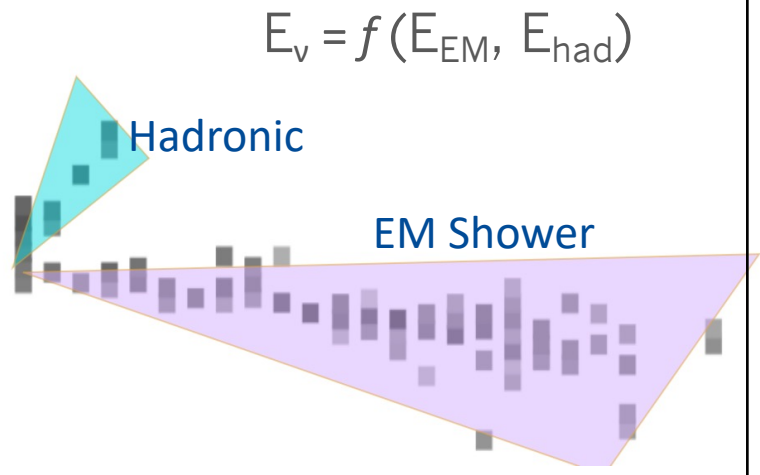
$\nu_\mu$



- Muon energy is estimated from the track length and hadronic energy from visible hadronic activity.

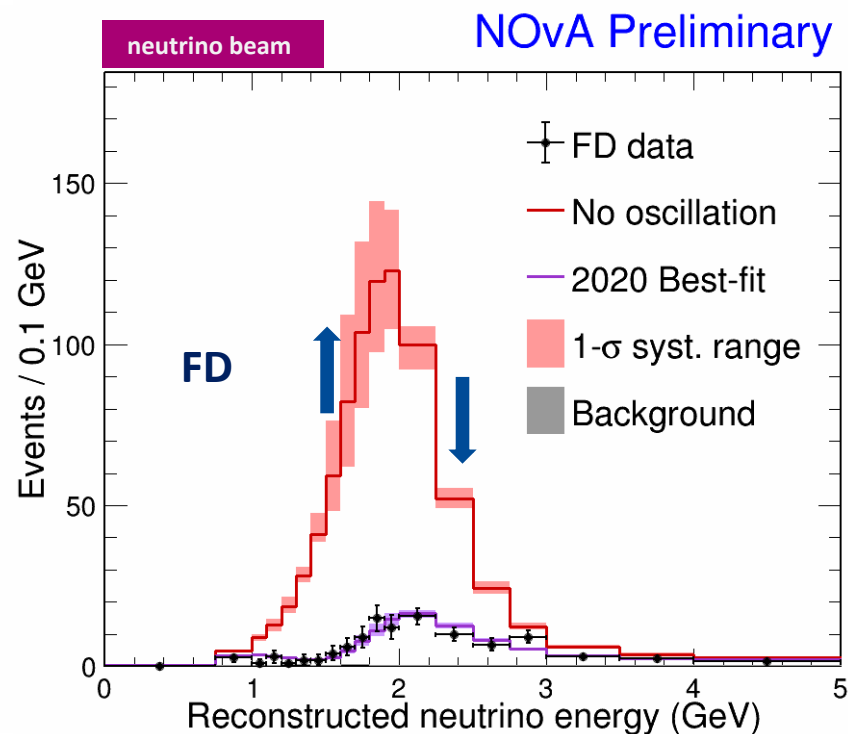
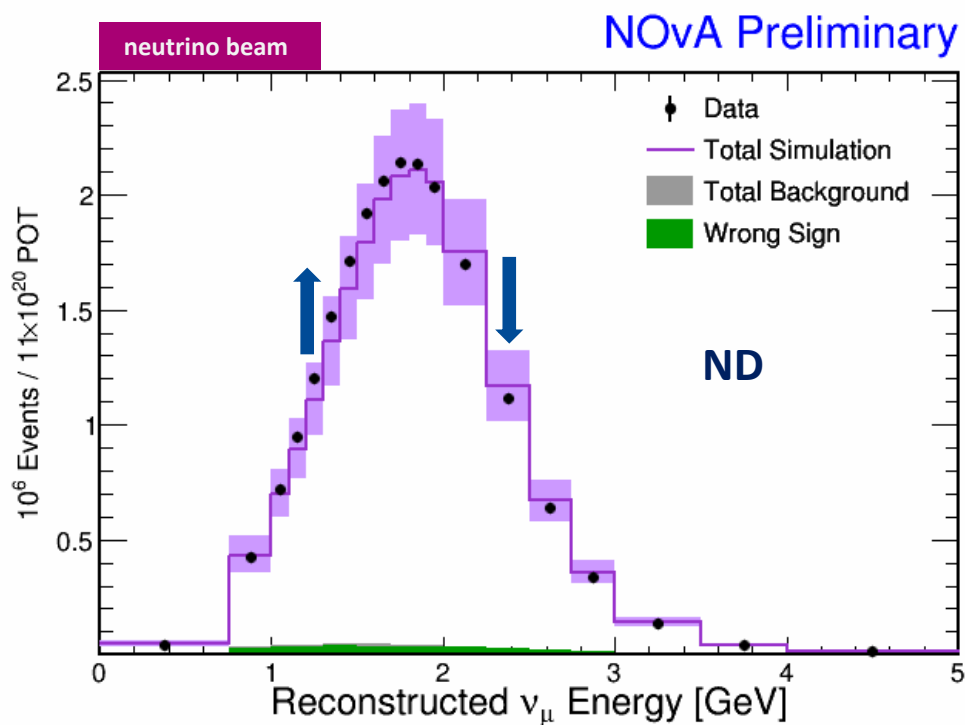
$\nu_e$

- Calorimetric energy estimation is done for EM and hadronic clusters separately.
- Overall neutrino energy resolution is  $\sim 11\%$ .



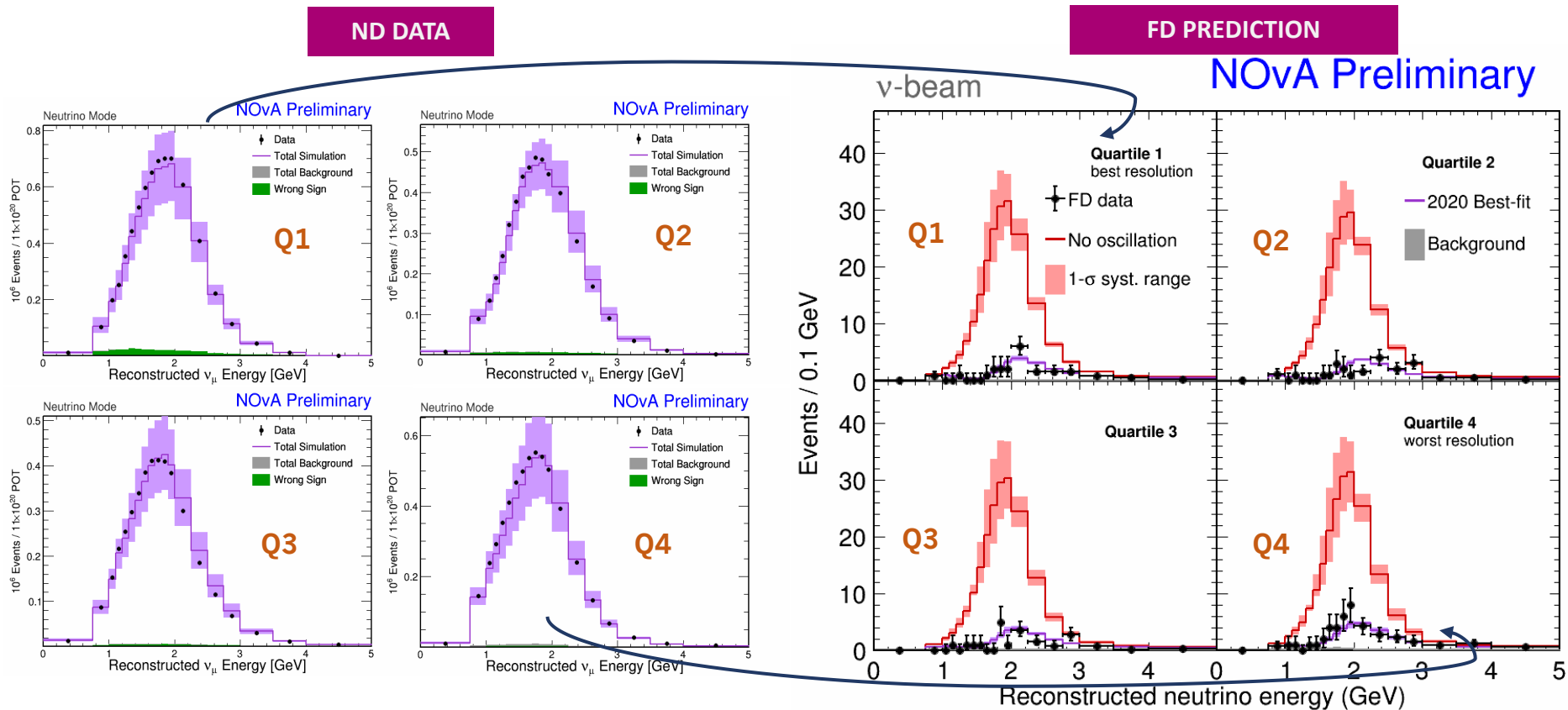
# Extrapolation : ND Data in Oscillation Analysis

- Data from the functionally identical ND is used to predict both  $\nu_\mu$  and  $\nu_e$  signal spectra at the FD.
- The differences between the two detectors in flux, acceptance and cross-sections are modeled using simulations and related systematic uncertainties.



# Extrapolation: Energy Resolution

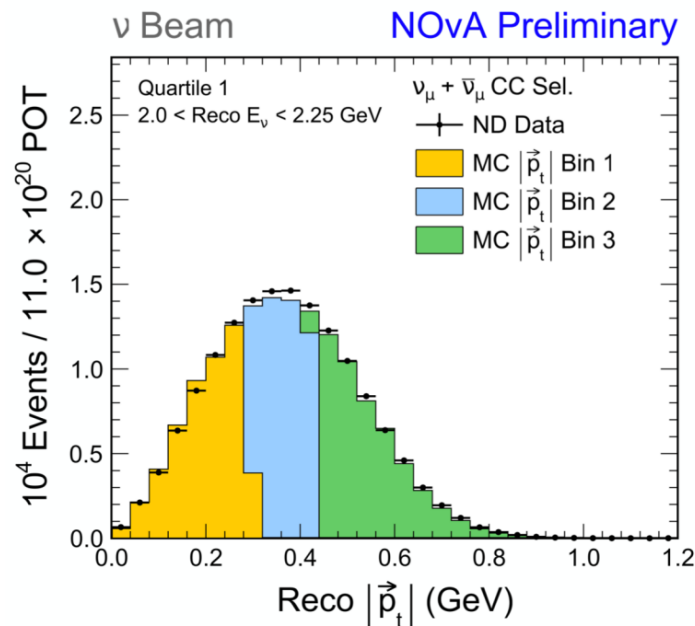
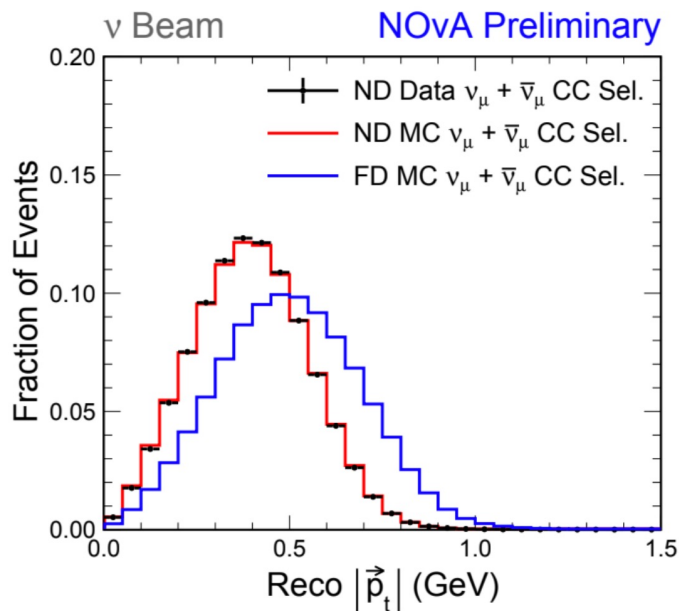
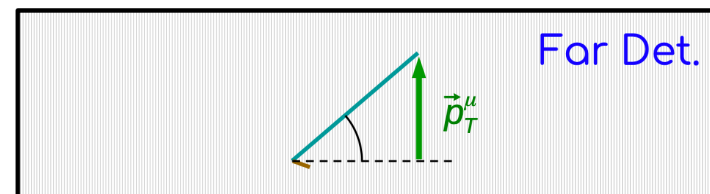
- Oscillation sensitivity for  $\nu_\mu$  disappearance measurement depends on the shape of the spectrum.
- Dividing the  $\nu_\mu$  sample in **quartiles (Q1-Q4)** of fraction of hadronic energy separates high-resolution events. This increases the sensitivity to the shape of the oscillation dip.





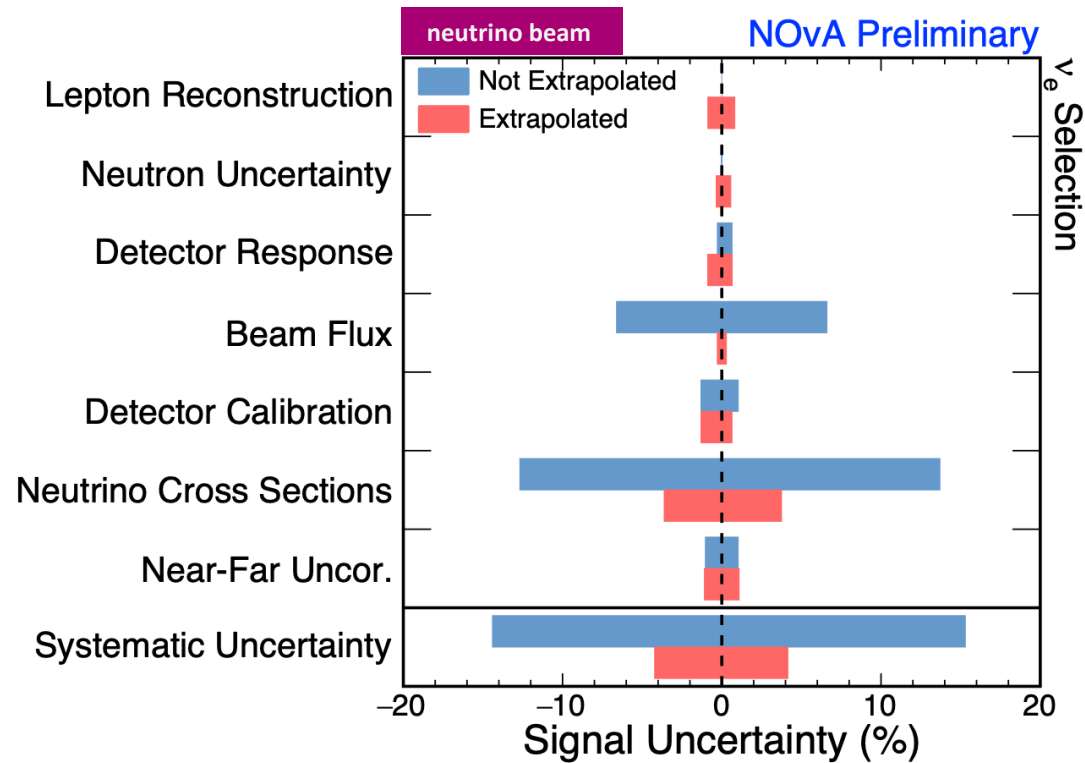
# Extrapolation: Detector Acceptance

- Due to a large difference in size of the detector, FD has a larger acceptance of the particles going in the transverse direction than the ND.
- This difference is evident in the lepton  $|\vec{p}_t|$  distributions.
- **Sub-divide the sample in 3 bins of lepton  $|\vec{p}_t|$**  and extrapolate ND data for each bin separately.

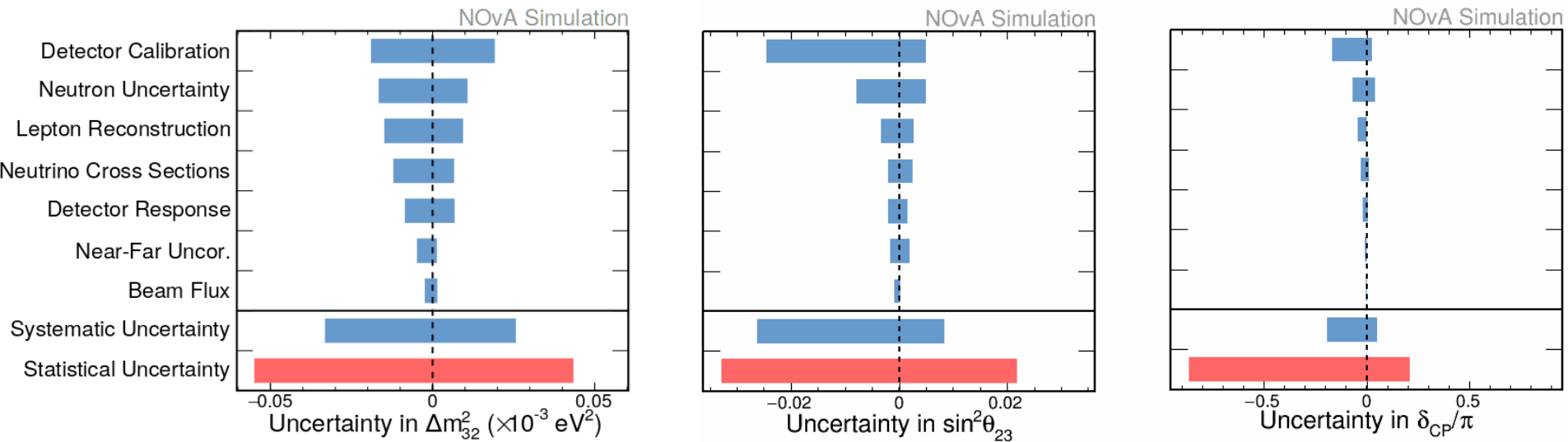


# Extrapolation : Constraining Systematics

ND Data constrains the total systematic uncertainties in the FD prediction from  $>15\%$  to  $\sim 5\%$ .

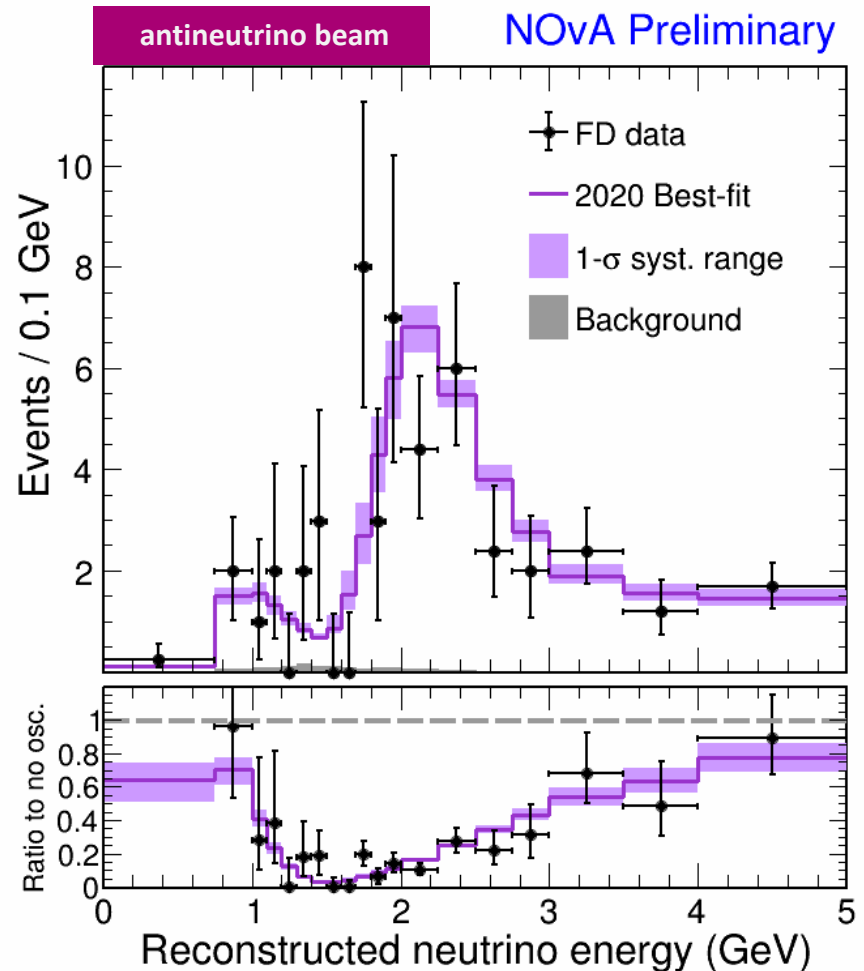
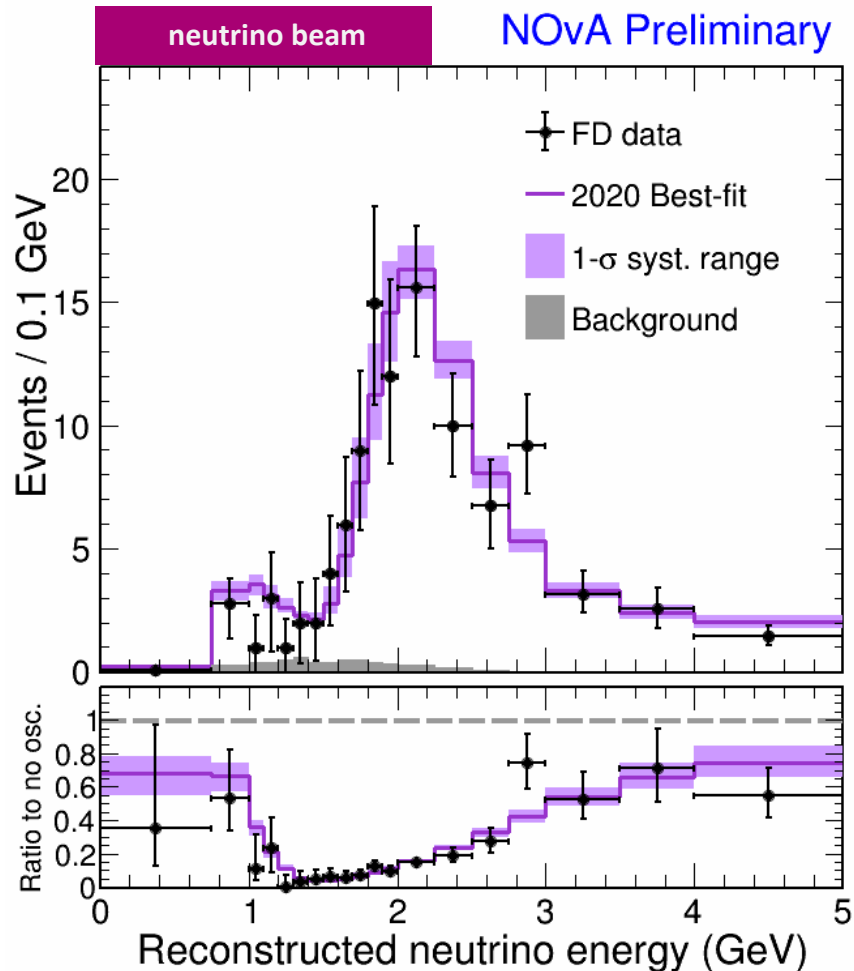


# Total uncertainty on oscillation parameters



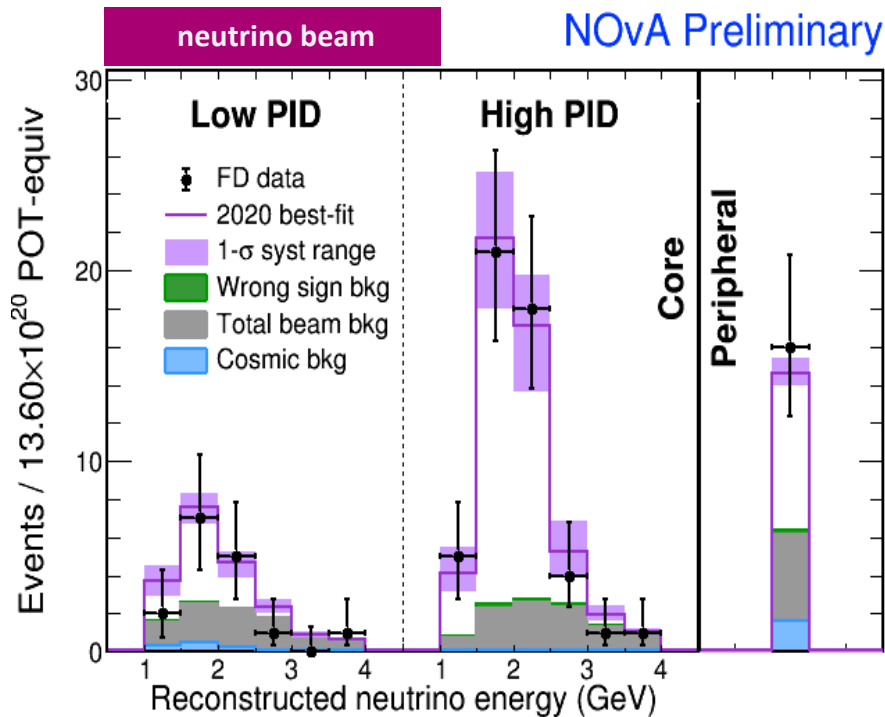
- Measurement of oscillation parameters at NOvA is significantly **statistically limited**.
- Largest source of systematic error comes from **detector calibration** which directly impacts the energy scale of the events.
- NOvA Test Beam Program currently underway to improve detector calibration.

# FD Data Samples : $\nu_\mu$ disappearance

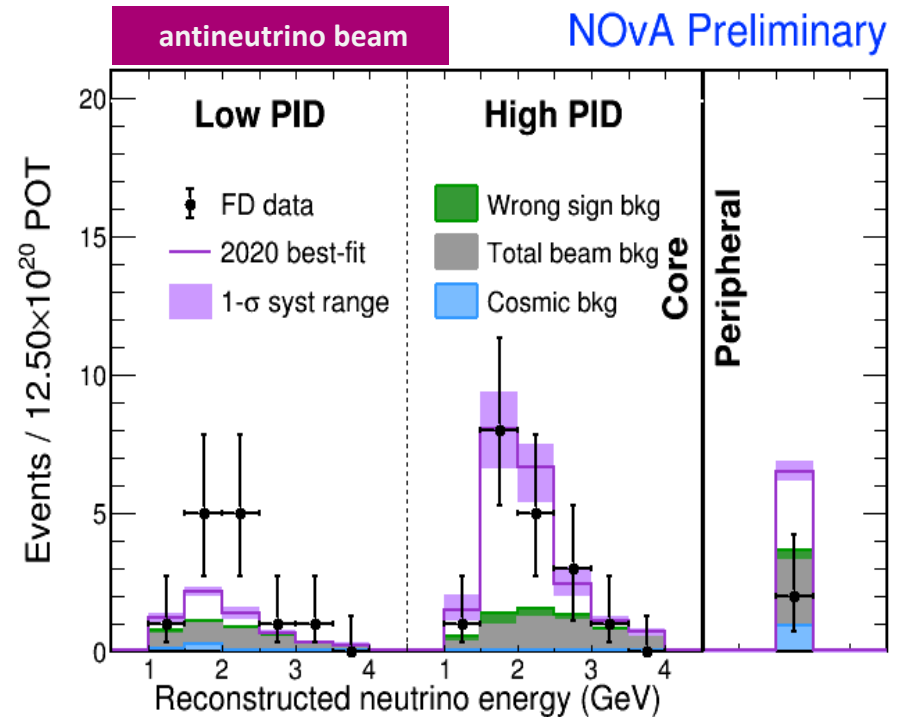


- The predictions (with the systematics band) are varied with the oscillation probabilities until the best-fit values with data are obtained.
- Applying 3-flavor oscillations describes these data well:  $p=0.705$ .

# FD Data Samples : $\nu_e$ appearance



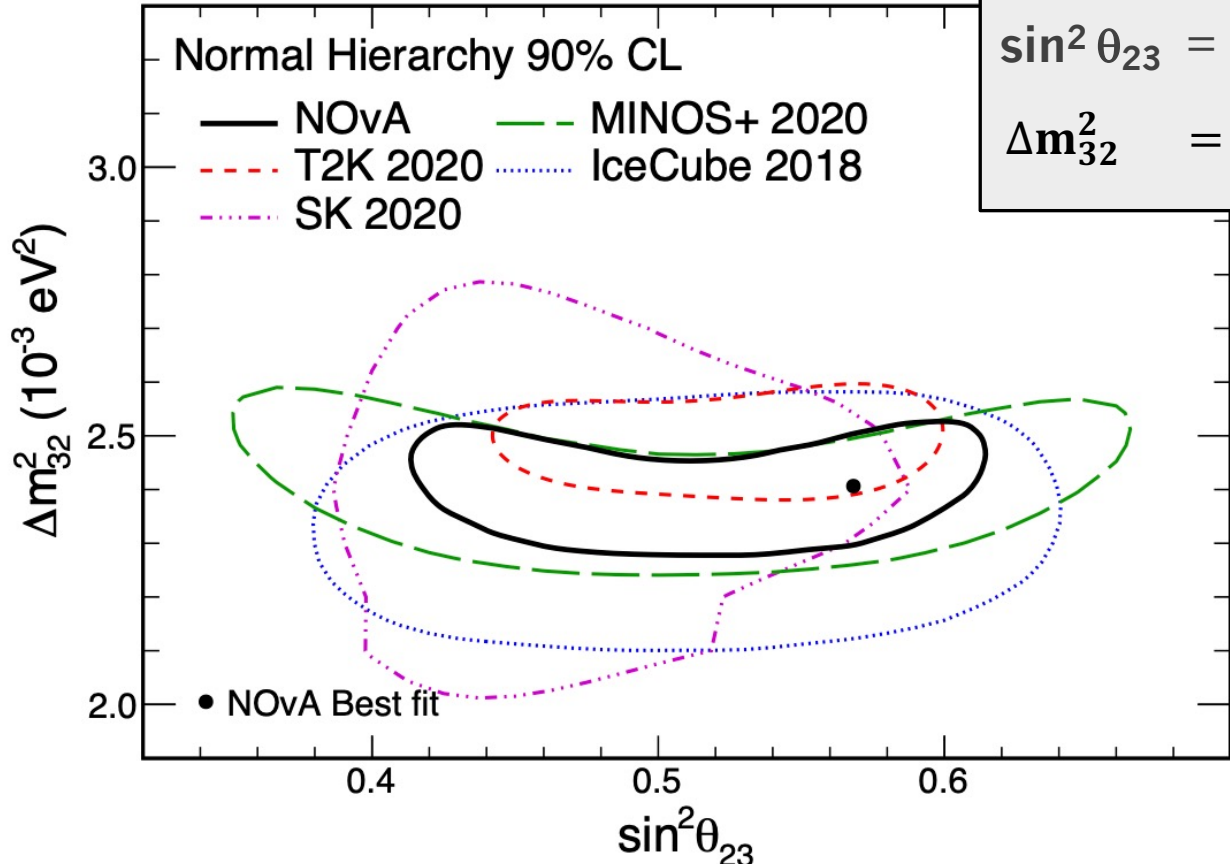
Observed 82 events (27 bkgd)



Observed 33 events (14 bkgd) :  $> 4\sigma$   $\bar{\nu}_e$  appearance

- Separating in bins of Particle ID enhances oscillation sensitivity which is dependent on a better rejection of background events.
- Peripheral sample include high PID events at the edges of the detector which might not be well contained.

# Mass splitting and mixing angle



**NOvA's Best Fit:**

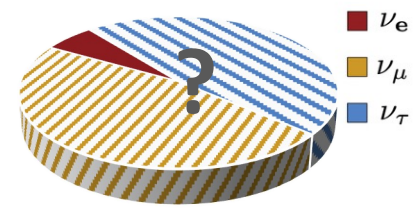
$\sin^2 \theta_{23} = 0.57^{+0.03}_{-0.04}$

$\Delta m^2_{32} = (+2.41 \pm 0.07) \times 10^{-3} \text{ eV}^2$

**~ 2.9% Precision**

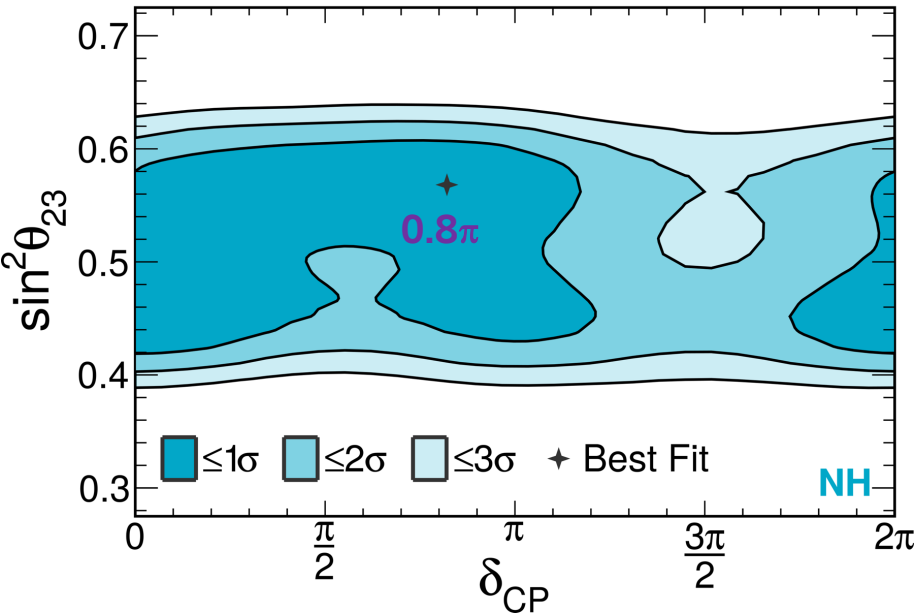
- Maximal Mixing is disfavored at  $1.1\sigma$
- Upper Octant is preferred at  $1.2\sigma$

Is the amount of  $\nu_\mu = \nu_\tau$ ?  $\theta_{23} = 45^\circ$ ?



# $\delta_{CP}$ measurement

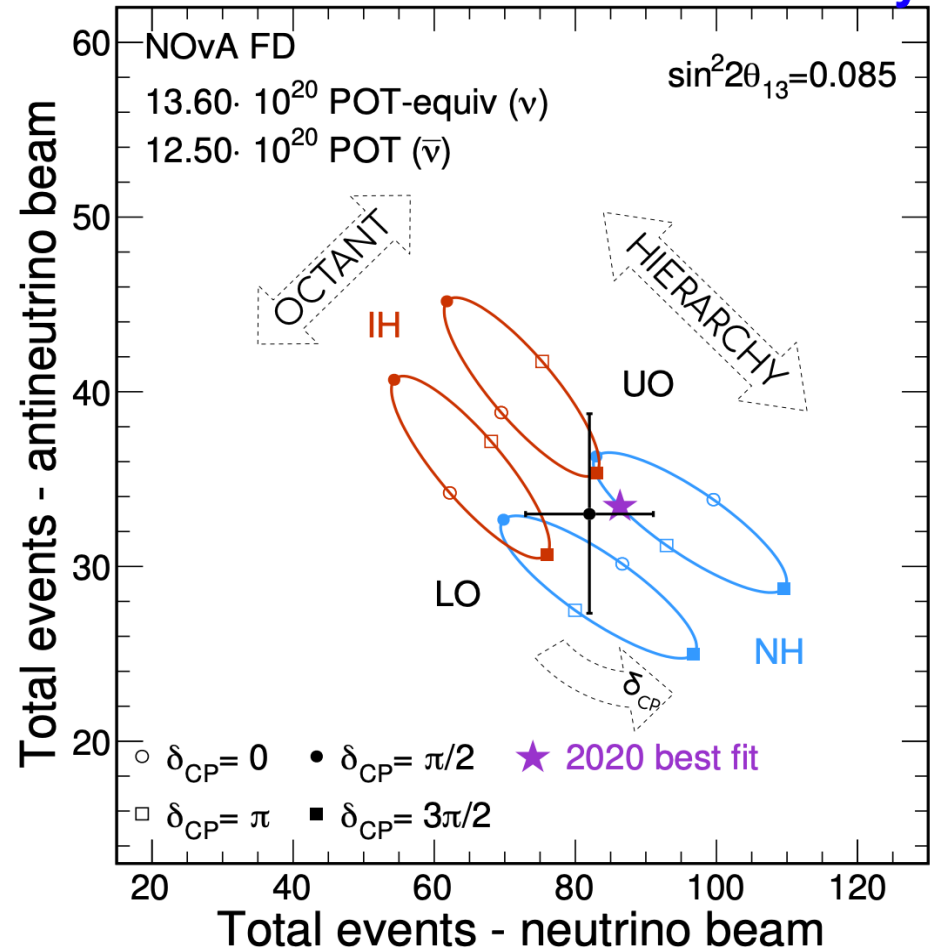
NOvA Preliminary



- Excludes  $\{IH, \delta_{CP} = \pi/2 \text{ at } > 3\sigma\}$  &  $\{NH, \delta_{CP} = 3\pi/2 \text{ at } \sim 2\sigma\}$

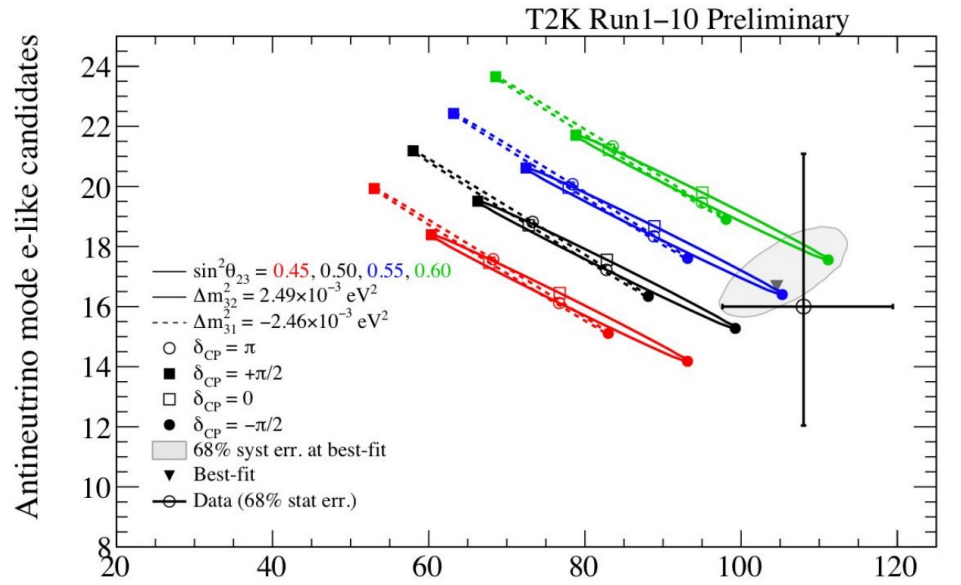
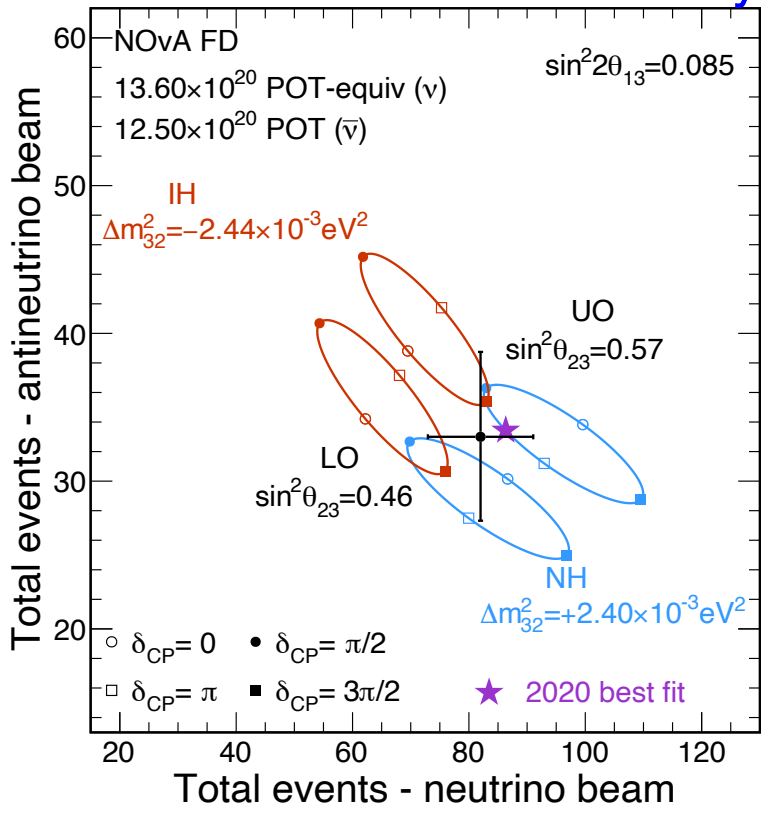
- Normal Ordering is preferred at  $1.0\sigma$

NOvA Preliminary



# NOvA and T2K

## NOvA Preliminary



Neutrino mode e-like candidates  
[P. Dunne's talk at Neutrino 2020](#)

➤ T2K sees an asymmetry in their  $\nu_e$  and  $\bar{\nu}_e$  appearance and their best fit is consistent with large CP violation for Normal Hierarchy.



# NOvA and T2K

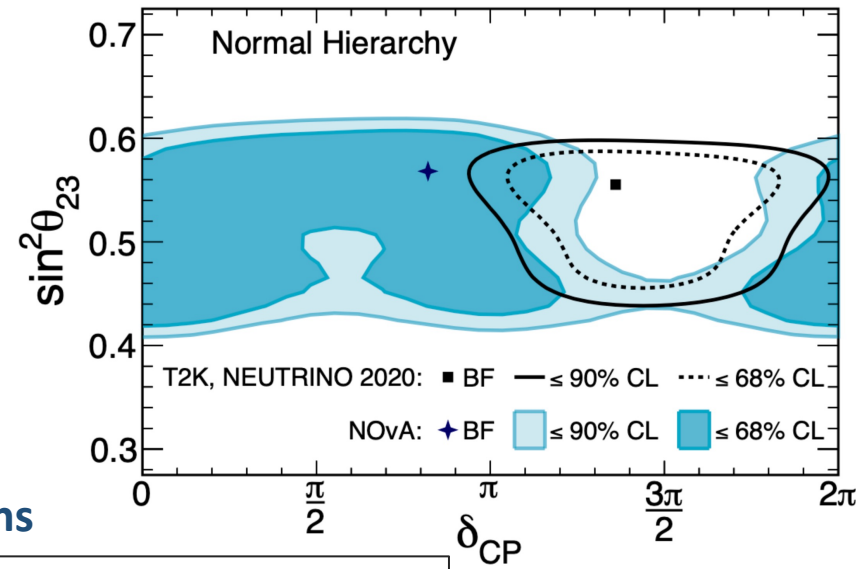
➤ Latest results from the two experiments have generated a lot of excitement in the community.

## ■ Non-standard Interactions

**CP-Violating Neutrino Non-Standard Interactions in Long-Baseline-Accelerator Data**  
 Peter B. Denton<sup>1,\*</sup> Julia Gehrlein<sup>1,†</sup> and Rebekah Pestes<sup>1,2,‡</sup>

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**Non-standard neutrino interactions as a solution to the NOvA and T2K discrepancy**  
 Sabya Sachi Chatterjee<sup>1,\*</sup> and Antonio Palazzo<sup>2,3,†</sup>



## ■ Global preference for Inverted Hierarchy

**Back to (Mass-)Square(d) One:  
 The Neutrino Mass Ordering in Light of Recent Data**  
 Kevin J. Kelly<sup>1,\*</sup> Pedro A. N. Machado<sup>1,†</sup> Stephen J. Parke<sup>1,‡</sup>  
 Yuber F. Perez-Gonzalez<sup>1,2,3,§</sup> and Renata Zukanovich Funchal<sup>4,¶</sup>

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**The fate of hints: updated global analysis of  
 three-flavor neutrino oscillations** [NuFit 5.0]

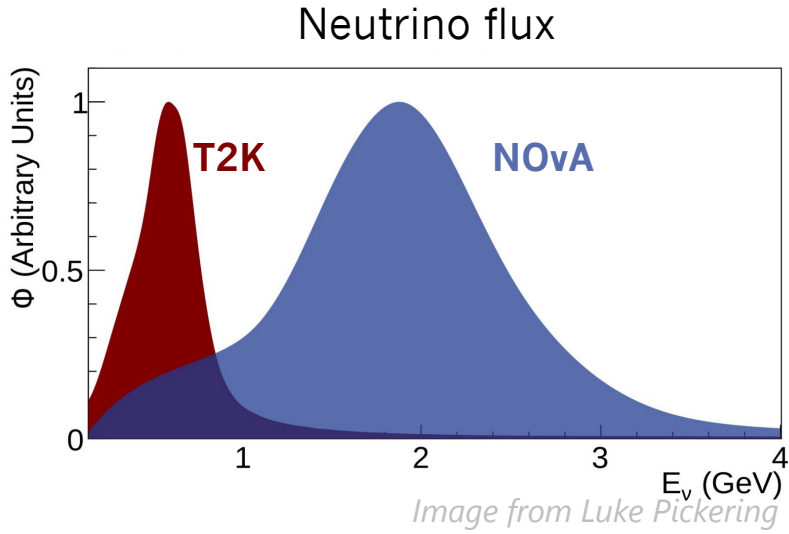
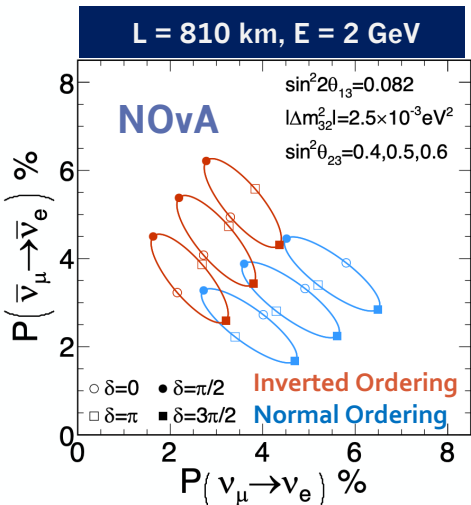
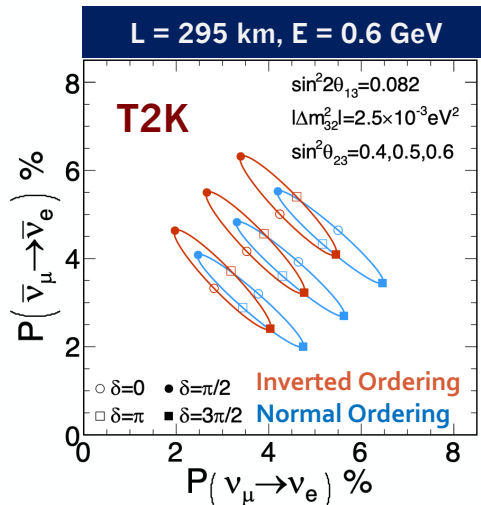
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Ivan Esteban,<sup>a</sup> M. C. Gonzalez-Garcia,<sup>a,b,c</sup> Michele Maltoni,<sup>d</sup> Thomas Schwetz,<sup>e</sup>  
 Albert Zhou<sup>e</sup>

## ■ Energy-dependent parameters

**Energy-Dependent Neutrino Mixing Parameters at Oscillation Experiments**  
 K. S. Babu,<sup>1</sup> Vedran Brdar,<sup>2,3</sup> André de Gouvêa,<sup>2</sup> and Pedro A. N. Machado<sup>3</sup>

# NOvA - T2K Joint Fit



- The NOvA and T2K collaborations are working on a combined joint-fit of their data.
- The joint-fit will provide a significantly **tighter statistical constraint**.
- **Complementary features** in NOvA and T2K experiments will be important for **breaking degeneracies** in the individual measurements.
  - Longer baseline sees larger matter effects and has greater sensitivity to mass hierarchy.
  - T2K is mostly dominated by CCQE (and MEC) events while NOvA has a larger component of RES and DIS events.
  - NOvA and T2K use distinct methods to estimate neutrino energy and different approaches to incorporate Near Detector data to constrain systematics.

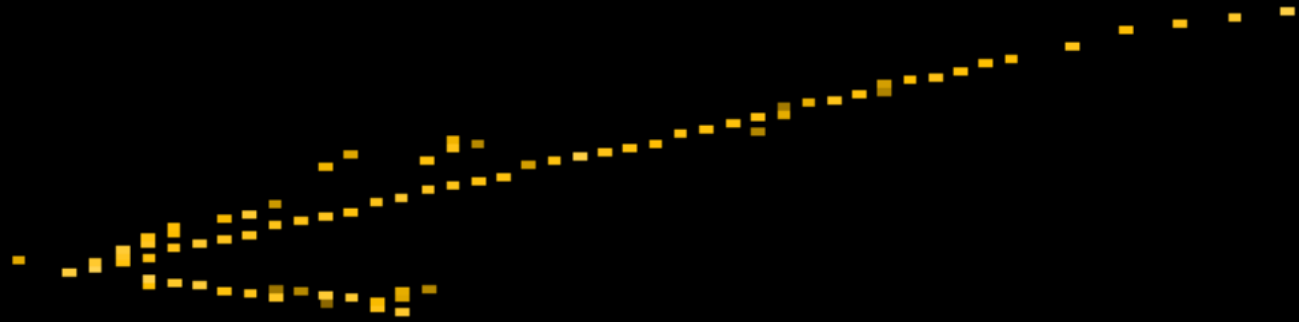
**First Result in \* 2022 \***



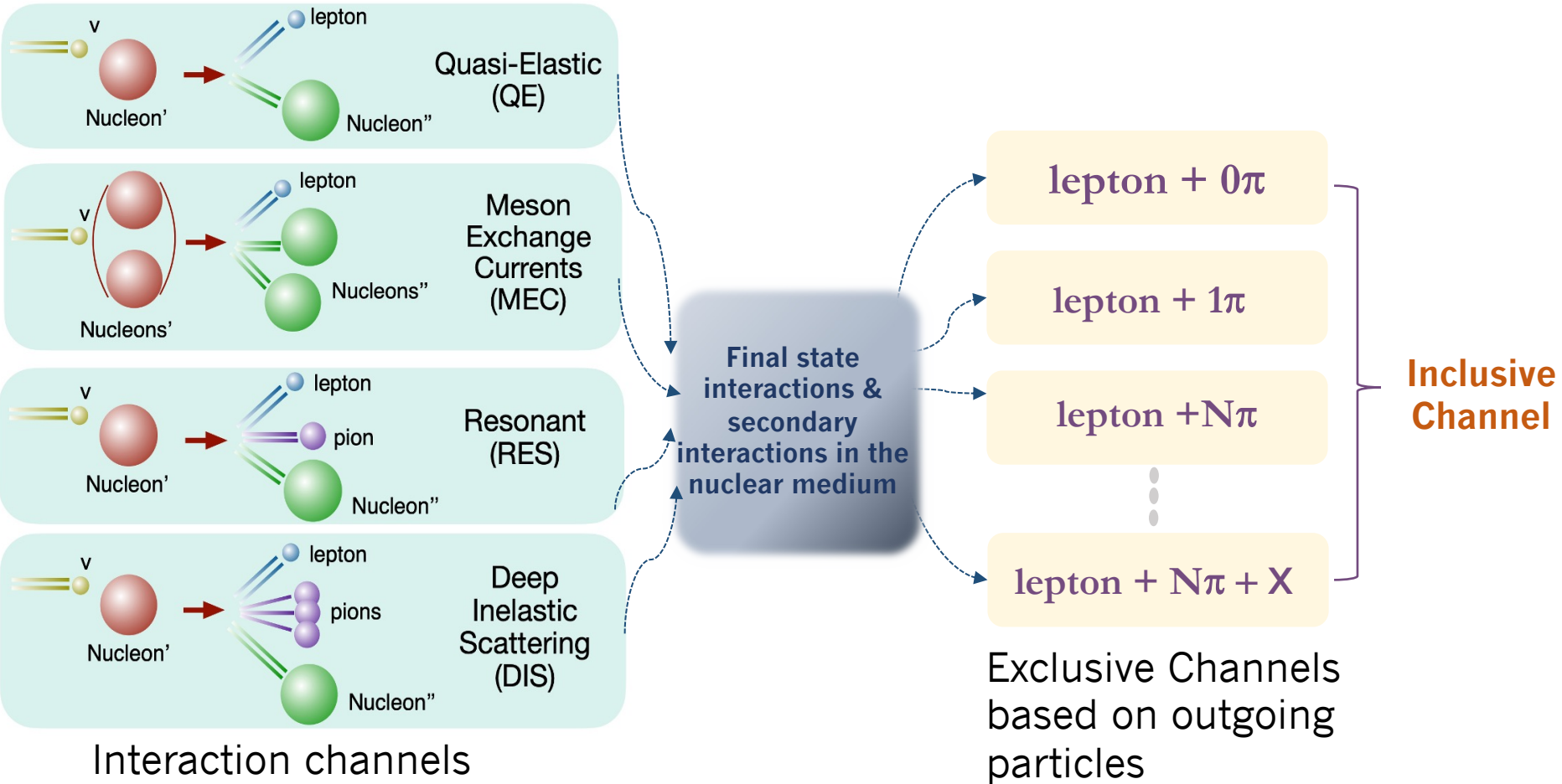
$\nu_{\mu}$  CC

$\nu_e$  CC

# Neutrino-nucleon cross-section measurements



# Cross-section measurements



# Cross-section measurements at NOvA

## Previous results

- ✓ NC Coherent  $\pi^0$  flux averaged cross-section [Phys. Rev. D 102, 012004 \(2020\)](#)
- ✓  $\nu_\mu$  CC  $\pi^0$  single differential cross-section [seminar](#) *[publication in progress]*

## Recent results

[discussed today]

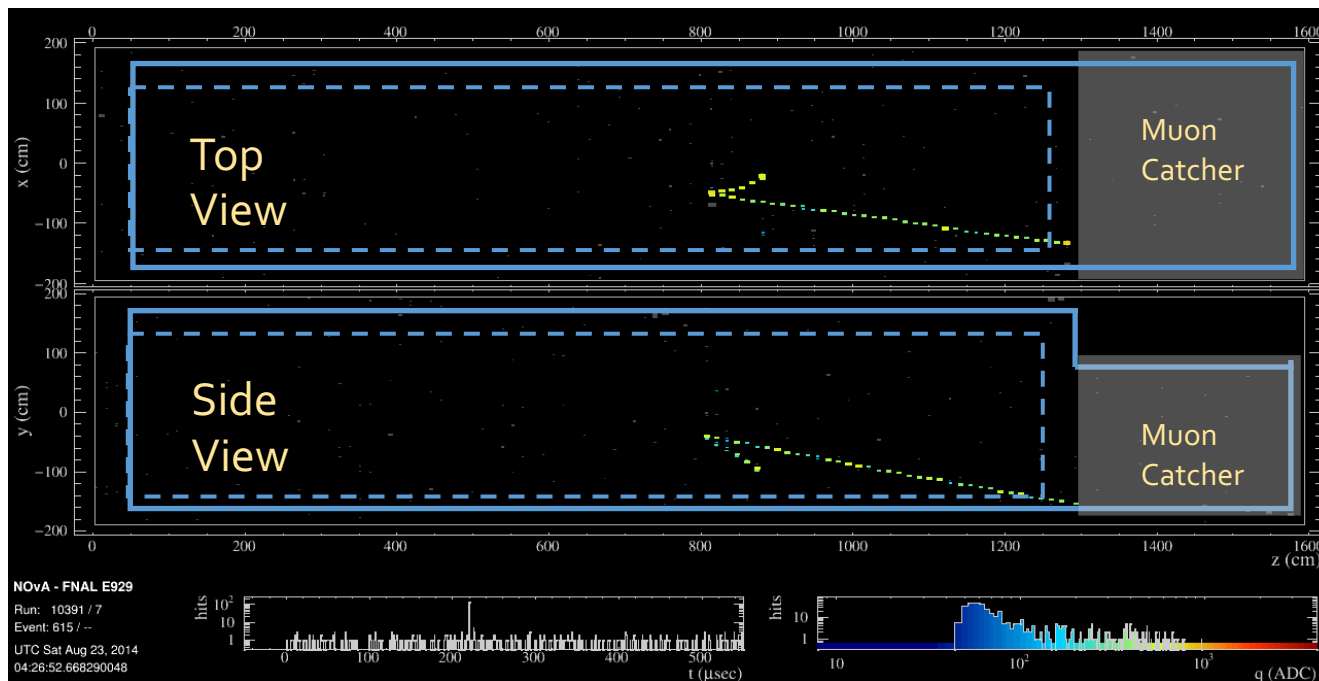
- ✓  $\nu_\mu$  CC inclusive double differential cross-section [seminar](#) *[publication in progress]*
- ✓  $\nu_e$  CC inclusive double differential cross-section [seminar](#) *[publication in progress]*

## Coming soon

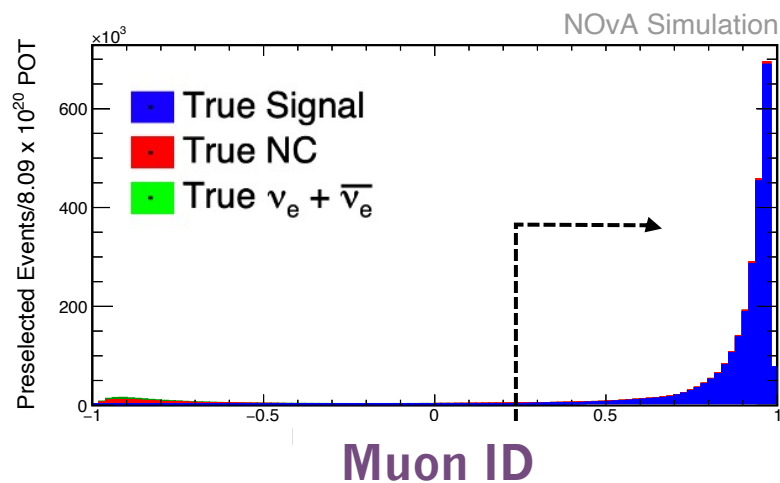
- ⚙  $\bar{\nu}_\mu$  CC  $\pi^0$
- ⚙  $\bar{\nu}_\mu$  CC inclusive
- ⚙  $\bar{\nu}_e$  CC inclusive
- ⚙  $\nu$  on e
- ⚙ Low hadronic energy

... and many more  
exclusive state topologies!

# $\nu_\mu$ CC Inclusive : Selection



dashed line : Fiducial Volume  
solid line : Containment Volume



➤ MuonID derived using a Boosted Decision Tree with muon  $dE/dx$  and scattering input observables.

➤ MuonID cut optimized on total cross section fractional uncertainty, provides a neat separation of signal from background.

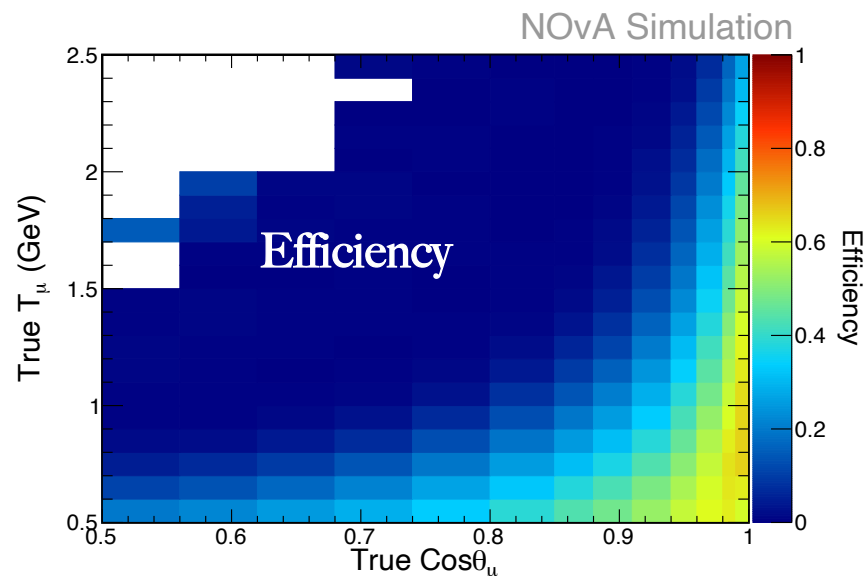
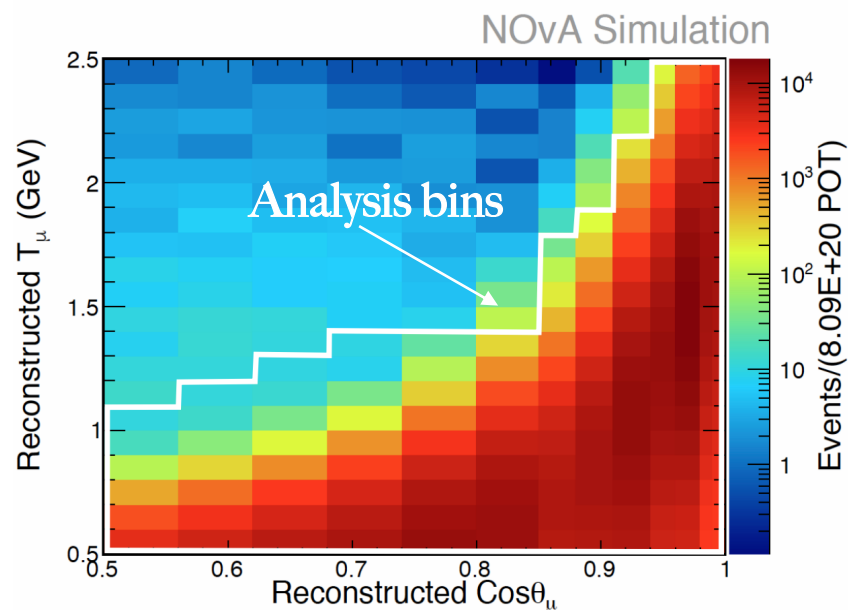
# $\nu_\mu$ CC Inclusive : Selection

More than **1M**  $\nu_\mu$  CC events are selected in **172 analysis bins**.

Resulting sample has **86% purity** and **~90% efficiency** with respect to the preselection cut.

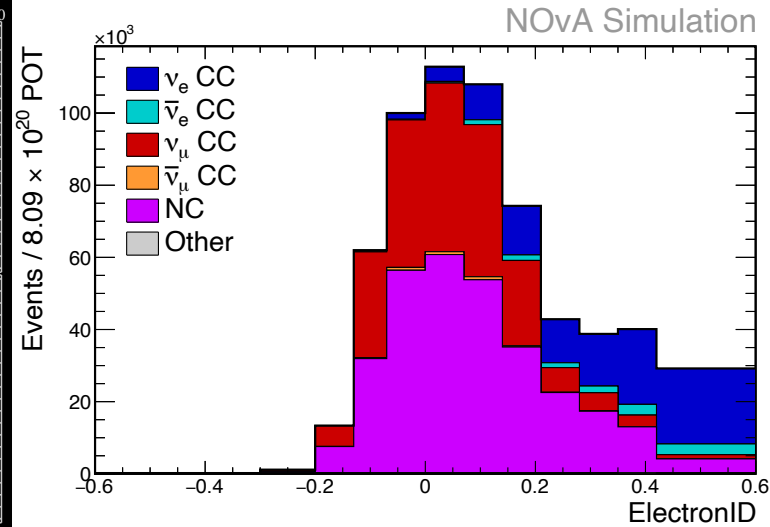
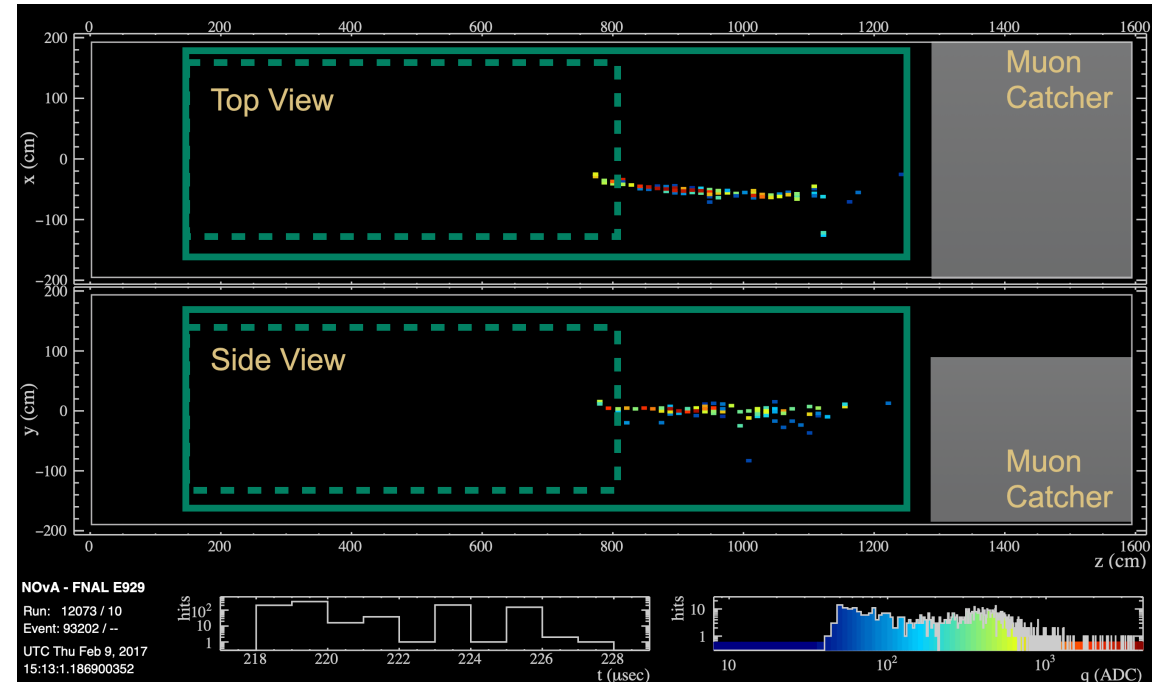
Efficiency drops with the increase in muon energy due to longitudinal containment.

Efficiency drops as the muon angle increases due to transverse containment.



# $\nu_e$ CC Inclusive : Selection

dashed line : Fiducial Volume  
solid line : Containment Volume



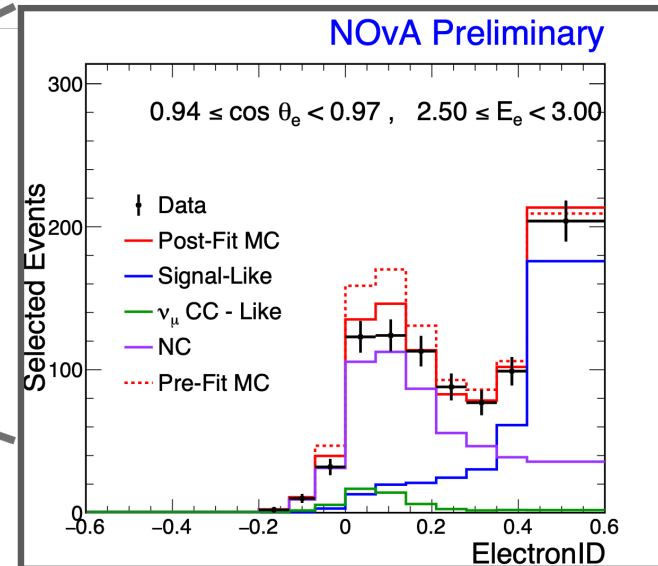
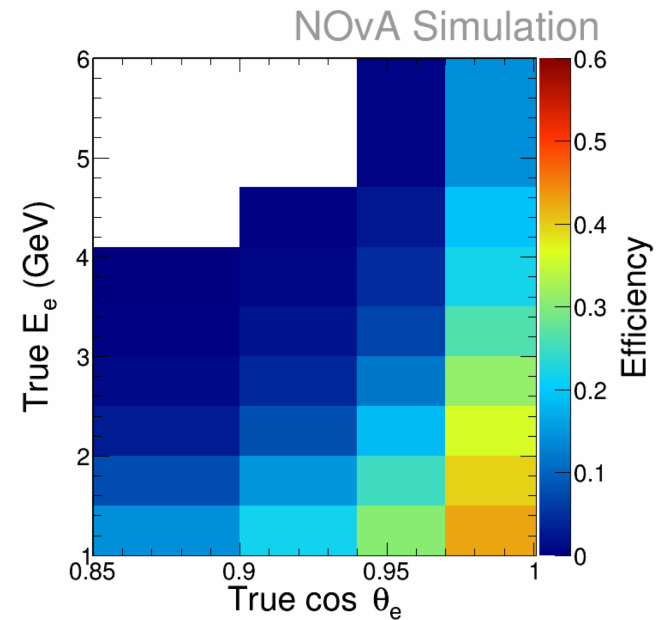
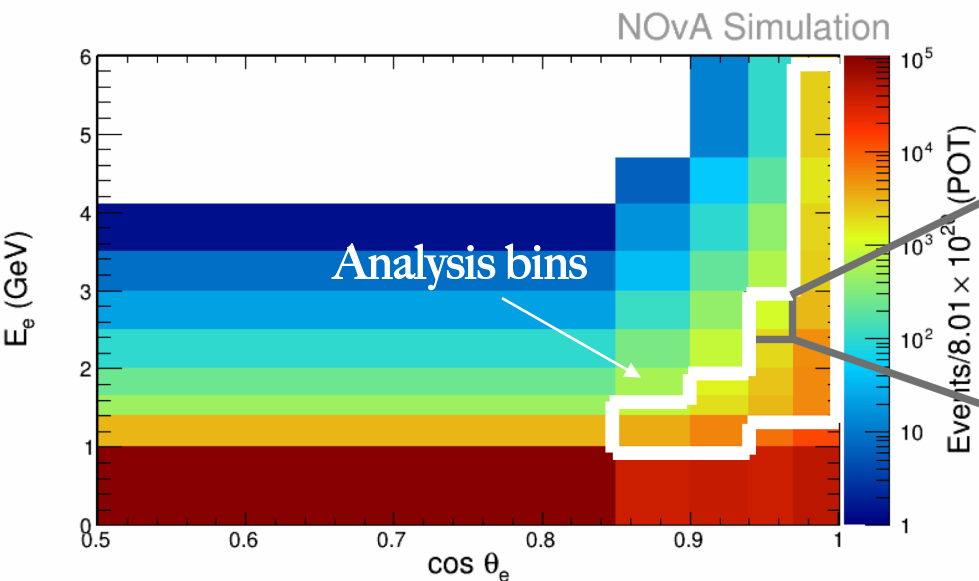
**ElectronID is not as strongly discriminating as MuonID.**

- Analysis uses a Boosted Decision Tree to distinguish electrons from other particles using output of a particle id from a deep convolutional network and EM shower properties.
- ElectronID is used to generate templates in which the fit is done.



# $\nu_e$ CC Inclusive : Selection

- Around **10k events** in the final selection.
- Background estimation in each electron kinematic bin is done via a template fit of the ElectronID distribution.
- Uncertainties in templates shape are accounted for using a covariance matrix.

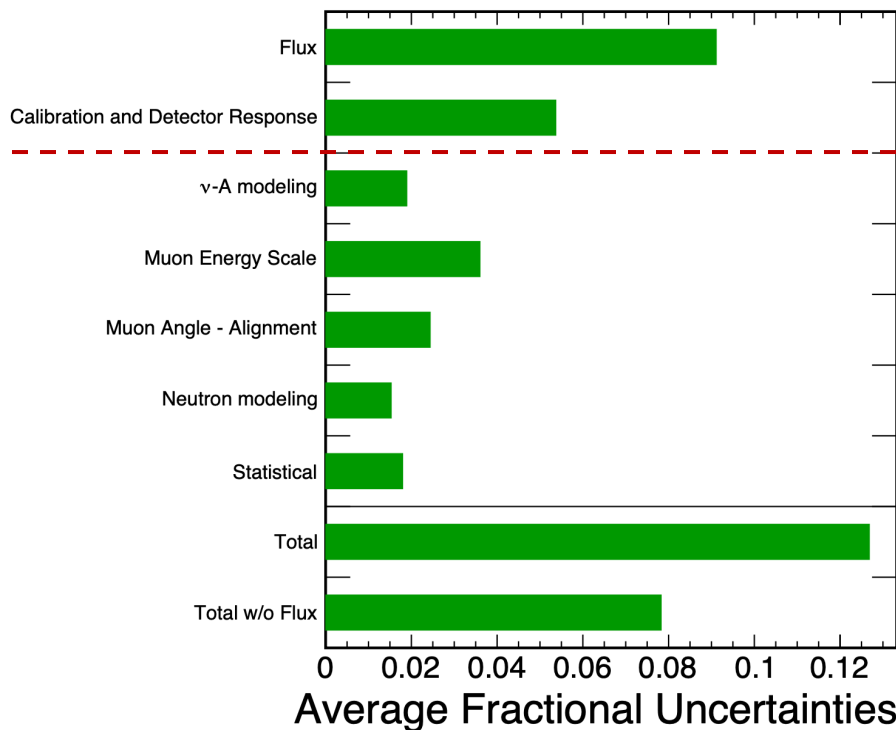


# Systematics

## $\nu_\mu$ CC Inclusive

- Flux contributes a normalization uncertainty of  $\sim 9\%$ .
- Interaction modeling uncertainties are sub-dominant.
- Measurements have typical total uncertainties around **12%** in each bin.

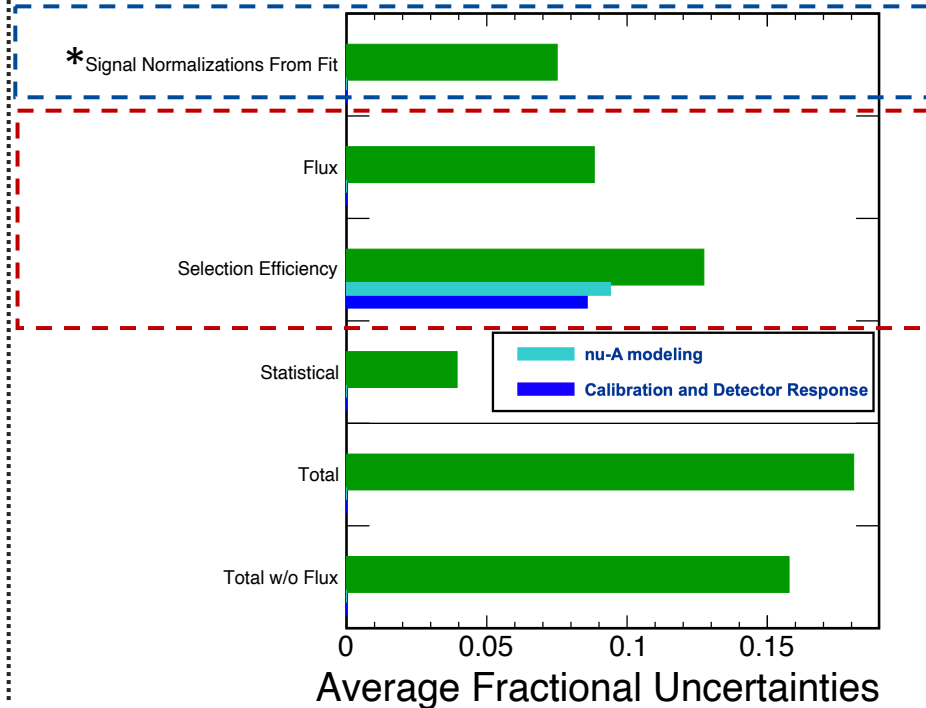
NOvA Preliminary



## $\nu_e$ CC Inclusive

- \*Uncertainty from the template fit.
- Interaction modeling uncertainties play a substantial role due to significant background events.
- Measurements have typical total uncertainties  **$\sim 15\%$  to  $20\%$**  in each bin

NOvA Preliminary



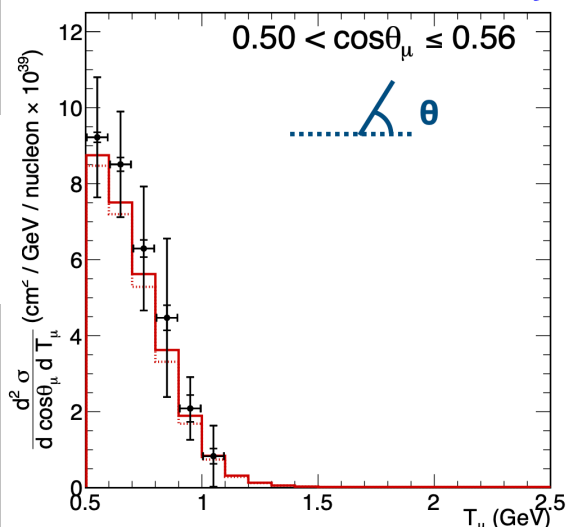
# $\nu_\mu$ CC Inclusive : Results

- Data (Stat. + Syst.)
- GENIE 2.12.2 - NOvA Tune
- ⋯ GENIE 2.12.2 - Untuned

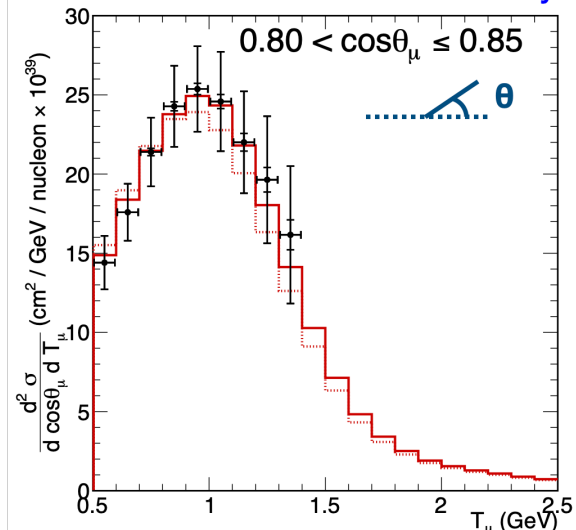
- At very forward angles (low  $Q^2$ ) the untuned GENIE 2 overshoots data significantly.

Generator	Total p-value
GENIE 2.12.2 - Tuned	0.93
GENIE 2.12.2 - Untuned	0.24

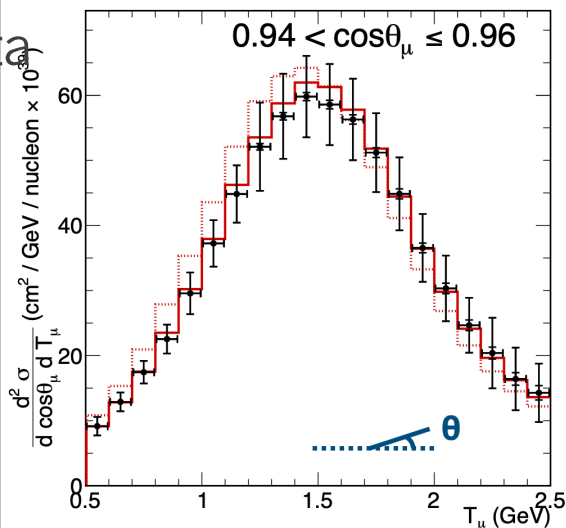
NOvA Preliminary



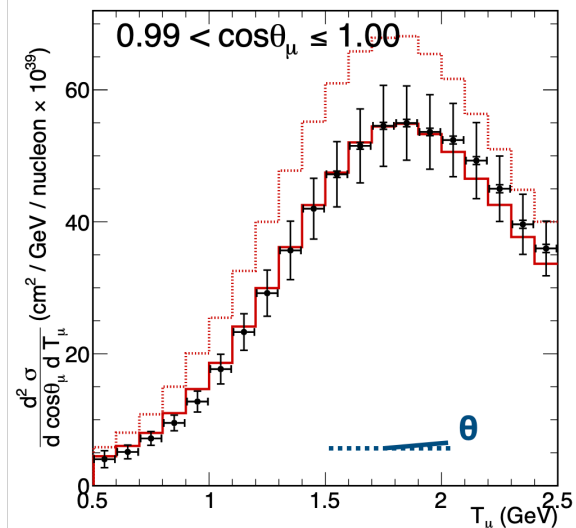
NOvA Preliminary



NOvA Preliminary



NOvA Preliminary

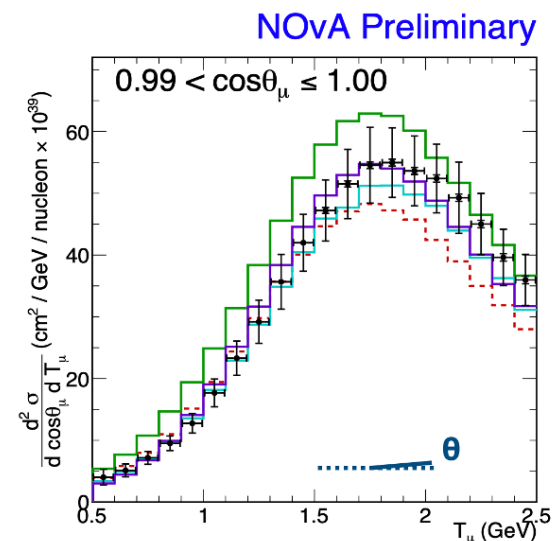
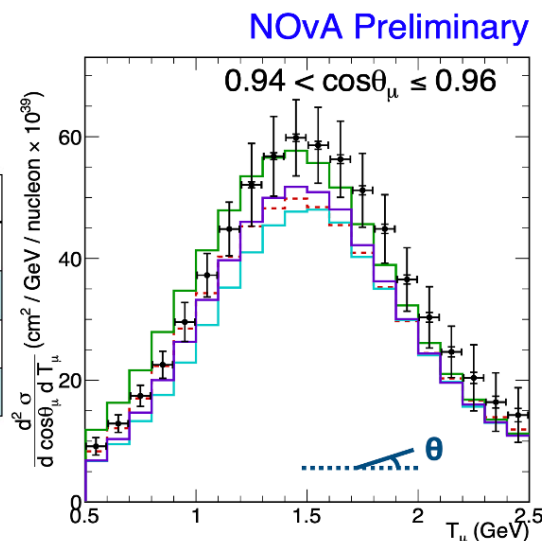
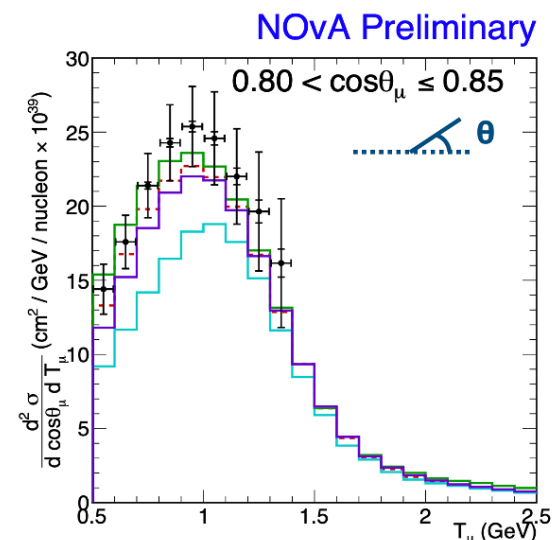
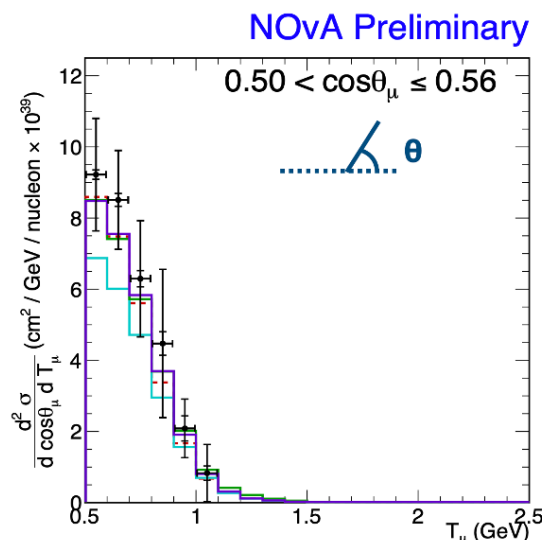


# $\nu_\mu$ CC Inclusive : Results

- Data (Stat. + Syst.)
- - - GENIE 3.00.06
- GiBUU 2019
- NEUT 5.4.0
- NuWro 2019

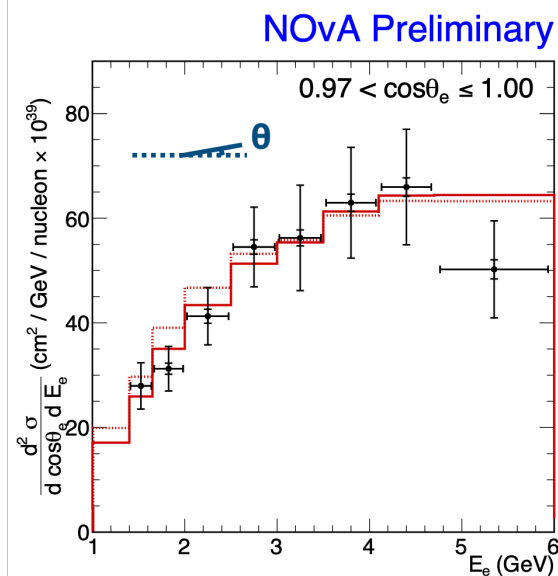
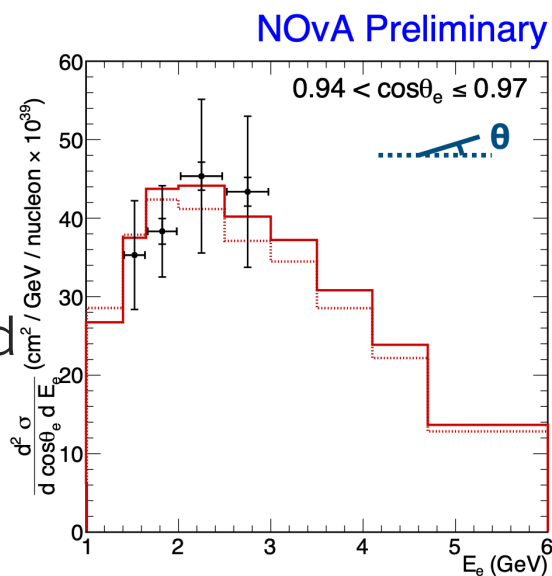
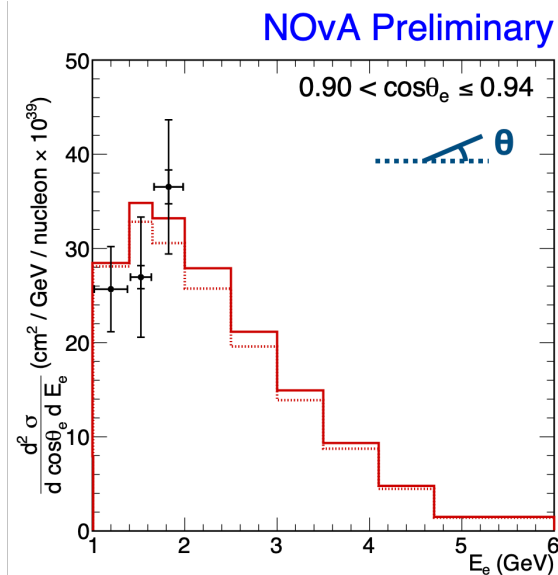
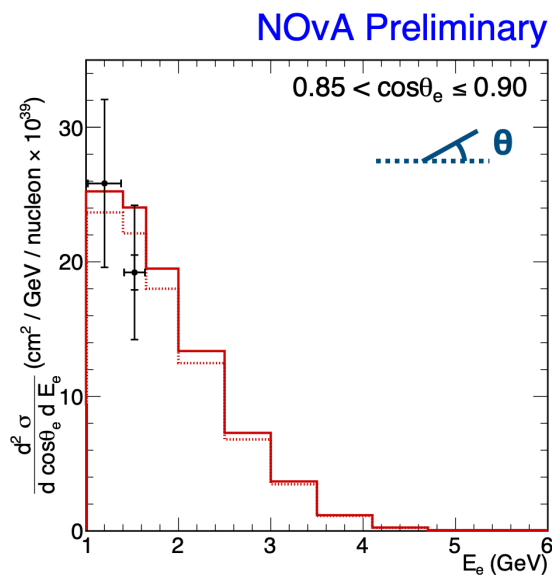
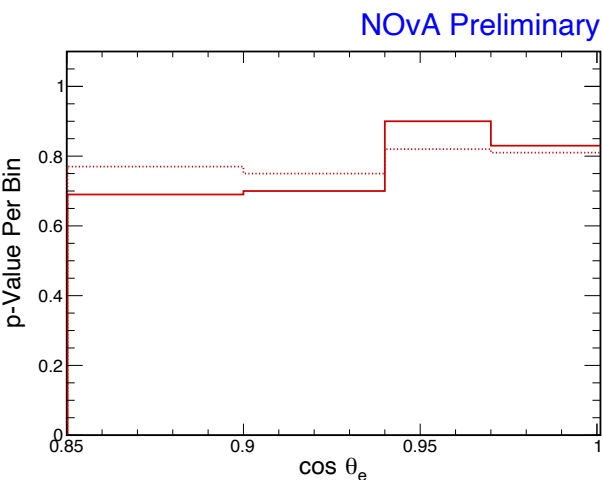
- Out of the box generator comparisons reproduces the shape of our data.

Generator	Total p-value
GENIE 3.00.06*	0.26
GiBUU 2019	0.03
NEUT 5.4.0	0.52
NuWro 2019	0.22



# $\nu_e$ CC Inclusive : Results

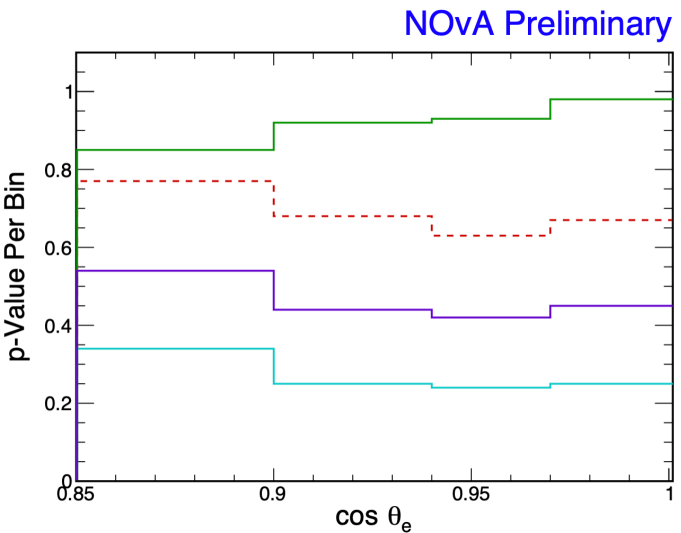
—●— Data (Stat. + Syst.)  
 — GENIE 2.12.2 - NOvA Tune  
 ····· GENIE 2.12.2 - Untuned



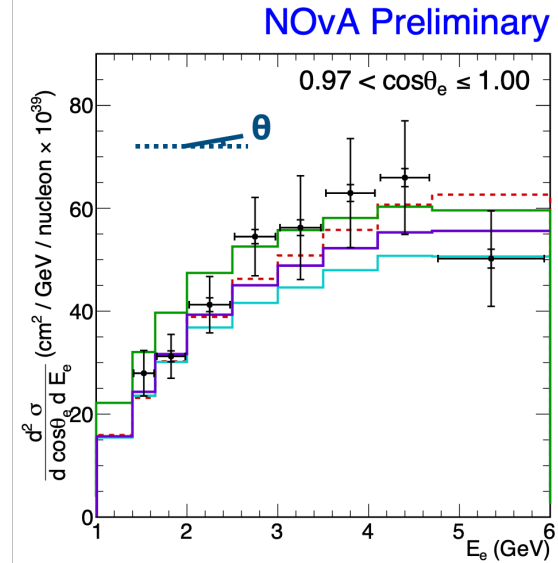
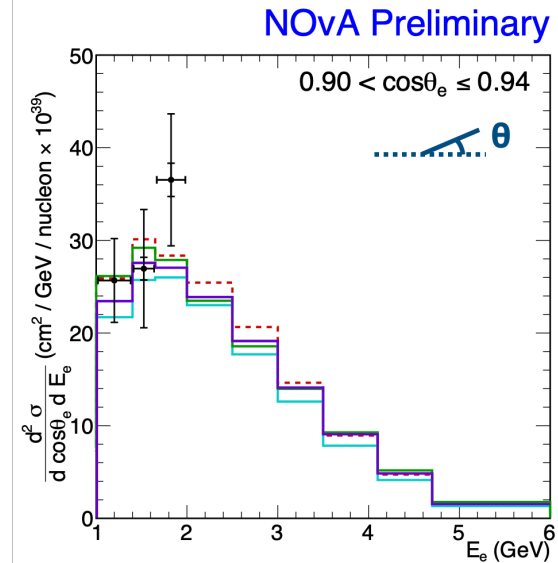
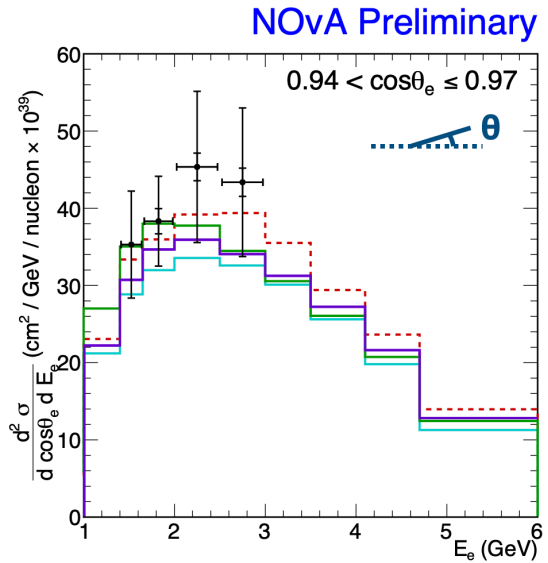
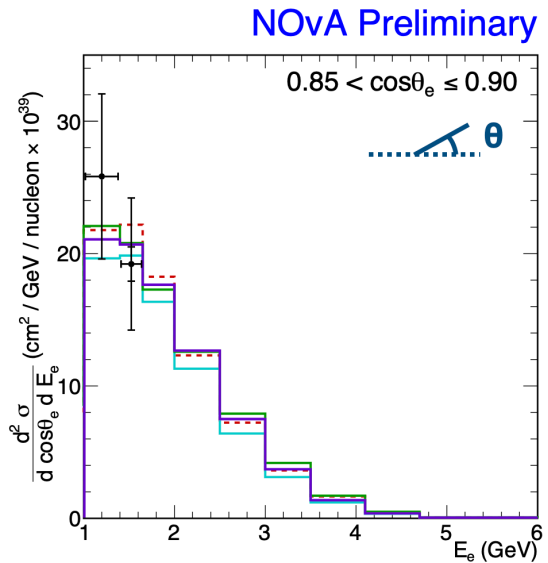
- Both tuned and untuned genie provide a good description of the observed data.

# $\nu_e$ CC Inclusive : Results

- Data (Stat. + Syst.)
- - - GENIE 3.00.06\*
- GiBUU 2019
- NEUT 5.4.0
- NuWro 2019



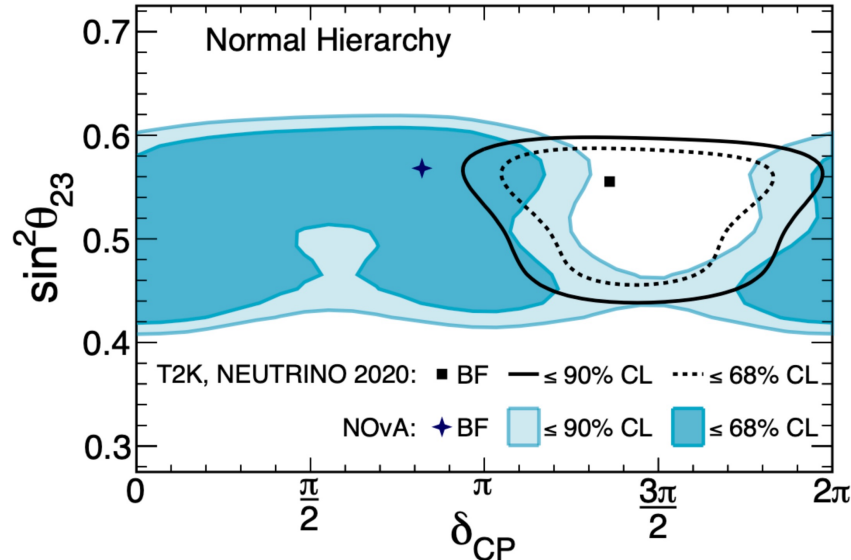
- GiBUU and NuWro provide a better description of data when their normalization is enhanced by  $\sim 15 - 20\%$ .
- Cross-section modeling systematics are large for this measurement.



# Summary & Outlook



# Summary

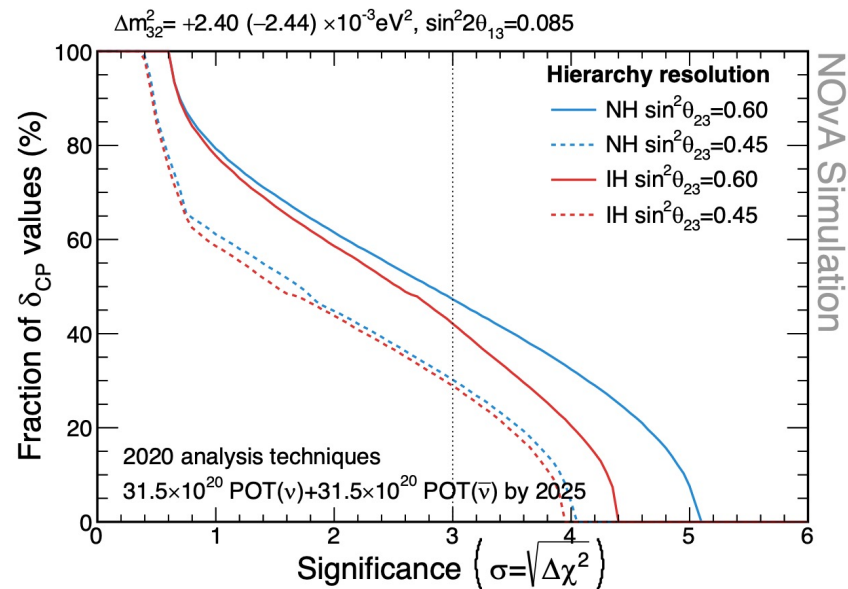


- NOvA data **disfavors asymmetry** between  $\nu_e$  and  $\bar{\nu}_e$  appearance while T2K sees a large asymmetry and their best fit is consistent with maximal CP violation.
- Precision measurements of  $\Delta m_{32}^2$  (**3%**) and  $\sin^2 \theta_{23}$  (**6%**).
- Presented the **first** measurement of the double-differential  $\nu_e$  CC inclusive cross section (15-20% uncertainties) and the double-differential  $\nu_\mu$  CC inclusive cross section with  $\sim 1$ M events (12% uncertainty).
- **Detector energy scale and response systematics** dominate systematic uncertainty budget for most measurements.



# Outlook

- NOvA is expected to run until 2026 and at least double its dataset for a total of 60-70 x 10<sup>20</sup> POT.
  - A 3  $\sigma$  determination of mass ordering is possible for 30-50% values of dCP.
- NOvA-T2K joint analysis** is converging rapidly and is on-track to report first measurements **this year!**
- Many more cross-section analyses with ND data that use the antineutrino mode beam and measure exclusive final state are coming out soon!



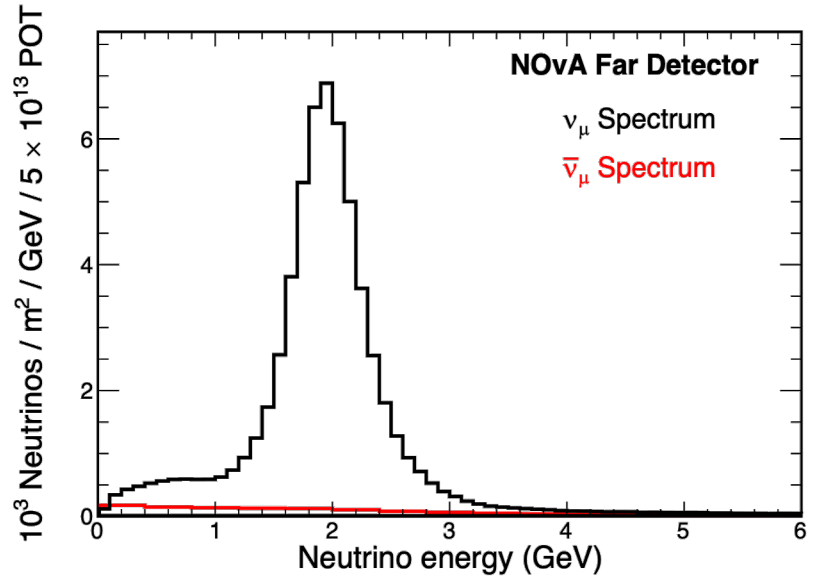
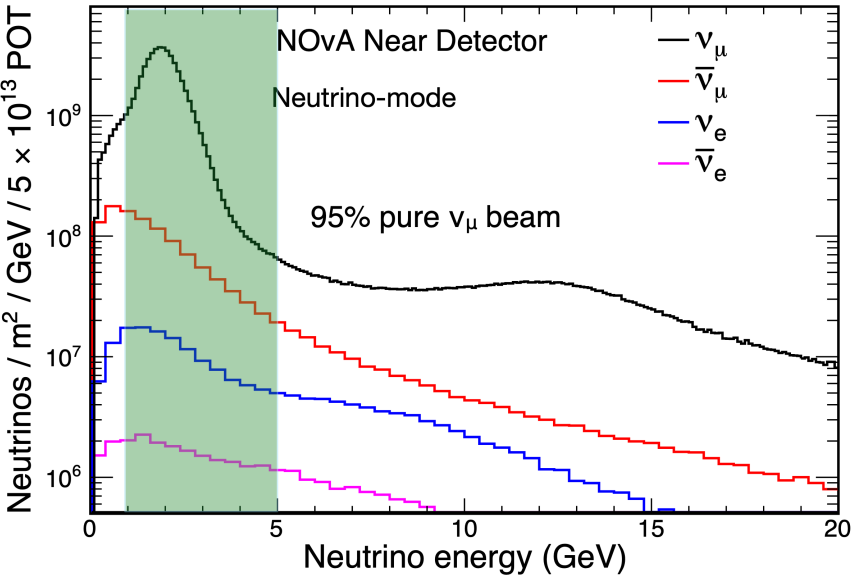
- NOvA Test Beam** is a miniature NOvA detector placed at Fermilab Test Beam Facility to study hadronic response and better understand the detector energy scale and response systematics.

backup

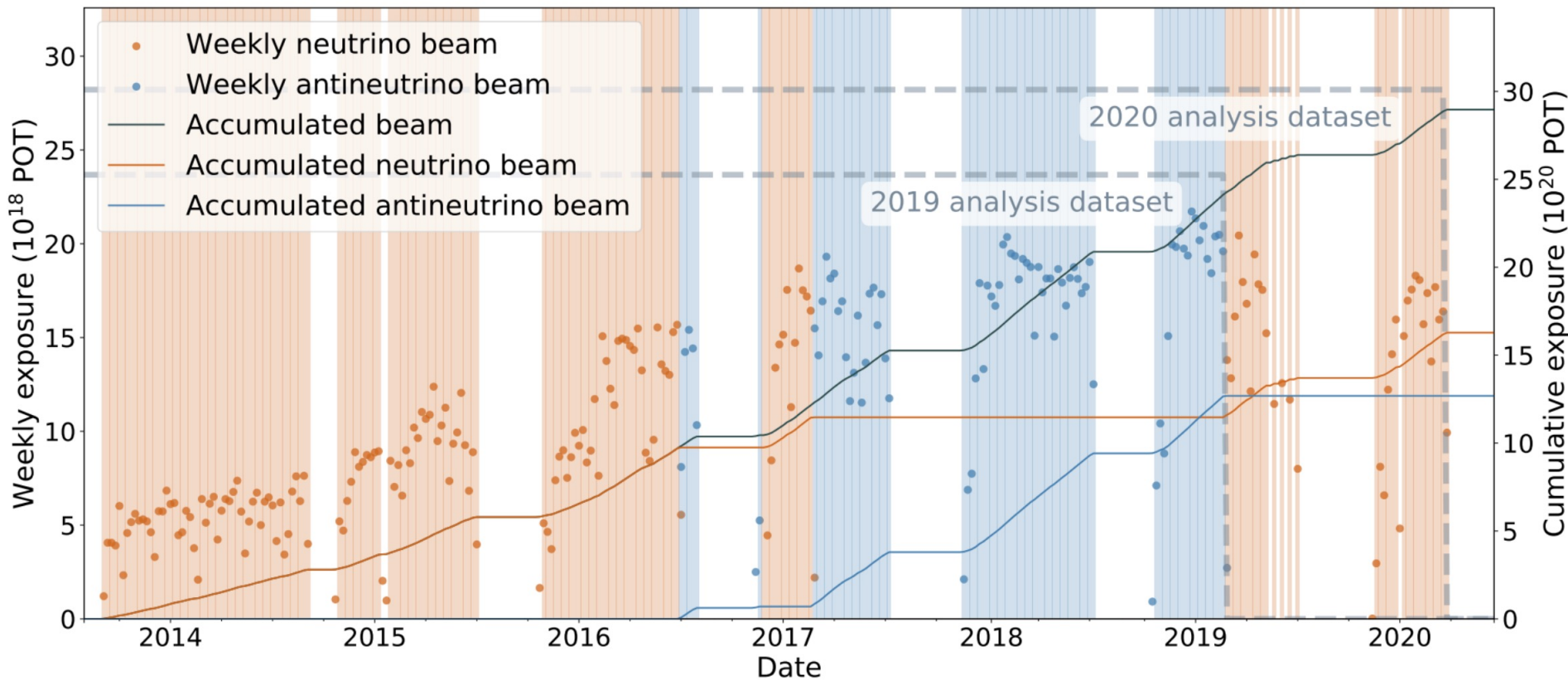


# NOvA Flux

NOvA Simulation

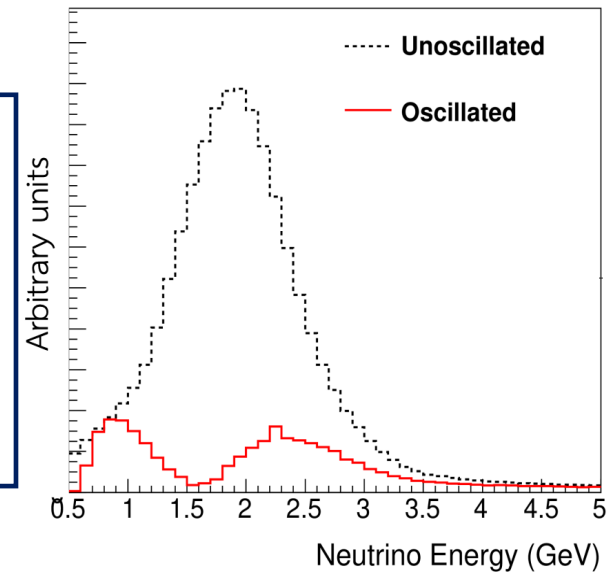
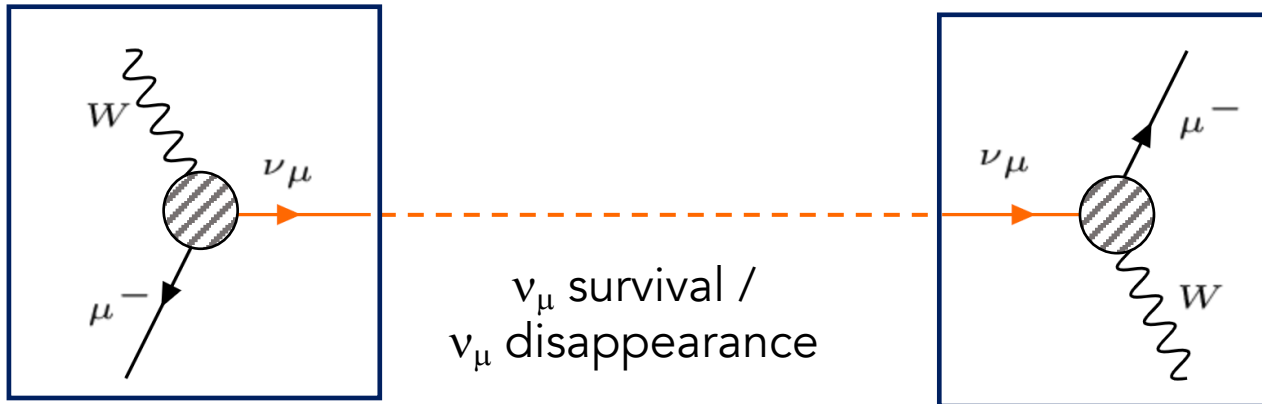


# NuMI Beamline



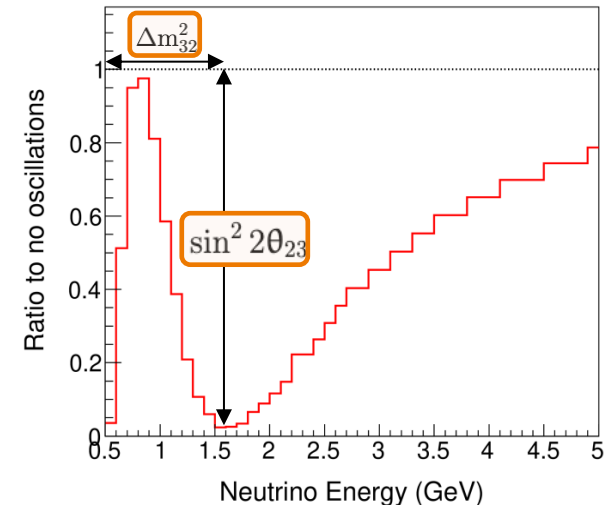
- The data analyzed for this analysis:
  - 13.6×10<sup>20</sup> POT neutrino beam
  - 12.5×10<sup>20</sup> POT anti-neutrino beam

# $\nu_\mu$ disappearance

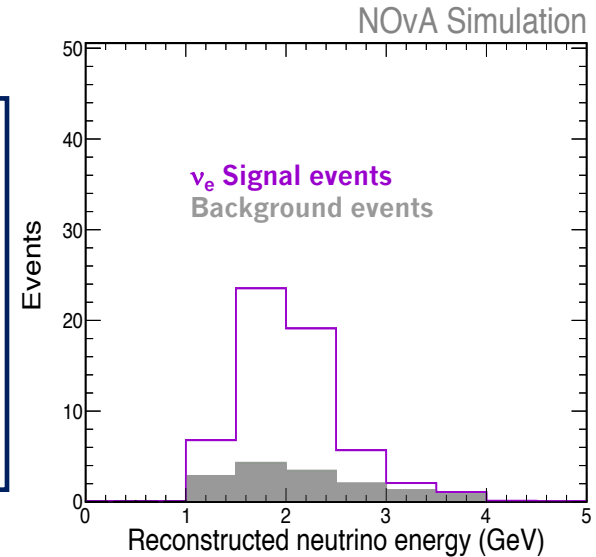
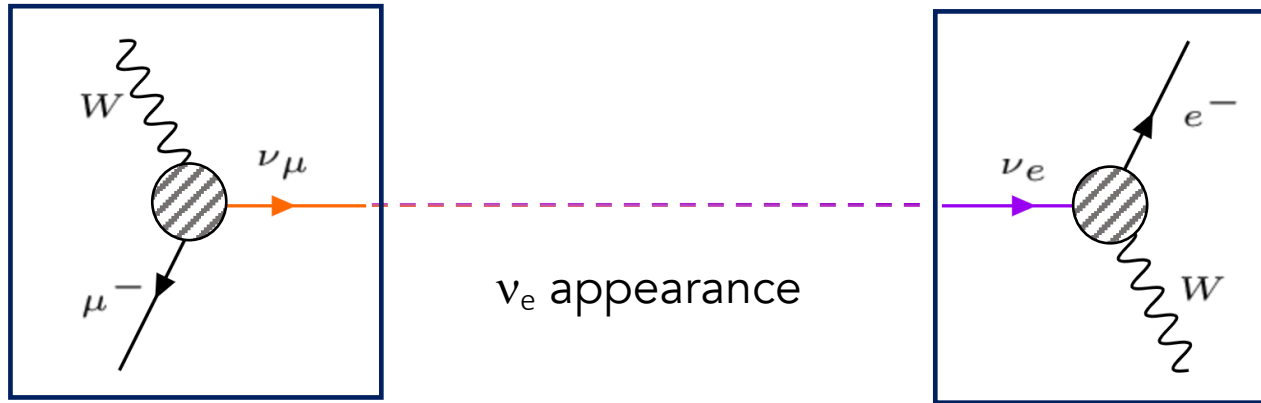


$$P\left(\nu_\mu^{(-)} \rightarrow \nu_\mu^{(-)}\right) \approx 1 - \sin^2 2\theta_{23} \sin^2\left(\frac{\Delta m_{32}^2 L}{4E}\right)$$

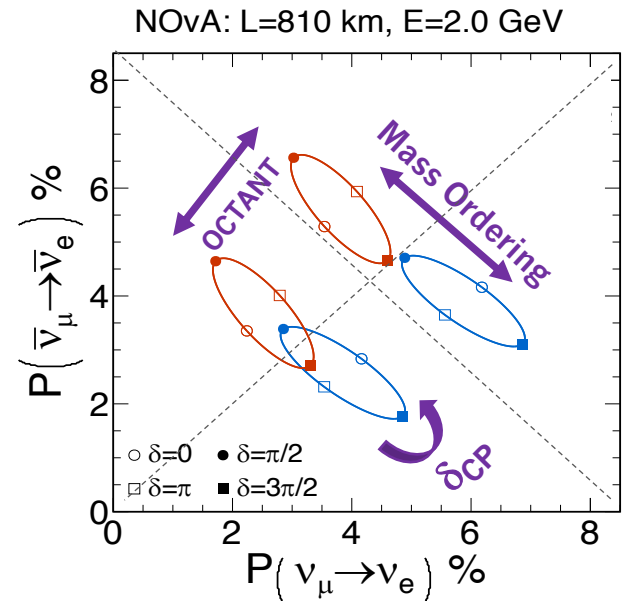
- Leading order dependence on  $\Delta m_{32}^2$  and  $\sin^2 \theta_{23}$
- Aims to answer the question of **maximal mixing in  $\theta_{23}$**



# LBL Oscillations : $\nu_e$ appearance



- Comparing the rate of  $\nu_e$  appearance with  $\bar{\nu}_e$  appearance provides a measurement of  $\delta_{CP}$  and mass ordering.
- $\delta_{CP}$  and mass ordering have inverse dependence on probability of  $\nu_e$  and  $\bar{\nu}_e$  appearance** while changing the octant is symmetric for the two beam modes.



# differential cross-section

$$\frac{d\sigma}{dx_i} = \frac{\sum_j U_{ij}^{-1} (N_j^{\text{sel}} \times P_j)}{\epsilon_i N_T \langle \Phi \rangle \Delta x_i}$$

Smearing

Selected  
events

Purity

Efficiency

# of targets

Integrated  
flux

$\nu_\mu$  CC Inclusive

$$\left( \frac{d^2\sigma}{d\cos\theta_\mu dT_\mu} \right)_i = \sum_k \left( \frac{\sum_j U_{ijk}^{-1} (N^{\text{sel}}(\cos\theta_\mu, T_\mu, E_{\text{avail}})_j P(\cos\theta_\mu, T_\mu, E_{\text{avail}})_j)}{N_t \Phi \epsilon(\cos\theta_\mu, T_\mu, E_{\text{avail}})_{ik} \Delta\cos\theta_{\mu_i} \Delta T_{\mu_i}} \right)$$

- Purity and efficiency corrections are applied in 3D space that consists of muon energy, direction and  $E_{\text{avail}}$ .
- Including  $E_{\text{avail}}$  in these corrections reduces potential model dependence on the final-state hadronic predictions.

# differential cross-section

$$\frac{d\sigma}{dx_i} = \frac{\sum_j U_{ij}^{-1} (N_j^{\text{sel}} \times P_j)}{\epsilon_i N_T \langle \Phi \rangle \Delta x_i}$$

Smearing

Selected  
events

Purity

Efficiency

# of targets

Integrated  
flux

$\nu_e$  CC Inclusive

$$\left( \frac{d^2\sigma}{d\cos\theta_e dE_e} \right)_i = \sum_j \left( \frac{U_{ij}^{-1} (N^{\text{sel}}(\cos\theta_e, E_e)_j - N^{\text{bkg}}(\cos\theta_e, E_e)_j)}{N_t \Phi \epsilon(\cos\theta_e, E_e)_{ik} \Delta\cos\theta_{e_i} \Delta E_{e_i}} \right)$$

- Analysis is performed in a 3D template space  $\Rightarrow$   $(\cos\theta_e, E_e, \text{ElectronID})$





# Neutrino Interactions : GENIE

- Using the latest GENIE v3.0.6
- Built a Custom-Model-Configuration (CMC) from the available collections of model
- 'Theory-driven' models with tune to external free-nucleon data were chosen as NOvA's nominal interaction model

## GENIE N1810j\_0211a \*

QE	MEC/2p2h	RES	DIS	FSI
<ul style="list-style-type: none"> <li>Valencia 1p1h</li> <li>Z-expansion axial form factor</li> </ul>	<ul style="list-style-type: none"> <li>Valencia MEC</li> </ul>	<ul style="list-style-type: none"> <li>Berger-Sehgal</li> </ul>	<ul style="list-style-type: none"> <li>Bodek-Yang</li> </ul>	<ul style="list-style-type: none"> <li>hN Semi Classical Cascade</li> </ul>



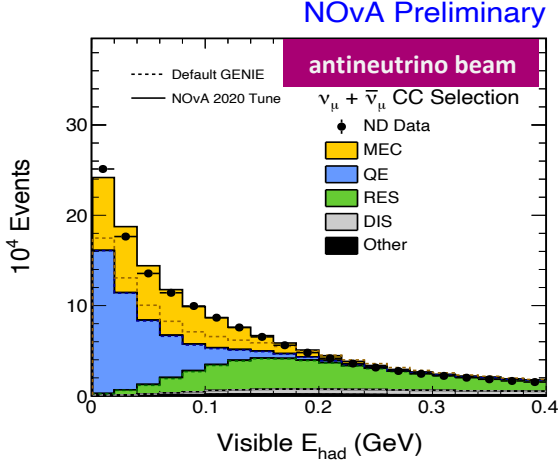
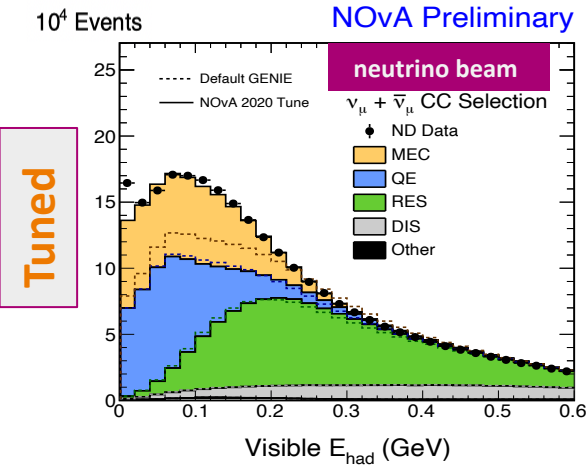
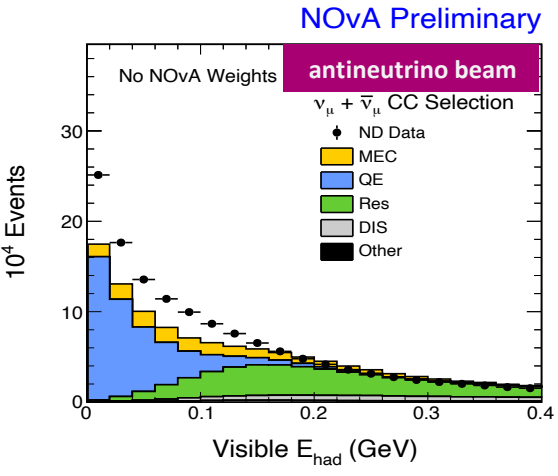
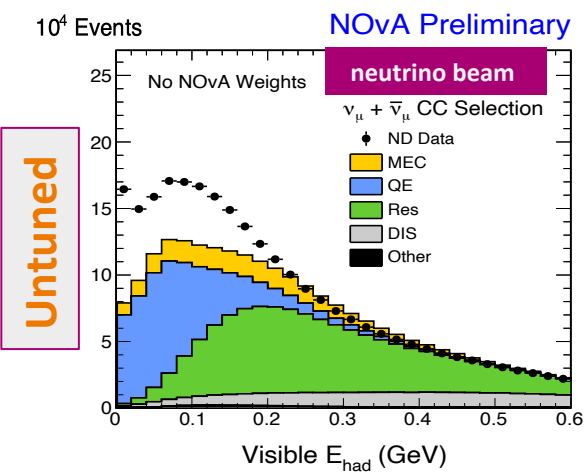
\*We call our "tune" N1810j\_0211a. It is built by starting with G1810b\_0211a and substituting the Z-expansion QE axial form factor for the dipole one. This combination was not available in the 3.0.6 release, but it may be available in future versions

# NOvA ND Tune

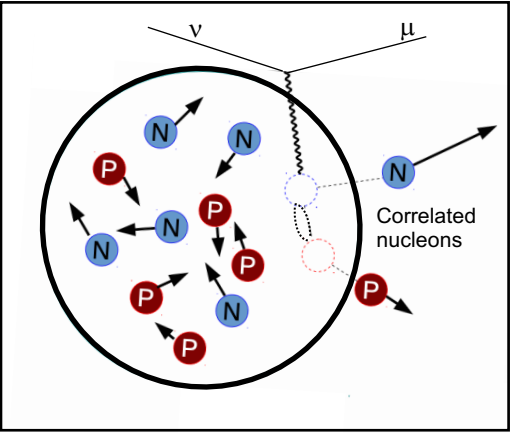
- We use **GENIE**(v3.0.6) for neutrino interaction generation.
- Nuclear effects are still not well-modeled. Out-of-the-box GENIE does not describe ND data well.

- Tuned GENIE by varying the MEC and FSI components.

- Any remaining differences between data and MC are covered by systematic uncertainty band and are extrapolated to FD Simulation as ND Data Constraints.



# 2p2h Tune



- “2p2h” or MEC (meson exchange current) interaction occurs when a neutrino interacts with a correlated pair of nucleons.
- NOvA tunes the MEC component of interaction simulation by doing a double gaussian fit to its data.

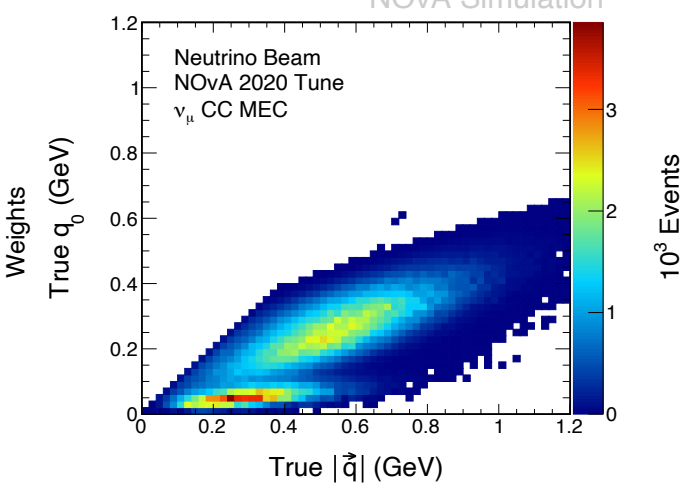
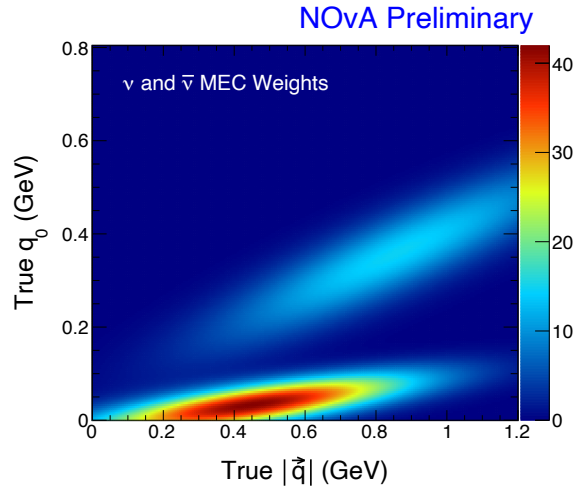
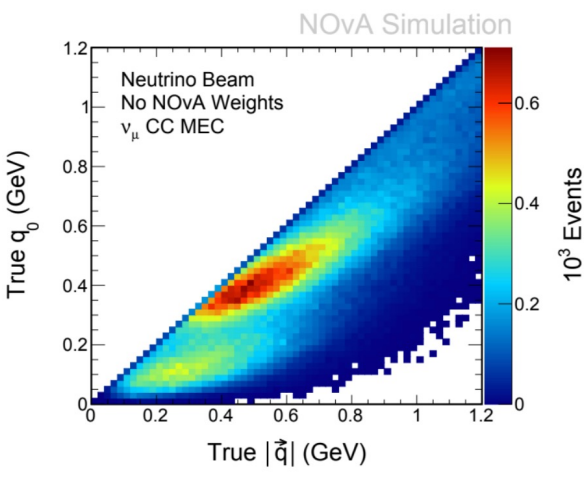
**Valencia 2p2h**

**x**

**MEC Weights**

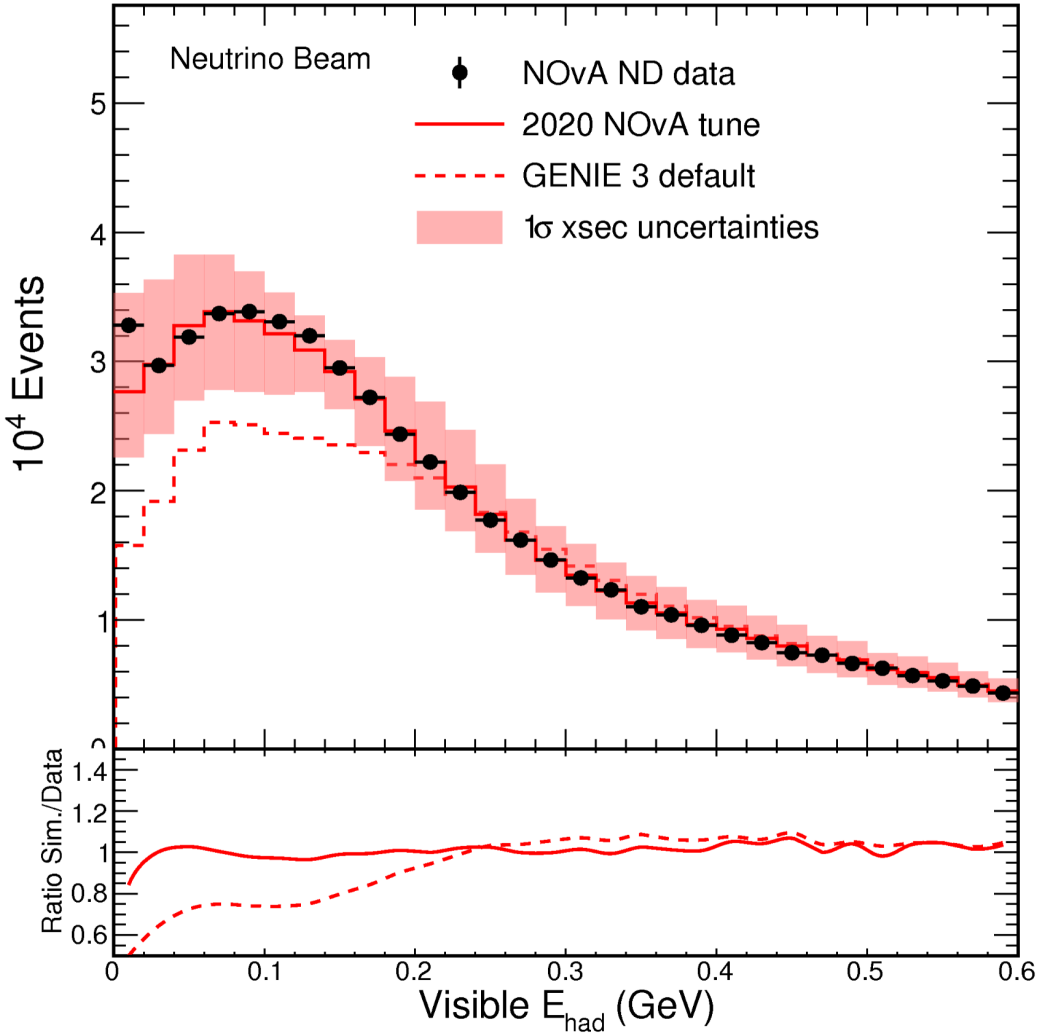
**=**

**NOvA Tune**



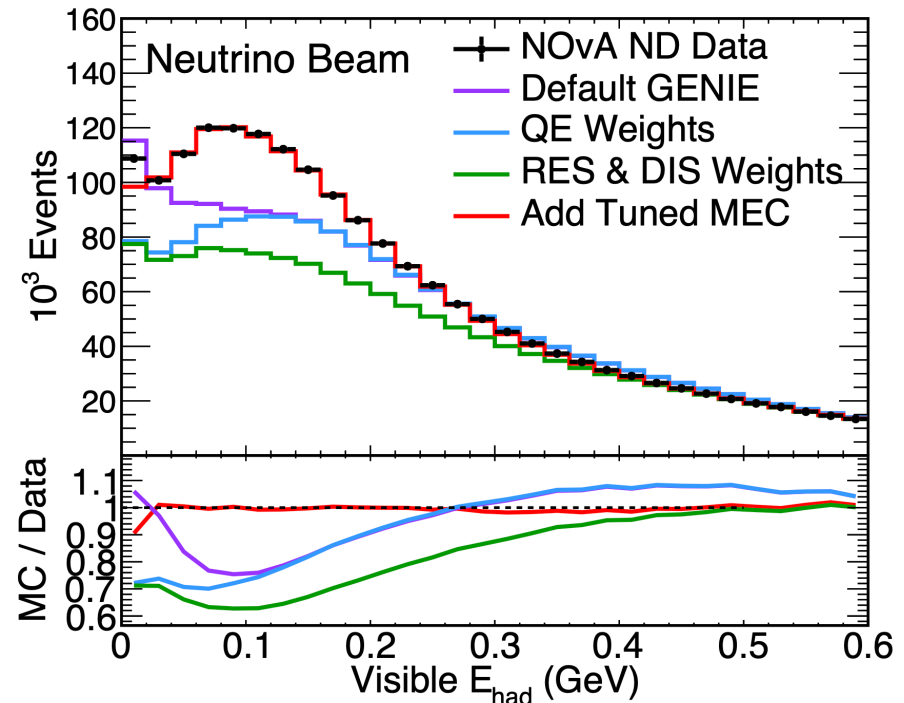
# NOvA 2020 Tune & Systematics

NOvA preliminary



# NOvA 2019 Tune

- This analysis uses the same tune as the NOvA 2019 oscillation results
- **From external theory:**
  - Valencia RPA model of nuclear charge screening applied to QE.
  - Same model is applied to RES interactions
- **From NOvA ND Data:**
  - 10% increase in non-resonant inelastic scattering at high  $W$  ( $> 1.7 \text{ GeV}^2/c^2$ )
  - Add in MEC interactions
    - Start from Empirical MEC and tune to NOvA data in bins of momentum transfer
- Details of tune procedure are in [arXiv:2006.08727](https://arxiv.org/abs/2006.08727).

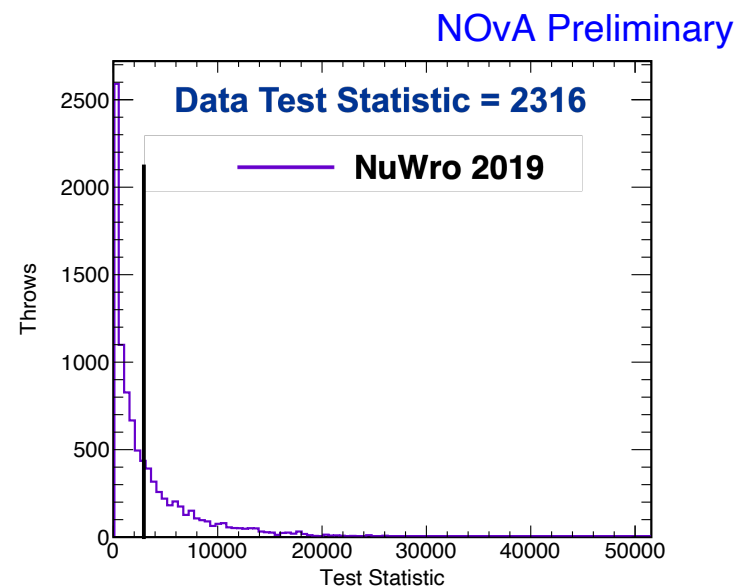
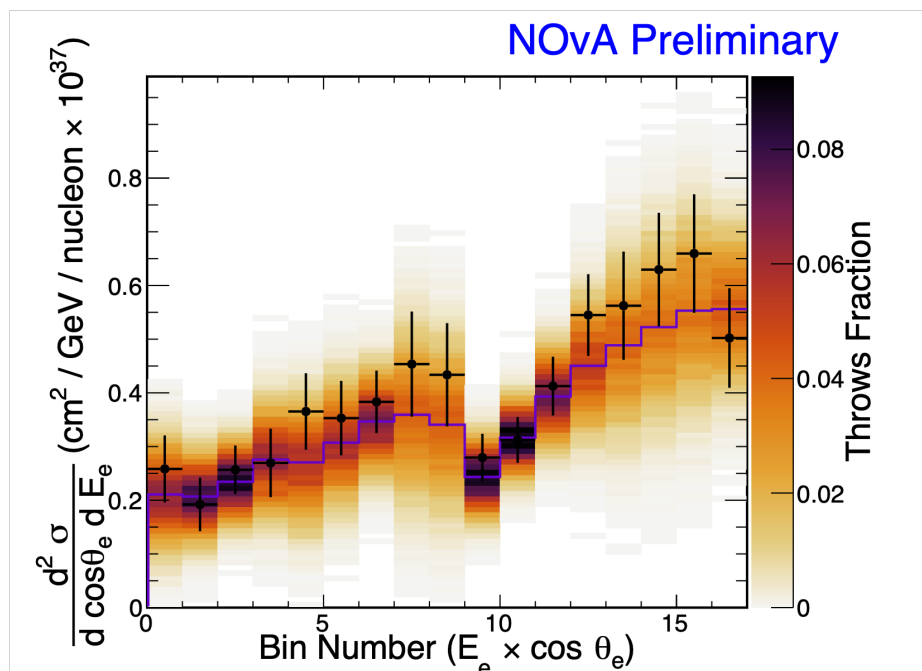


# Generator comparisons

	QE/MEC Initial State	QE	MEC	Res	DIS	FSI
<b>GENIE v2.12.2</b>	RFG	L-S	Empirical (NOvA tune)	R-S	PYTHIA 6	hA
<b>GENIE v3.00.06</b>	LFG	Valencia (Nieves, et al)	Valencia (Nieves, et al)	B-S	PYTHIA 6	hN
<b>NEUT 5.4.0</b>	LFG	Valencia (Nieves, et al)	Valencia (Nieves, et al)	B-S	PYTHIA 5	Oset (low mom. pions) + ext. data
<b>NuWro 2019</b>	LFG	L-S + RPA	Valencia (Nieves, et al)	NuWro	PYTHIA 6	Oset (pions) + NuWro (nucleons)
<b>GiBUU 2019</b>	LFG	GiBUU Model				BUU equations

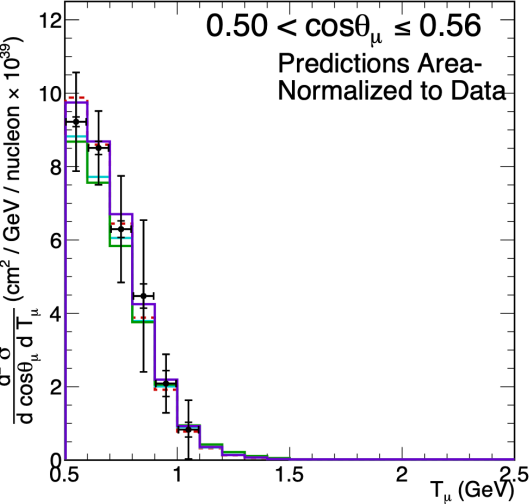
# Generator comparisons

- Use the covariance matrix to calculate 50,000 throws from each generator prediction (RooFit).
- Compare test statistics of throws to data to find p-values.

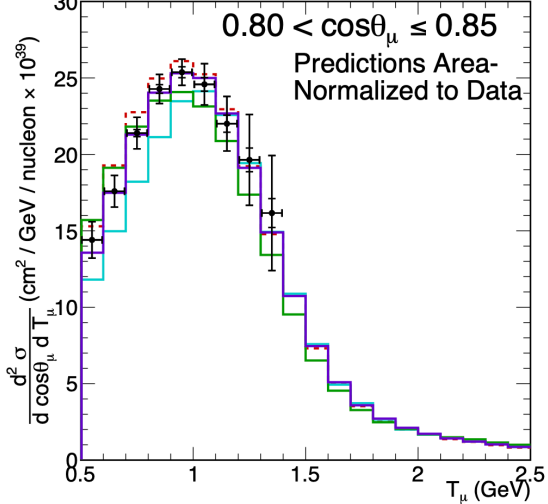


# Shape only comparisons : $\nu_{\mu}$ CC

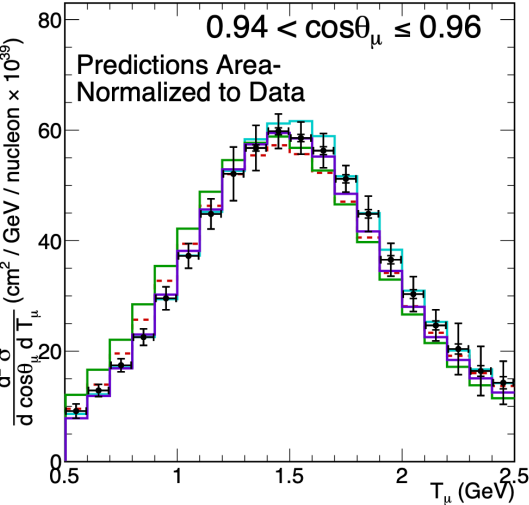
NOvA Preliminary



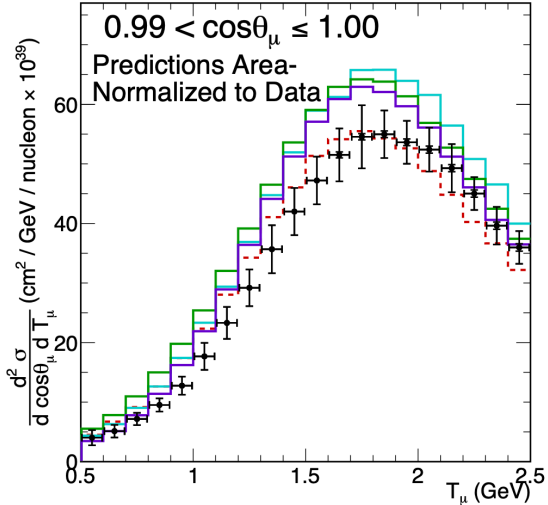
NOvA Preliminary



NOvA Preliminary



NOvA Preliminary

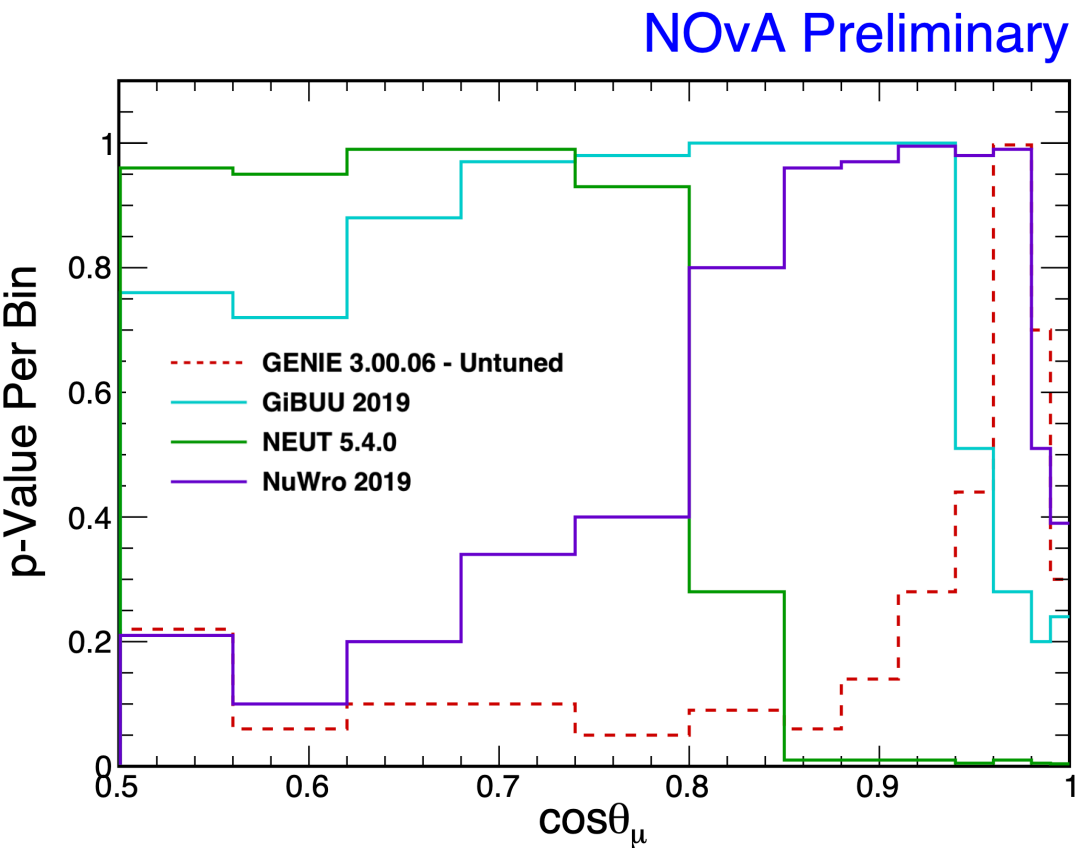


—●— **Data (Stat. + Syst.)**  
- - - **GENIE 3.00.06**  
— **GiBUU 2019**  
— **NEUT 5.4.0**  
— **NuWro 2019**

Generator	Total p-value	Norm.
<b>GENIE 3.00.06*</b>	0.31	1.15
<b>GiBUU 2019</b>	0.38	1.28
<b>NEUT 5.4.0</b>	0.004	1.02
<b>NuWro 2019</b>	0.54	1.15



# Shape only p-values : $\nu_\mu$ CC



- Shapes from GiBUU agree much better with a normalization increase of ~28%.
- Generally the shapes predicted by the generators still need improvement in the most forward angles.

Generator	Total p-value	Norm.
<b>GENIE 3.00.06*</b>	0.31	1.15
<b>GiBUU 2019</b>	0.38	1.28
<b>NEUT 5.4.0</b>	0.004	1.02
<b>NuWro 2019</b>	0.54	1.15

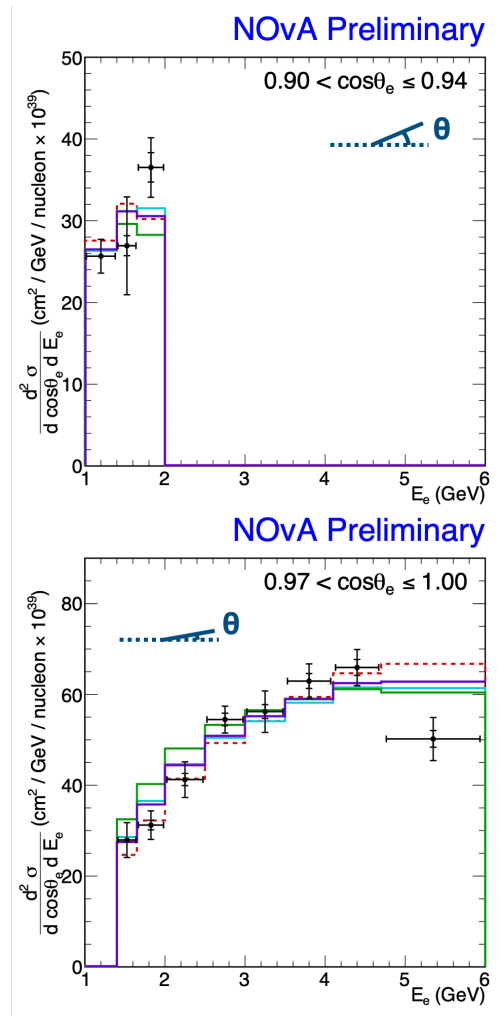
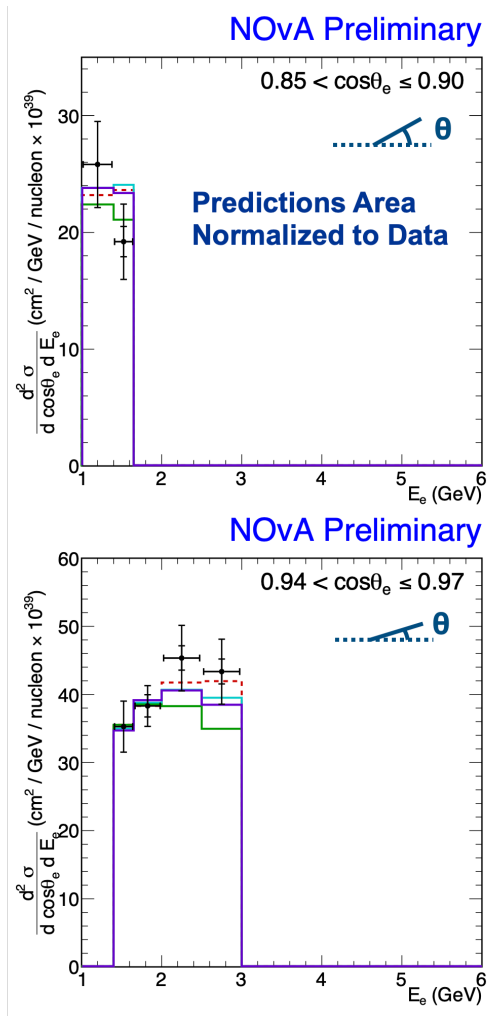
# Shape only comparisons : $\nu_e$ CC

- Data (Stat. + Syst.)
- - - GENIE 3.00.06 - Untuned
- GiBUU 2019
- NEUT 5.4.0
- NuWro 2019

Area normalized predictions to Data

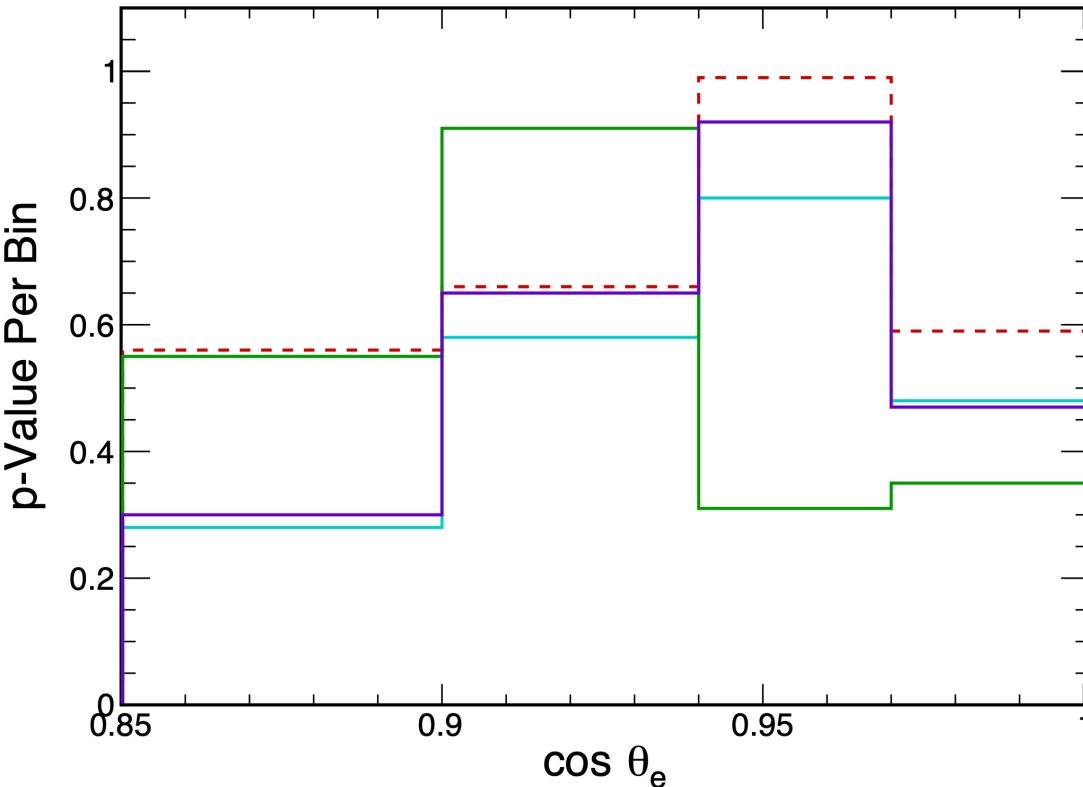
GiBUU and NuWro are shifted up 20% and 14%, respectively

p-values ranging from 0.4 to 0.95



# Shape only p-values : $\nu_e$ CC

NOvA Preliminary



- GENIE 3.00.06 - Untuned
- GiBUU 2019
- NEUT 5.4.0
- NuWro 2019

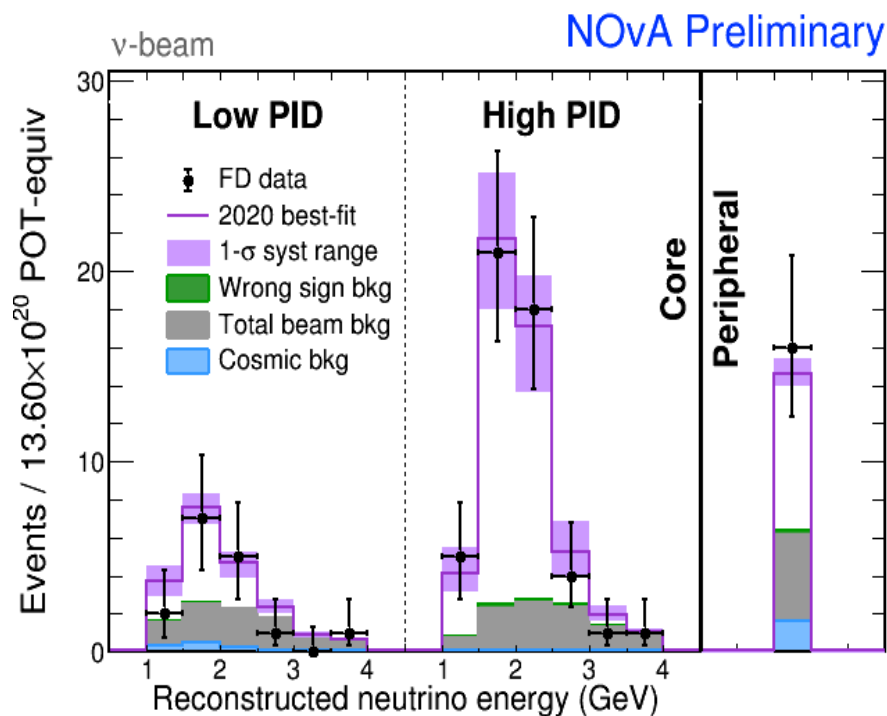
- Area normalized predictions to Data
- GiBUU and NuWro are shifted up 20% and 14%, respectively
- p-values ranging from 0.4 to 0.95

# Test Beam

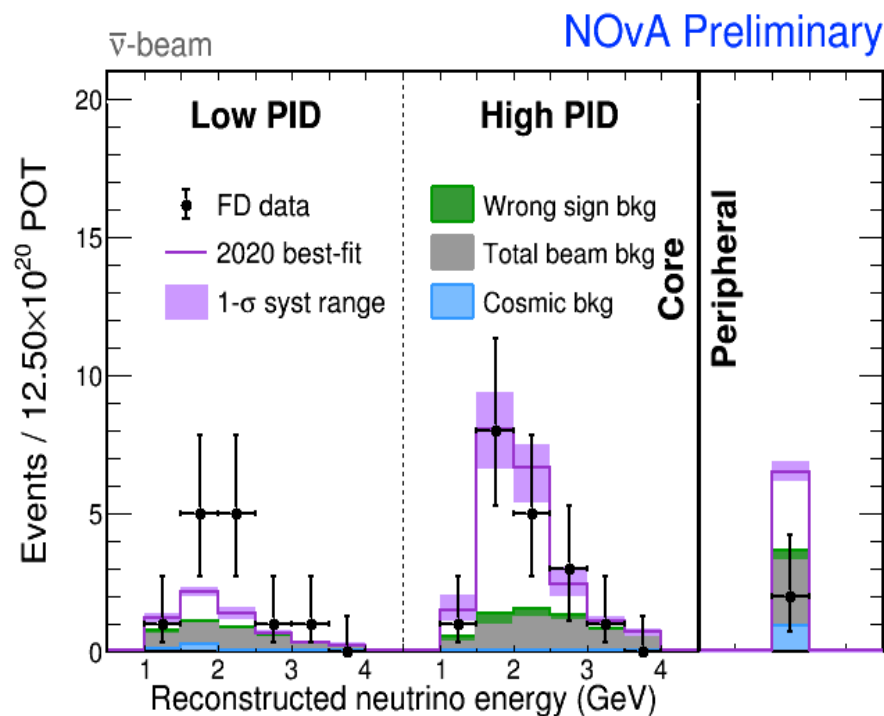
- Our systematics are dominated by detector energy scale and response. Data-MC disagree at 5% level for proton candidates.
- Feature : NOvA Test Beam at Fermilab Test Beam Facility to study hadronic response.
- A tertiary beam is created from 64 GeV secondary beam particles to study charged particles in 0.2 – 2 GeV range relevant to NOvA.
- **Currently taking data!**



# FD Data Samples : $\nu_e$ appearance



Observed	: 82
Total Predicted	: 85.8
Total Background	: 26.8
Wrong-Sign	: 1.0
Beam Bkg	: 22.7
Cosmics	: 3.1



Observed	: 33
Total Predicted	: 33.2
Total Background	: 14.0
Wrong-Sign	: 2.3
Beam Bkg	: 1.6
Cosmics	: 3.1

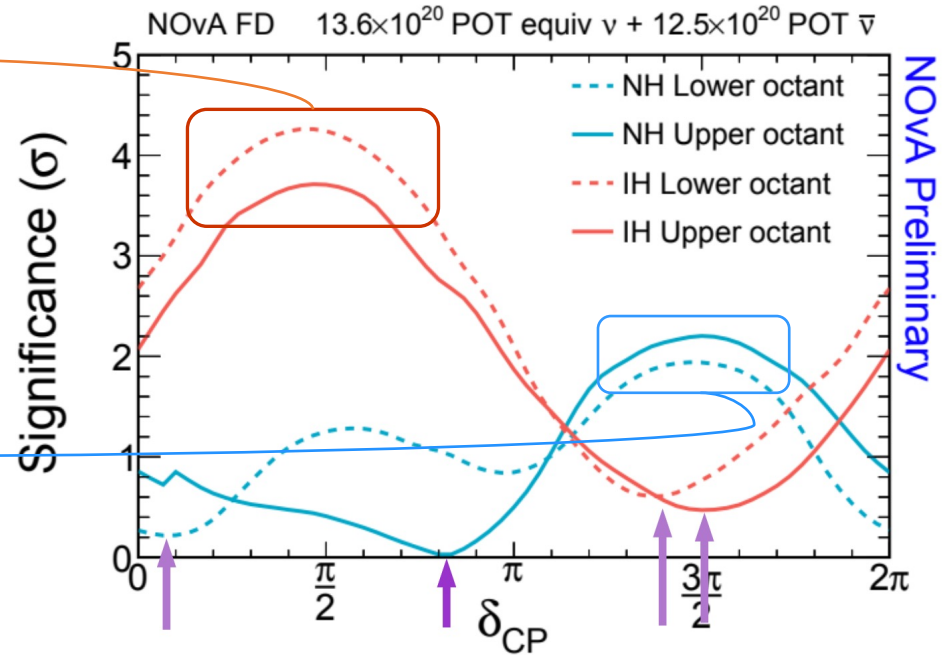
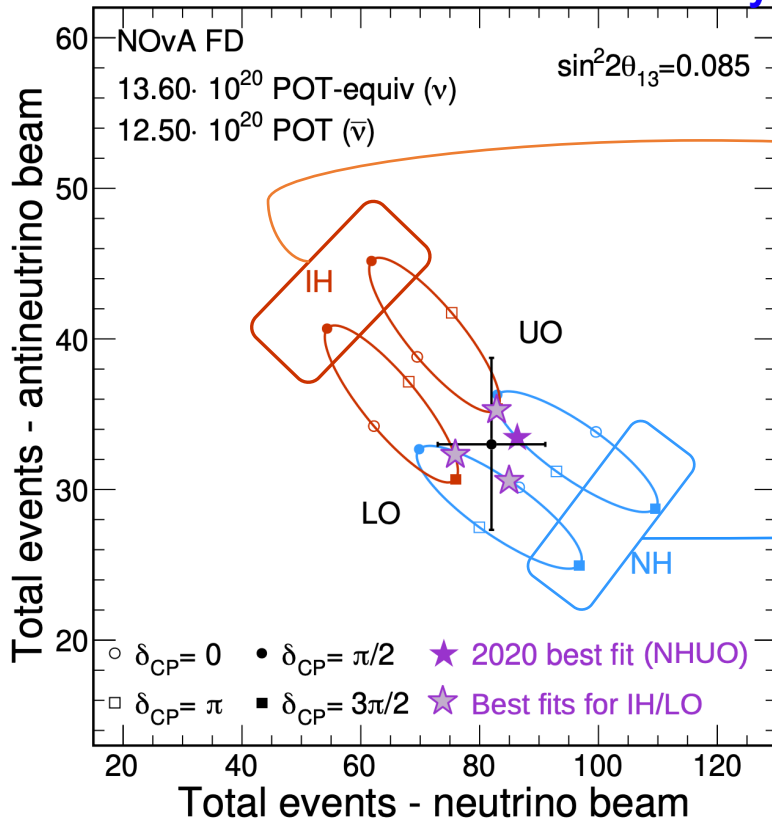
# FD Data Samples : $\nu_{\mu}$ disappearance

## Muon Neutrino Disappearance

	Neutrino	Antineutrino
Observed	211	105
No Disappearance	1156	488
Best Fit Expectation	222	105
Background	8.2	2.1

# Extracting Oscillation Parameters

NOvA Preliminary



➤ NOvA data disfavors strong asymmetry in rates of  $\nu_e$  and  $\bar{\nu}_e$  appearance.