
Matter-Wave Turbulence

Beyond kinetic scaling



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Center for
Quantum
Dynamics



Thanks & credits to...

...my work group in Heidelberg:

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Philipp Struck (→ Konstanz)
Kristan Temme (→ Vienna)

...my collaborators

Jürgen Berges (Darmstadt) • ...



RUPRECHT-KARLS-
UNIVERSITÄT
HEIDELBERG

LGFG BaWue

DAAD

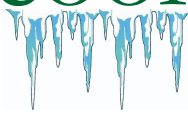
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ULTRACool...



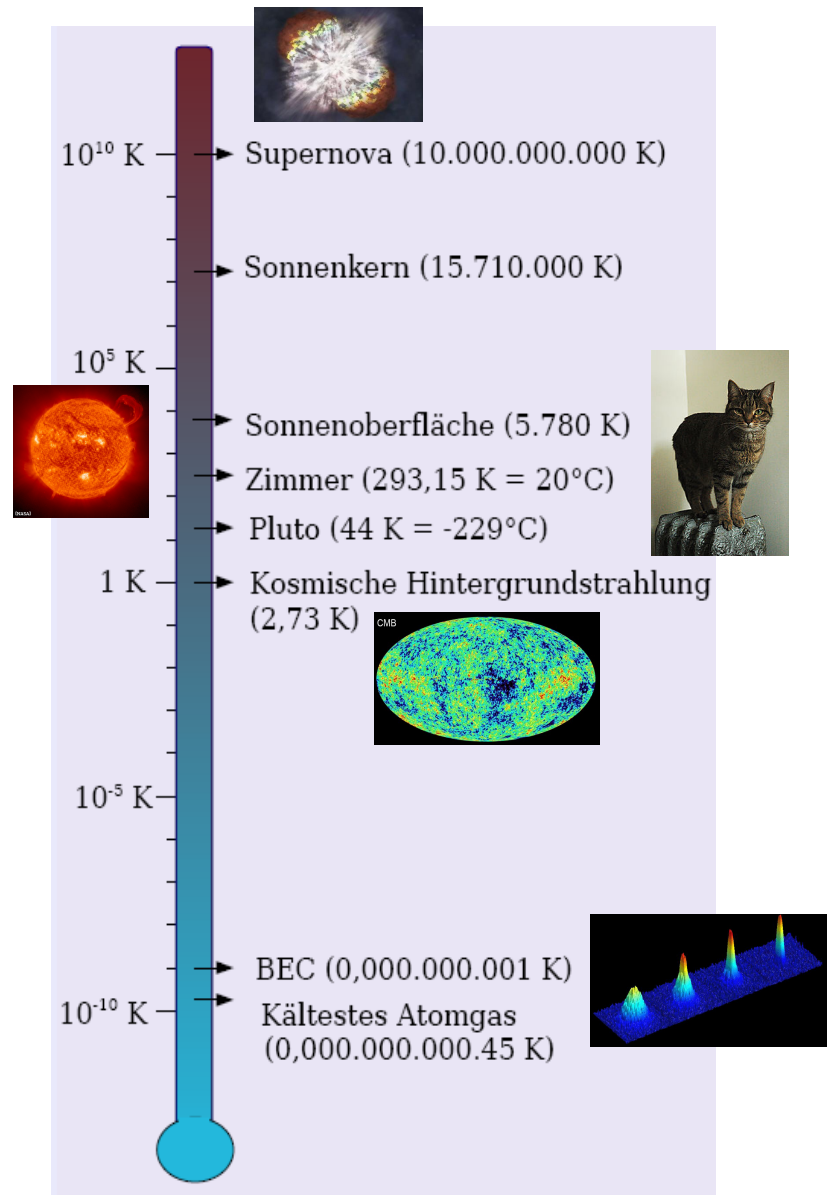
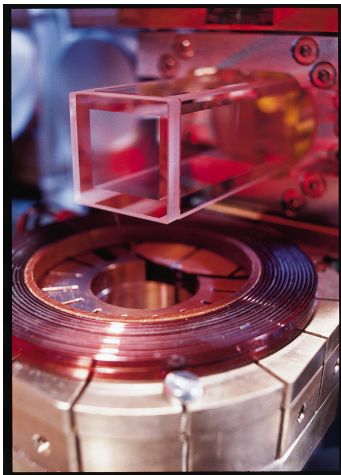
... atoms @ nanokelvins -

trapped only a few mm away from

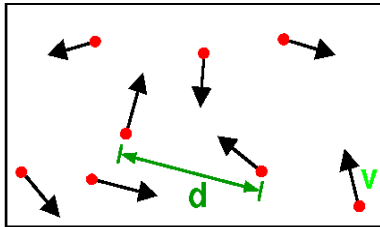
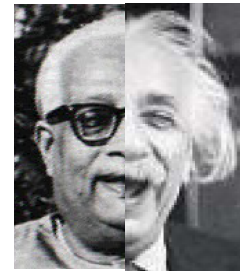
glass cell @ room temperature

(vacuum of 10^{-12} Torr,
i.e. 10^{-15} bar,
or 10^{-10} Pa,

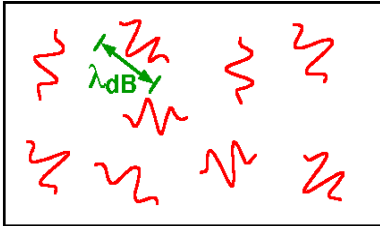
\approx atmospheric
pressure on the moon)



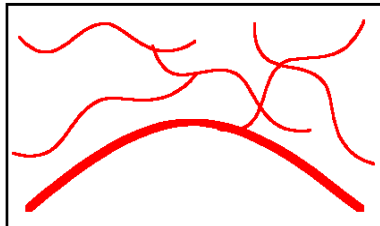
Bose-Einstein condensation



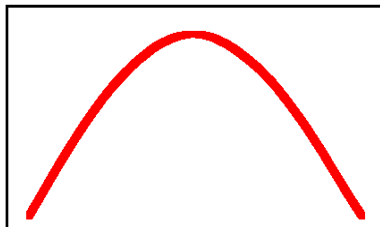
High Temperature T :
 thermal velocity v
 density d^{-3}
 "Billiard balls"



Low Temperature T :
 De Broglie wavelength
 $\lambda_{dB} = h/mv \propto T^{-1/2}$
 "Wave packets"



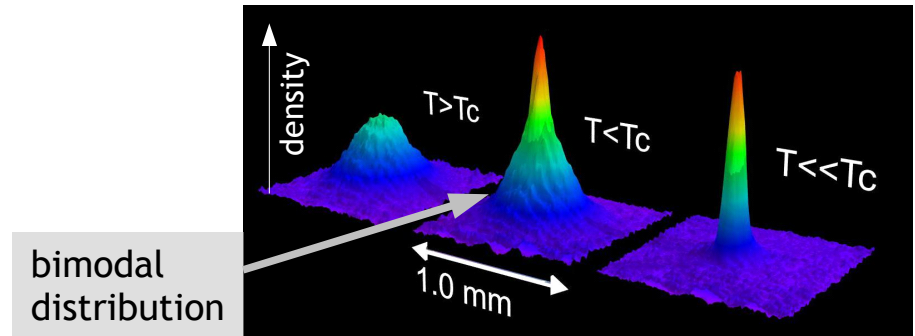
$T = T_{crit}$:
 Bose-Einstein Condensation
 $\lambda_{dB} \approx d$
 "Matter wave overlap"



$T = 0$:
 Pure Bose condensate
 "Giant matter wave"

Experimental picture after free expansion of the trapped cloud:

Bose-Einstein condensation (BEC)



Vortices in a Na condensate

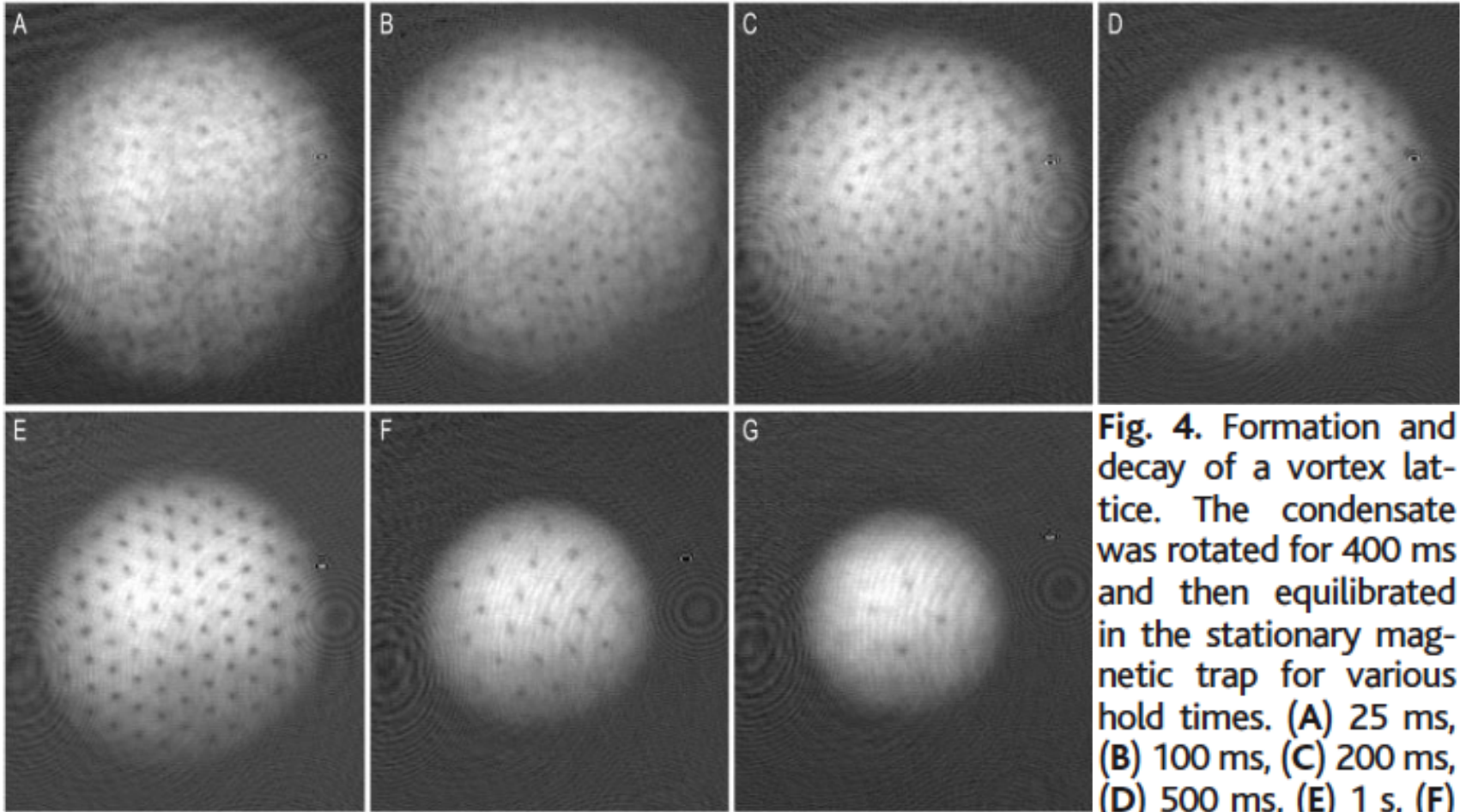
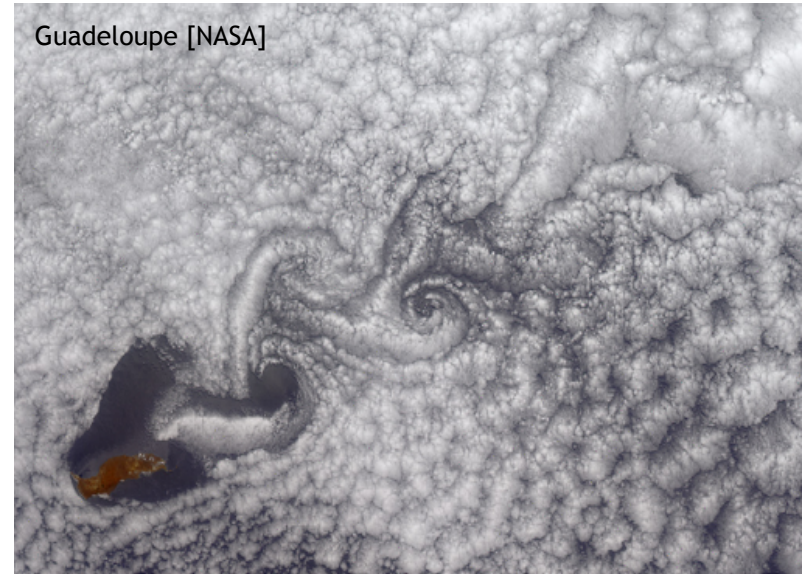


Fig. 4. Formation and decay of a vortex lattice. The condensate was rotated for 400 ms and then equilibrated in the stationary magnetic trap for various hold times. (A) 25 ms, (B) 100 ms, (C) 200 ms, (D) 500 ms, (E) 1 s, (F) 5 s, (G) 10 s.

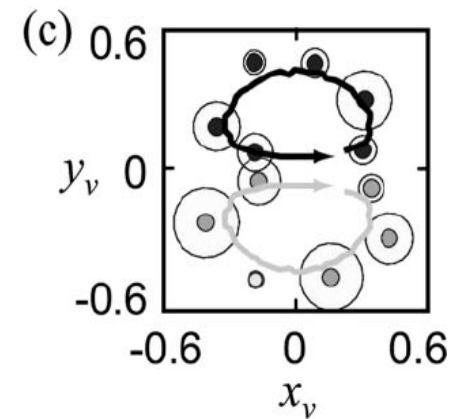
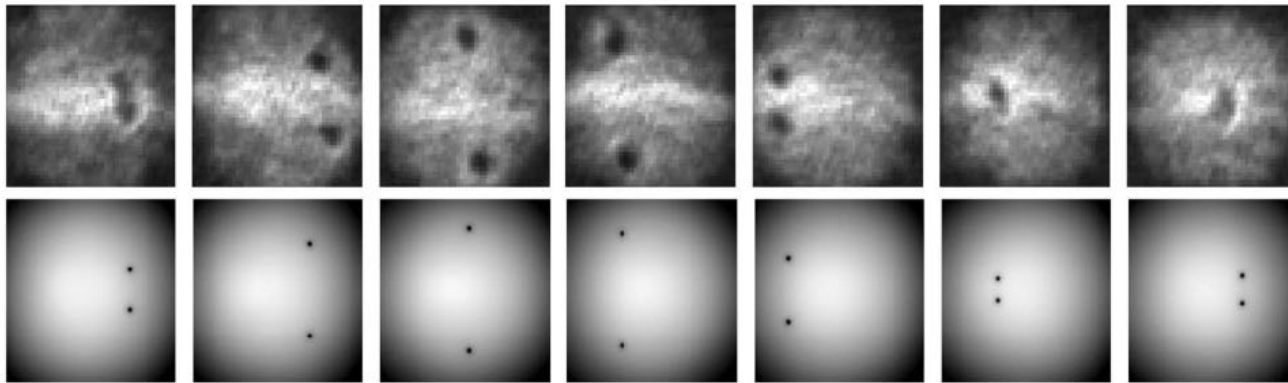
J. R. Abo-Shaeer, C. Raman, J. M. Vogels, W. Ketterle
20 APRIL 2001 VOL 292 SCIENCE



Vortex pairs



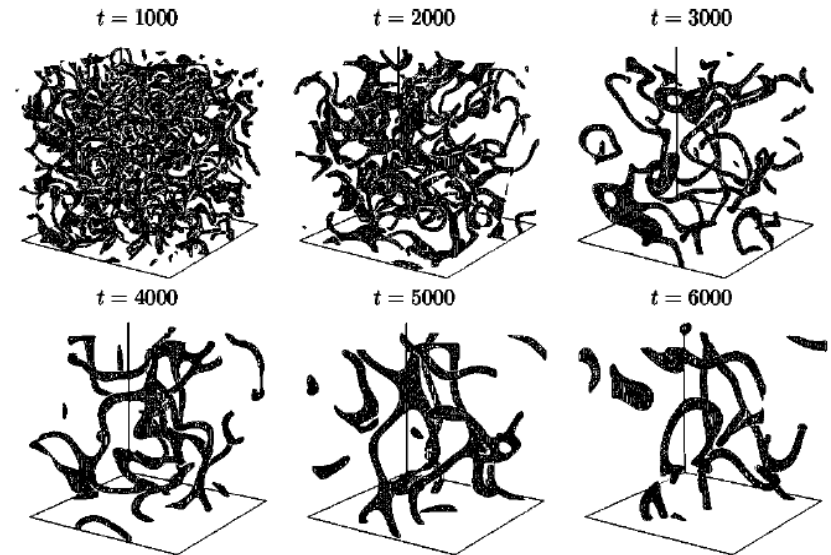
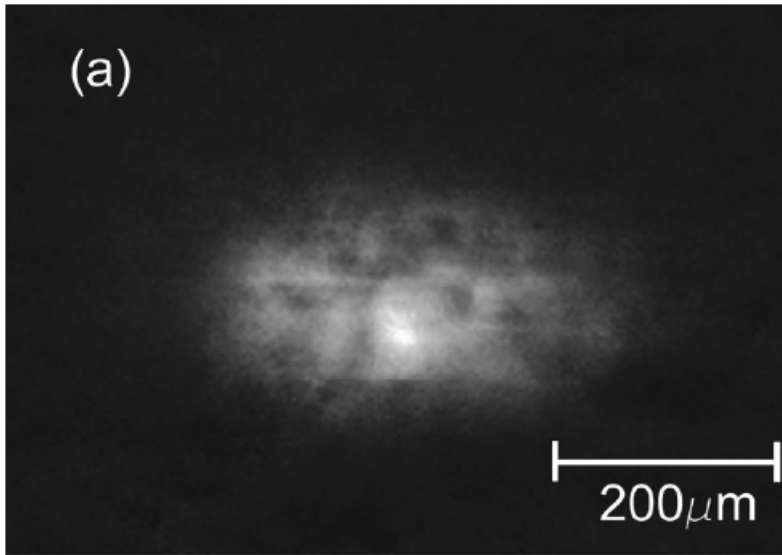
Tucson [AZ]



[T.W. Neely et al. PRL 104 (10)]



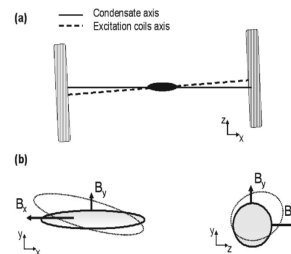
Vortex tangles in Bose Einstein Condensates



[N. Berloff & B. Svistunov, PRA (02)]

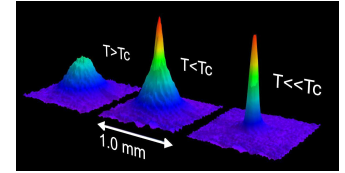


[E.A.L. Henn et al. PRL 103 (09)]



Quantum field theory of a Bose-Einstein gas

How to describe a condensate?



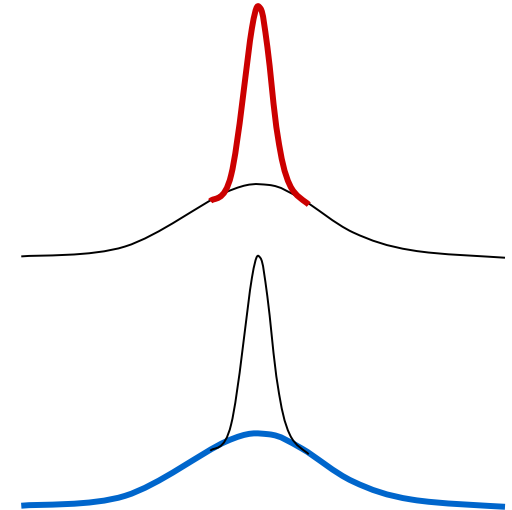
For **bosons**: $[\hat{\Phi}_{t,x}, \hat{\Phi}_{t,x'}^\dagger] = \delta(x - x')$

- Matter wave **mean field** [$x = (x_0, \mathbf{x}) = (t, \mathbf{x})$]

$$\phi_{\mathbf{x}} = \langle \hat{\Phi}_{\mathbf{x}} \rangle, \quad |\phi_{\mathbf{x}}|^2 = n_c(\mathbf{x}) = \text{condensate density},$$

- Density of **non-condensed atoms** ($\hat{\Phi} = \phi + \tilde{\Phi}$, $\phi = \langle \hat{\Phi} \rangle$)

$$\langle \tilde{\Phi}_{\mathbf{x}}^\dagger \tilde{\Phi}_{\mathbf{x}} \rangle = n_{\text{nc}}(\mathbf{x}) \equiv n(\mathbf{x}) - n_c(\mathbf{x}),$$



How to describe a condensate?

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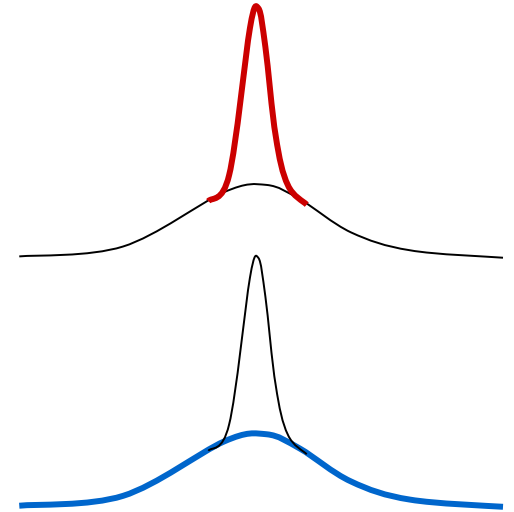
$$\langle \tilde{\Phi}_x^\dagger \tilde{\Phi}_x \rangle = n_{nc}(x) \equiv n(x) - n_c(x),$$

- Total one-body **density matrix**

$$G_{11}(x, y) = \langle \tilde{\Phi}_x^\dagger \tilde{\Phi}_y \rangle \Rightarrow \text{spatial Fourier transform: momentum distribution } n(\mathbf{p}, t)$$

$$\Rightarrow \text{1}^{\text{st}}\text{-order phase coherence } n^{(1)}(s) = n(\mathbf{R} + \mathbf{s}/2, \mathbf{R} - \mathbf{s}/2)$$

$$(n(\mathbf{x}, \mathbf{y}; t) = \langle \hat{\Phi}^\dagger(\mathbf{x}, t) \hat{\Phi}(\mathbf{y}, t) \rangle)$$



Gross-Pitaevskii dynamics

Many-body Hamiltonian:

$$H = \int d\mathbf{x} \hat{\Phi}_x^\dagger \left[-\frac{\hbar^2}{2m} \Delta + V_{\text{ext}}(\mathbf{x}) \right] \hat{\Phi}_x + \frac{1}{2} \int d\mathbf{x} d\mathbf{y} \hat{\Phi}_x^\dagger \hat{\Phi}_y^\dagger V_{\mathbf{x}-\mathbf{y}} \hat{\Phi}_y \hat{\Phi}_x$$



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$$H = \int dx \hat{\Phi}_x^\dagger \left[-\frac{\hbar^2}{2m} \Delta + V_{\text{ext}}(\mathbf{x}) \right] \hat{\Phi}_x + \frac{1}{2} \int dx dy \hat{\Phi}_x^\dagger \hat{\Phi}_y^\dagger V_{\mathbf{x}-\mathbf{y}} \hat{\Phi}_y \hat{\Phi}_x$$

Gross-Pitaevskii, i.e. **Classical Field Equation** for “matter waves”, from vNE:

$$i\partial_t \langle \mathcal{O} \rangle_t = \langle [\mathcal{O}, H] \rangle_t$$
$$\Rightarrow i\hbar \partial_t \phi_{\mathbf{x}} = \left[-\frac{\hbar^2}{2m} \Delta + V_{\text{ext}}(\mathbf{x}) + \overbrace{g|\phi_{\mathbf{x}}|^2}^{\text{from } \int V \langle \hat{\Phi}^\dagger \hat{\Phi} \hat{\Phi} \rangle} \right] \phi_{\mathbf{x}},$$

$$[\hat{\Phi}_{t,\mathbf{x}}, \hat{\Phi}_{t,\mathbf{x}'}^\dagger] = \delta(\mathbf{x} - \mathbf{x}')$$



Gross-Pitaevskii dynamics

Many-body Hamiltonian:

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Gross-Pitaevskii, i.e. **Classical Field Equation** for “matter waves”, from vNE:

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$$\Rightarrow i\hbar \partial_t \phi_{\mathbf{x}} = \left[-\frac{\hbar^2}{2m} \Delta + V_{\text{ext}}(\mathbf{x}) + \overbrace{g|\phi_{\mathbf{x}}|^2}^{\text{from } \int V \langle \hat{\Phi}^\dagger \hat{\Phi} \hat{\Phi} \rangle} \right] \phi_{\mathbf{x}},$$

typical scattering length: $a \simeq 5 \text{ nm}$
 typical bulk density: $n \simeq 10^{14} \text{ cm}^{-3}$
 \Rightarrow diluteness parameter: $na^3 \simeq 10^{-5}$

$$\frac{4\pi\hbar^2 a}{m}$$

(GPE valid for $\sqrt{na^3} \ll 1$)

\Leftrightarrow small condensate depletion)



Beyond Gross-Pitaevskii

$$H = \int dx \hat{\Phi}_x^\dagger \left[-\frac{\hbar^2}{2m} \Delta + V_{\text{ext}}(x) \right] \hat{\Phi}_x + \frac{1}{2} \int dx dy \hat{\Phi}_x^\dagger \hat{\Phi}_y^\dagger V_{x-y} \hat{\Phi}_y \hat{\Phi}_x$$

Hartree-Fock-Bogoliubov (HFB) from $i\partial_t \langle \mathcal{O} \rangle_t = \langle [\mathcal{O}, H] \rangle_t$ and factorisation:

$$\langle \hat{\Phi}_x^\dagger \hat{\Phi}_y^\dagger \hat{\Phi}_z \hat{\Phi}_u \rangle_t \rightarrow \langle \hat{\Phi}_x^\dagger \hat{\Phi}_y^\dagger \rangle_t \langle \hat{\Phi}_z \hat{\Phi}_u \rangle_t + \langle \hat{\Phi}_x^\dagger \hat{\Phi}_z \rangle_t \langle \hat{\Phi}_y^\dagger \hat{\Phi}_u \rangle_t + \langle \hat{\Phi}_x^\dagger \hat{\Phi}_u \rangle_t \langle \hat{\Phi}_y^\dagger \hat{\Phi}_z \rangle_t$$

analogously for 3rd-order correlation function.



Beyond mean field

Dynamical Field Theory



$$[\Phi(t, \mathbf{x}), \Phi^\dagger(t, \mathbf{y})] = \delta(\mathbf{x} - \mathbf{y}) \quad (\text{Bose})$$

Our focus:

time dependence of the lowest-order “correlation” functions:

$$G_{ab}(\mathbf{x}, \mathbf{y}) = \langle \mathcal{T} \Phi_a(\mathbf{x}) \Phi_b(\mathbf{y}) \rangle \quad \mathbf{x} = (t, \mathbf{x})$$



Dynamical field theory



Alternative to BBGKY hierarchy:
Schwinger-Dyson...

$$G_{ab}^{-1}(x, y) = G_{0,ab}^{-1}(x, y; \phi) - \Sigma_{ab}(x, y; \phi, G)$$

or Kadanoff-Baym eqs.:

$$\int_z G_{0,ac}^{-1}(x, z) G_{cb}(z, y) = \delta_{ab} \delta_C(x - y) + \int_z \Sigma_{ac}(x, z) G_{cb}(z, y)$$



Dynamical field theory



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with classical propagator defined by action

$$S[\varphi] = \frac{1}{2} \int_{xy} \varphi_a(x) i G_{0,ab}^{-1}(x, y) \varphi_b(y) - \frac{g}{8} \int_x \varphi_a(x) \varphi_a(x) \varphi_b(x) \varphi_b(x)$$
$$= \delta(x - y) [-i\sigma_{ab}^2 \partial_{x_0} - H_{1B}(x) \delta_{ab}]$$



Dynamical field theory



Alternative to BBGKY hierarchy:
Schwinger-Dyson...

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From two-particle irreducible effective action: [Luttinger, Ward (60); Baym (62); Cornwall, Jackiw, Tomboulis (74)]

$$\Gamma[\phi, G] = S[\phi] + \frac{i}{2} \text{Tr}(\ln G^{-1} + G_0^{-1}(\phi)G) +$$



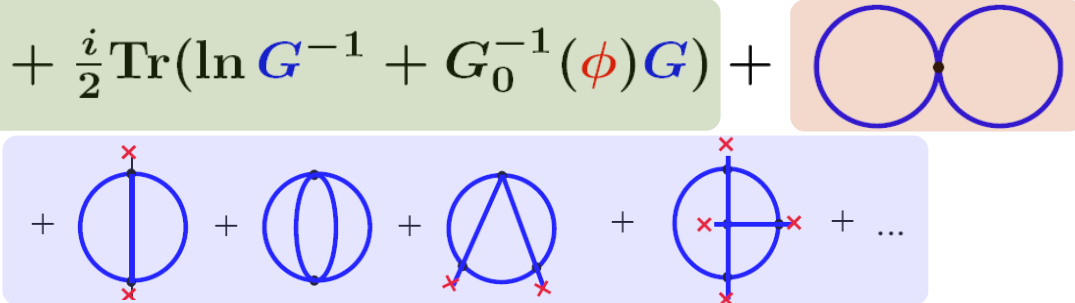
Dynamical field theory

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$$\frac{\delta \Gamma[\phi, G]}{\delta G(x, y)} = 0$$

$$\Sigma_{ab}(x, y; \phi, G) = 2i \frac{\delta \Gamma_2[\phi, G]}{\delta G_{ab}(x, y)}$$



Scaling from field dynamics

Many-body equilibration



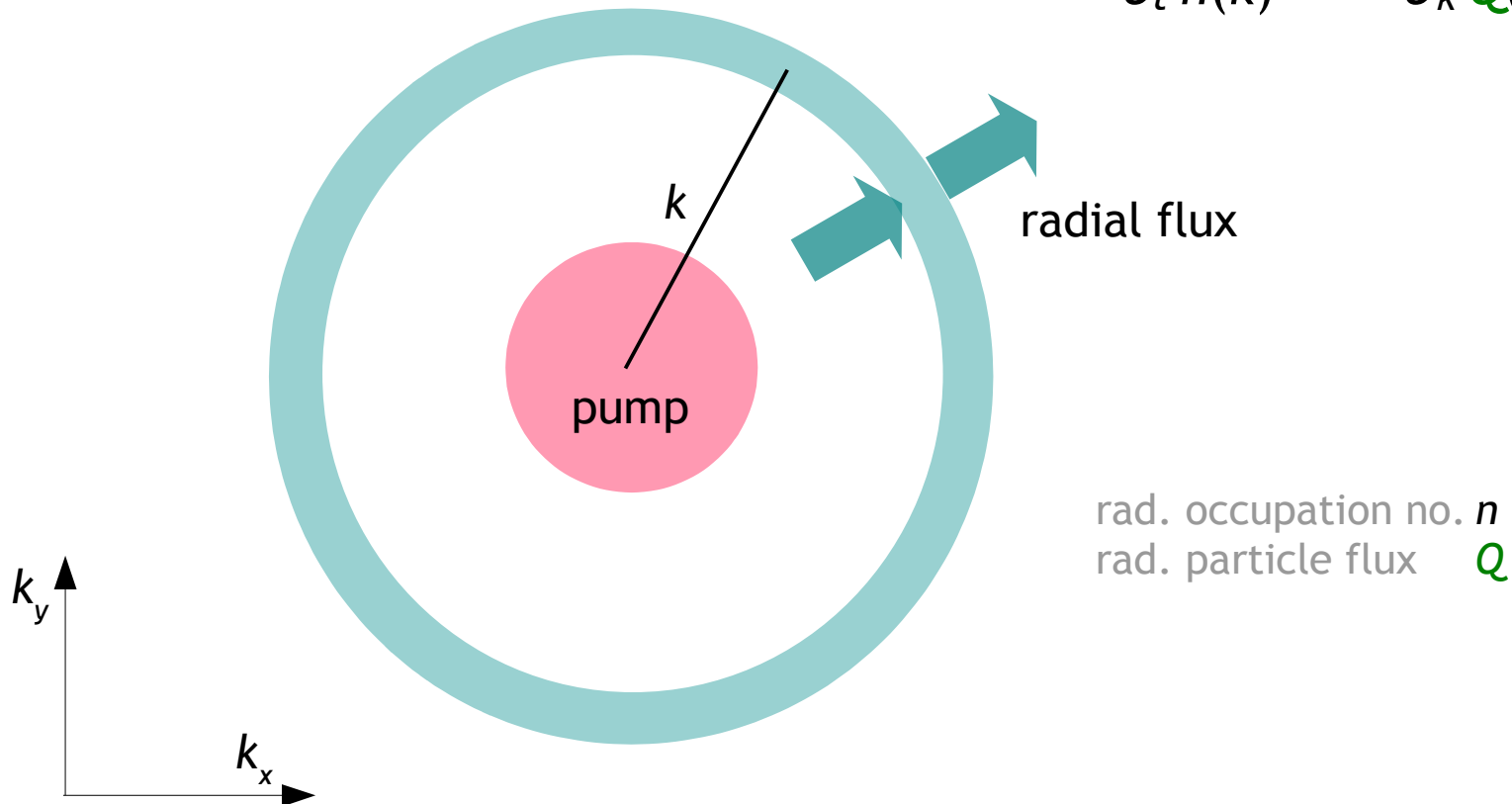
Many-body equilibration



Local radial flux only

Balance equation for radial flux

$$\partial_t n(k) = - \partial_k Q(k)$$

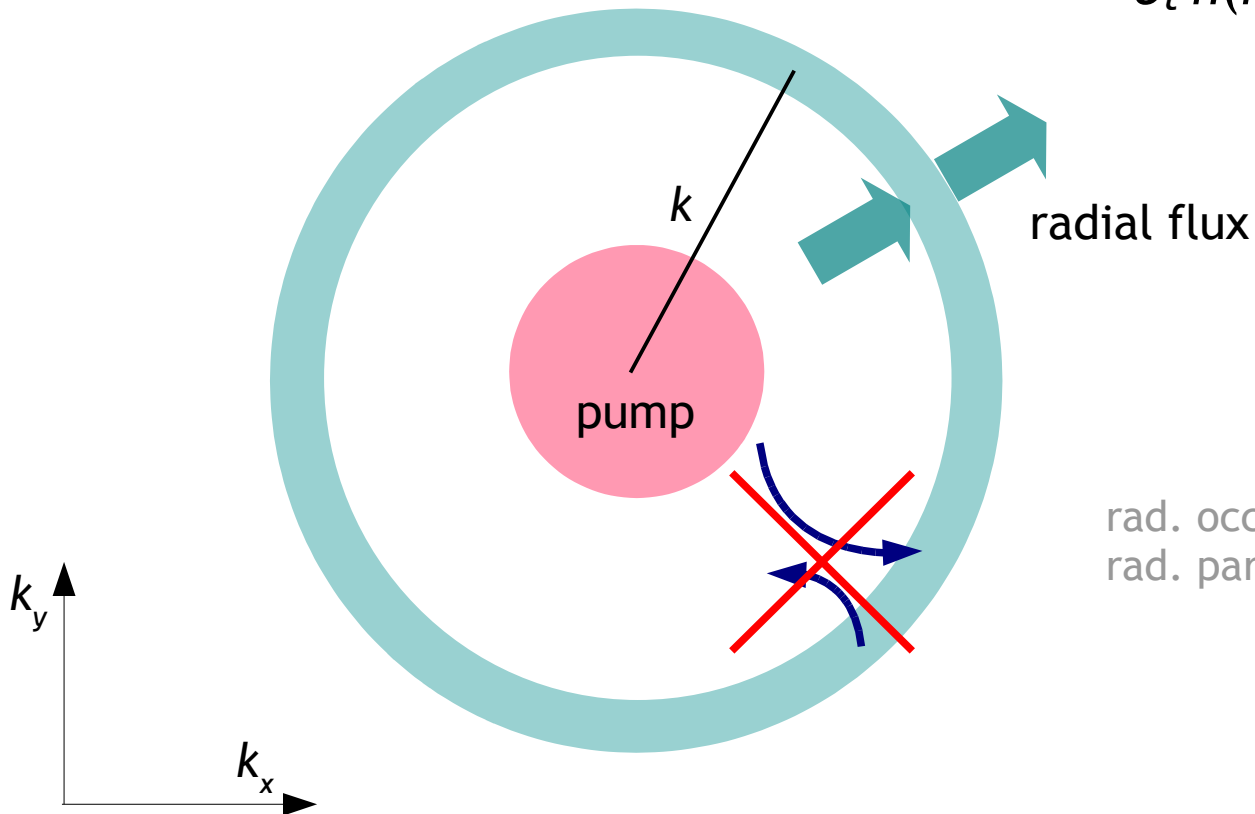


Local radial flux only

With kinetic (Boltzmann) eq.

Boltzmann scattering integral

$$\begin{aligned}\partial_t n(k) &= -\partial_k Q(k) \\ &\sim k^{d-1} J(k) \\ &\neq 0\end{aligned}$$



Local radial flux only

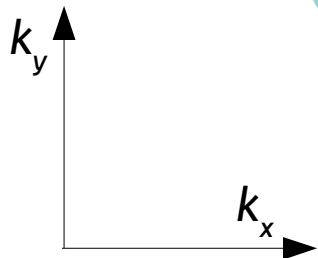
With kinetic (Boltzmann) eq.

Boltzmann
scattering integral

$$\begin{aligned}\partial_t n(k) &= -\partial_k Q(k) \\ &\sim k^{d-1} J(k)\end{aligned}$$

Scattering integral:

$$\begin{aligned}J(\mathbf{k}, t) &= g^2 \int d^d p d^d q d^d r |T_{\mathbf{k}\mathbf{p}\mathbf{q}\mathbf{r}}|^2 \delta(\mathbf{k} + \mathbf{p} - \mathbf{q} - \mathbf{r}) \\ &\quad \times \delta(\omega_{\mathbf{k}} + \omega_{\mathbf{p}} - \omega_{\mathbf{q}} - \omega_{\mathbf{r}}) \\ &\quad \times [(n_{\mathbf{k}} + 1)(n_{\mathbf{p}} + 1)n_{\mathbf{q}}n_{\mathbf{r}} - n_{\mathbf{k}}n_{\mathbf{p}}(n_{\mathbf{q}} + 1)(n_{\mathbf{r}} + 1)]\end{aligned}$$



rad. particle flux Q



Local radial flux only

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Scattering integral:

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$$\times \delta(\omega_{\mathbf{k}} + \omega_{\mathbf{p}} - \omega_{\mathbf{q}} - \omega_{\mathbf{r}}) \quad -2$$
$$\times [(n_{\mathbf{k}} + 1)(n_{\mathbf{p}} + 1)n_{\mathbf{q}}n_{\mathbf{r}} - n_{\mathbf{k}}n_{\mathbf{p}}(n_{\mathbf{q}} + 1)(n_{\mathbf{r}} + 1)] \quad -3\zeta$$

Radial flux density is k -independent, $Q(k) \equiv Q$, if:

$$d - 1 + 1 + 3d - d - 2 - 3\zeta = 0$$



Local radial flux only

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$$\partial_t n(k) = -\partial_k Q(k) \sim k^{d-1} J(k)$$

Scattering integral:

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Radial flux density is k -independent, $Q(k) \equiv Q$, if:

$$\cancel{d-1} + \cancel{1} + 3d - \cancel{d} - 2 - 3\zeta = 0 \Rightarrow \zeta = d - 2/3$$



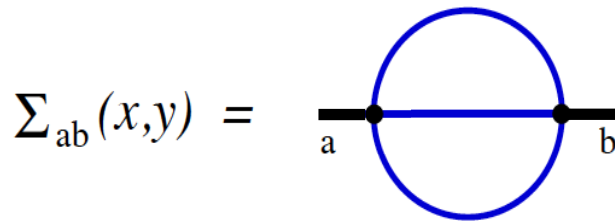
Dynamical field theory

Kinetic (Quantum-Boltzmann) eq.:

$$\partial_t n_{\mathbf{k}} = J(\mathbf{k}, t)$$

Dynamic (Schwinger-Dyson) eq.: (from 2PI effective action)

$$\partial_t n(p) = \Sigma_{ab}^{\rho}(p) F_{ba}(p) - \Sigma_{ab}^F(p) \rho_{ba}(p)$$



$$p = (p_0, \mathbf{p}):$$

$F_{ab}(x, y) = \frac{1}{2} \langle \{ \Phi_a(x), \Phi_b(y) \} \rangle_c$ Statistical function: Occupation

$\rho_{ab}(x, y) = i \langle [\Phi_a(x), \Phi_b(y)] \rangle_c$. Spectral function: Available modes

$$n_{\text{BE}}(\omega) = 1 / (e^{\beta(\omega - \mu)} - 1)$$

$$F_{ab}^{(\text{th})}(\omega, \mathbf{p}) = -i \left(n_{\text{BE}}(\omega) + \frac{1}{2} \right) \rho_{ab}^{(\text{th})}(\omega, \mathbf{p})$$



Dynamical field theory

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$$\partial_t n(p) = \Sigma_{ab}^{\rho}(p) F_{ba}(p) - \Sigma_{ab}^F(p) \rho_{ba}(p)$$

$$\Sigma_{ab}(x,y) = \text{---} \text{---} \text{---}$$

$$p = (p_0, \mathbf{p}):$$

$$\Gamma_2^{3\text{loop}}[\phi, G] = \text{---} \text{---} \text{---} + \text{---} \text{---} \text{---}$$



Scaling solutions

We look for **scaling solutions** fulfilling **stationarity condition** $J(p) = 0$

Scaling ansatz:

$$\rho_{ab}(s^z p_0, s\mathbf{p}) = s^{-2+\eta} \rho_{ab}(p_0, \mathbf{p})$$

$$F_{ab}(s^z p_0, s\mathbf{p}) = s^{-2-\kappa} F_{ab}(p_0, \mathbf{p})$$

Implies scaling of the single-particle momentum distribution:

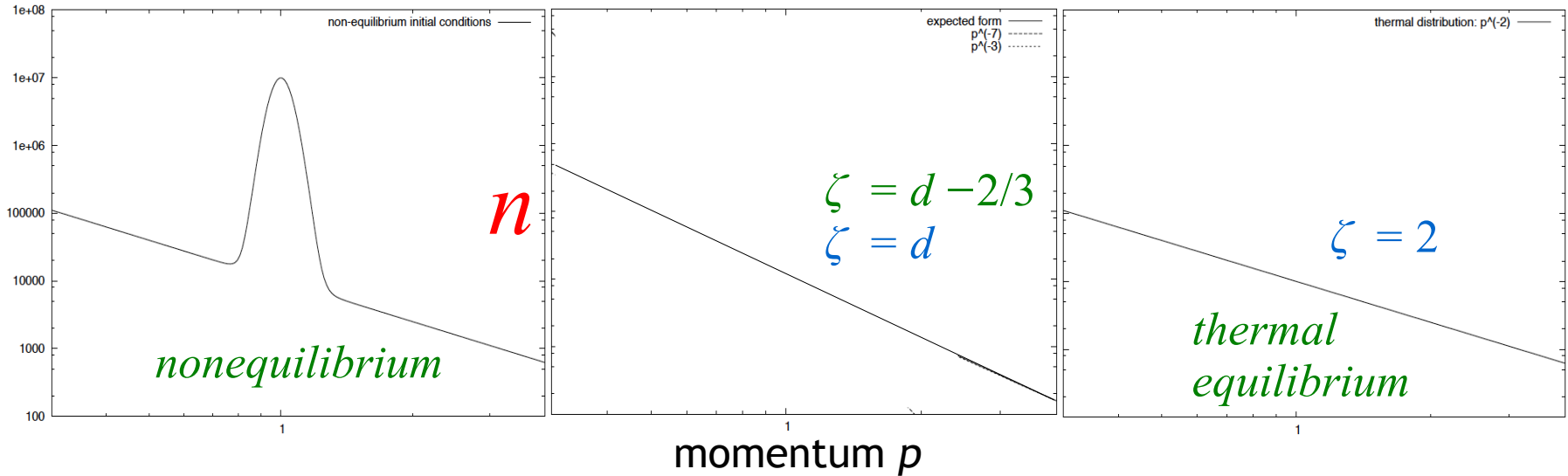
$$n(s\mathbf{p}) = s^{z-2-\kappa} n(\mathbf{p})$$

$= -\zeta$



Turbulent scaling in $d+1$ Dim.

$$n \sim k^{-\zeta}$$



C. Scheppach, J. Berges, TG PRA 81 (10) 033611

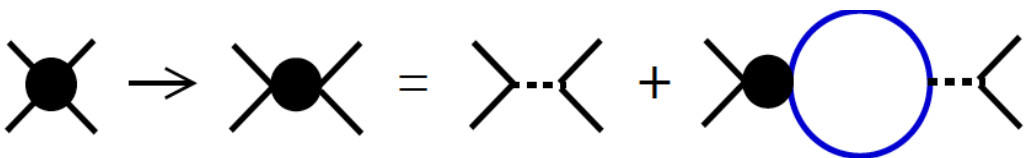


Strong turbulence

$$p = (p_0, \mathbf{p}):$$

$$J(p) := \Sigma_{ab}^{\rho}(p) F_{ba}(p) - \Sigma_{ab}^F(p) \rho_{ba}(p) \stackrel{!}{=} 0$$

$$\Sigma_{ab}(x,y) = \text{Diagram: a horizontal line with vertices 'a' and 'b'. A blue circle loop is attached to vertex 'a'.$$

2PI to NLO in $1/N$: Vertex: 

[Dynamics: J. Berges, (02); G. Aarts et al., (02); TG, Seco, Schmidt, Berges (05);
Kadan.Baym: “GW-Approximation”, Hedin (65)]



Local radial flux only

With kinetic (Boltzmann) eq.

$$T(k) \sim g / |1 + \text{const.} \times g k^{d-2} n(k)| \quad 2 - d + \zeta$$

Scattering integral:

$$J(\mathbf{k}, t) = g^2 \int d^d p d^d q d^d r |T_{\mathbf{k}\mathbf{p}\mathbf{q}\mathbf{r}}|^2 \delta(\mathbf{k} + \mathbf{p} - \mathbf{q} - \mathbf{r}) \quad 3d - d$$

$$\times \delta(\omega_{\mathbf{k}} + \omega_{\mathbf{p}} - \omega_{\mathbf{q}} - \omega_{\mathbf{r}}) \quad -2$$

$$\times [(n_{\mathbf{k}} + 1)(n_{\mathbf{p}} + 1)n_{\mathbf{q}}n_{\mathbf{r}} - n_{\mathbf{k}}n_{\mathbf{p}}(n_{\mathbf{q}} + 1)(n_{\mathbf{r}} + 1)] \quad -3\zeta$$

Radial flux density is k -independent, $Q(k) \equiv Q$, if:

$$\cancel{d-1} + \cancel{1} + 3d - \cancel{d} - 2 - 3\zeta + 2(2 - d + \zeta) = 0$$



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With kinetic (Boltzmann) eq.

$$T(k) \sim g / |1 + \text{const.} \times g k^{d-2} n(k)| \quad 2 - d + \zeta$$

Scattering integral:

$$J(\mathbf{k}, t) = g^2 \int d^d p d^d q d^d r |T_{\mathbf{k}\mathbf{p}\mathbf{q}\mathbf{r}}|^2 \delta(\mathbf{k} + \mathbf{p} - \mathbf{q} - \mathbf{r}) \quad 3d - d$$

$$\times \delta(\omega_{\mathbf{k}} + \omega_{\mathbf{p}} - \omega_{\mathbf{q}} - \omega_{\mathbf{r}}) \quad -2$$

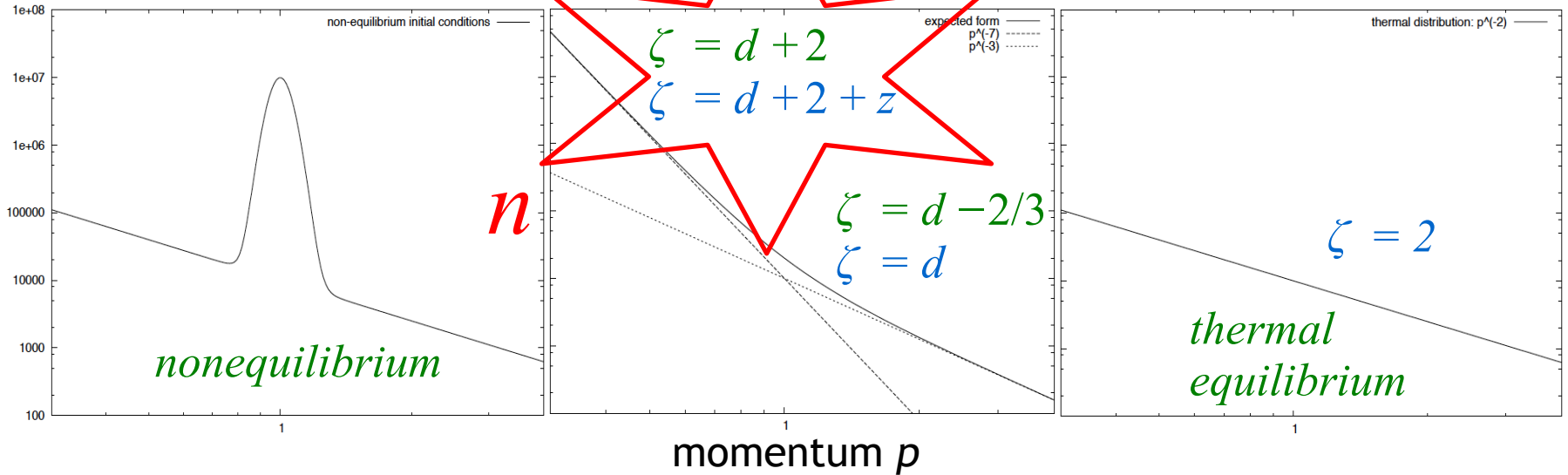
$$\times [(n_{\mathbf{k}} + 1)(n_{\mathbf{p}} + 1)n_{\mathbf{q}}n_{\mathbf{r}} - n_{\mathbf{k}}n_{\mathbf{p}}(n_{\mathbf{q}} + 1)(n_{\mathbf{r}} + 1)] \quad -3\zeta$$

Radial flux density is k -independent, $Q(k) \equiv Q$, if:

$$\cancel{d-1} + \cancel{1} + \cancel{3d} - \cancel{d} - 2 - \cancel{3\zeta} + 2(2 - \cancel{d} + \cancel{\zeta}) = 0 \Rightarrow \zeta = d + 2$$



Strong Turbulence in $d+1$ Dim. $n \sim k^{-\zeta}$

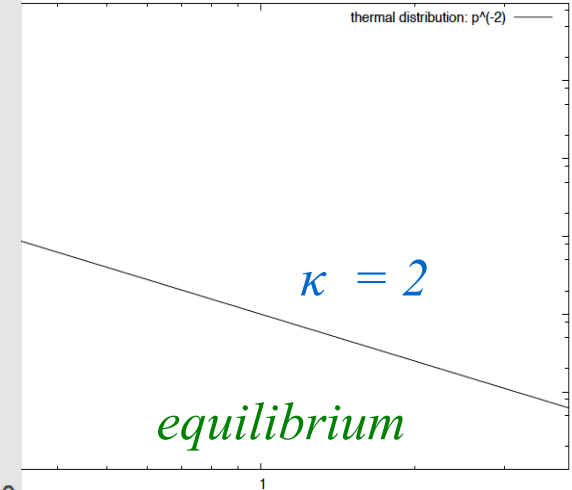
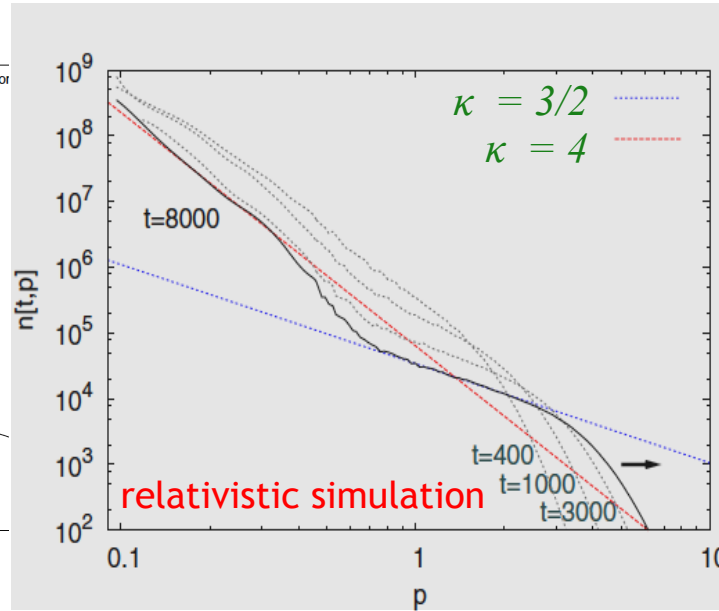
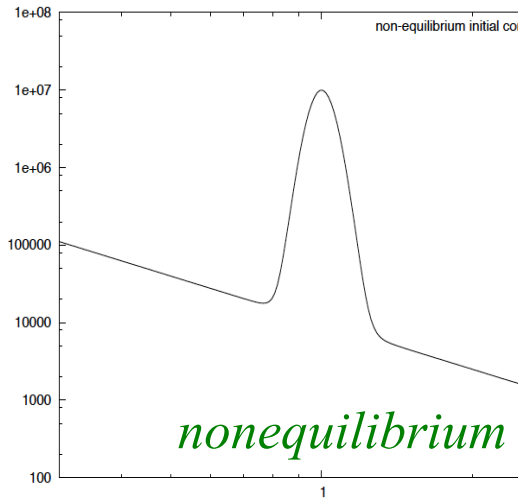


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Summary

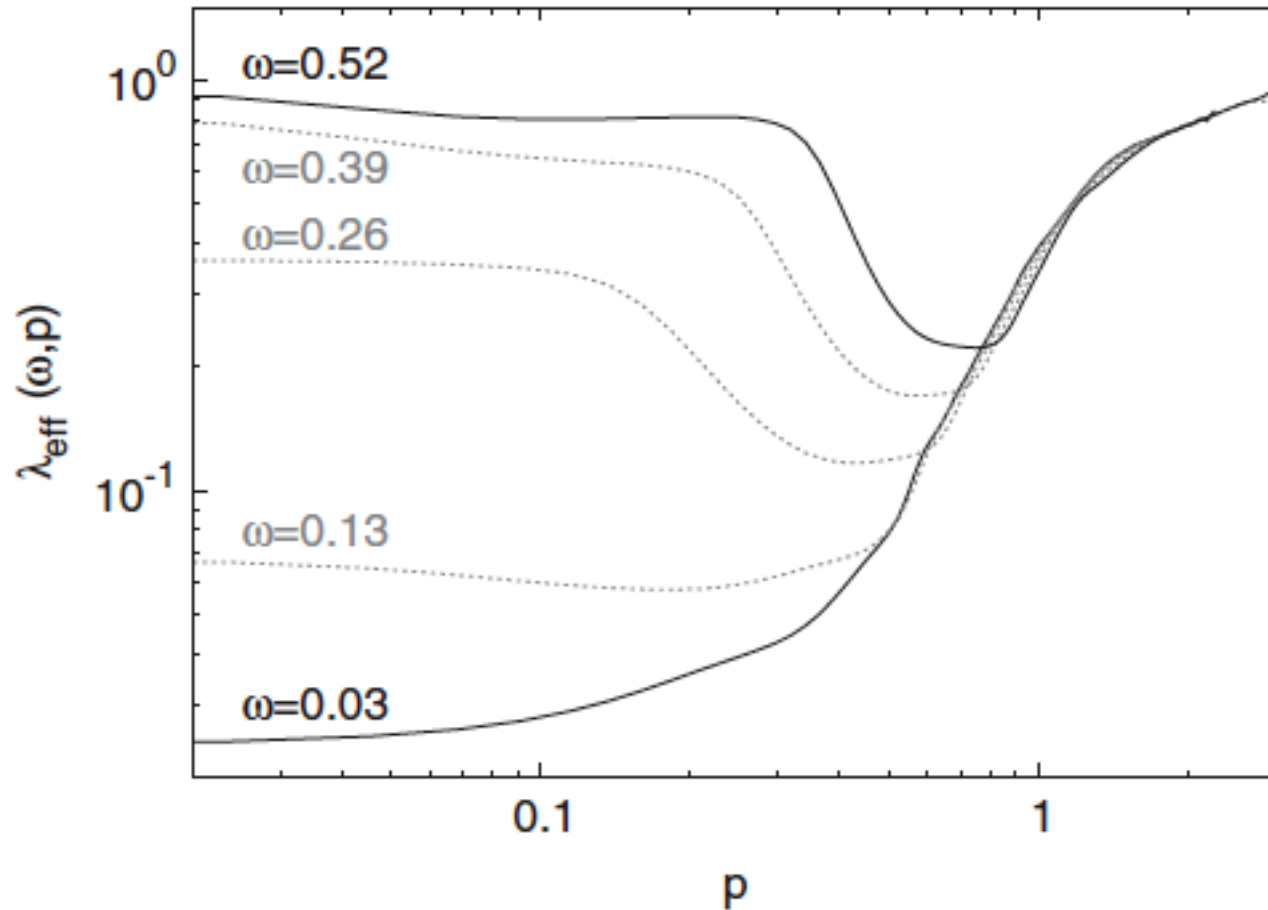
$$n(s\vec{p}) = s^{-\kappa} n(\vec{p})$$



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Effective coupling



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